

Why stop at two?

The search for HHH production

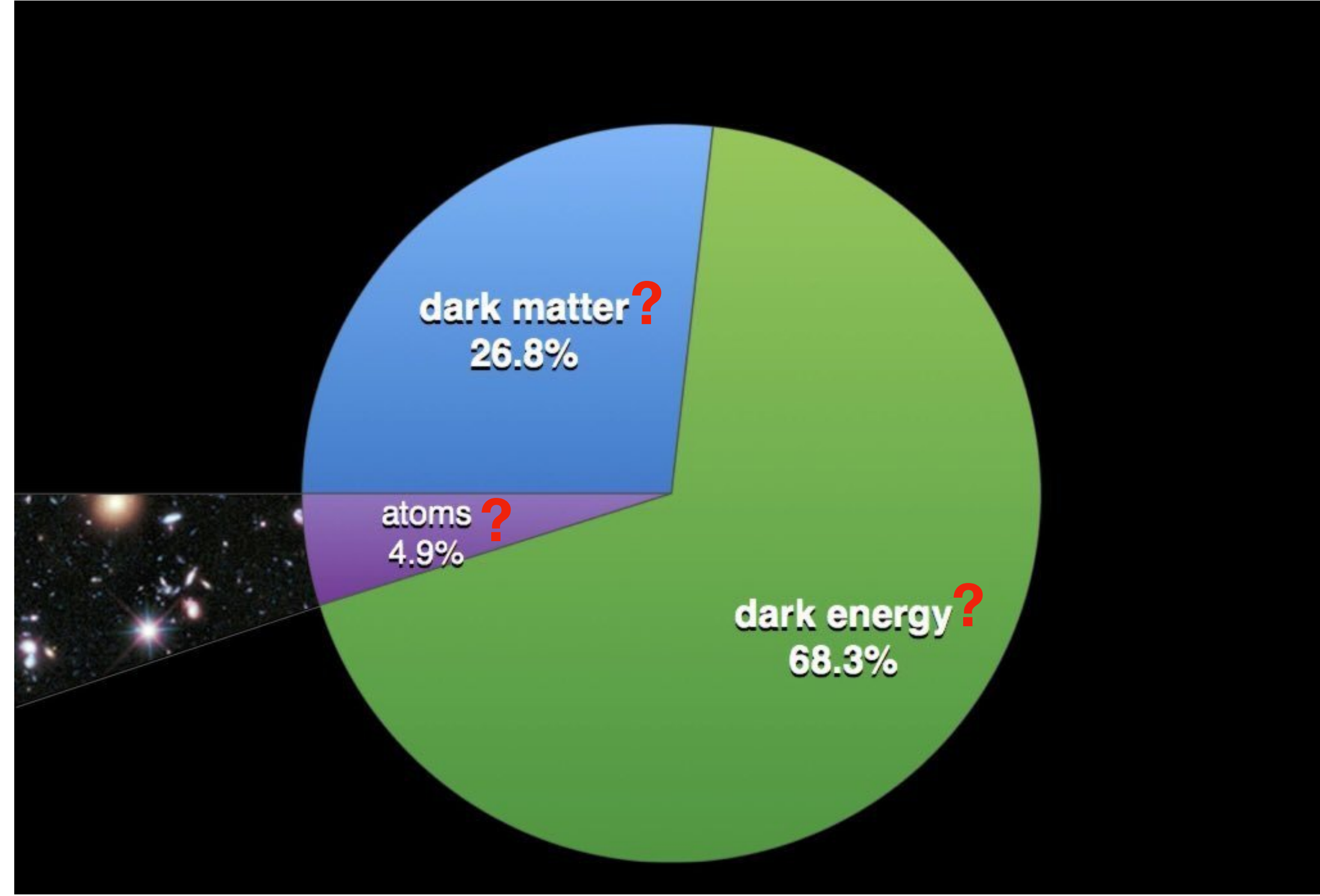
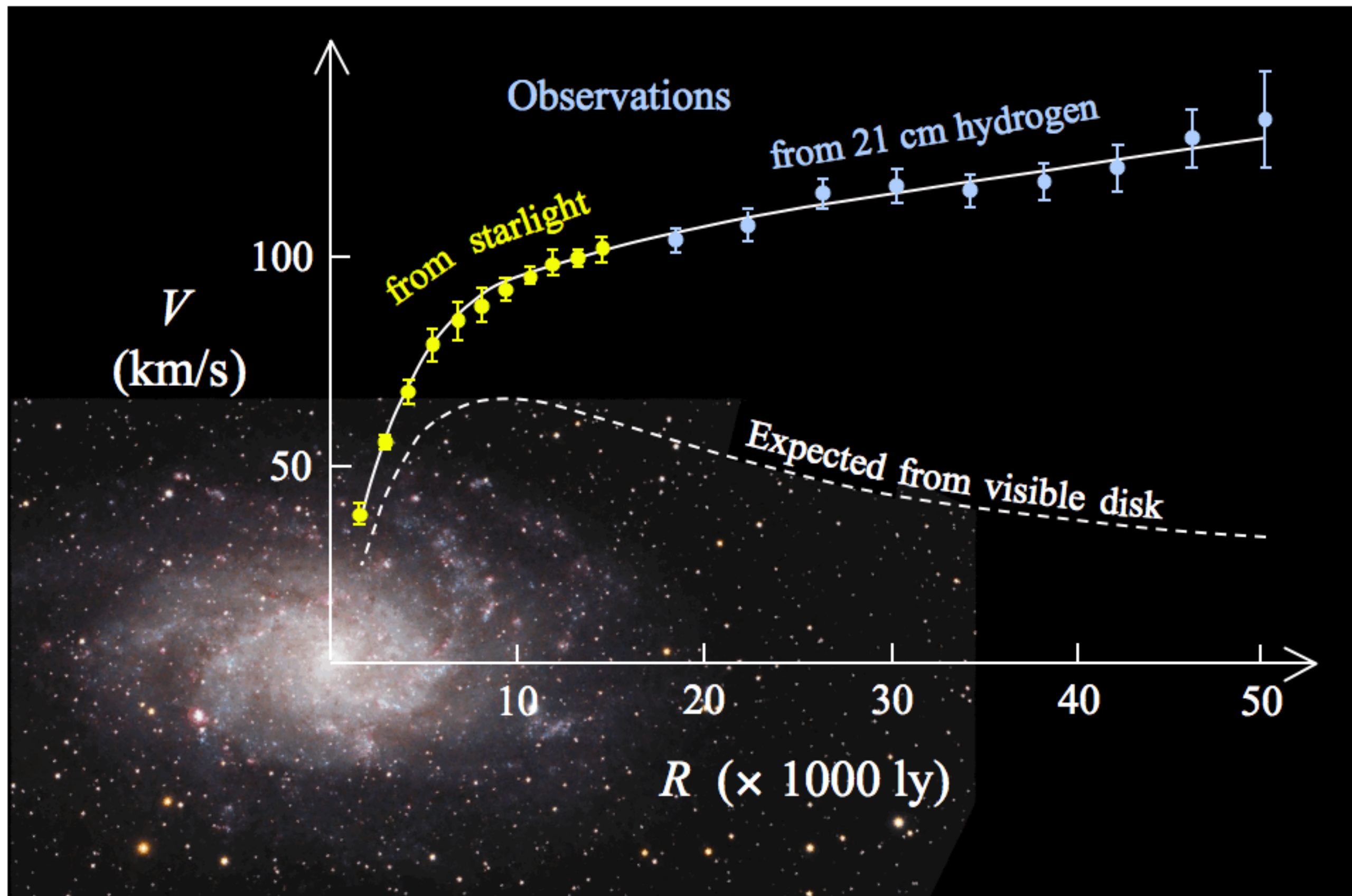
Kevin Nelson

SLAC Seminar

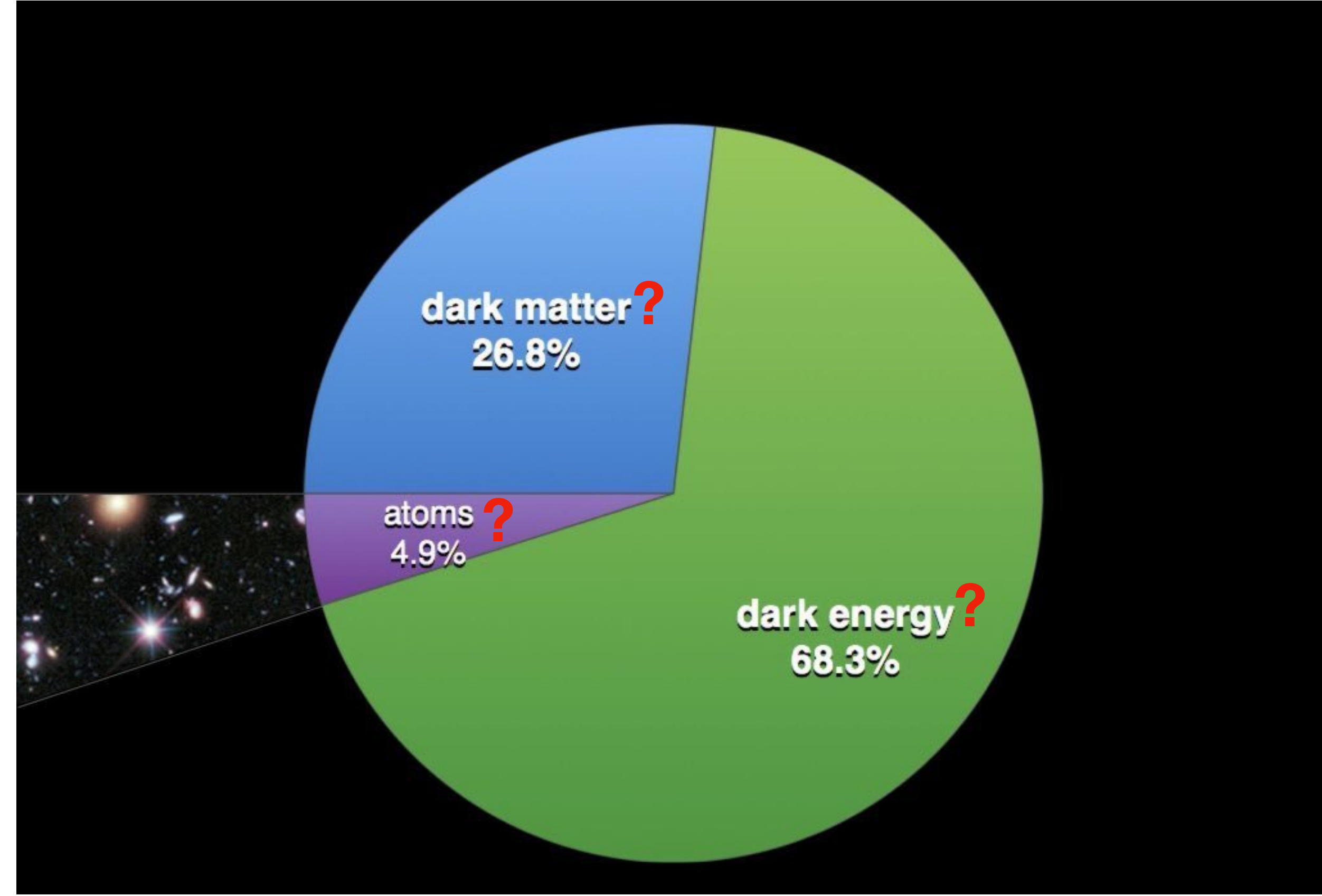
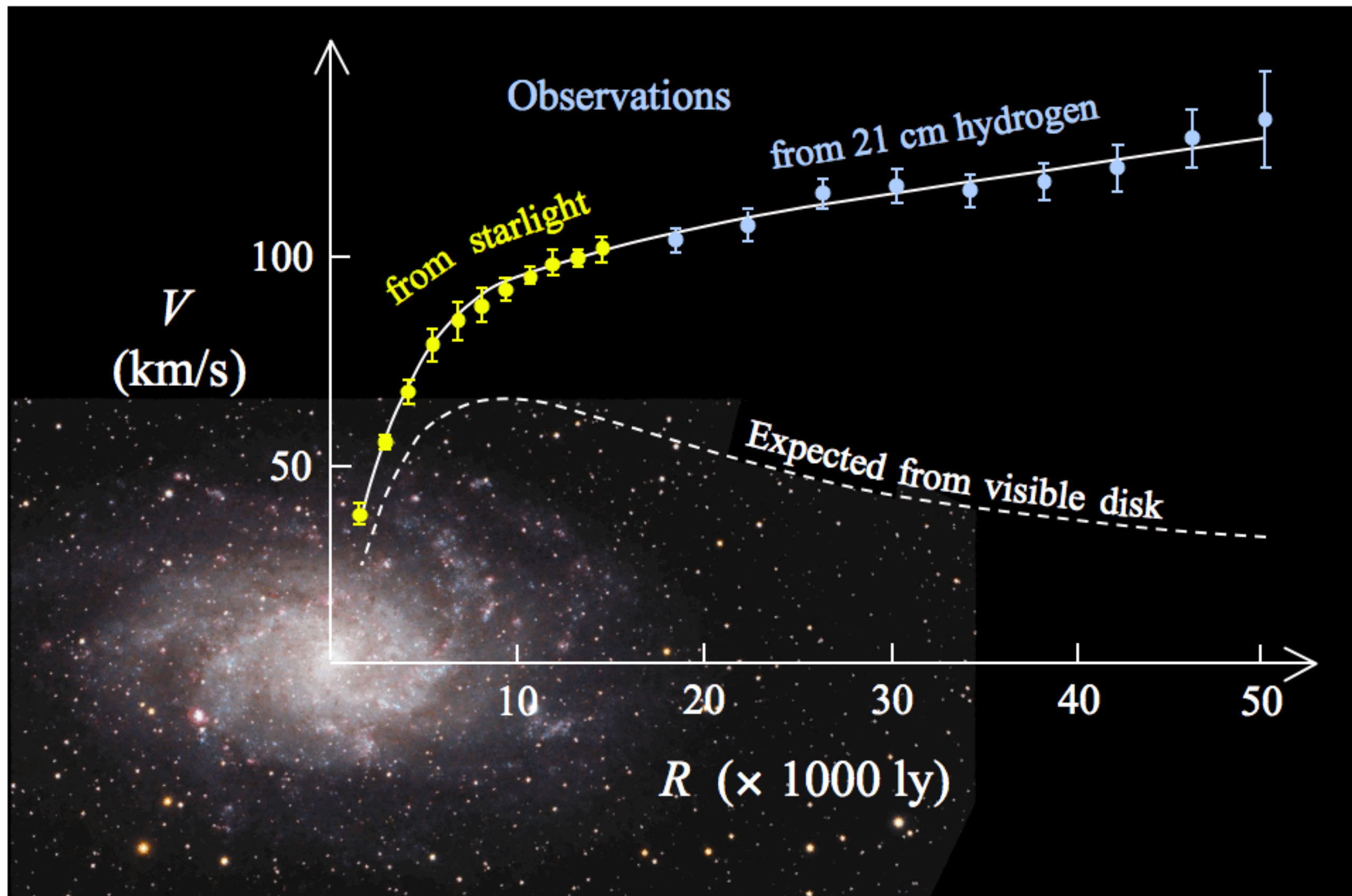
21 January, 2025



We've already seen "new physics"



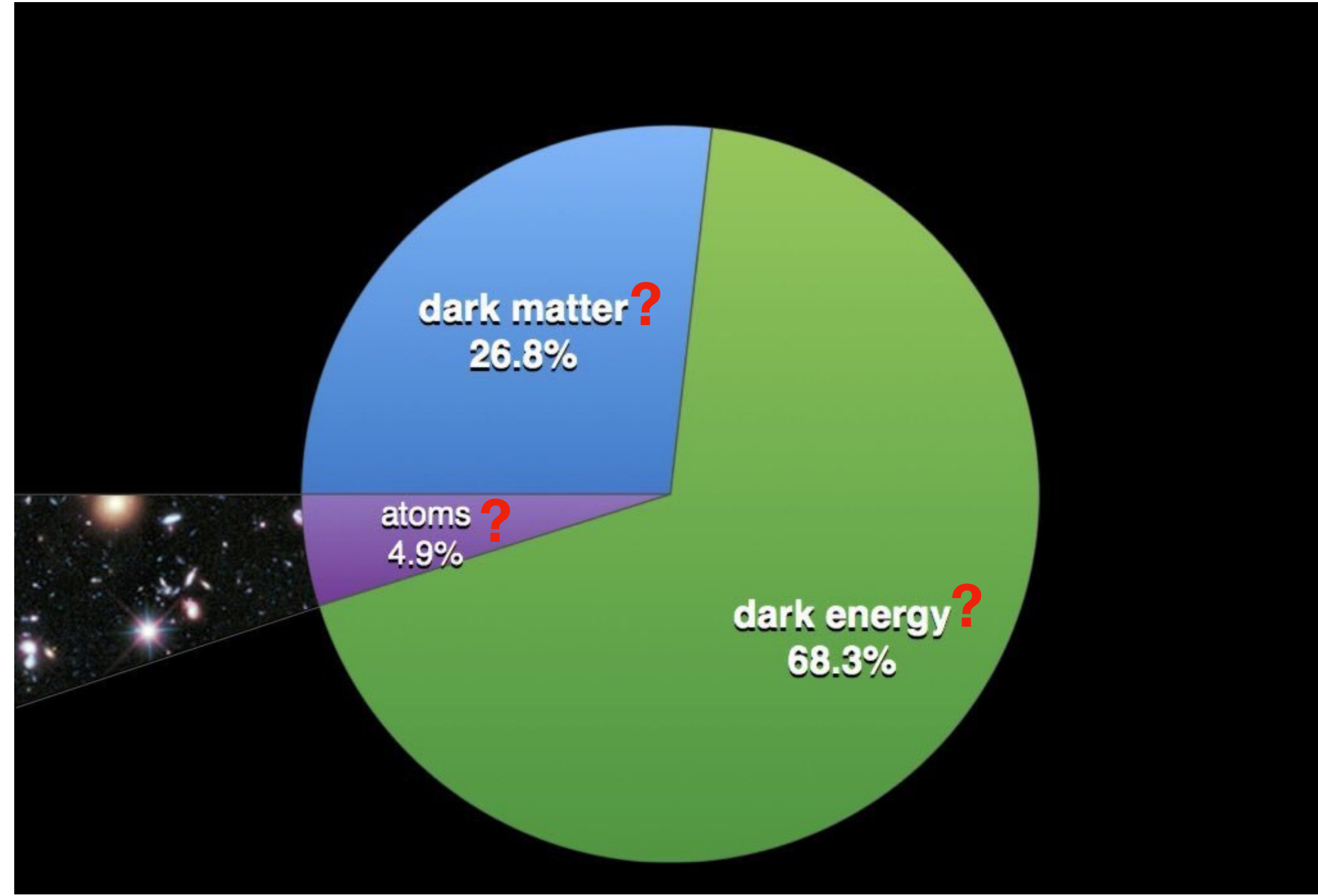
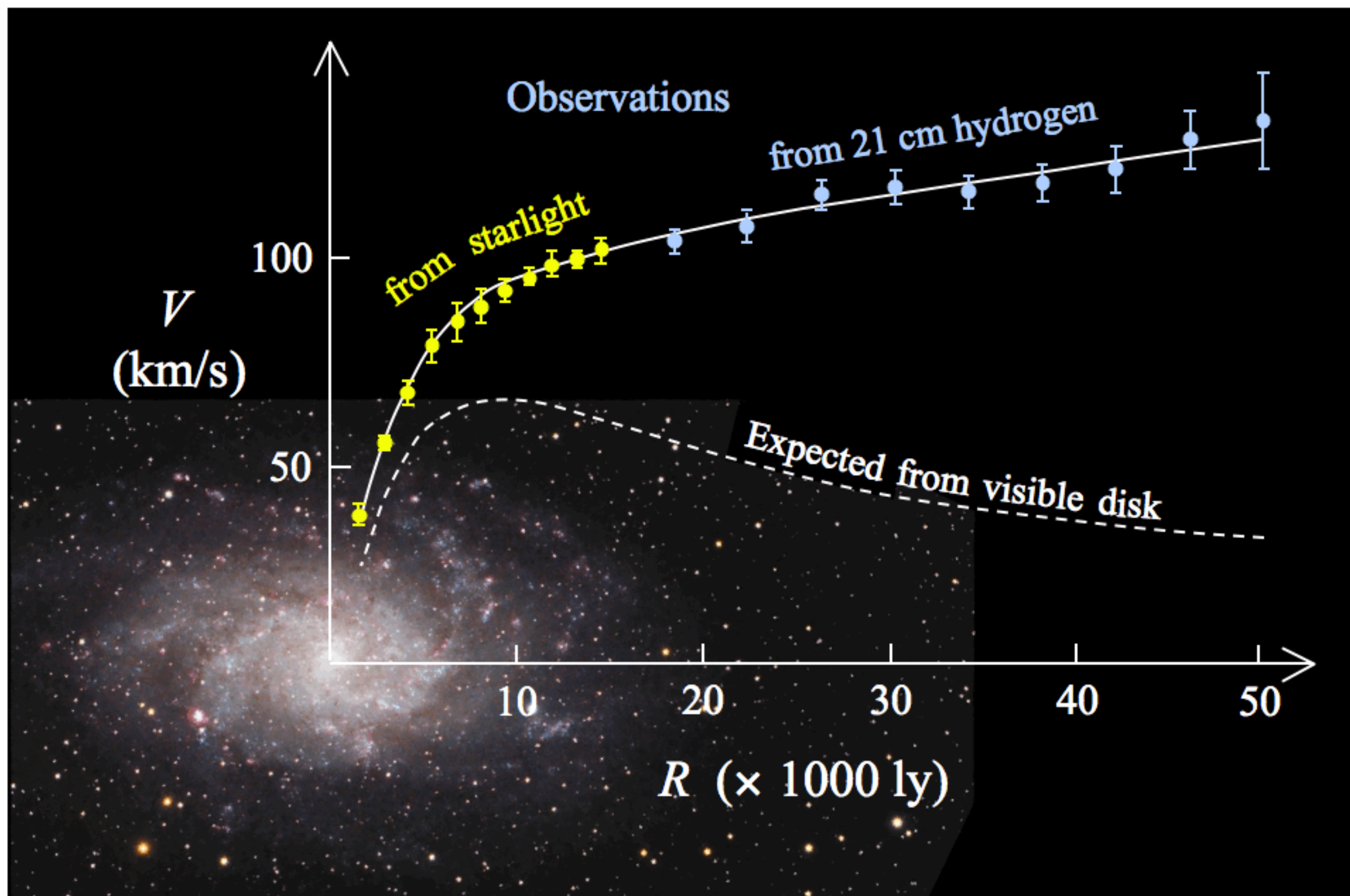
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- Can we describe the particle content of dark matter?
- Are there mediators between dark and luminous matter?
- Can we explain the matter-anti-matter asymmetry and baryogenesis?

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- **But we are left with some questions:**

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All of these questions are related to the search for HHH!

Can we use modern AI methods to find it?

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Yes, there is physics beyond the Standard Model.

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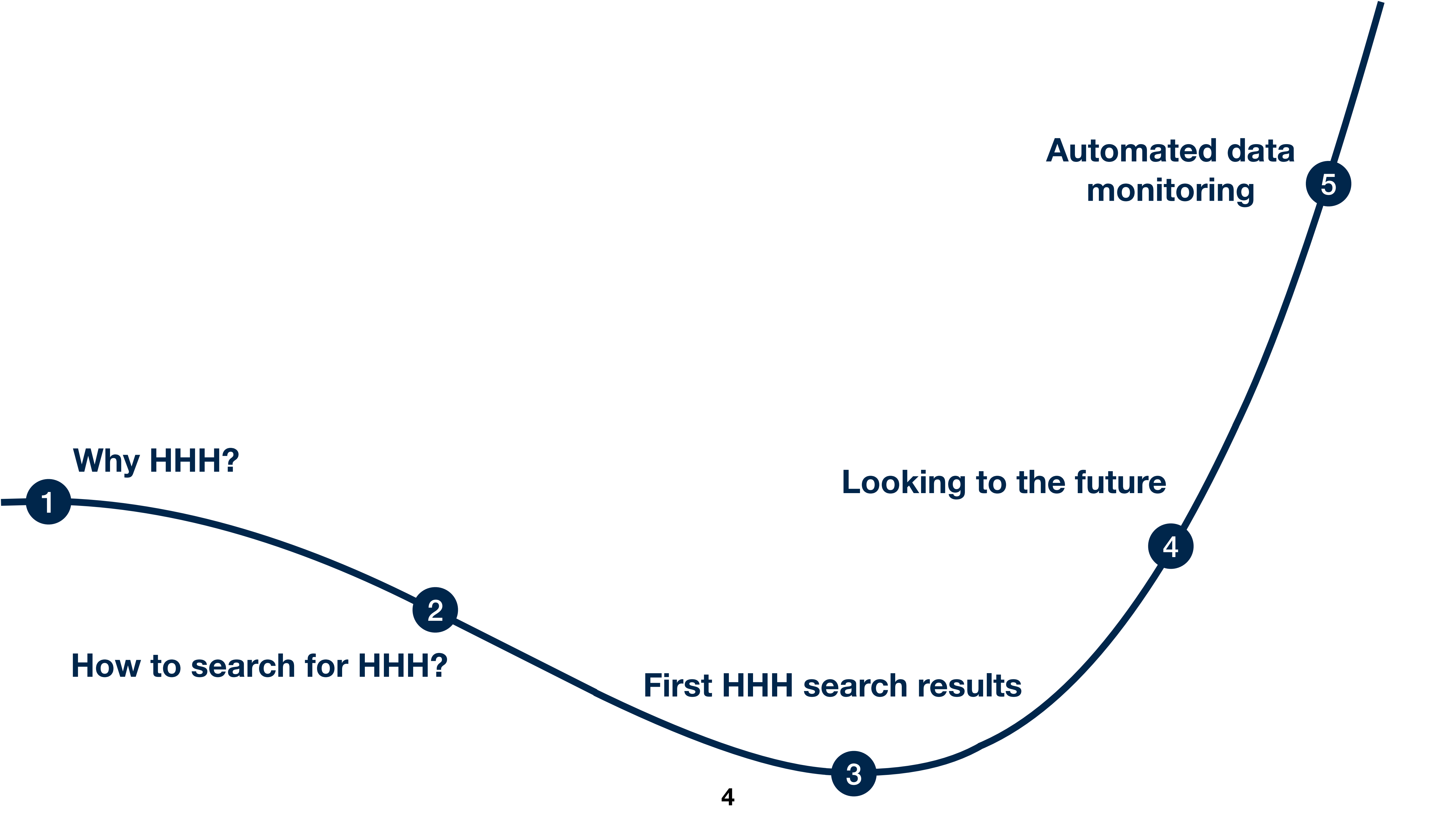
Is there physics beyond the standard model?



Yes, there is physics beyond the Standard Model. The Standard Model of particle physics is a highly successful theory that describes the electromagnetic, weak, and strong nuclear forces, along with the particles that make up matter, such as quarks, leptons, and bosons. However, it is known to be incomplete and has several limitations. Some of the major areas of physics beyond the Standard Model (BSM) include:

1. Gravity

The Standard Model does not include gravity. General relativity describes gravity on a large scale, but it is not compatible with quantum mechanics, which underlies the Standard Model. A theory of



1

Why HHH?

2

How to search for HHH?

3

First HHH search results

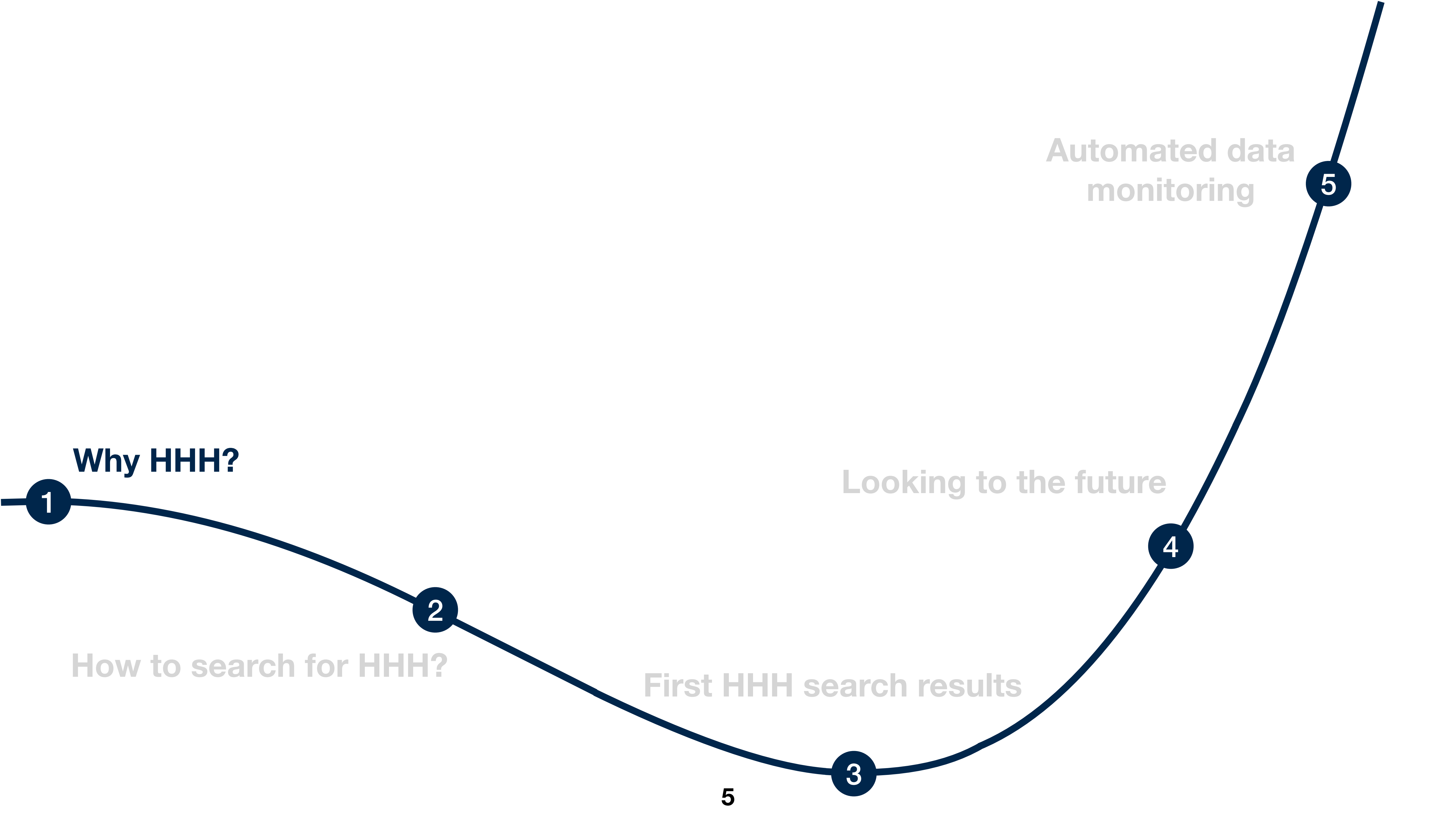
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Looking to the future

5

Automated data monitoring



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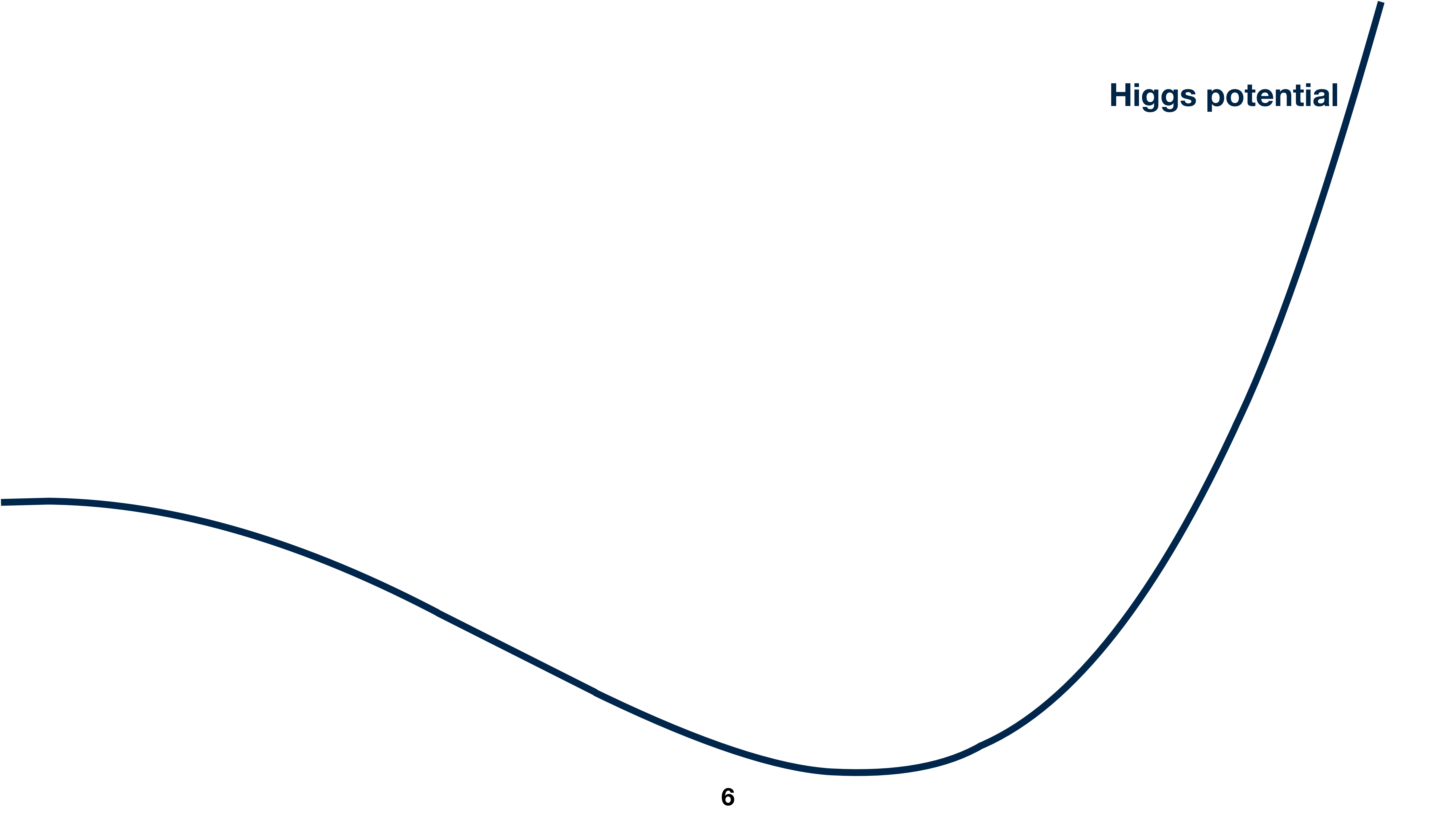
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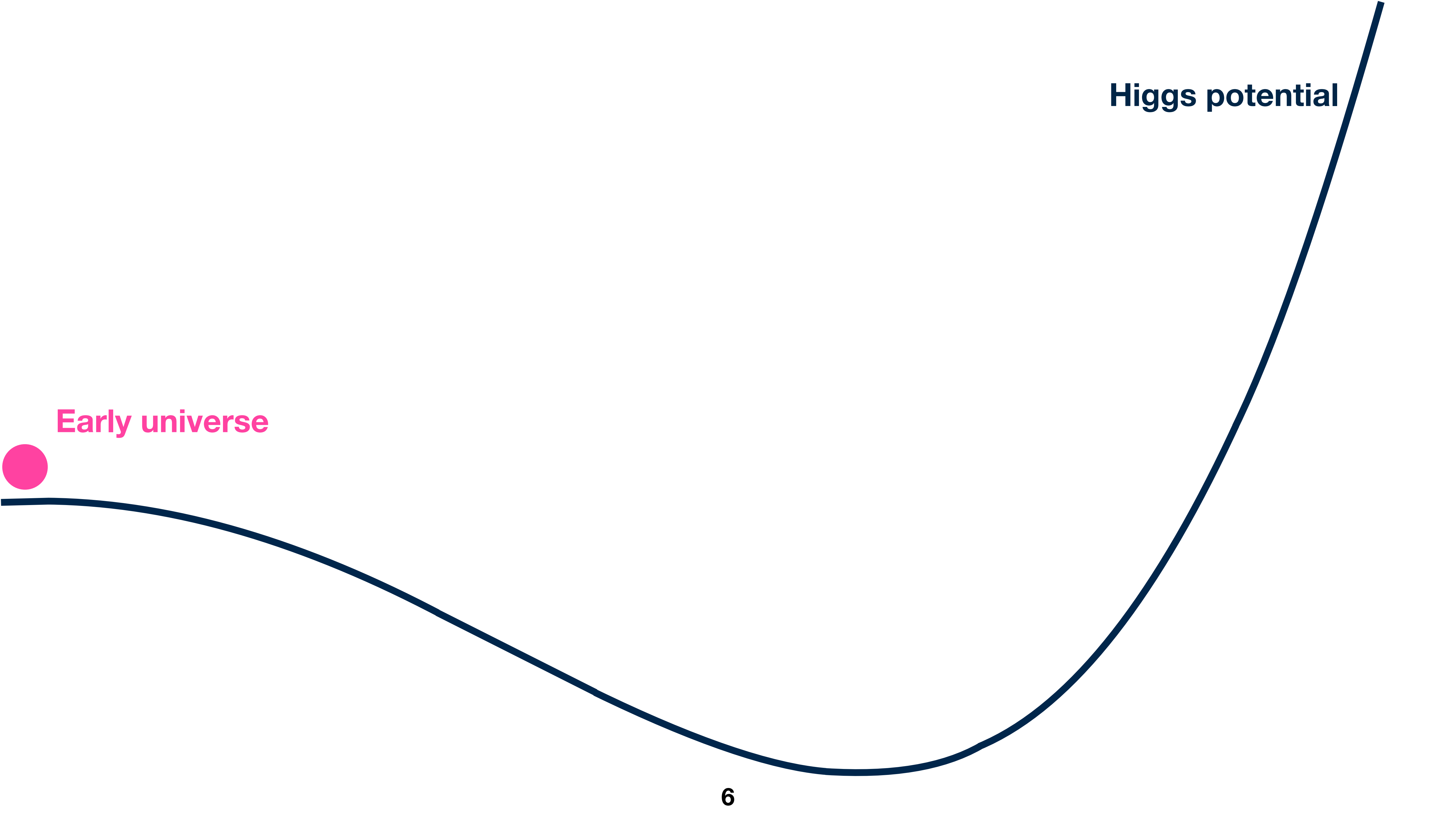
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Higgs potential

6

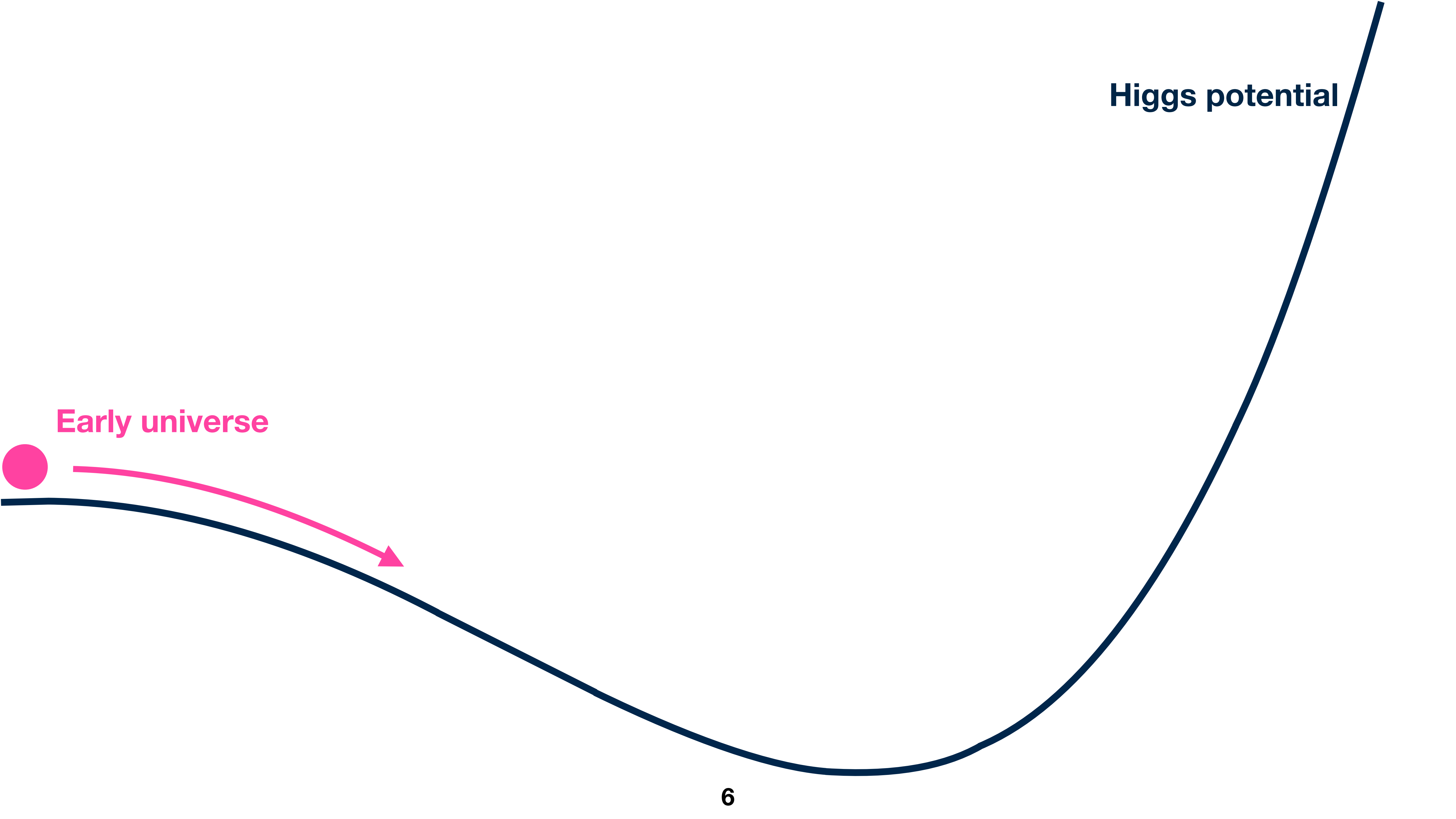




Early universe

Higgs potential

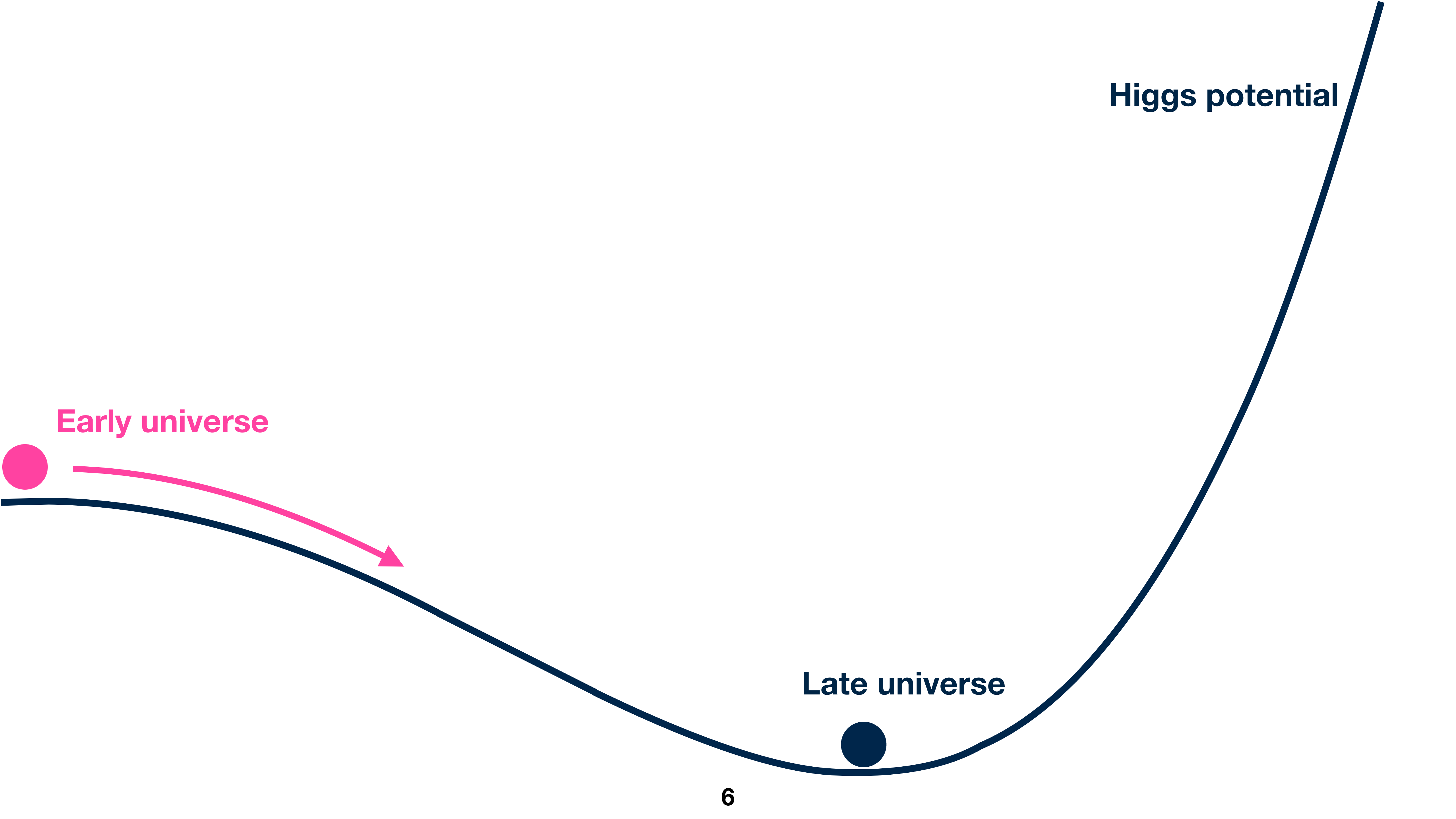
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Higgs potential

Early universe

6



Higgs potential

Early universe

Late universe

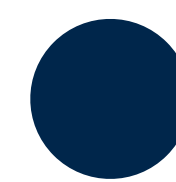
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Higgs potential

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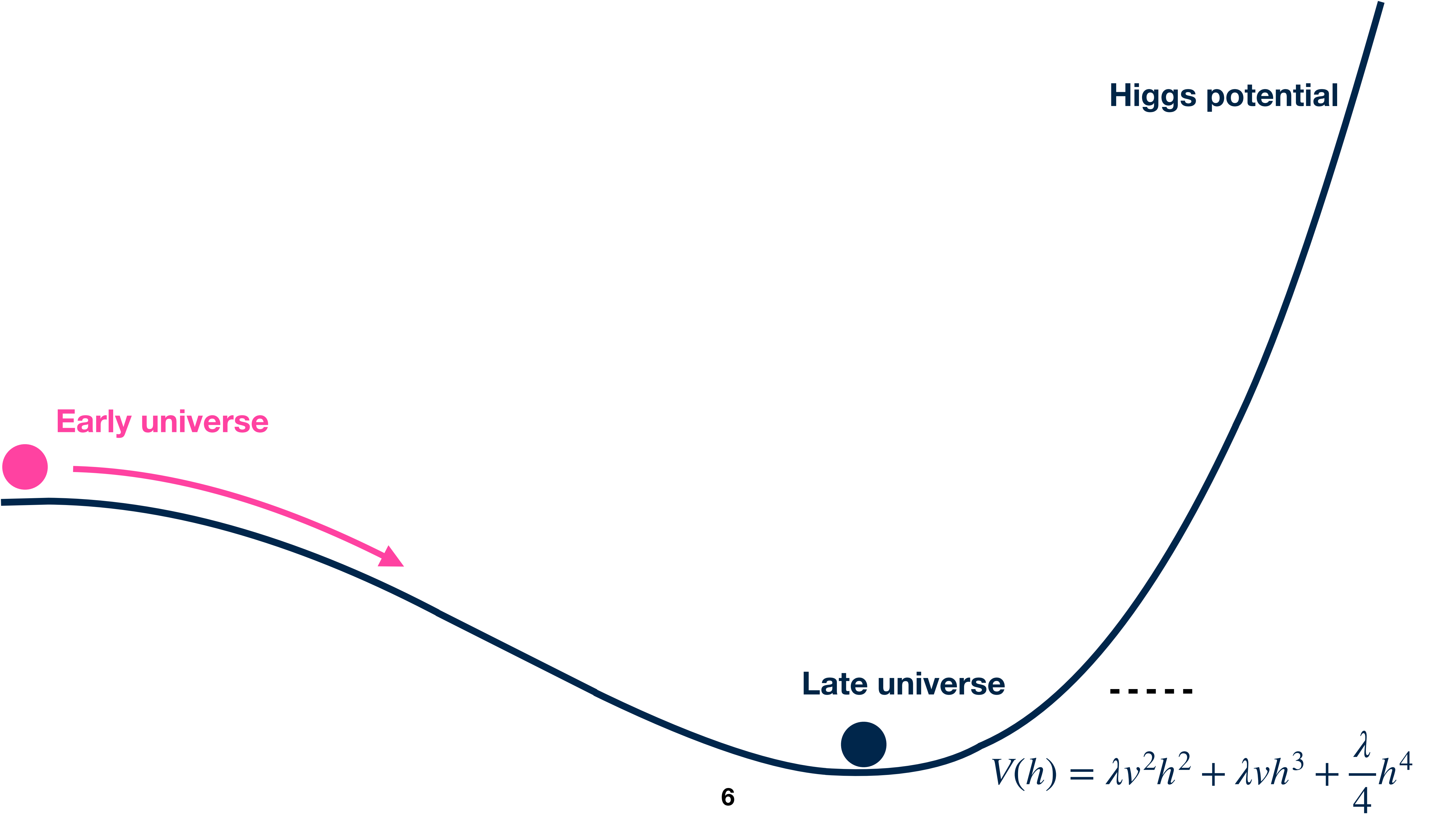


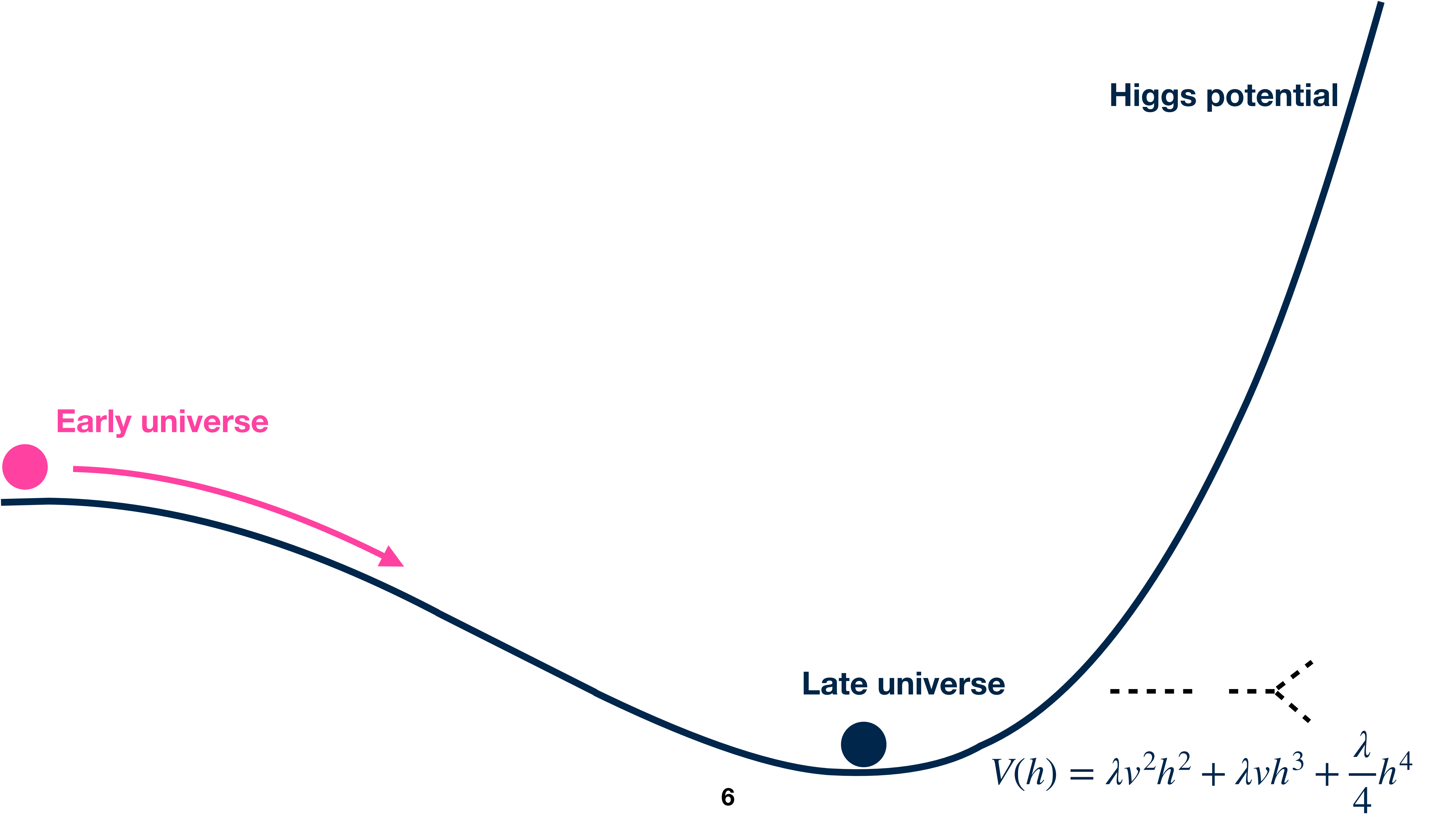
Late universe



6

$$V(h) = \lambda v^2 h^2 + \lambda v h^3 + \frac{\lambda}{4} h^4$$





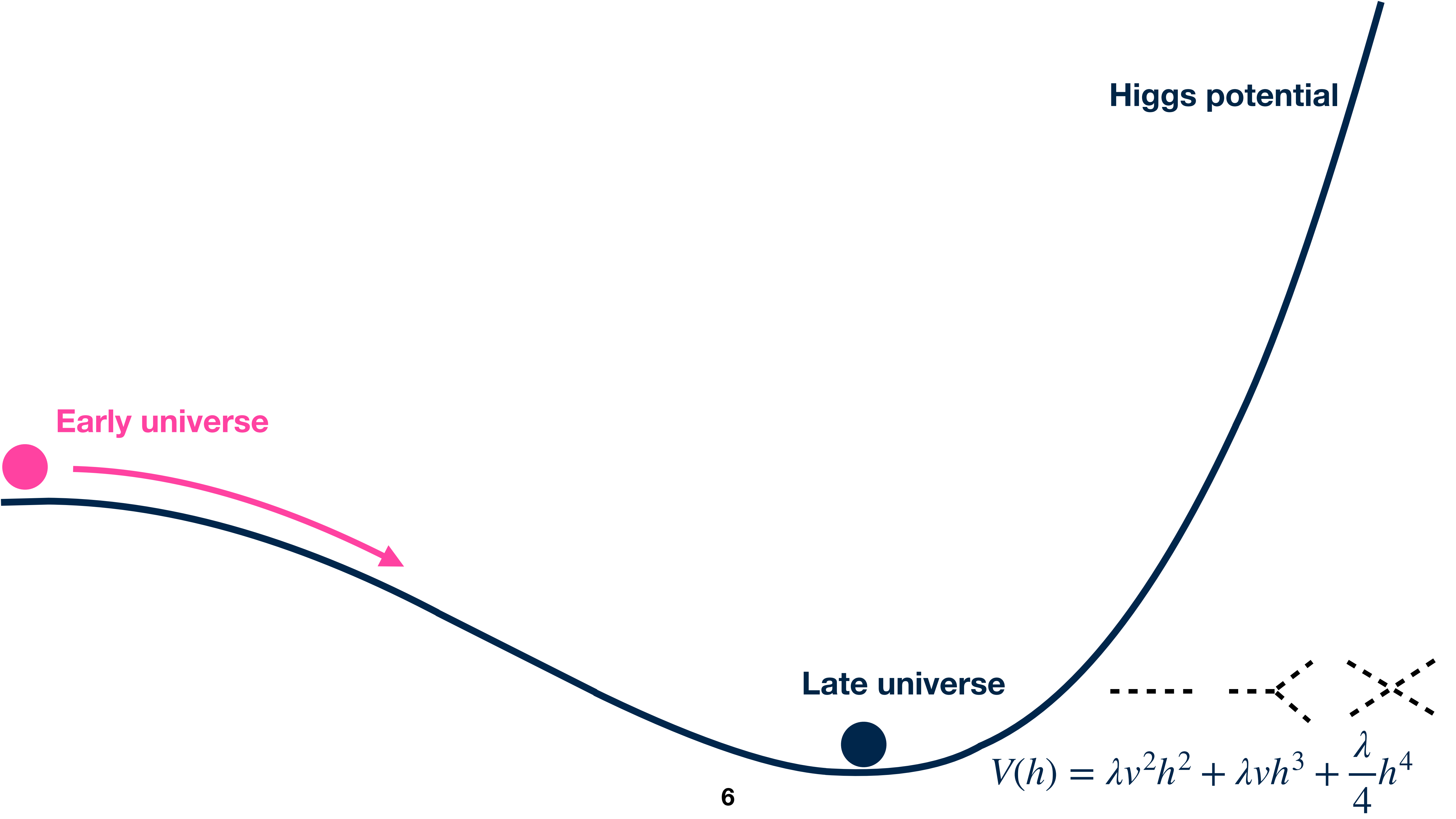
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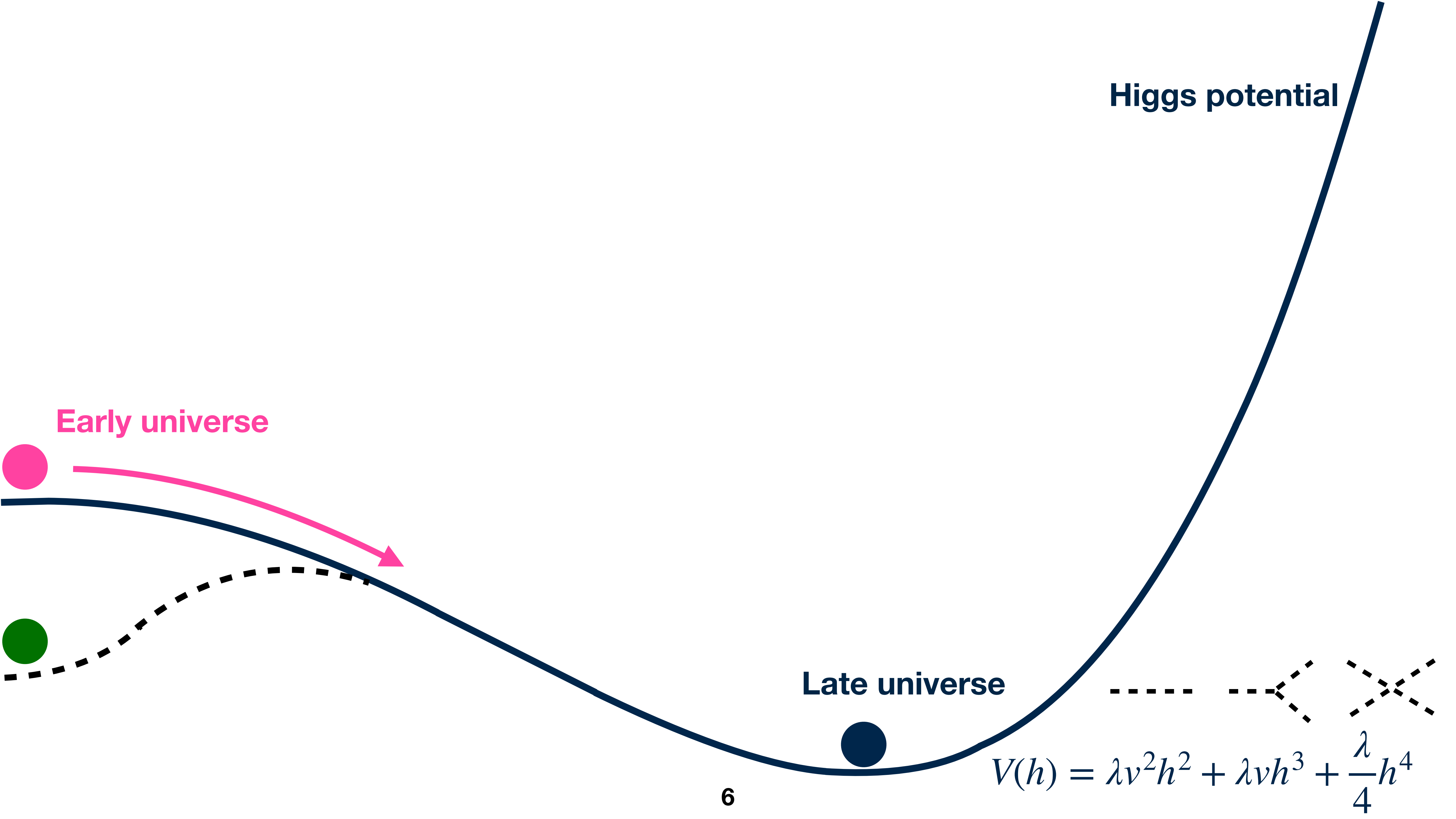
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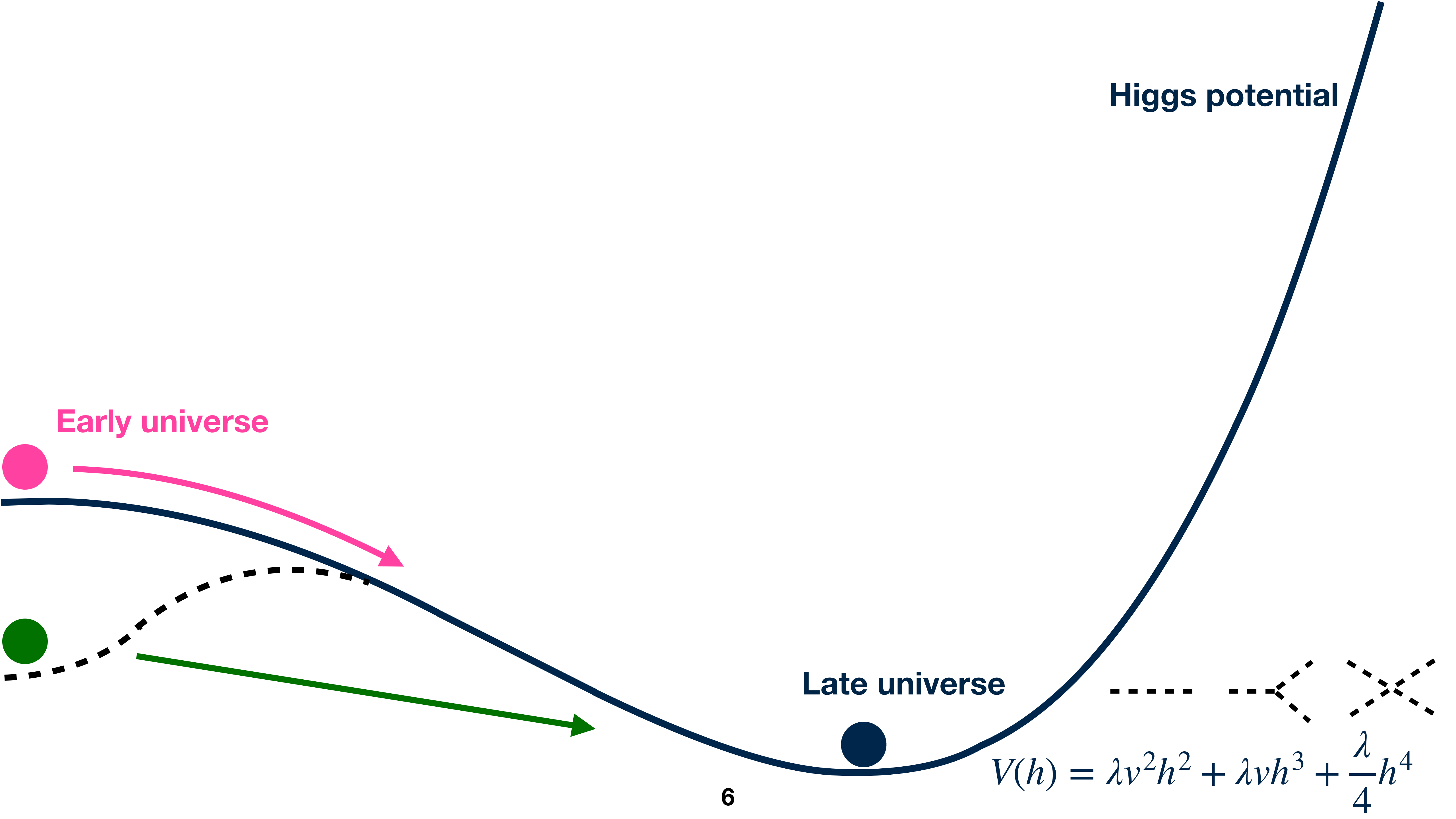


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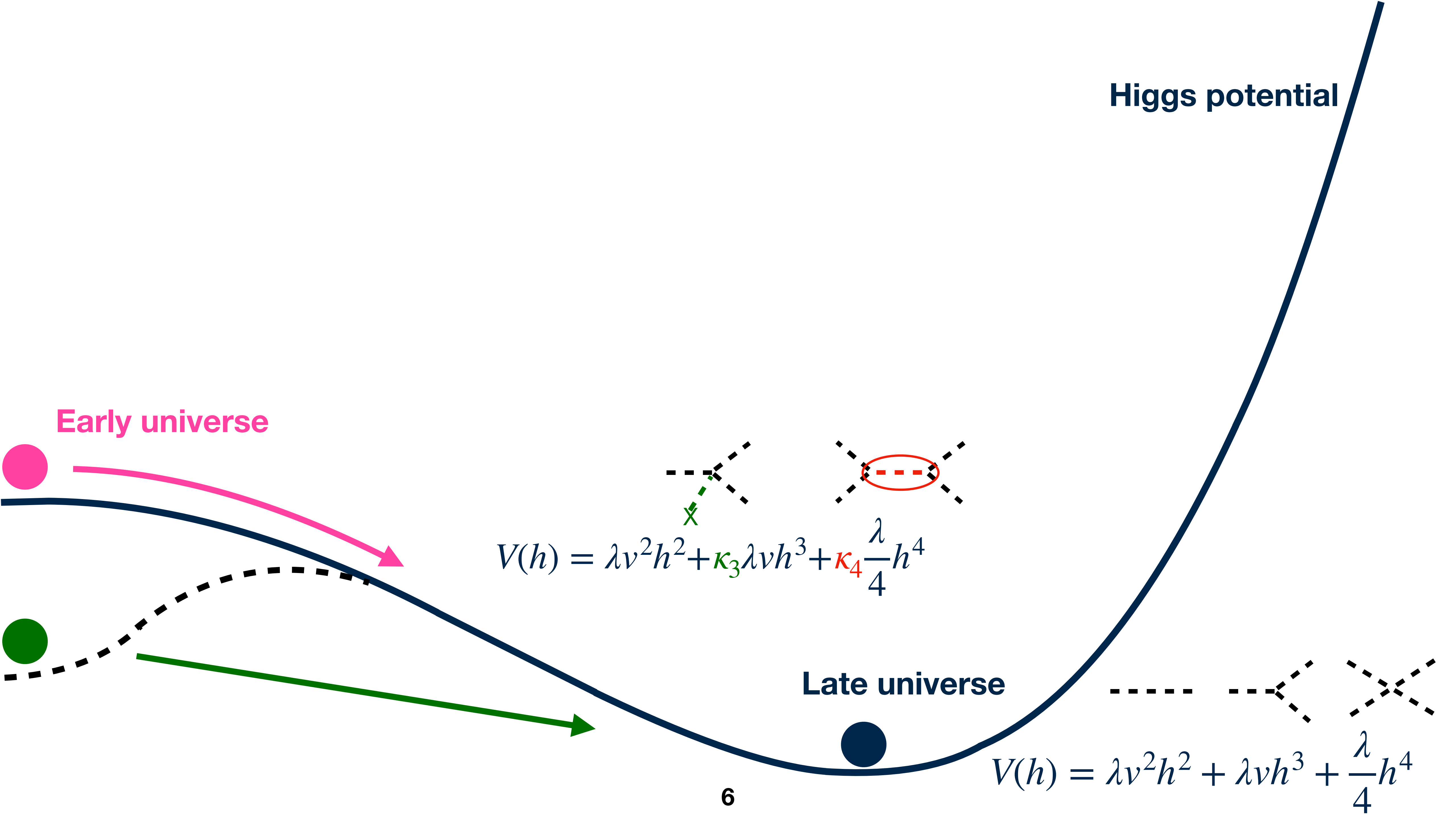
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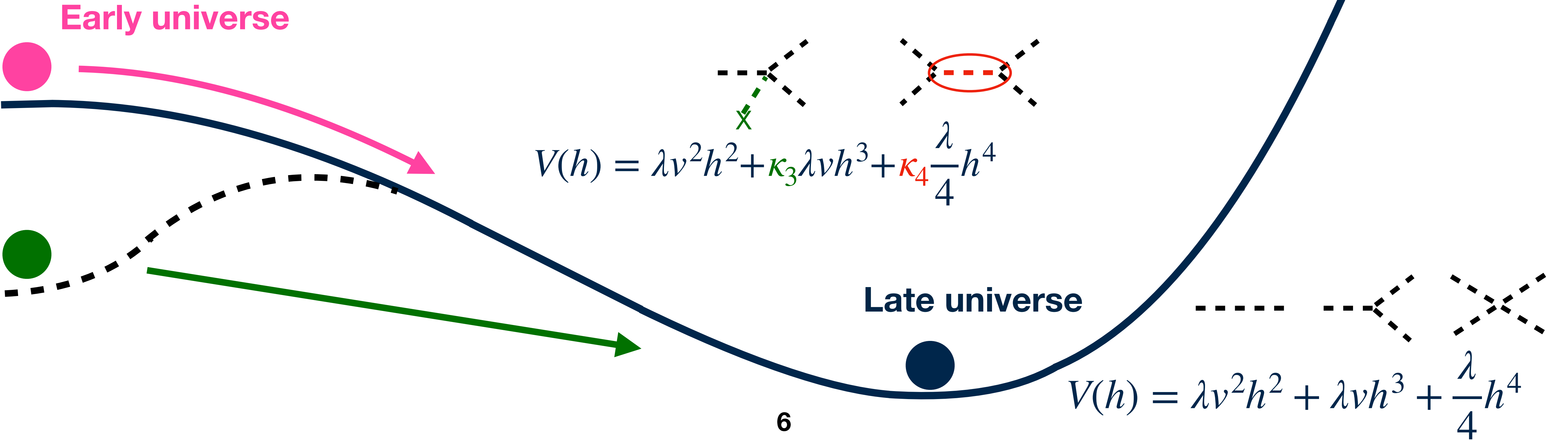


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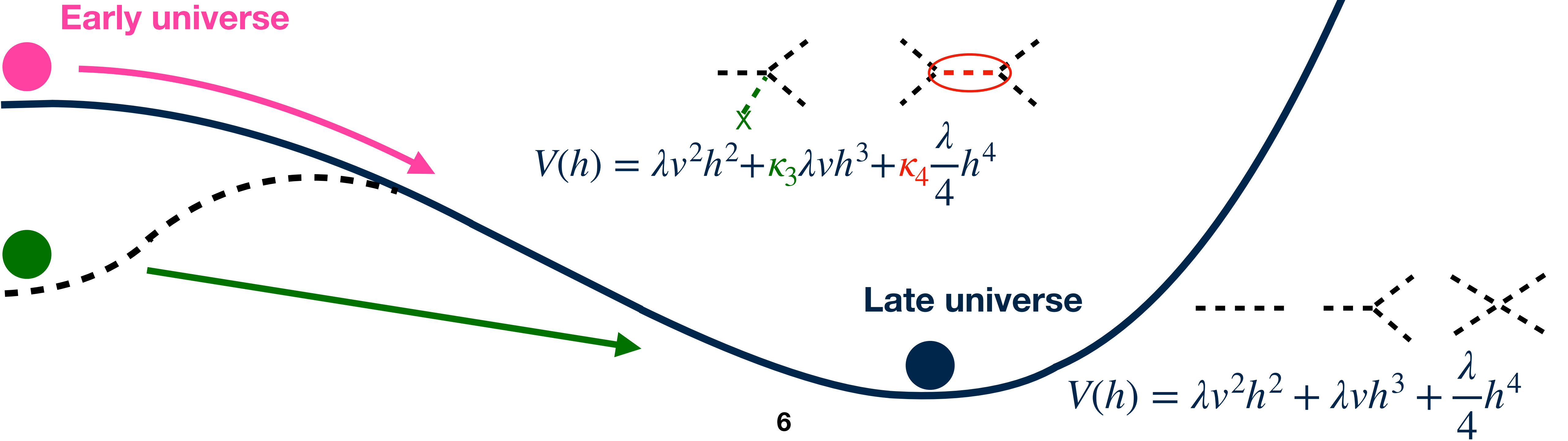
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Higgs potential



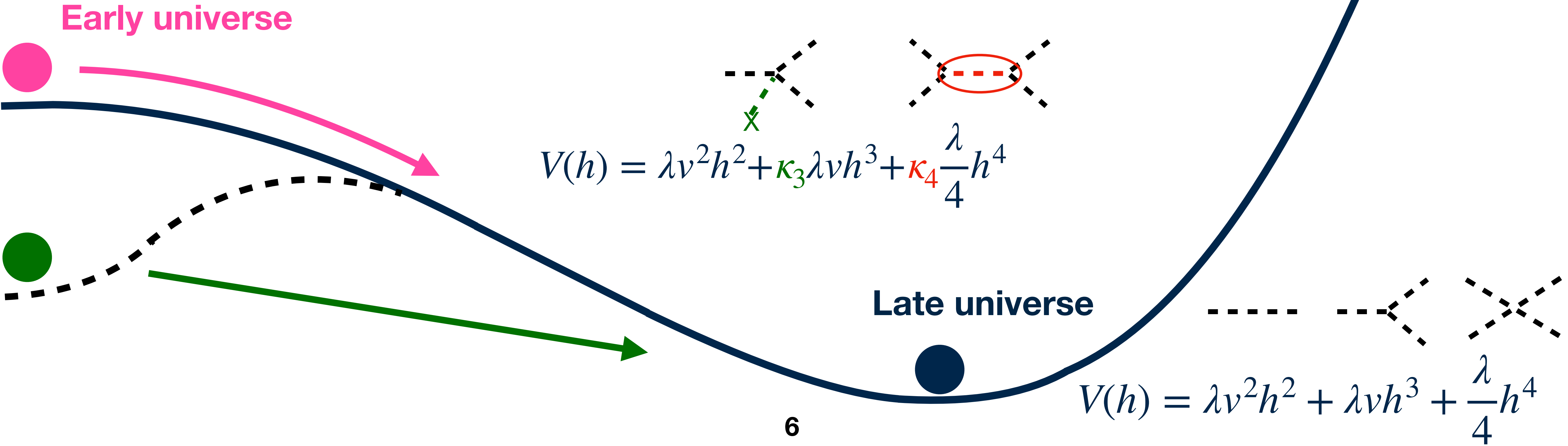
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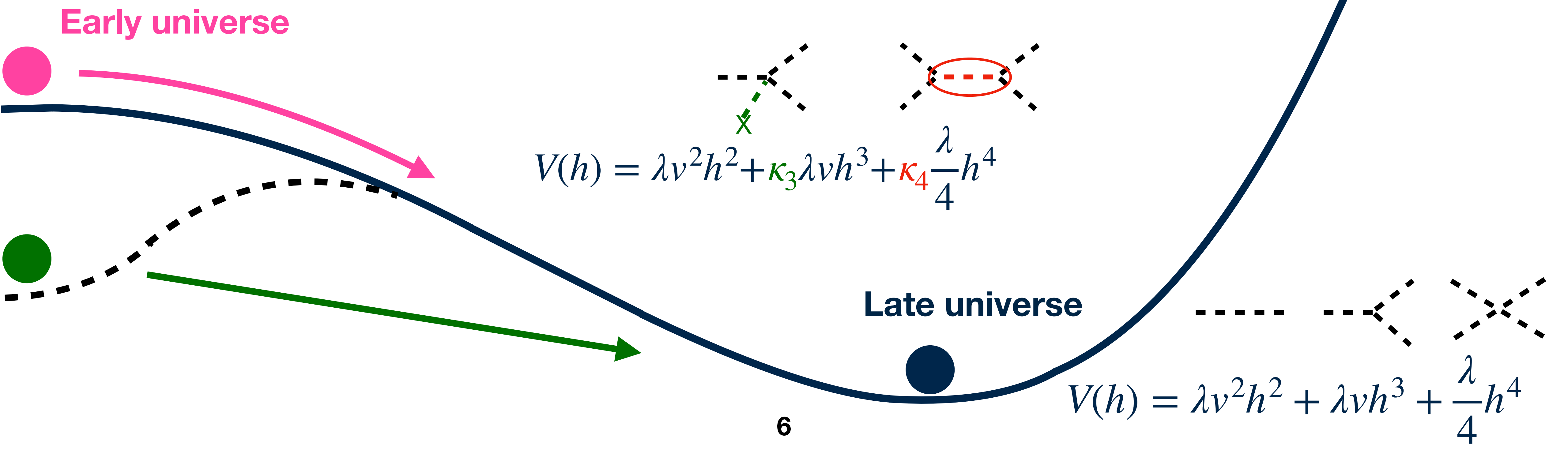
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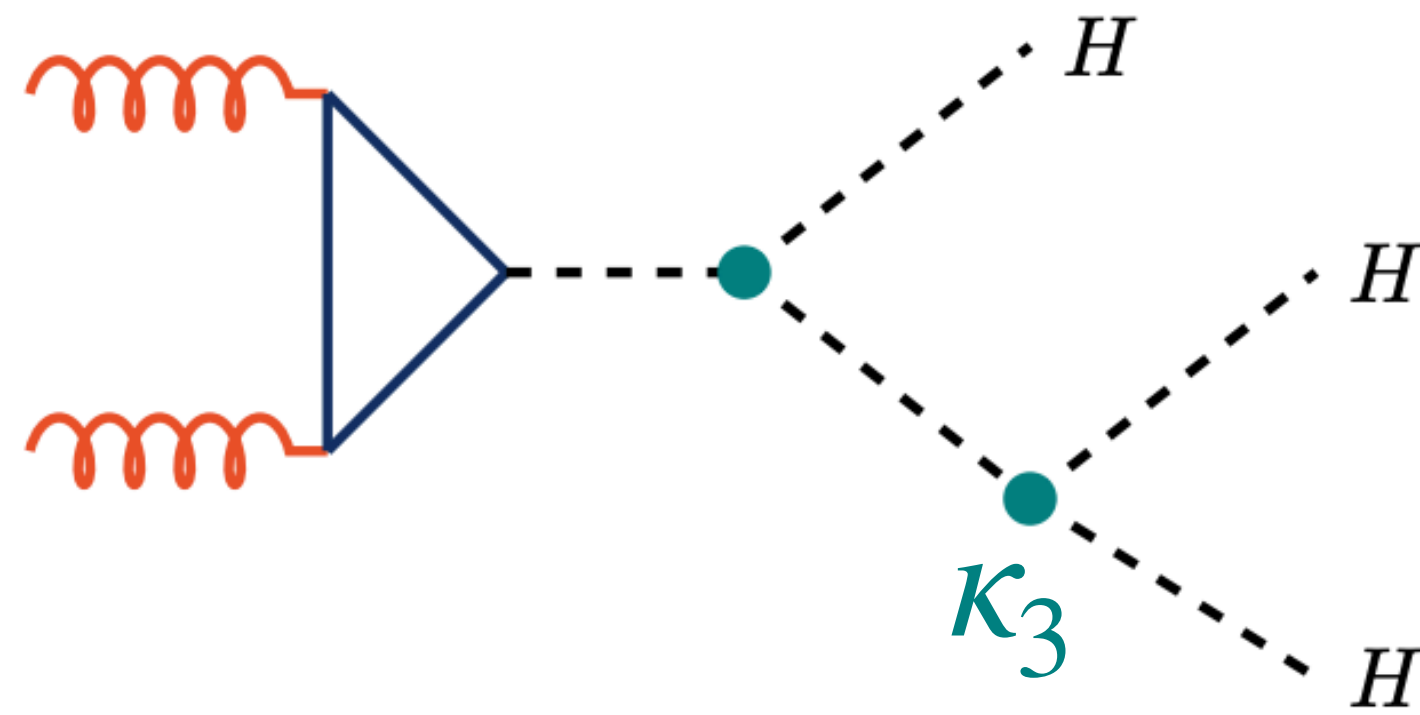


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- Effect on the late universe: modifications to Higgs self-coupling κ_3, κ_4

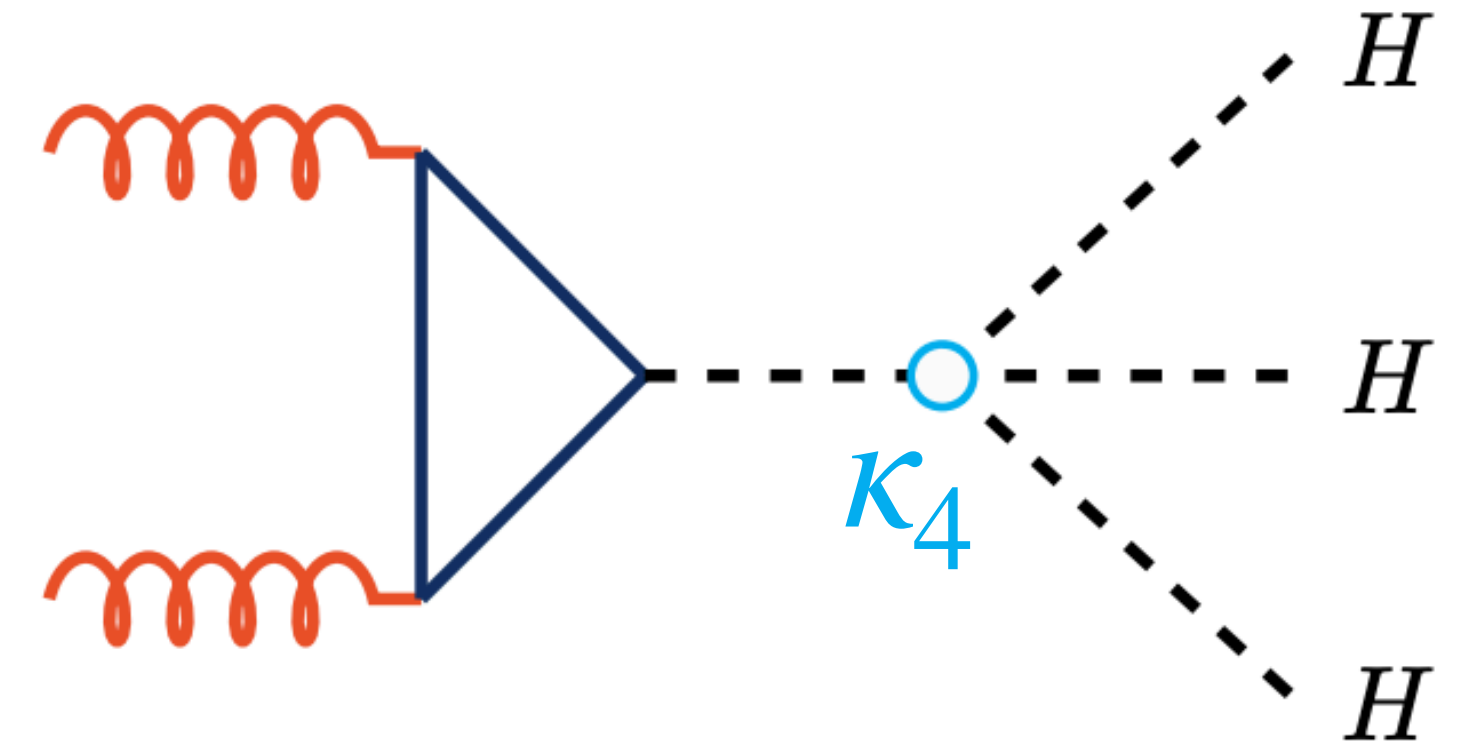
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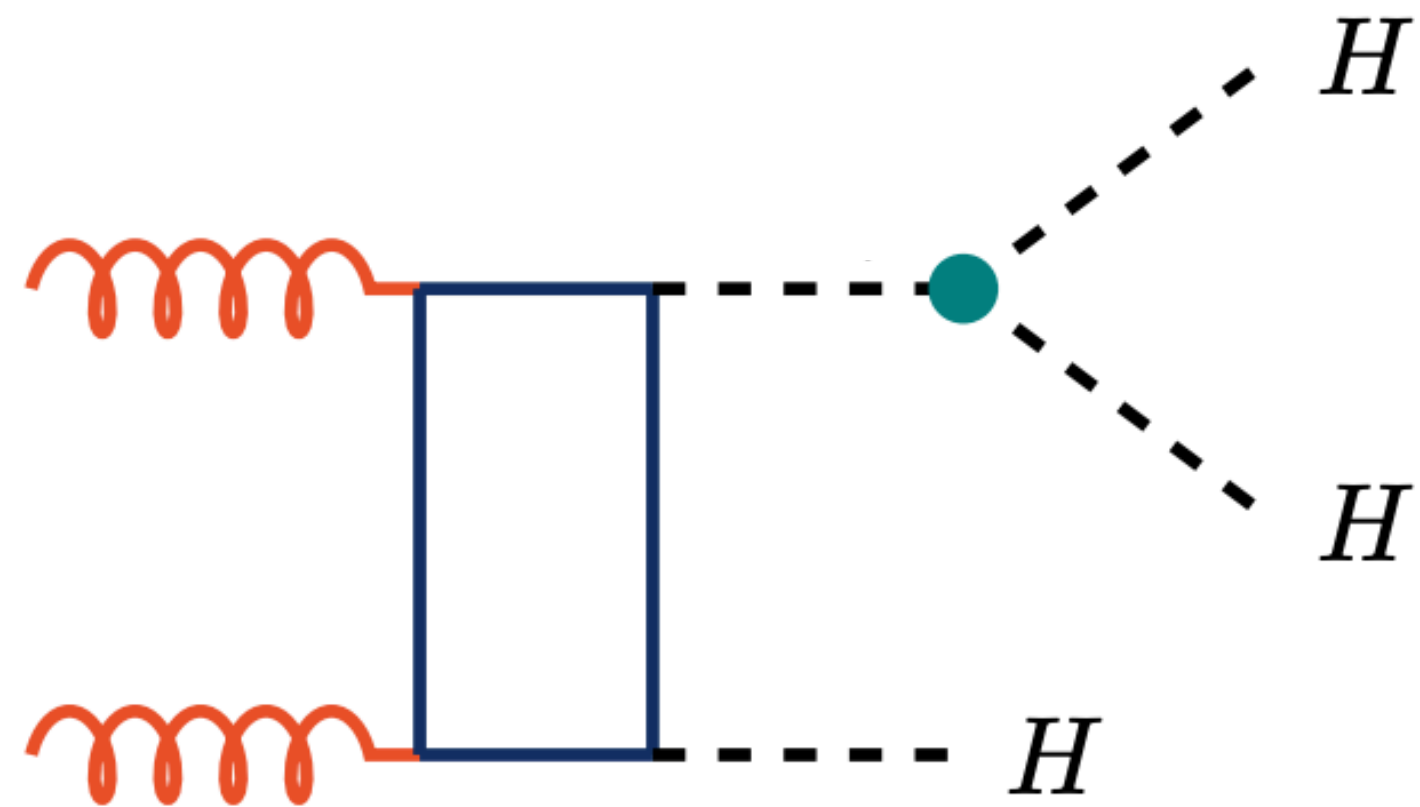
Why HHH?



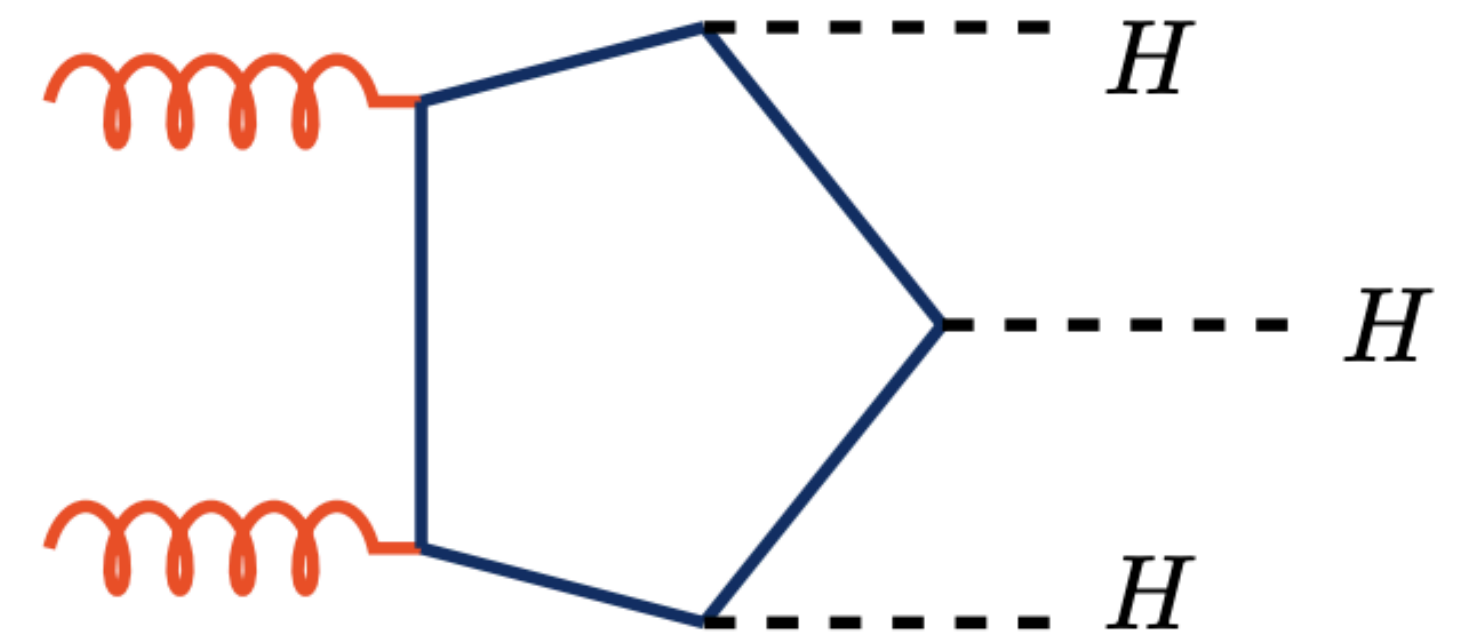
(a)



(b)



(c)

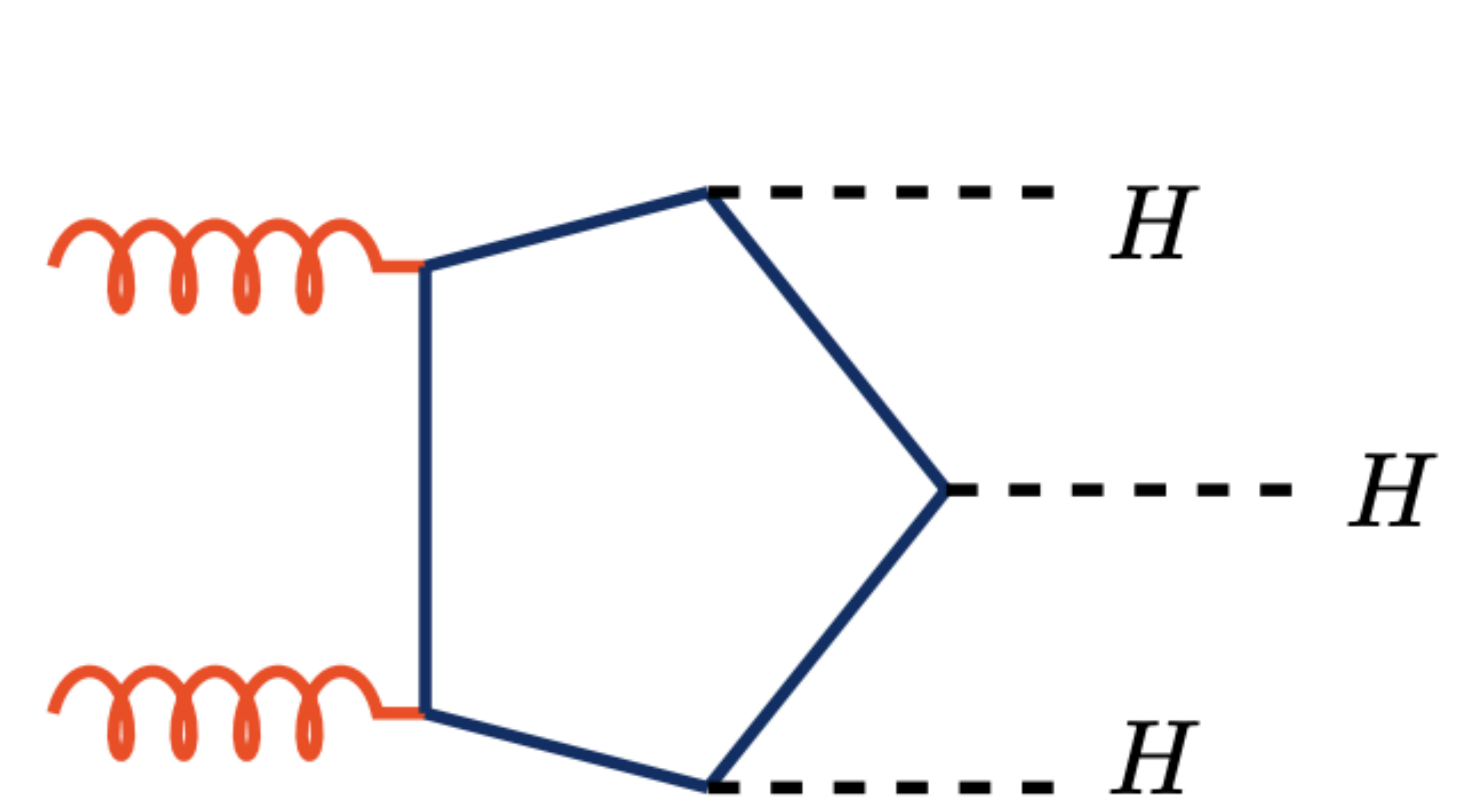
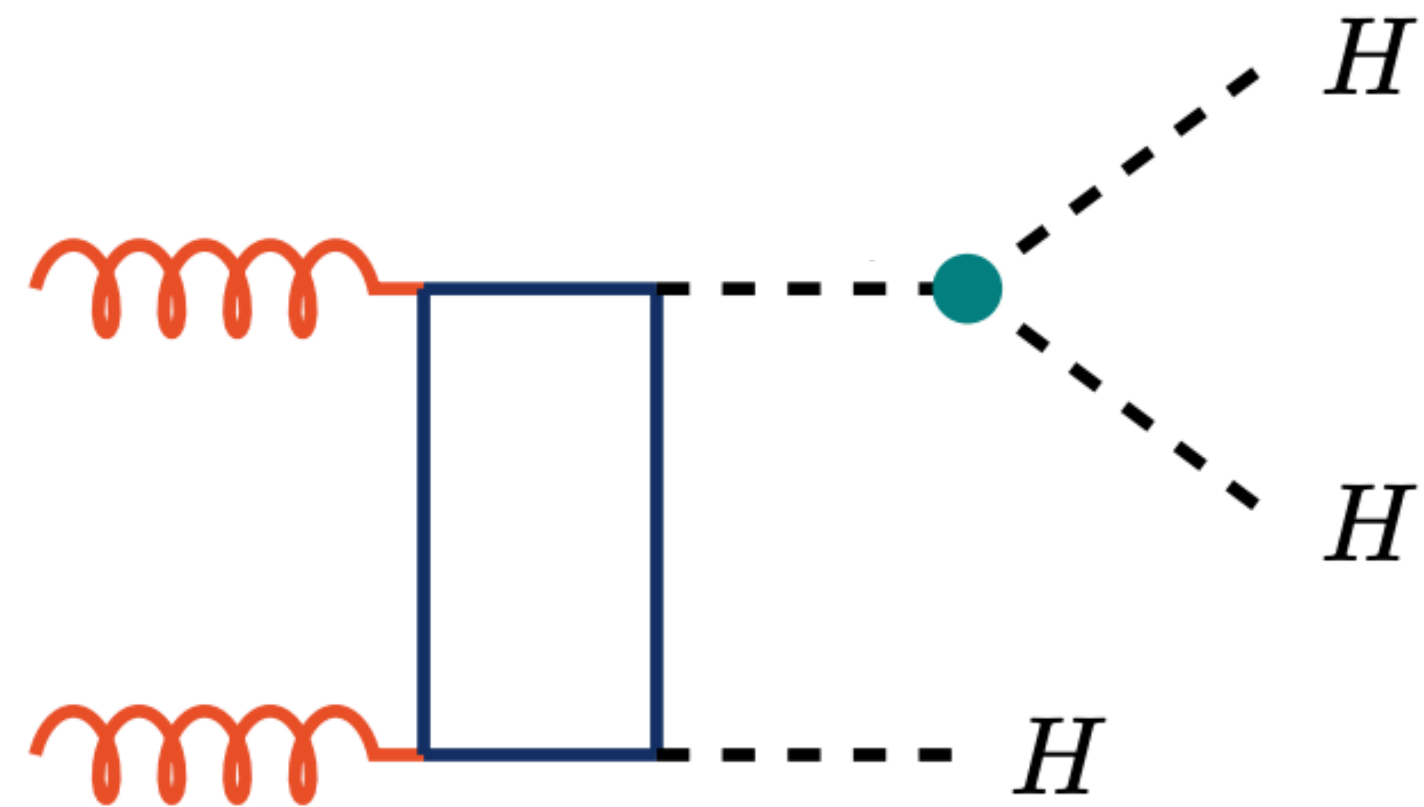
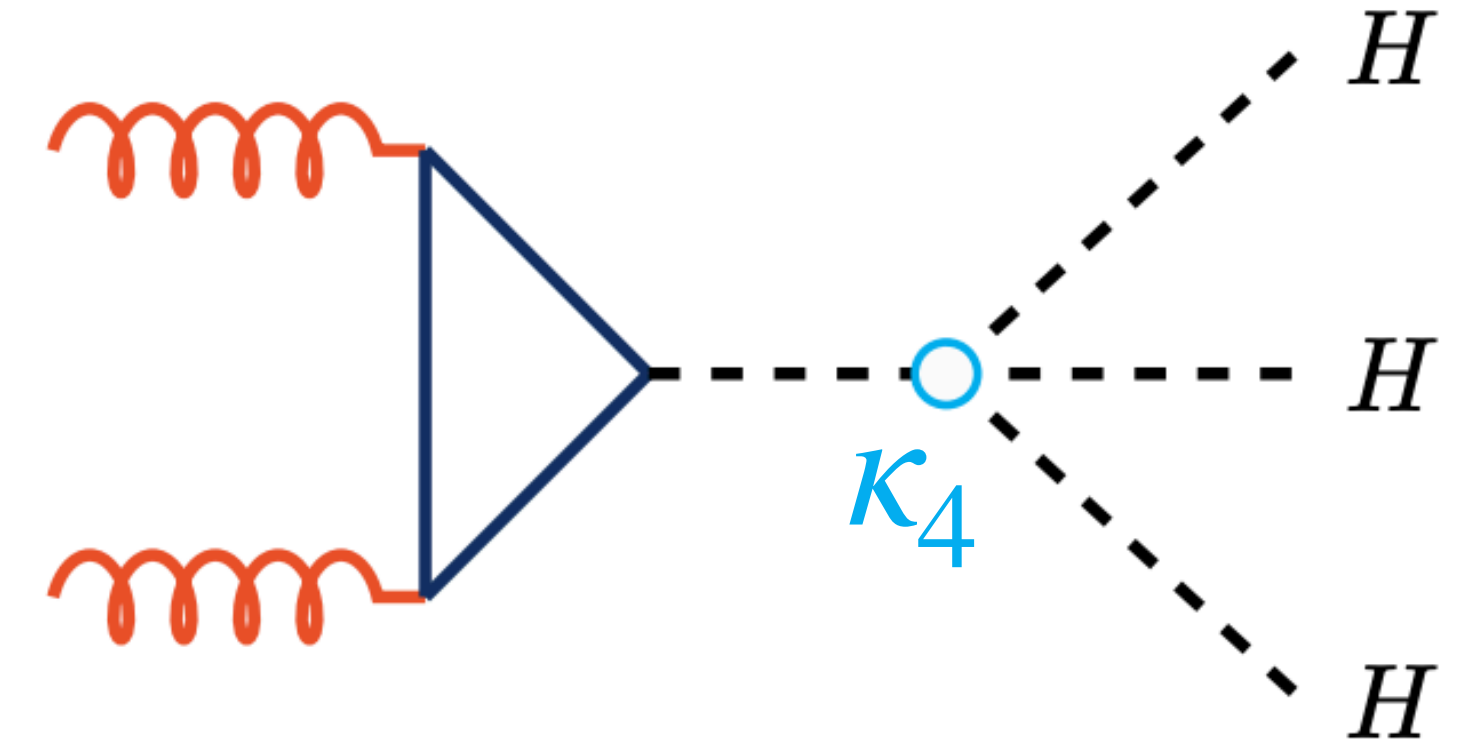
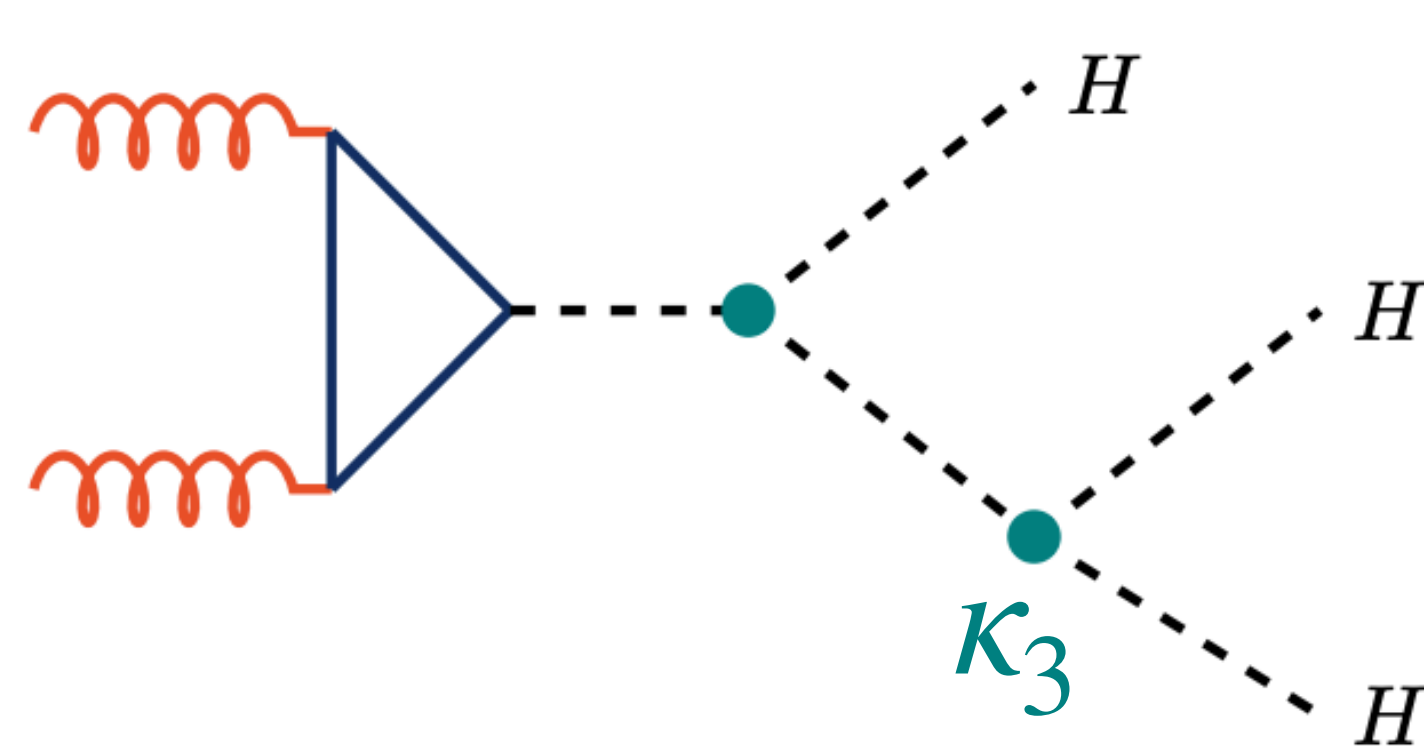


(d)

$$\sigma_{SM} = 0.079 \text{ fb}$$

Why HHH?

- Measuring the shape of the Higgs potential probes electroweak baryogenesis
 - Provides unique access to the quartic coupling κ_4
 - And can simultaneously limit κ_3



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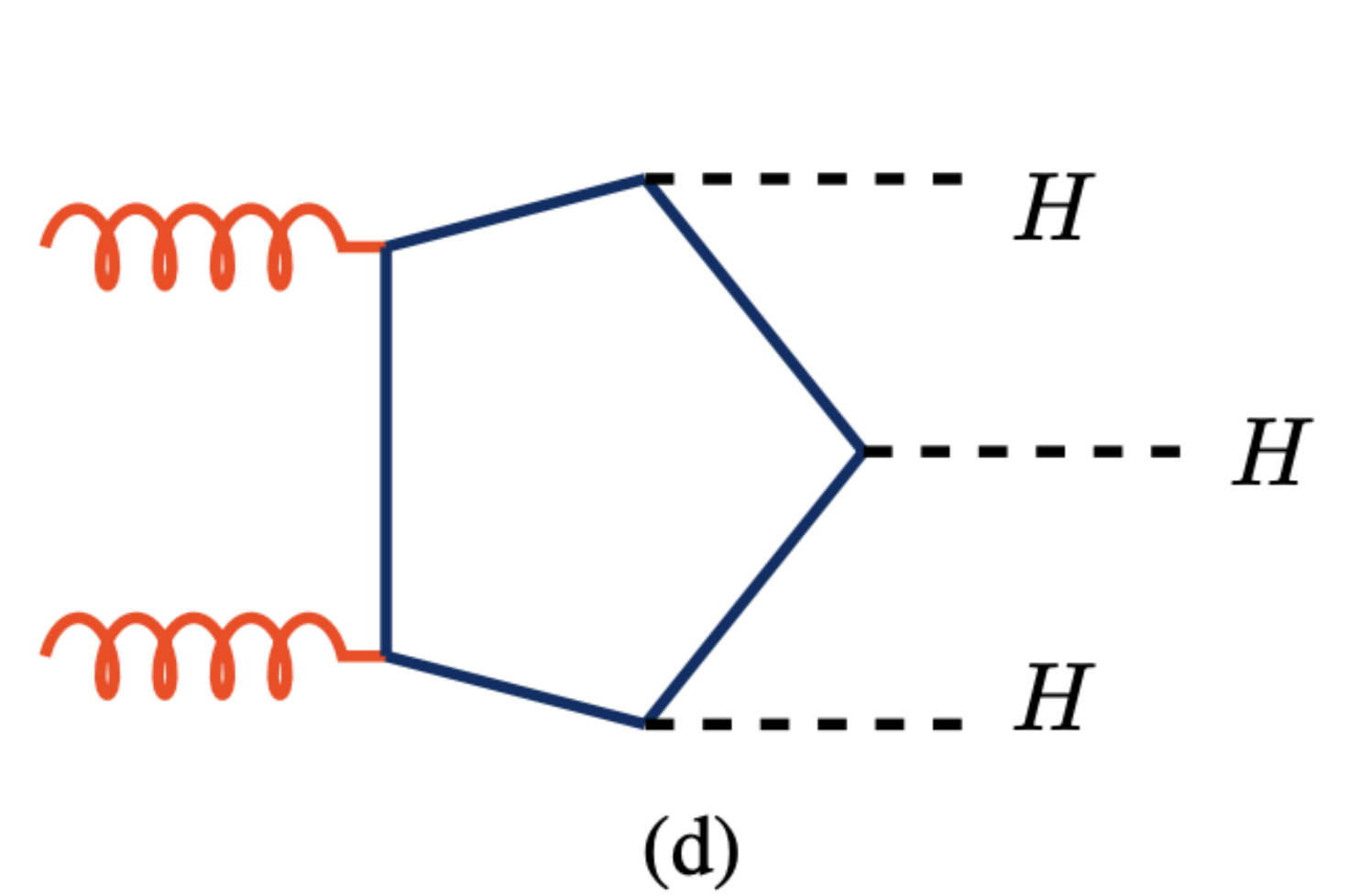
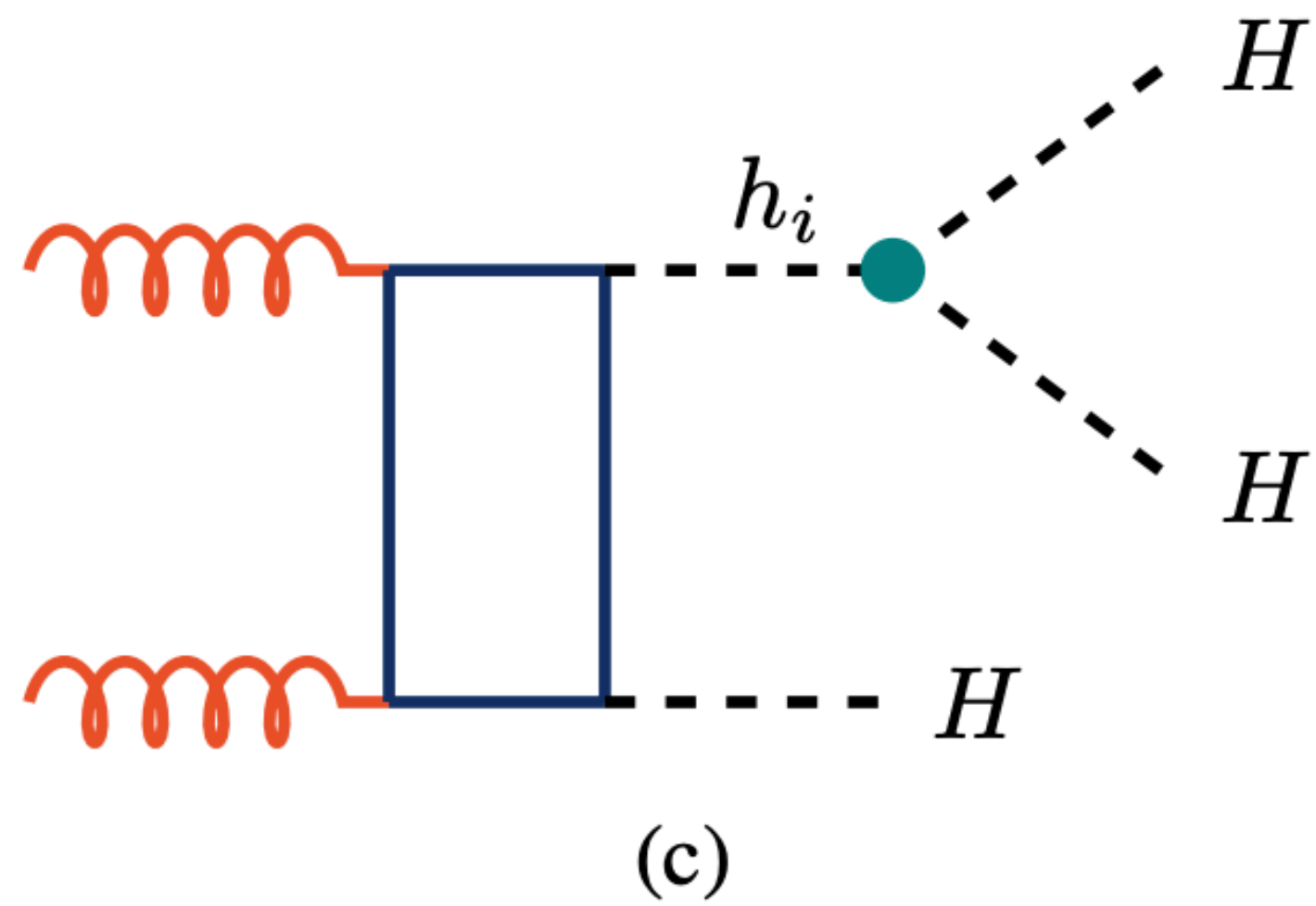
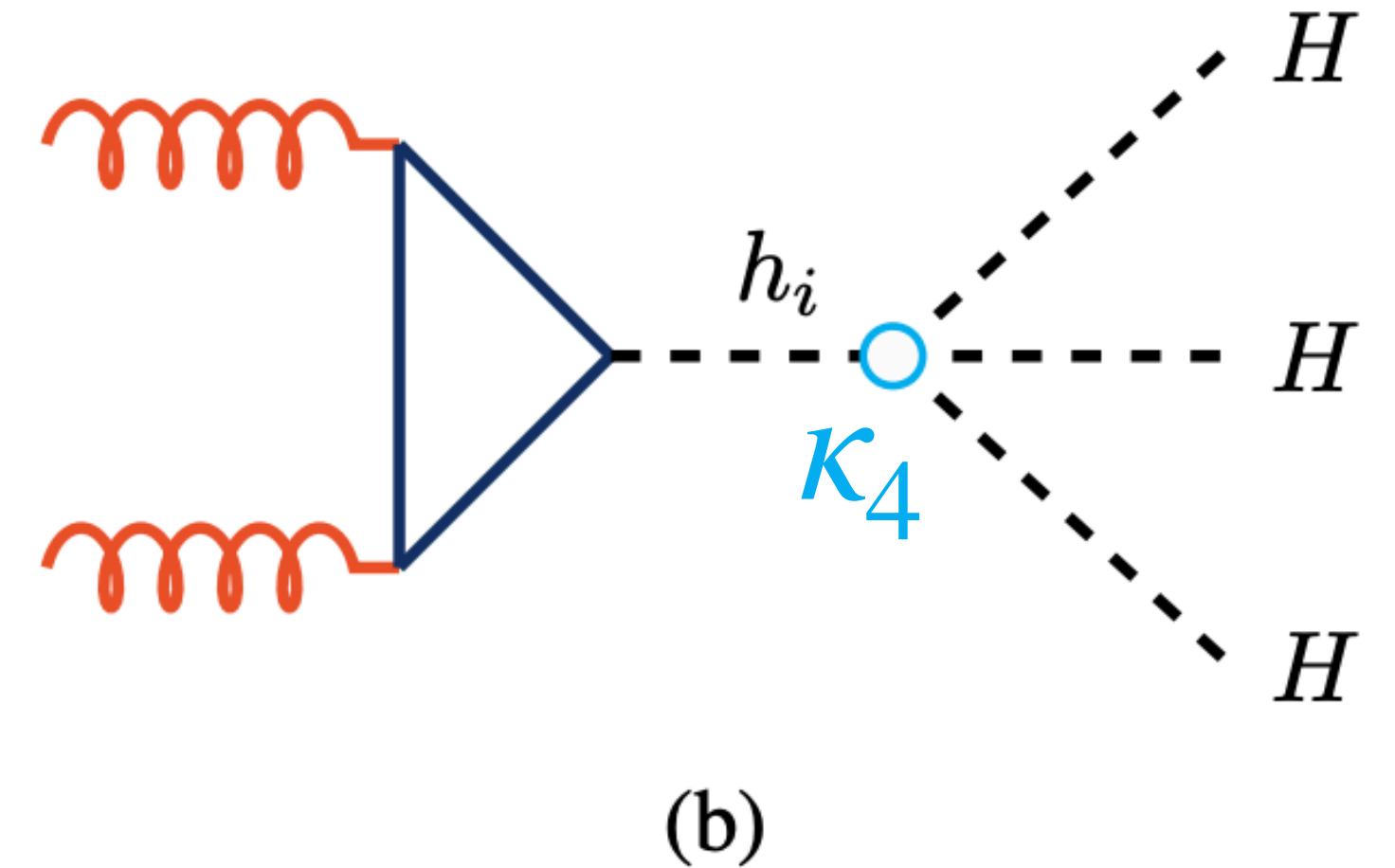
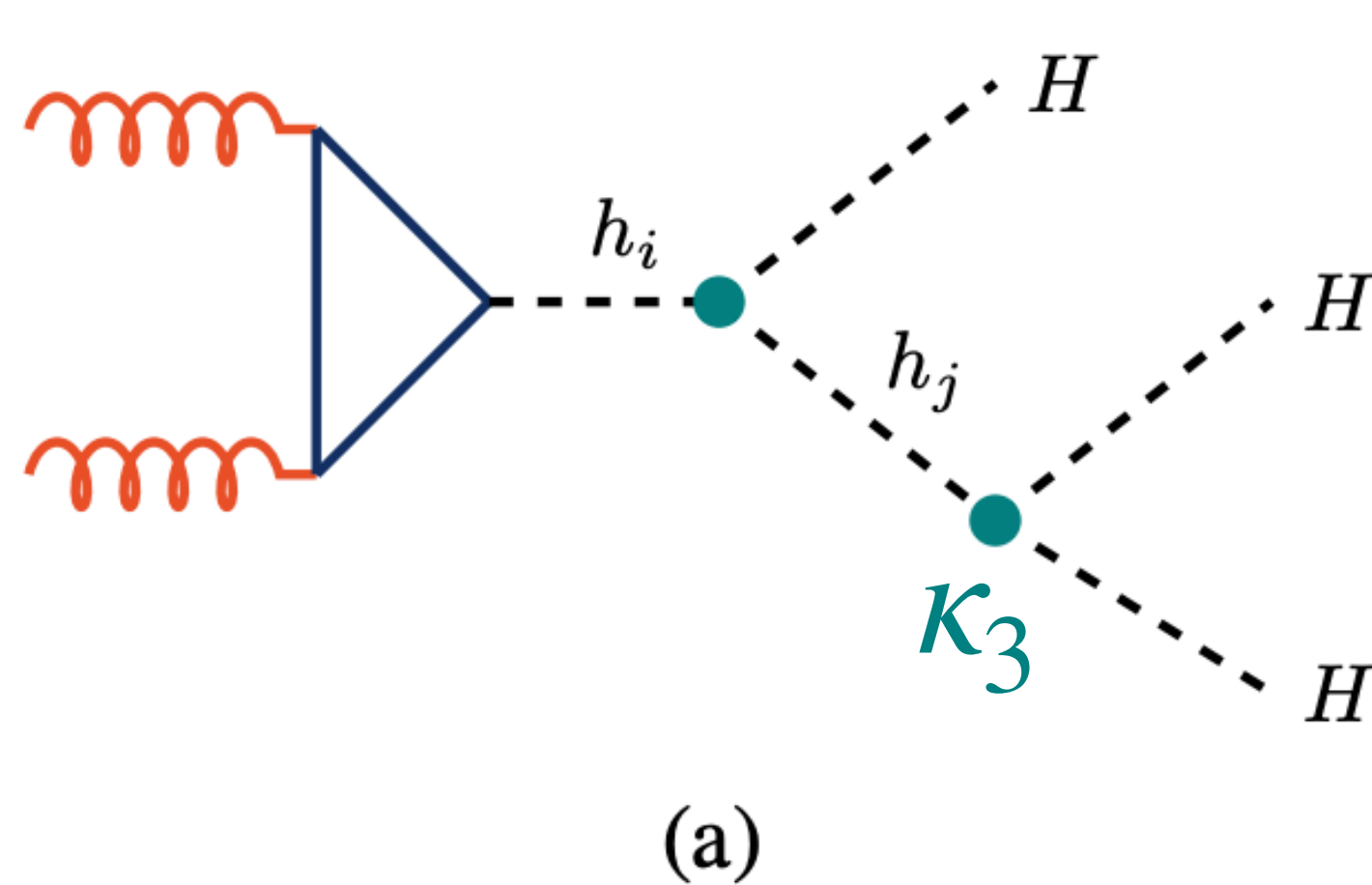
Why HHH?

- **Measuring the shape of the Higgs potential probes electroweak baryogenesis**

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- **BSM theories can predict enhanced HHH, even above HH production**

- h_i, h_j can be new scalars
- HHH production can be 100s fb when new scalars are present: [\[1908.08554\]](#), [\[2211.10557\]](#)
- Additional scalars can provide sources of dark matter, CP violation: [\[2202.02954\]](#)
 - CP violation in the scalar potential and/or between scalars and DM.



$$\sigma_{BSM} \approx 100 \text{ fb}$$

TRSM model



TRSM model

- Two Real Singlet Model: [\[1908.08554\]](#)



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- **Add two new real singlets to SM:**
 - Both can have a VeV, which leads to HHH

$$\Phi = \begin{pmatrix} 0 \\ \frac{\phi_h + v}{\sqrt{2}} \end{pmatrix}, \quad S = \frac{\phi_S + v_S}{\sqrt{2}}, \quad X = \frac{\phi_X + v_X}{\sqrt{2}}$$

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- 3 mixing angles
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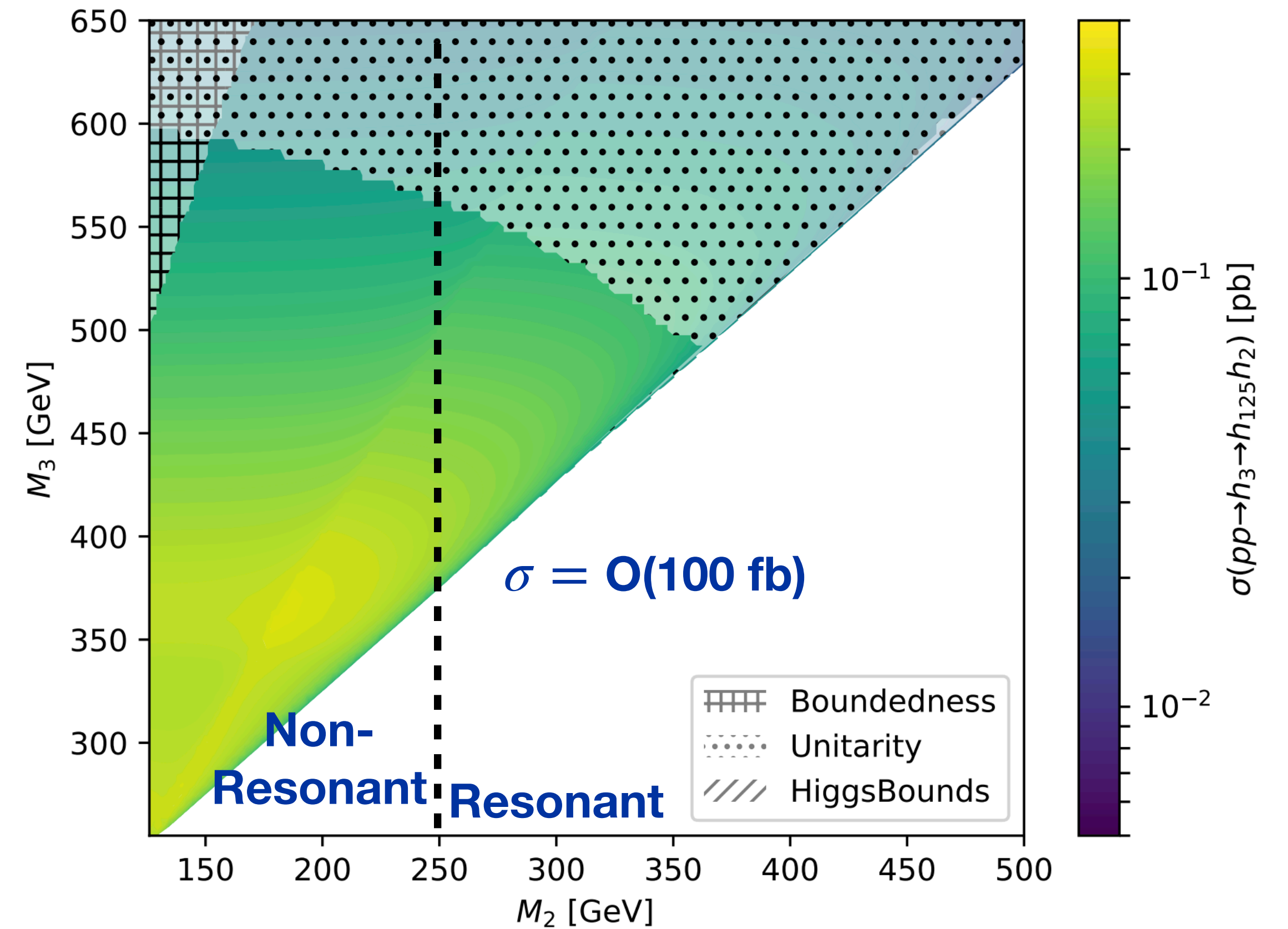
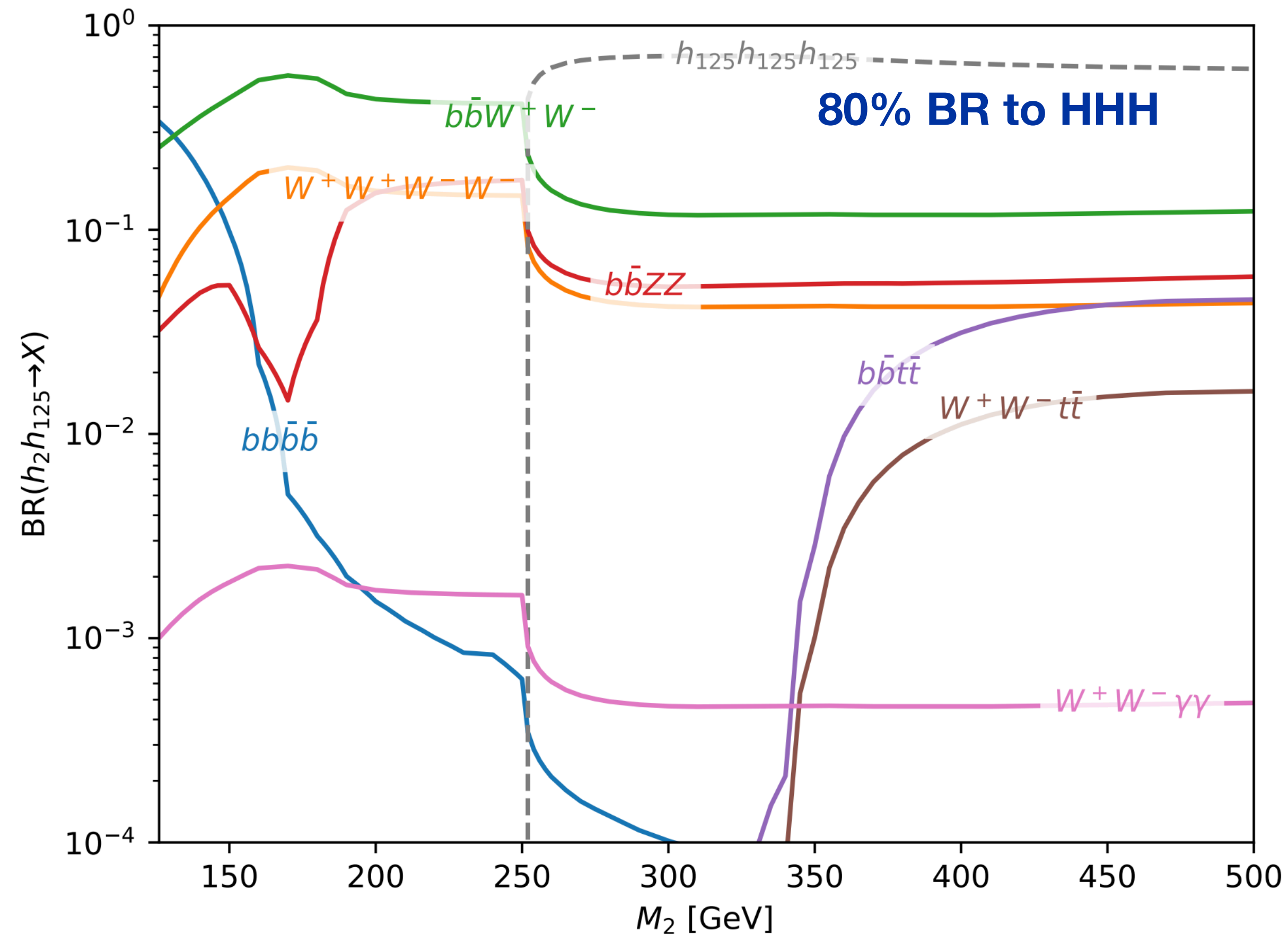
- **Scalar couplings, mixing from V:**

- Remember, any can have VeV

$$V = \mu_\Phi^2 \Phi^\dagger \Phi + \lambda_\Phi (\Phi^\dagger \Phi)^2 + \mu_S^2 S^2 + \lambda_S S^4 + \mu_X^2 X^2 + \lambda_X X^4 \\ + \lambda_{\Phi S} \Phi^\dagger \Phi S^2 + \lambda_{\Phi X} \Phi^\dagger \Phi X^2 + \lambda_{SX} S^2 X^2.$$

BSM searches

- In a **most favorable** scenario, $X \rightarrow SH$ is up to ~ 150 of fb and $Br(SH \rightarrow HHH)$ is up to 80%
- A large HHH signal which could be visible in Run 2 at the LHC



[1908.08554]



DM-CPV model

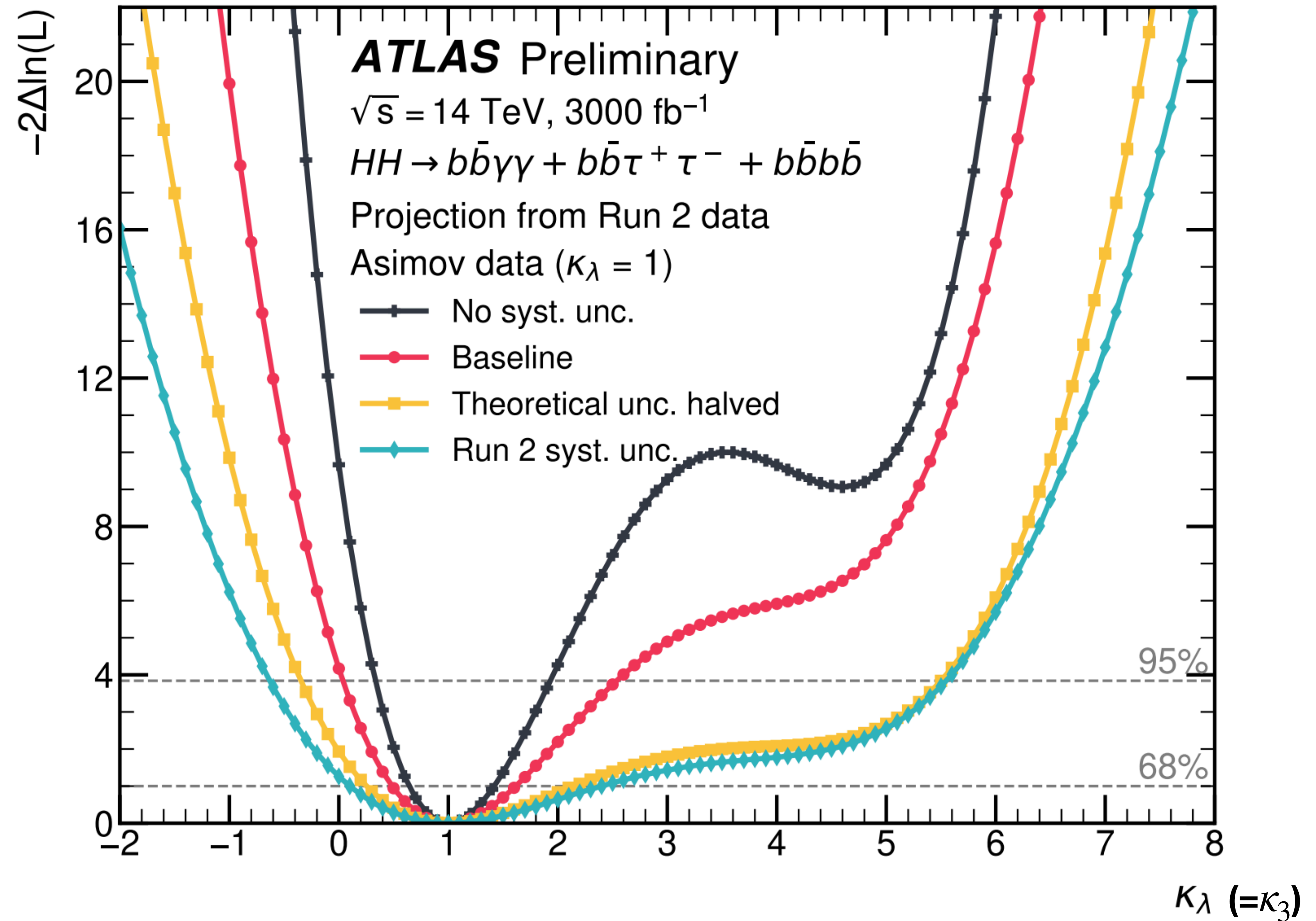
- A simple model of dark matter and CP violation: [\[2202.02954\]](#)
- **Similar phenomenology due to a complex scalar instead of 2 real scalars**
 - Also introduces a dark vector-like fermion χ
- **More free parameters and Z_2 symmetries are not imposed**
- **Produces cross sections for HHH up to 55fb. In most favorable scenarios kinematics are nearly identical to TRSM model.**
 - ATLAS tested generation of TRSM and DM-CPV signals and differences are negligible.



Measuring the Higgs potential

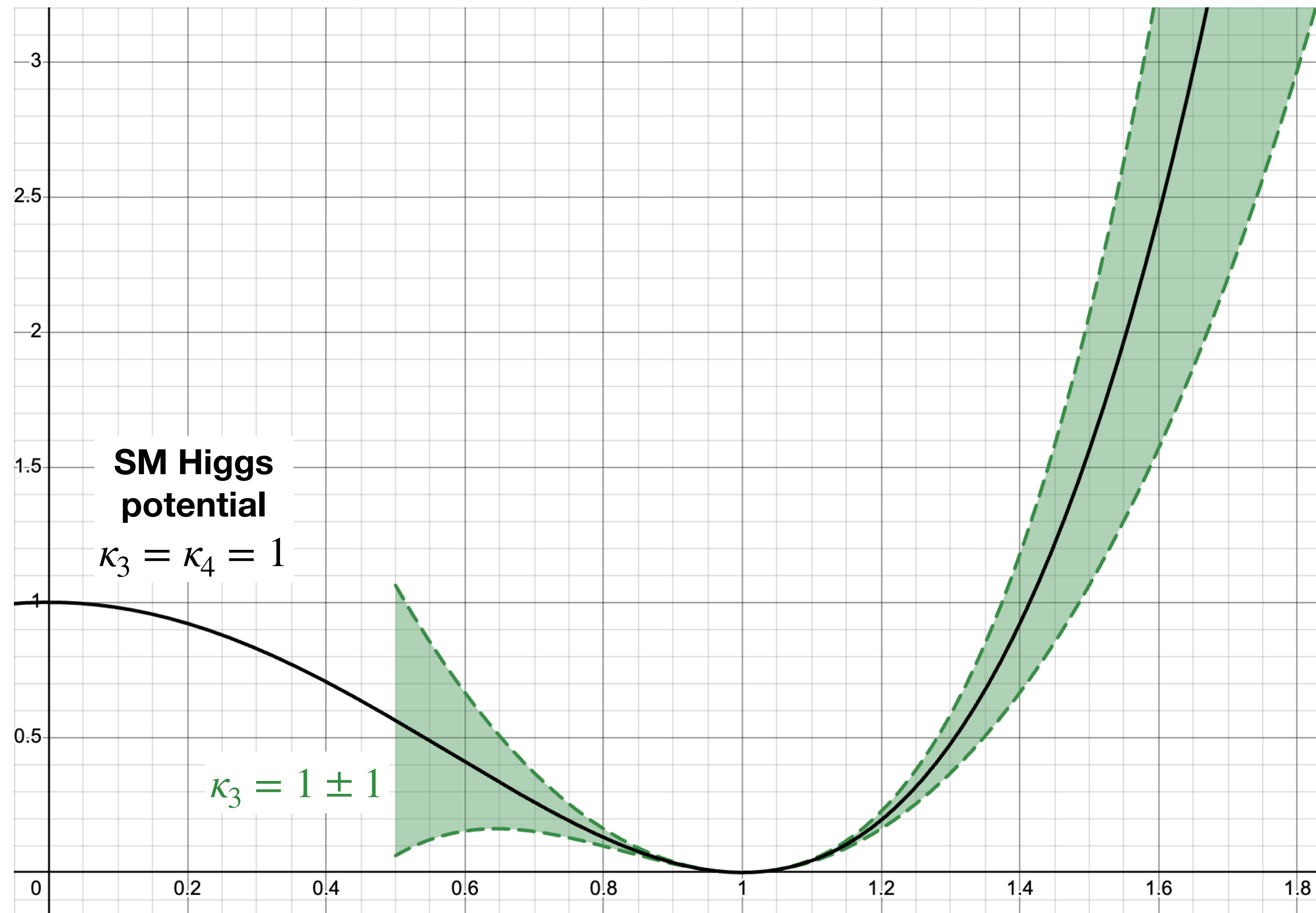
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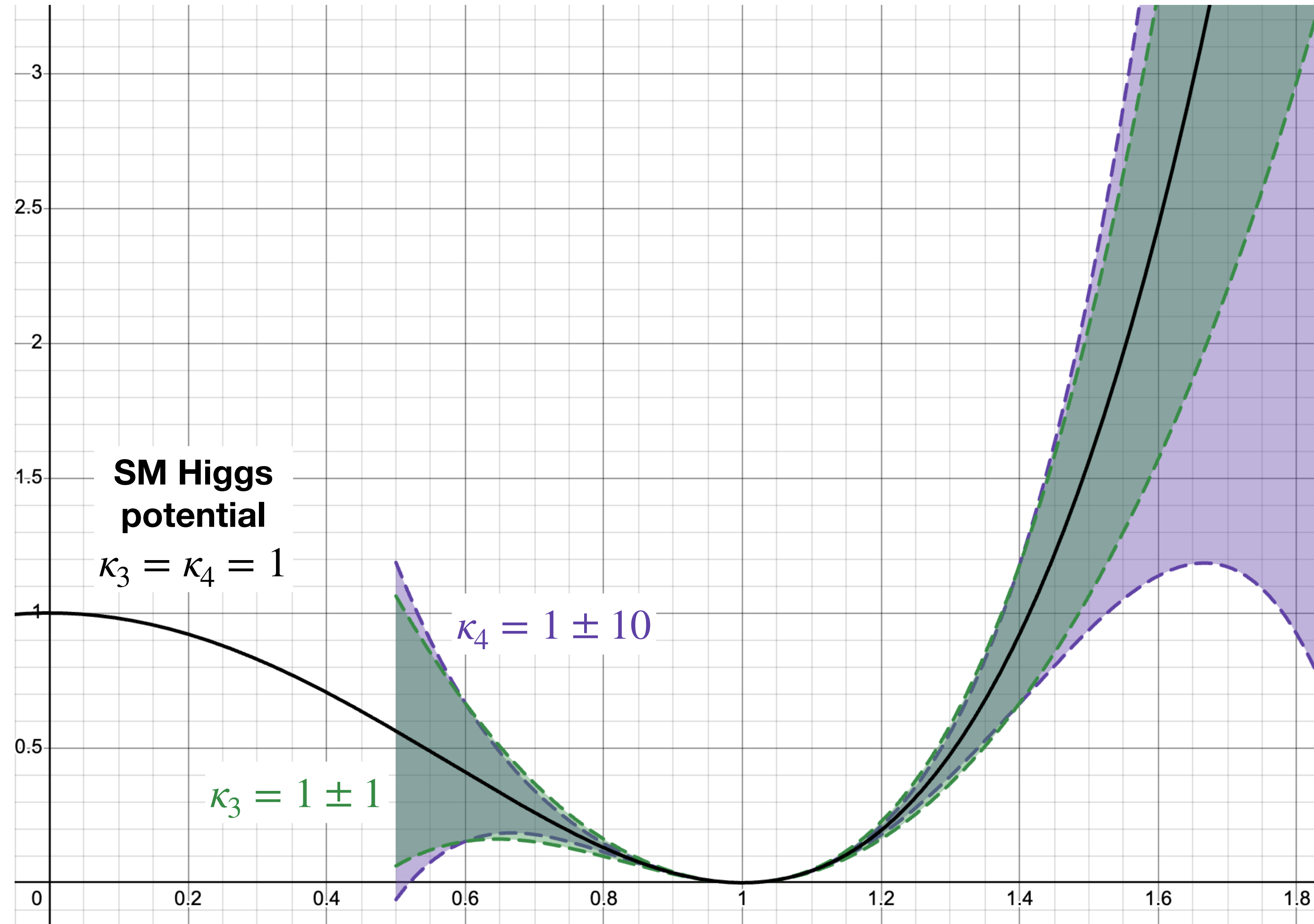
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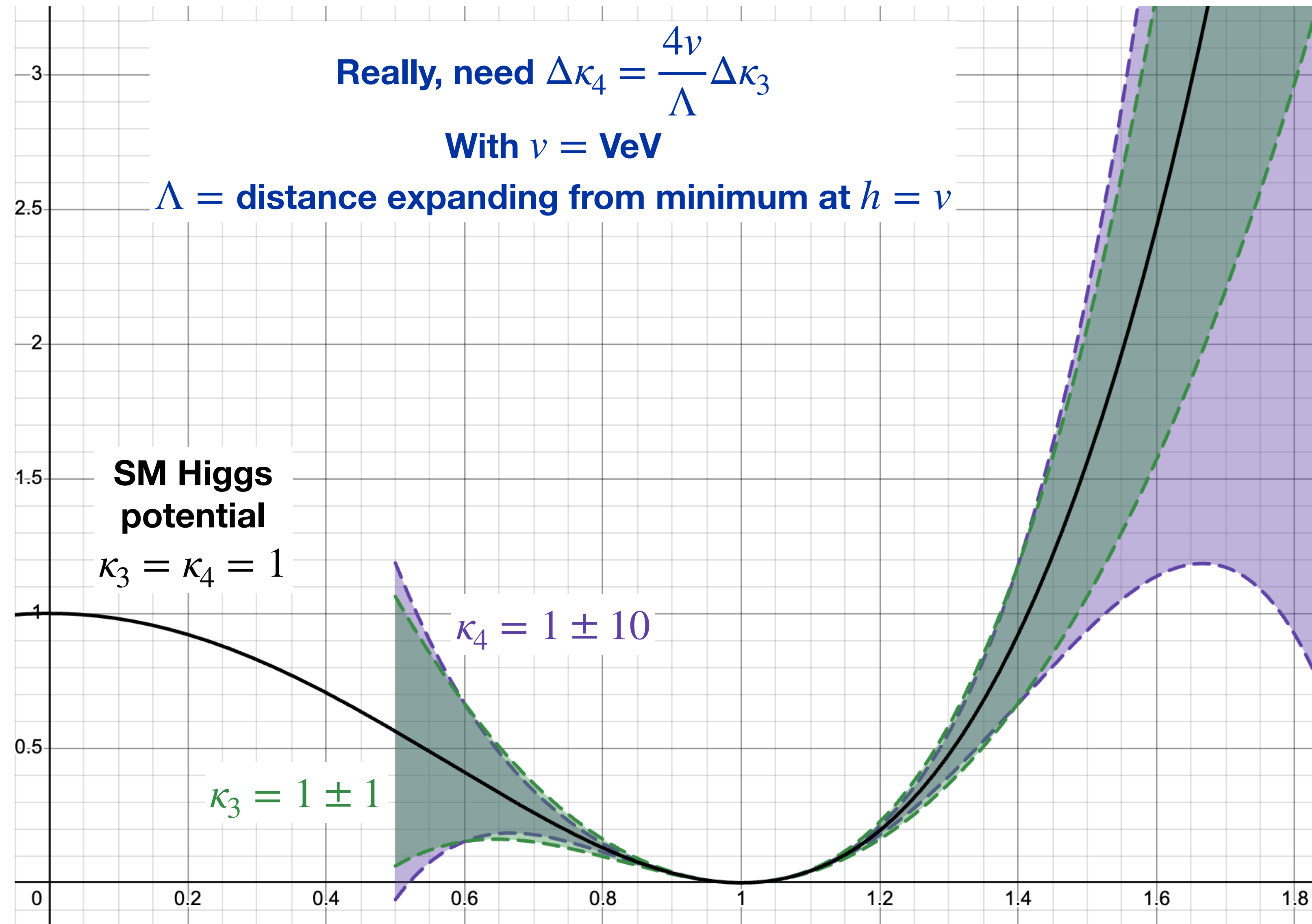
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- Luckily, the shape of the potential is less sensitive to κ_4 , at least near the minimum.
- To obtain similar constraints on the potential shape, need $\kappa_4 = 1 \pm 10$



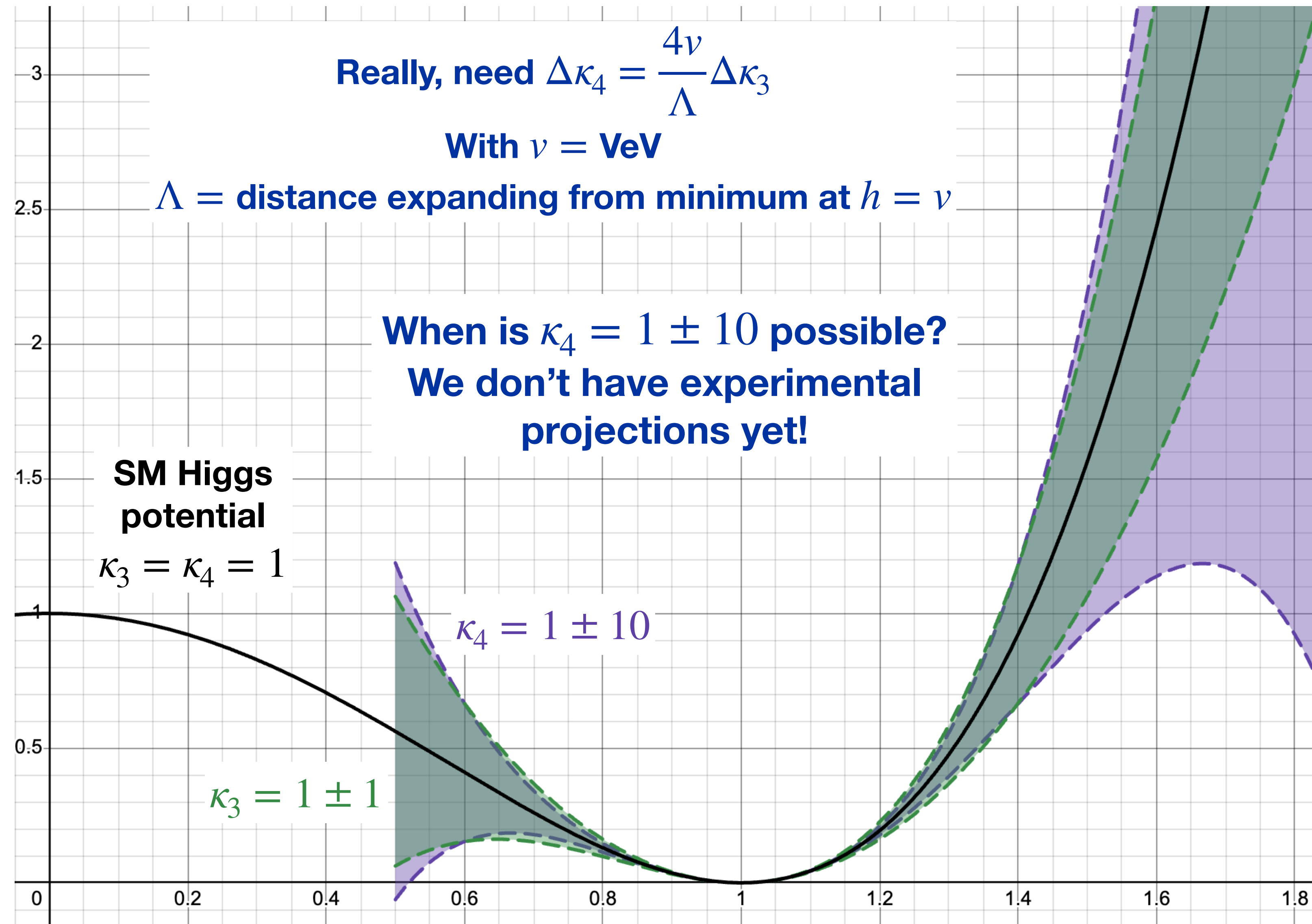
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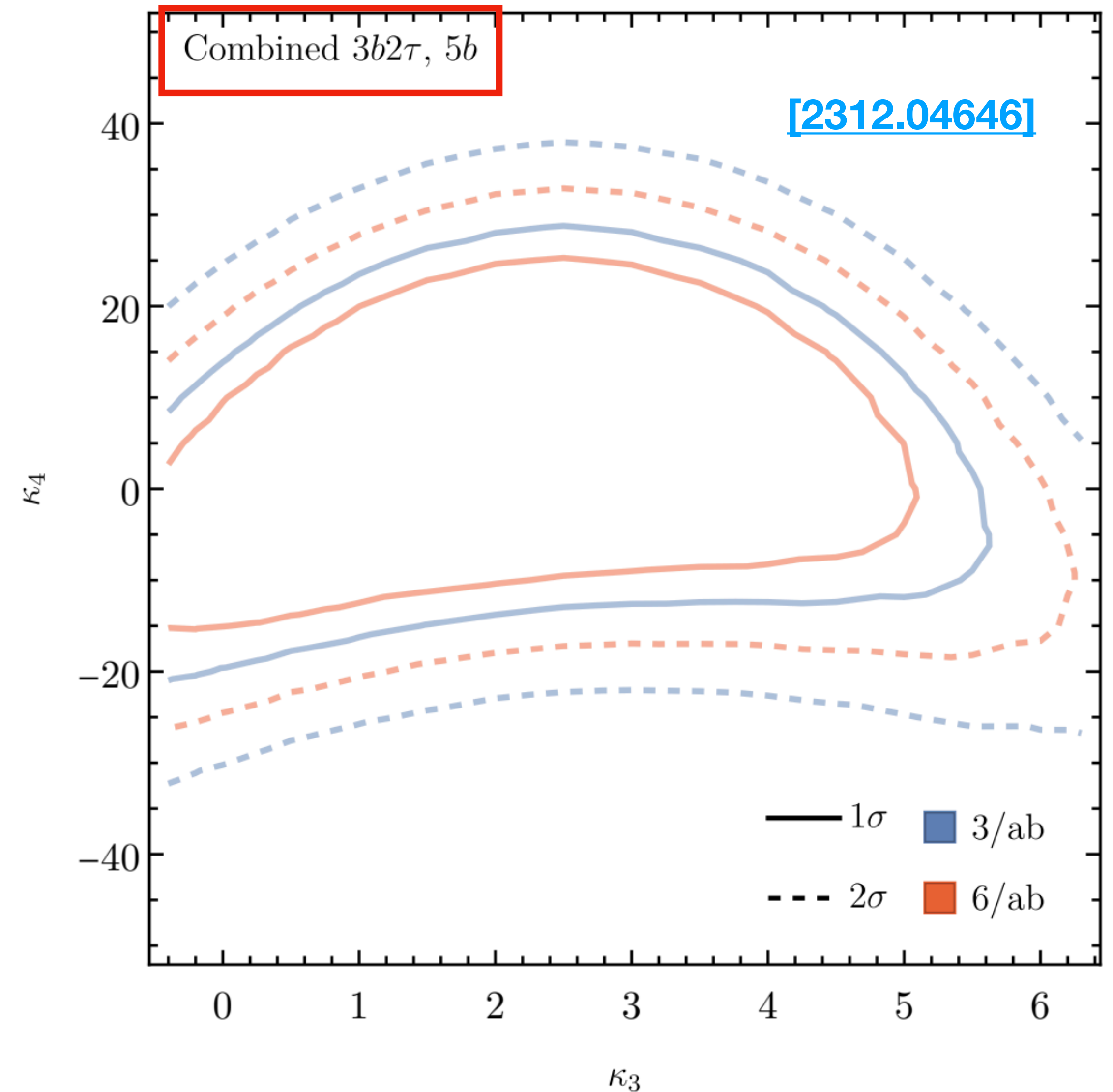
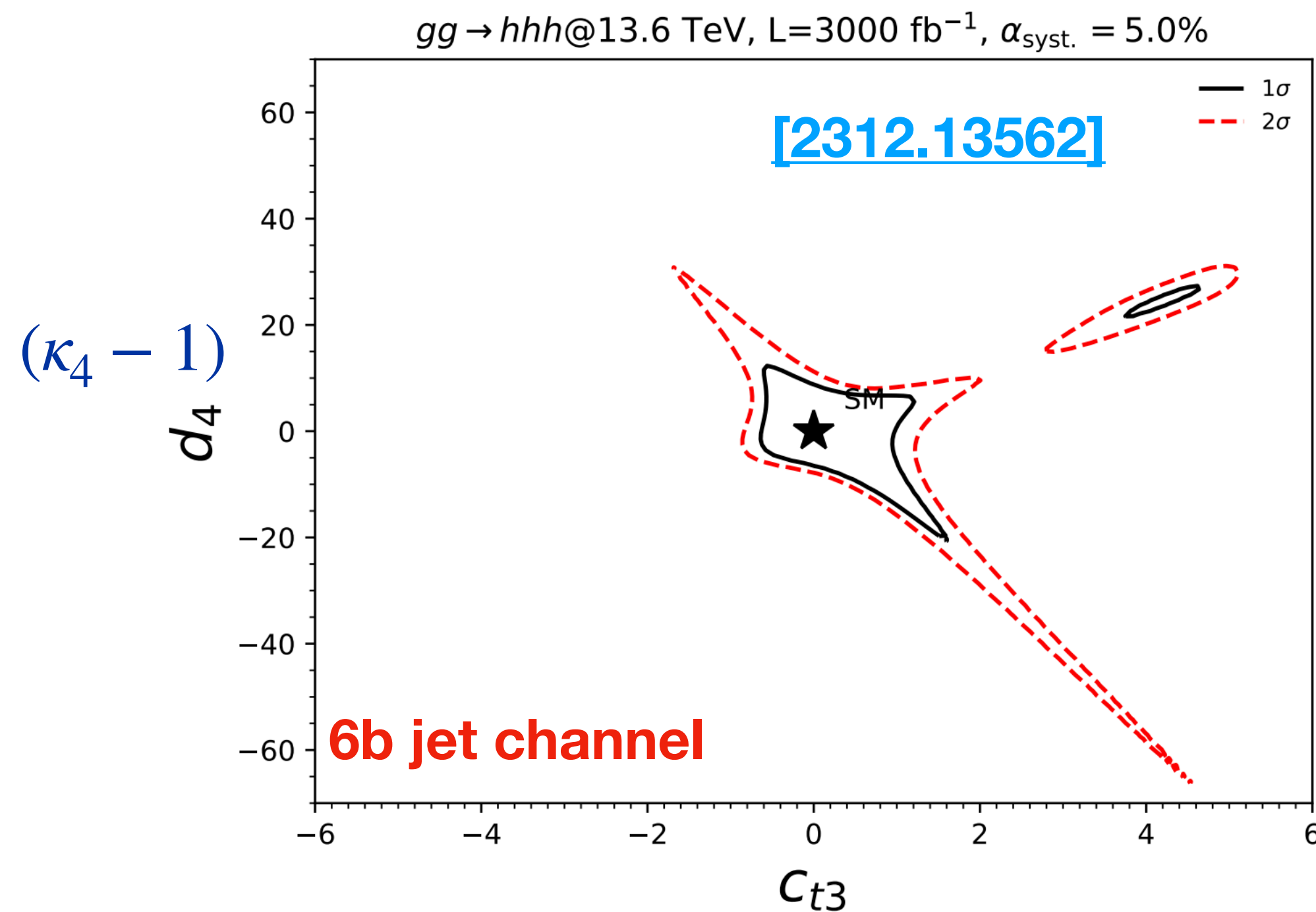
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(Phenomenological) HL-LHC Projections

- Some projections from the [HHH whitepaper](#), [\[2312.13562\]](#), and [\[2312.04646\]](#) suggest $\kappa_4 = 1 \pm 30$ is possible in HL-LHC. 95% CL.
- Take them with a grain of salt, as projections are not yet based in a published analysis.

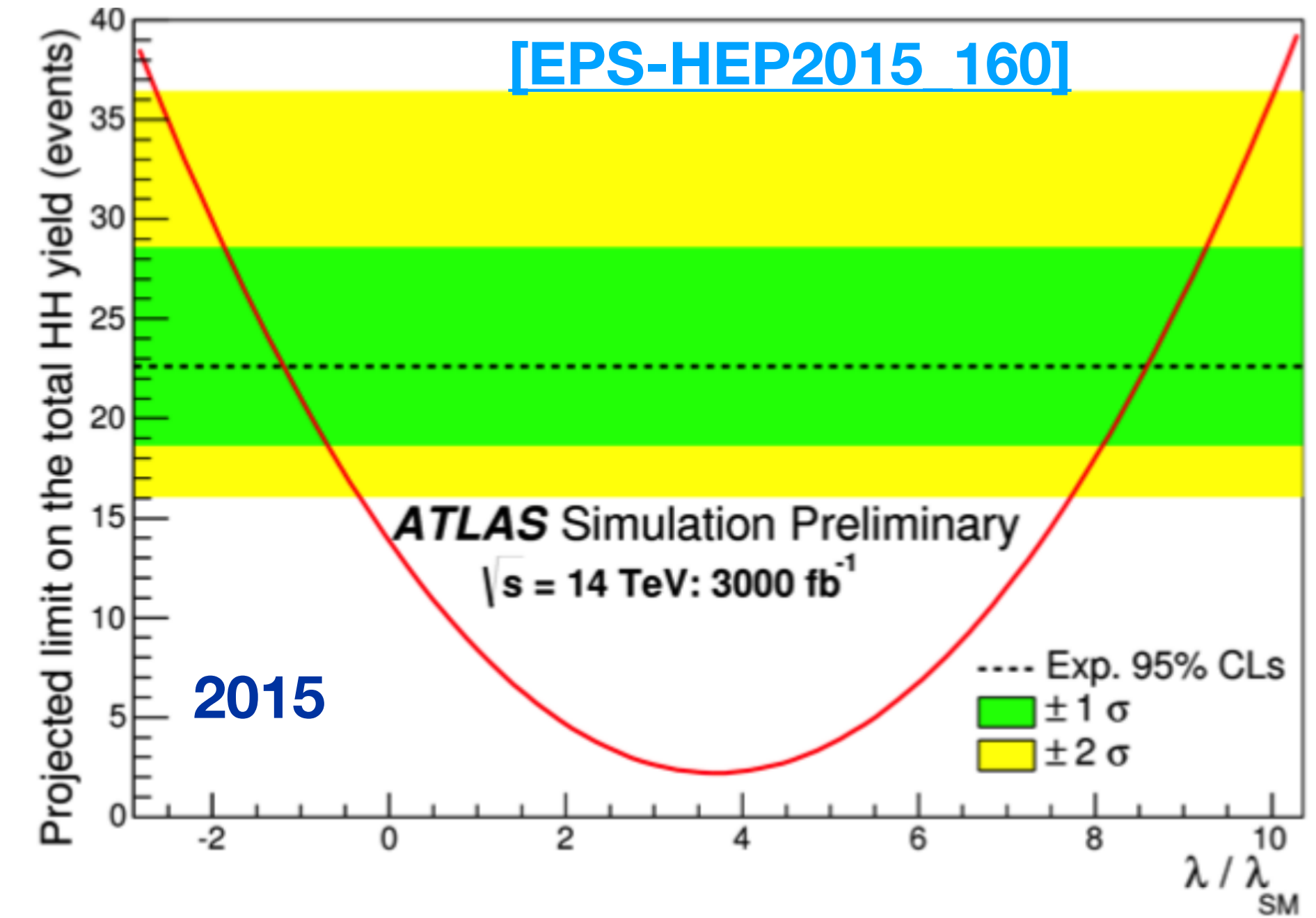


Performance of projections: HHistory



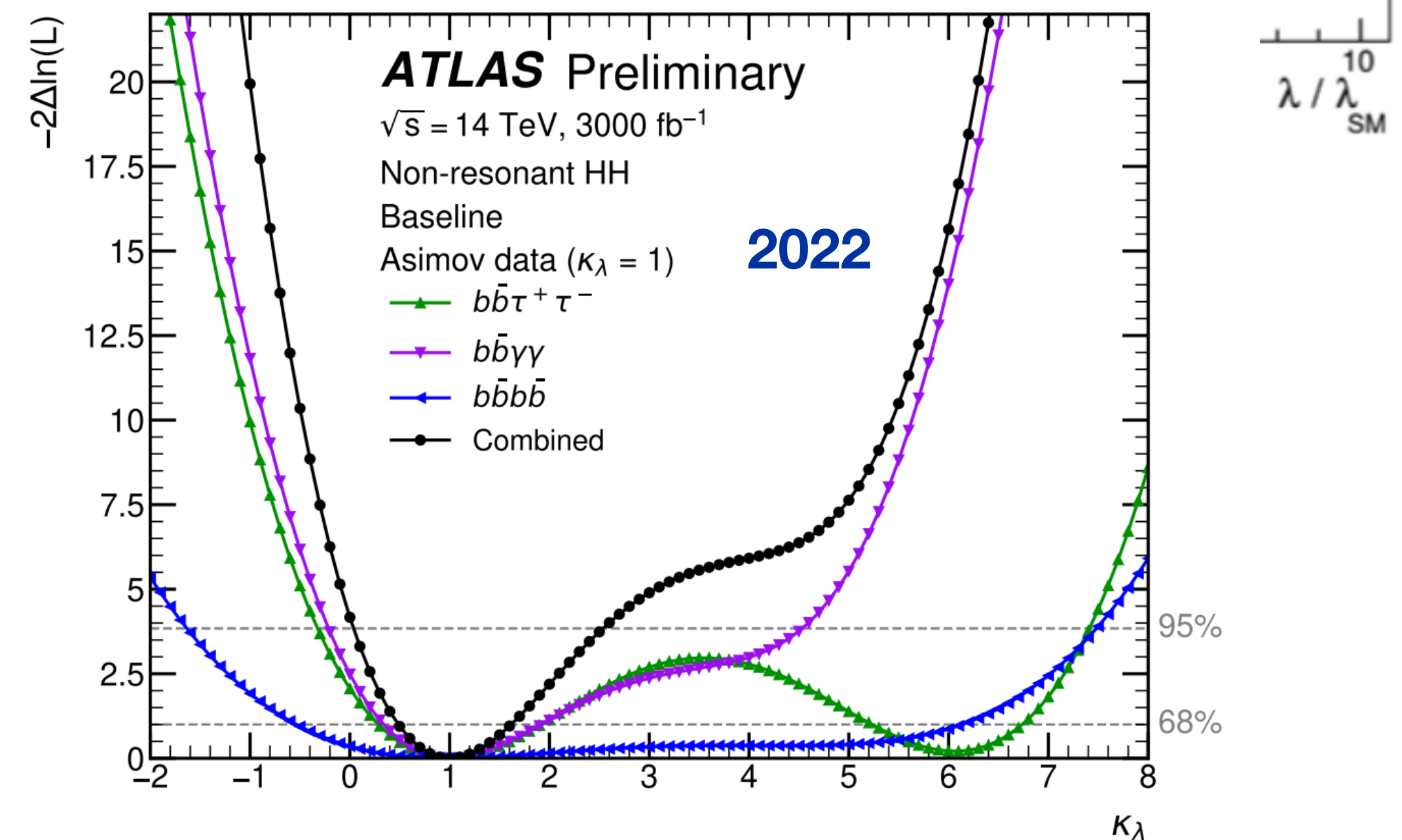
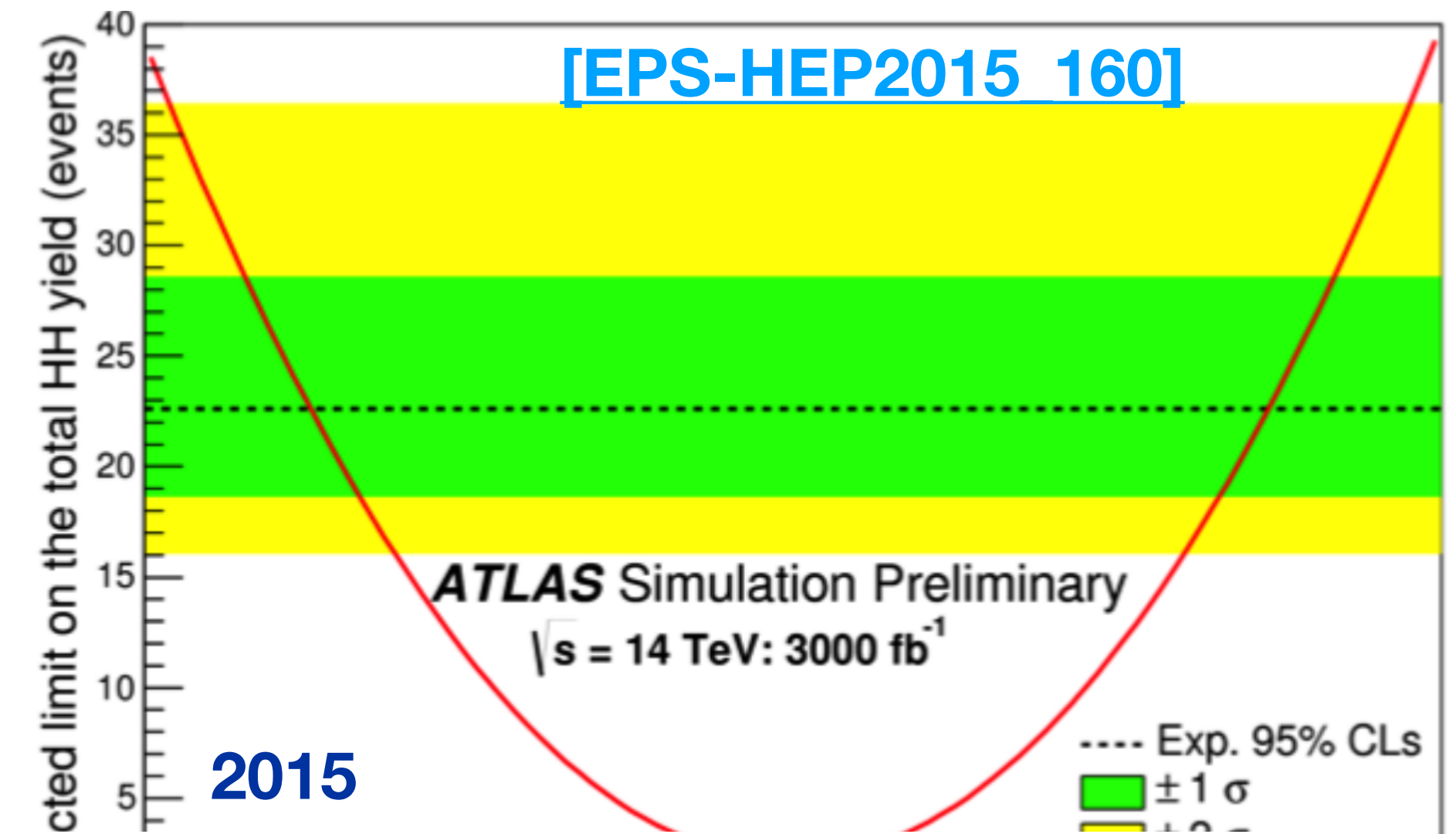
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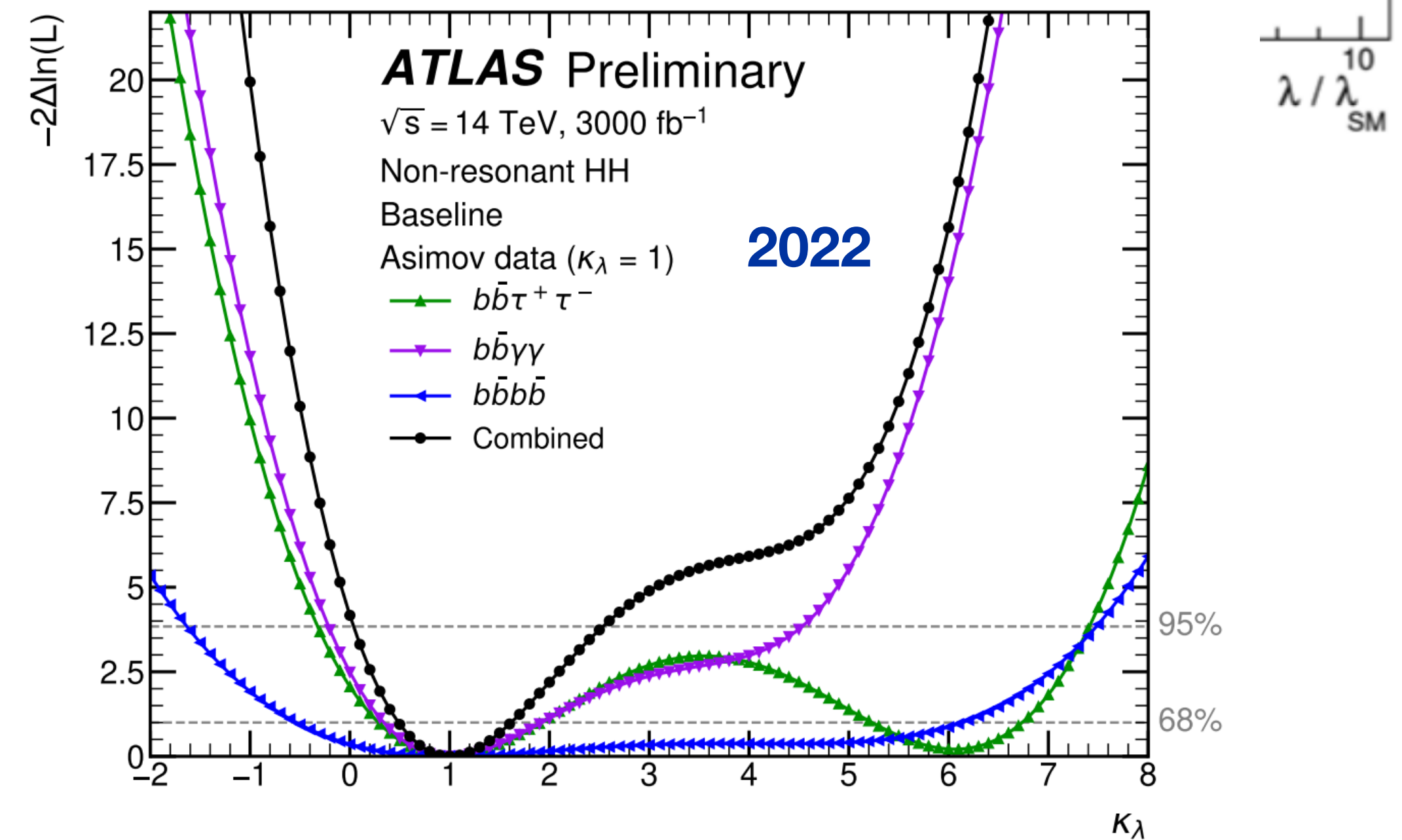
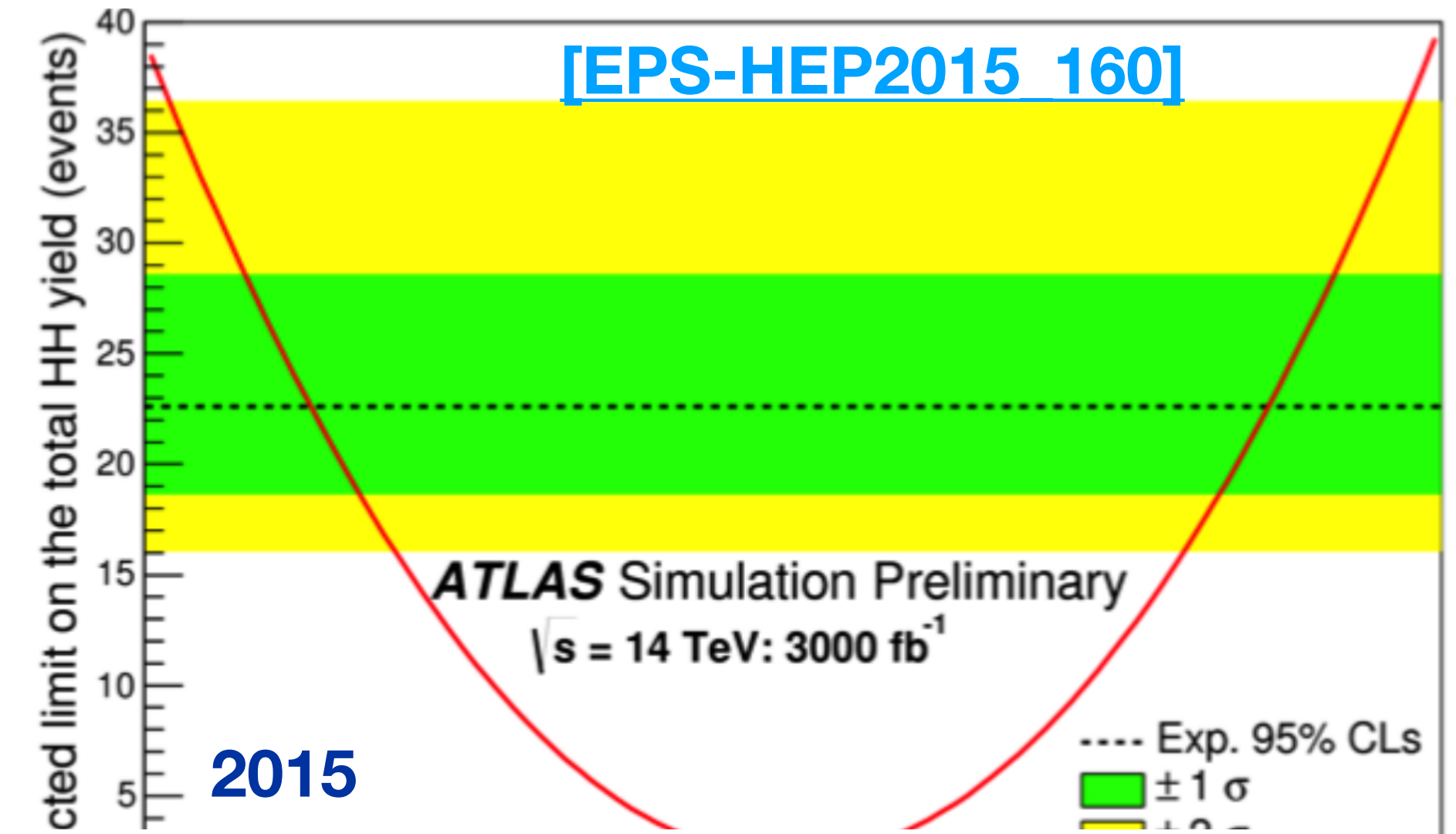
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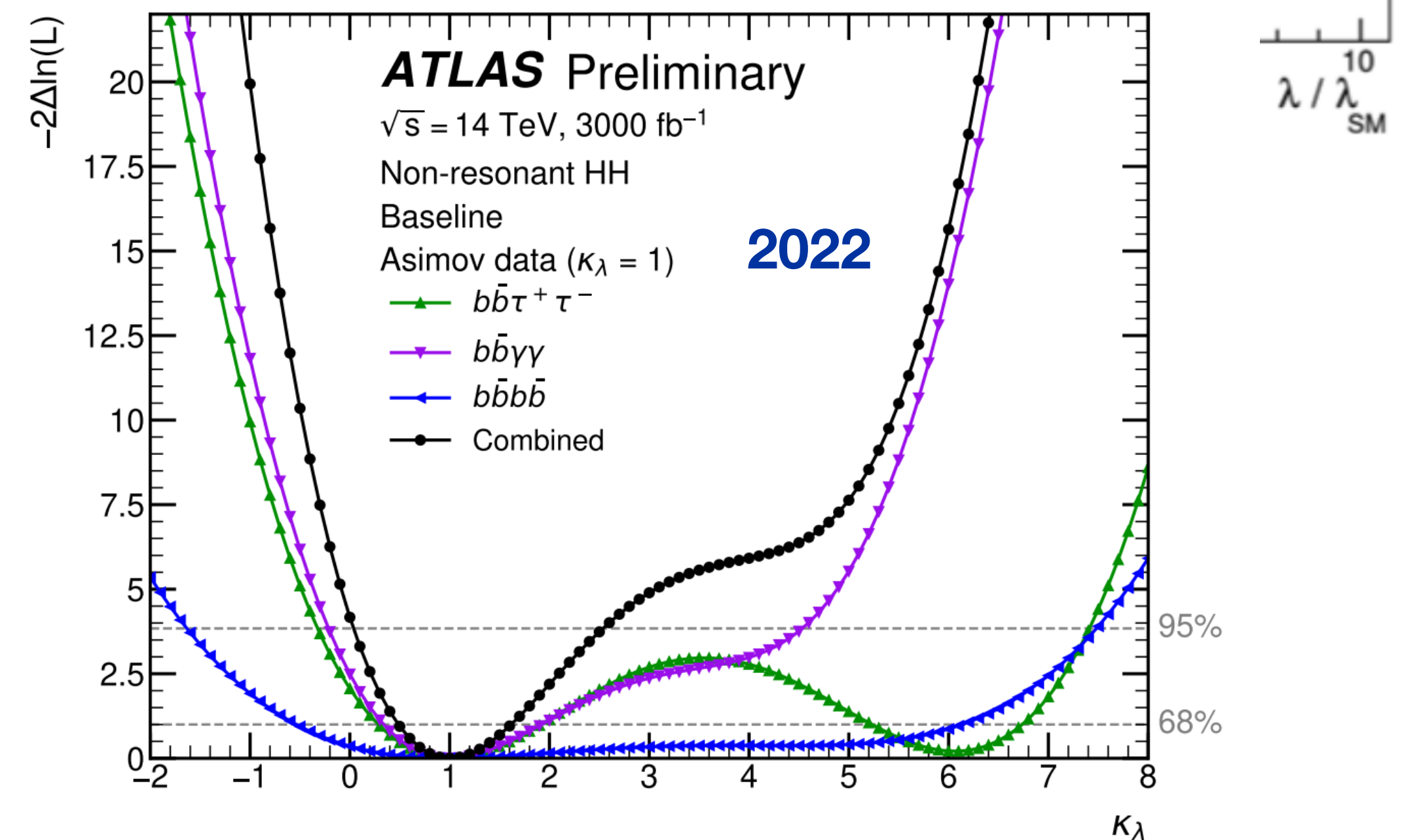
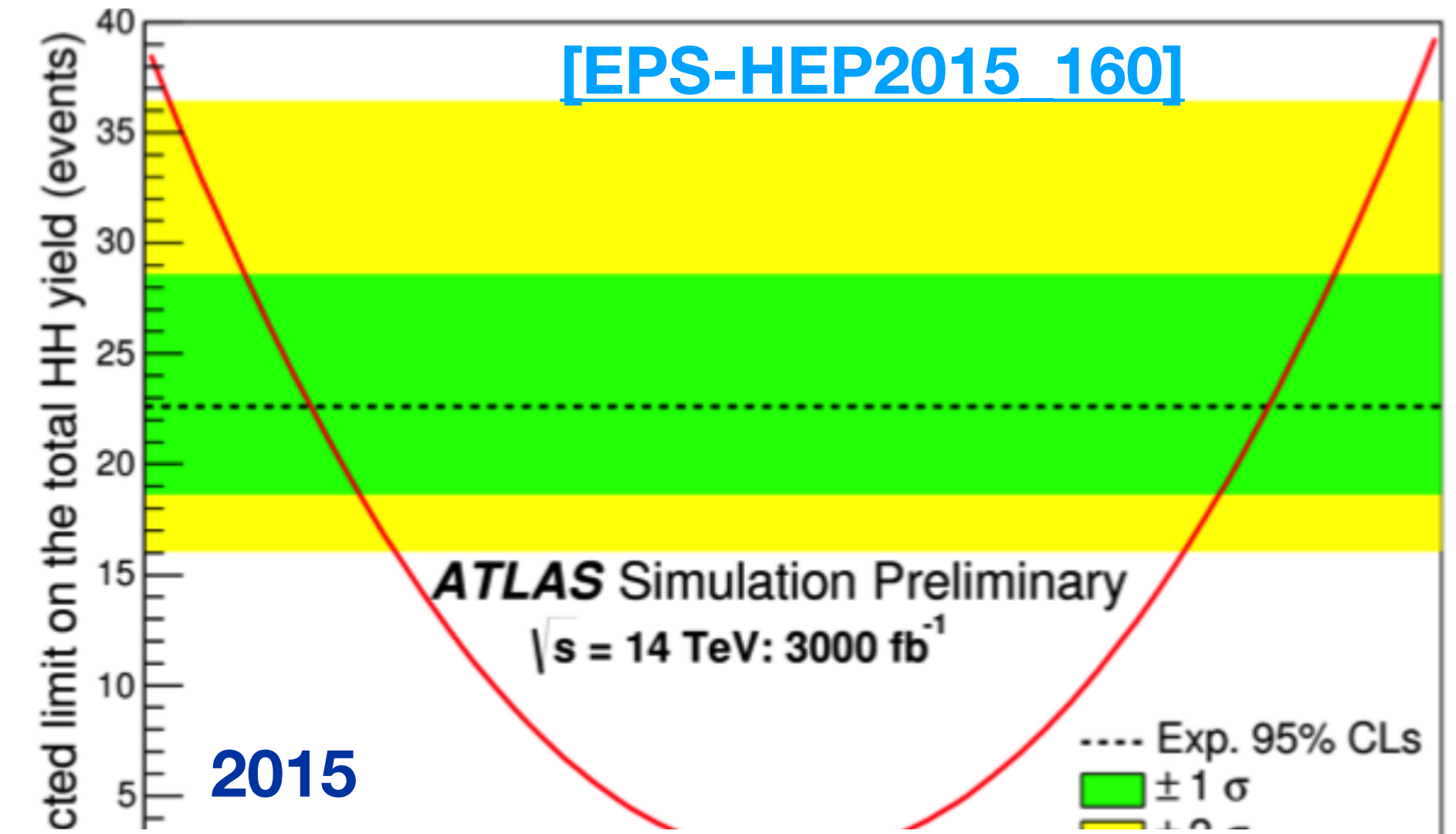


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- In just 7 years, phase space for projection reduces by 4x:
 - Detector performance improvements
 - Improvements in b-tagging
 - New analysis ideas: AI/ML, categorization
 - Reduction of theoretical uncertainties
 - Exploration of more channels

I argue both can be improved

HHH is behind in these aspects

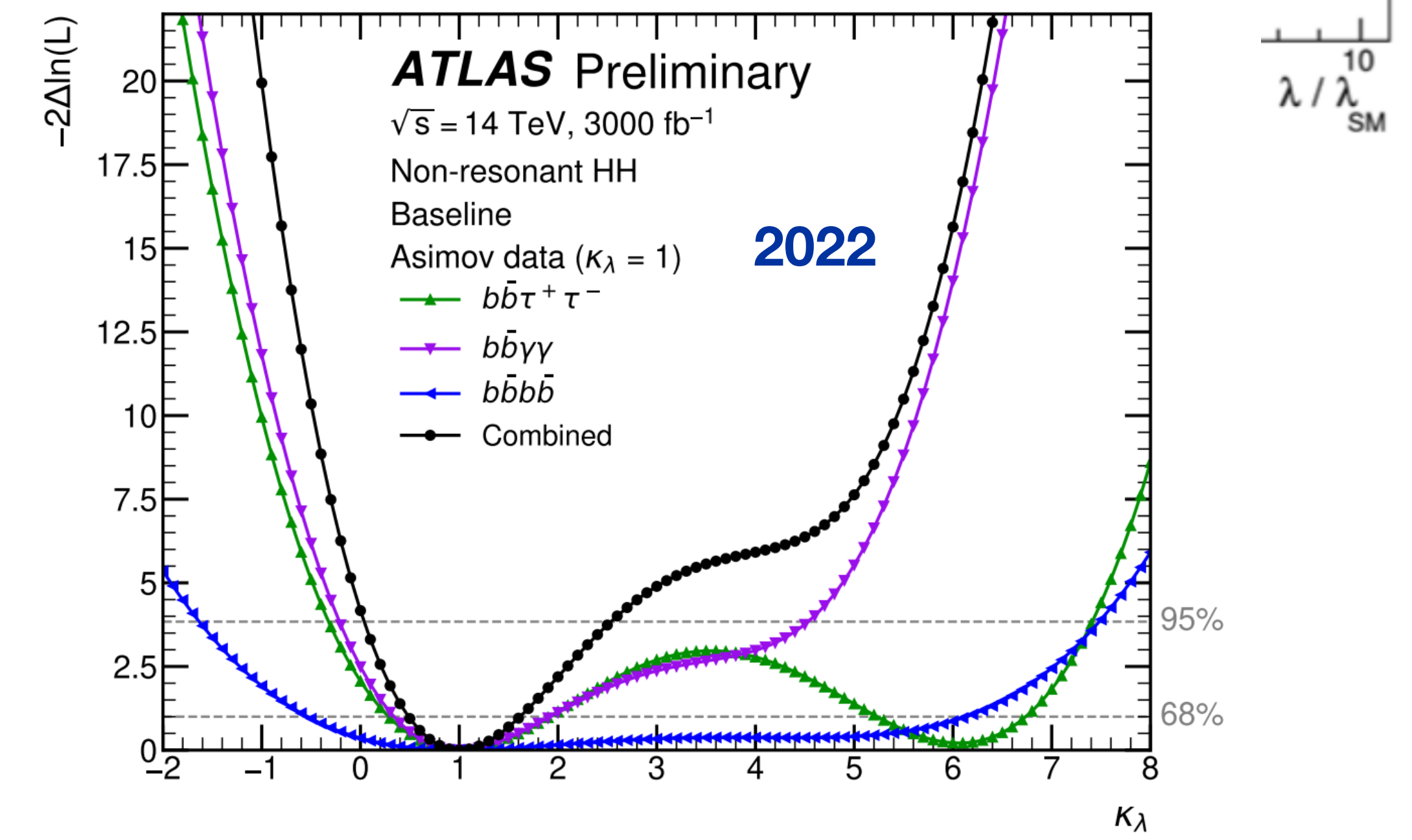
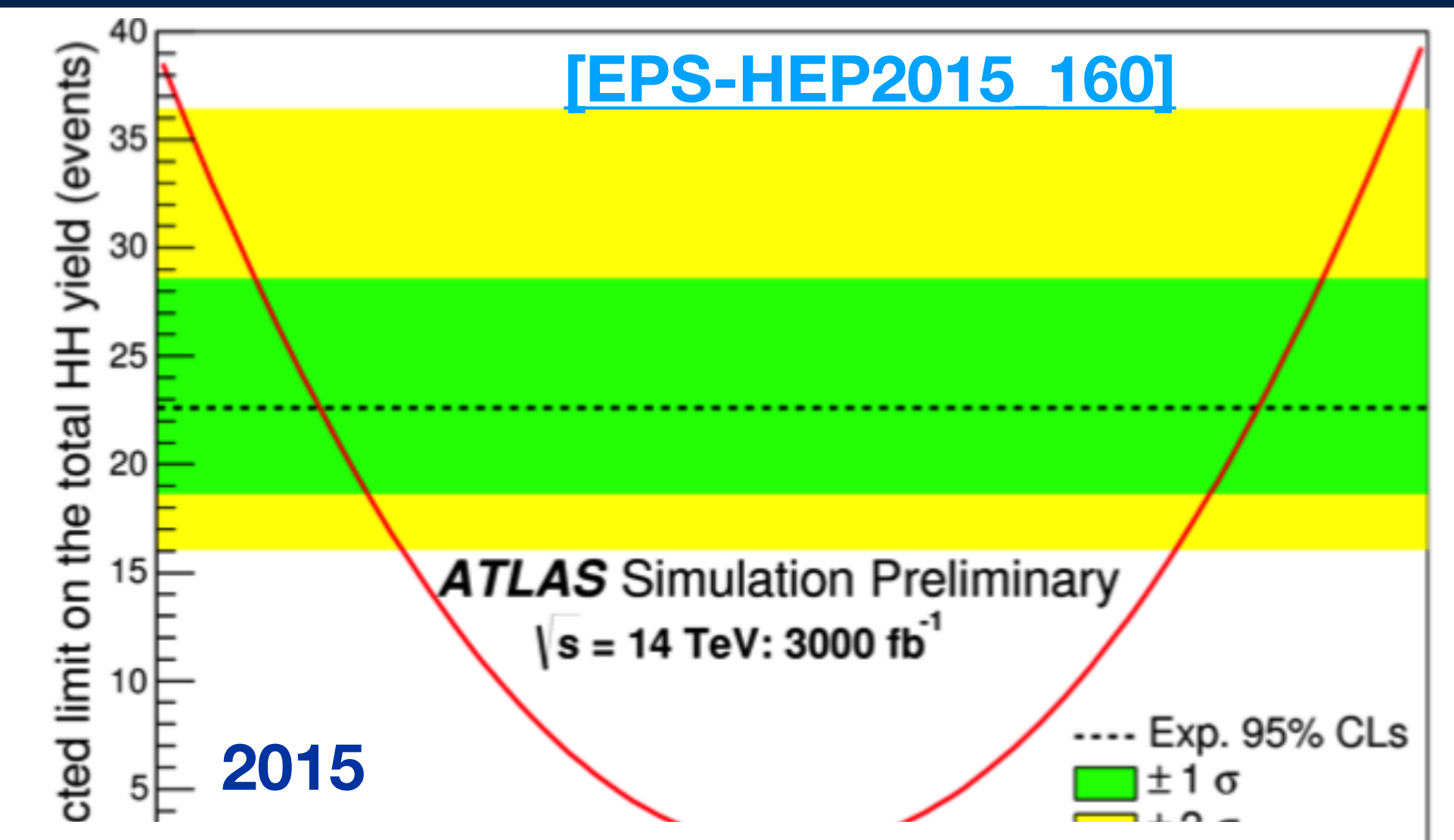


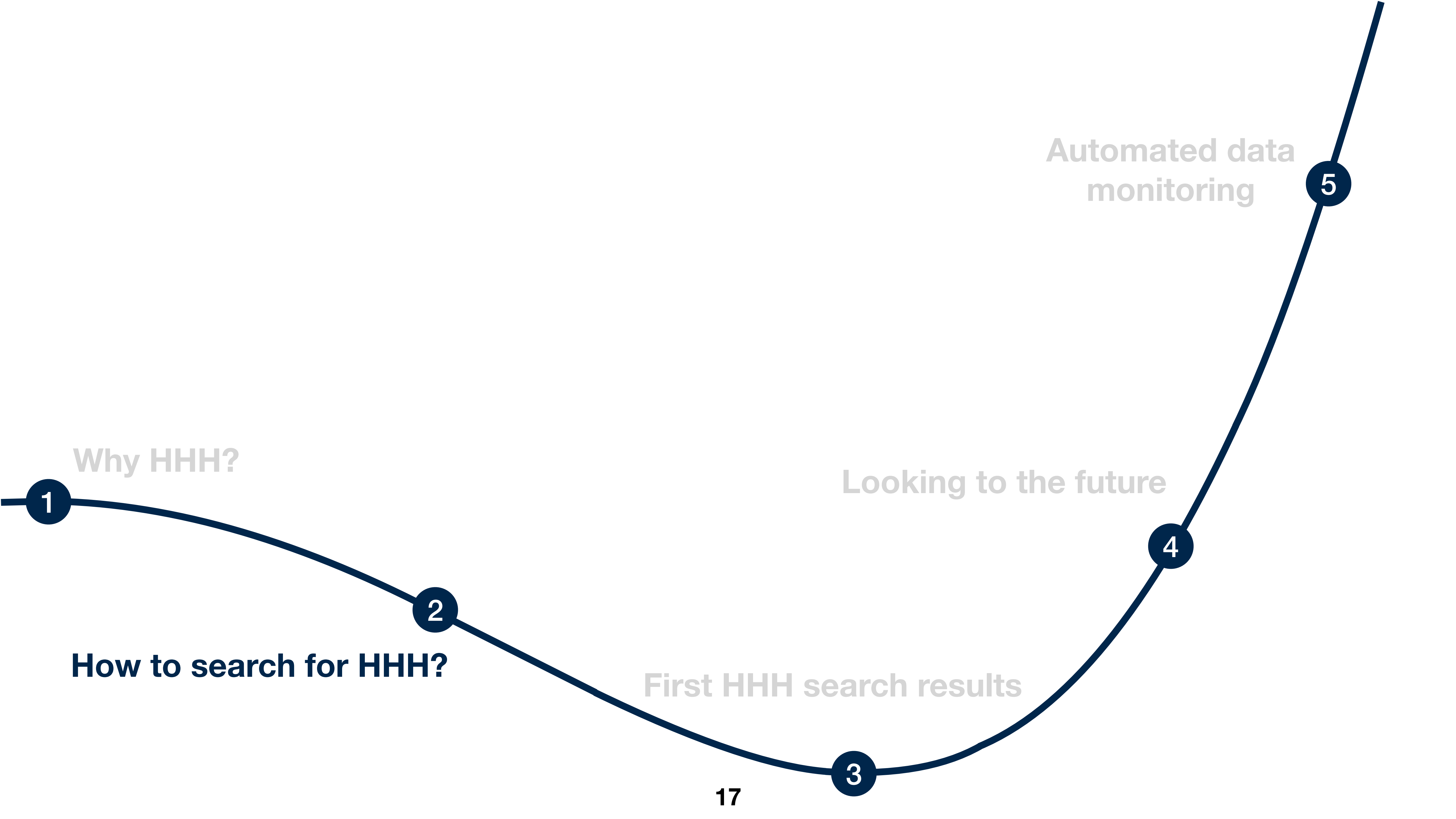
Performance of projections: HHHistory

- In 2015, ATLAS projected that the $b\bar{b}\gamma\gamma$ channel would be the most sensitive, and estimated limits on κ_3 : $-1.3 < \kappa_3 < 8.7$ in HL-LHC
- By 2022, the $b\bar{b}\gamma\gamma$ projection had improved to $-0.2 < \kappa_3 < 4.6$
- Other channels in combination improved projection to $0.1 < \kappa_3 < 2.5$
- In just 7 years, phase space for projection reduces by 4x:
 - Detector performance improvements
 - Improvements in b-tagging
 - New analysis ideas: AI/ML, categorization
 - Reduction of theoretical uncertainties
 - Exploration of more channels
- Is $\kappa_4 = 1 \pm 10$ within the realm of possibility? Probably not, but we don't know yet. HHH not guaranteed to evolve like of HH projections

I argue both can be improved

HHH is behind in these aspects





Why HHH?

1

How to search for HHH?

2

First HHH search results

3

17

Looking to the future

4

Automated data monitoring

5

Decay modes



Decay modes

- The first question is what decay mode to search for!



Decay modes

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- There are more options for HHH than HH

Reminder: HH branching fractions

	bb	WW	$\tau\tau$	ZZ	$\gamma\gamma$
bb	34%				
WW	25%	4.6%			
$\tau\tau$	7.4%	2.5%	0.39%		
ZZ	3.1%	1.2%	0.34%	0.076%	
$\gamma\gamma$	0.26%	0.10%	0.029%	0.013%	0.0005%

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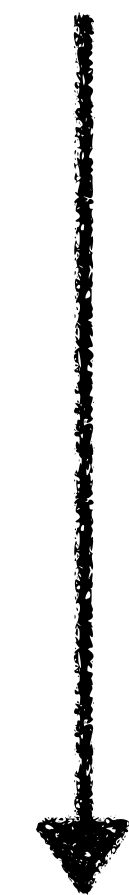
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*An understudied mode

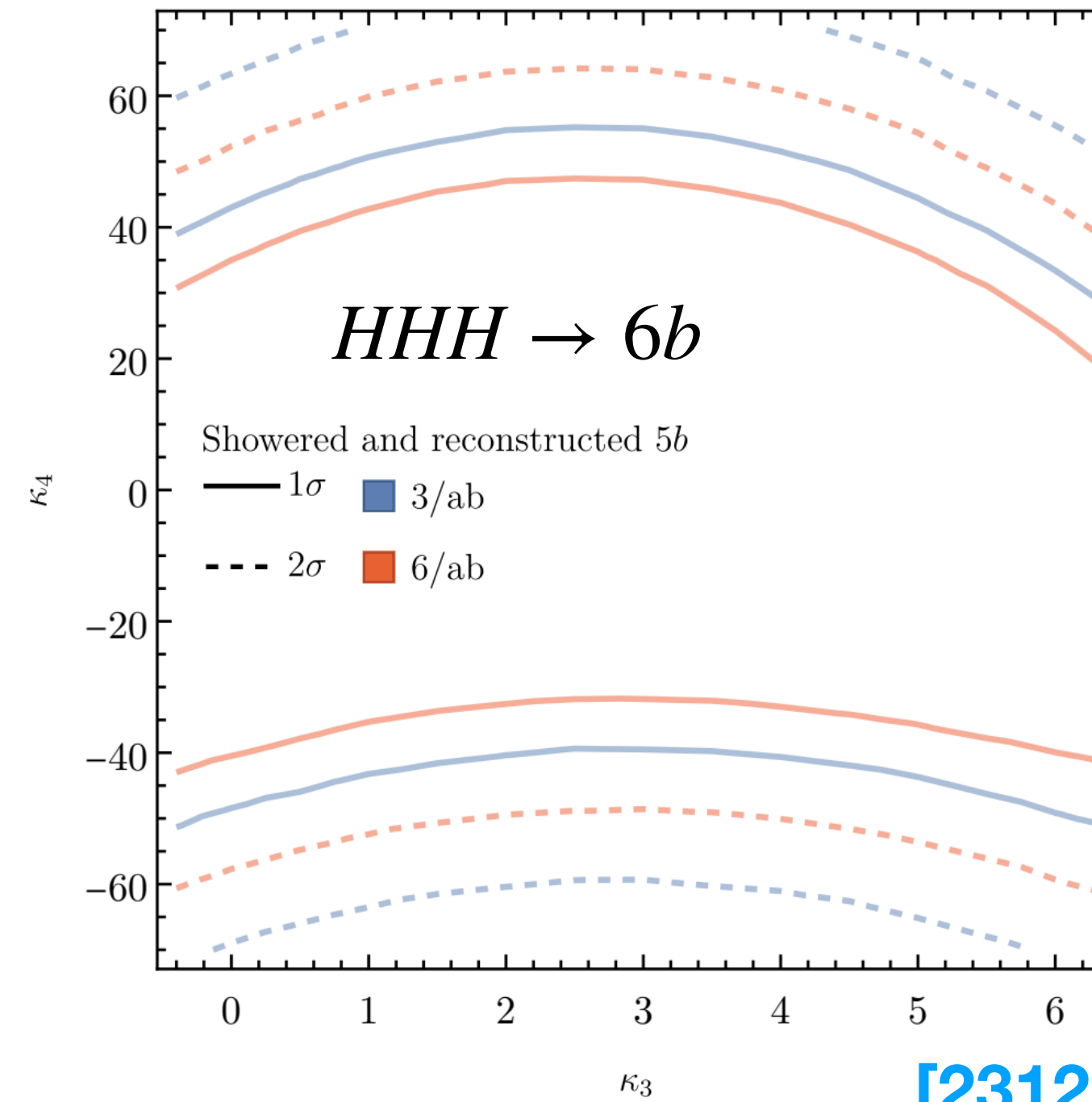
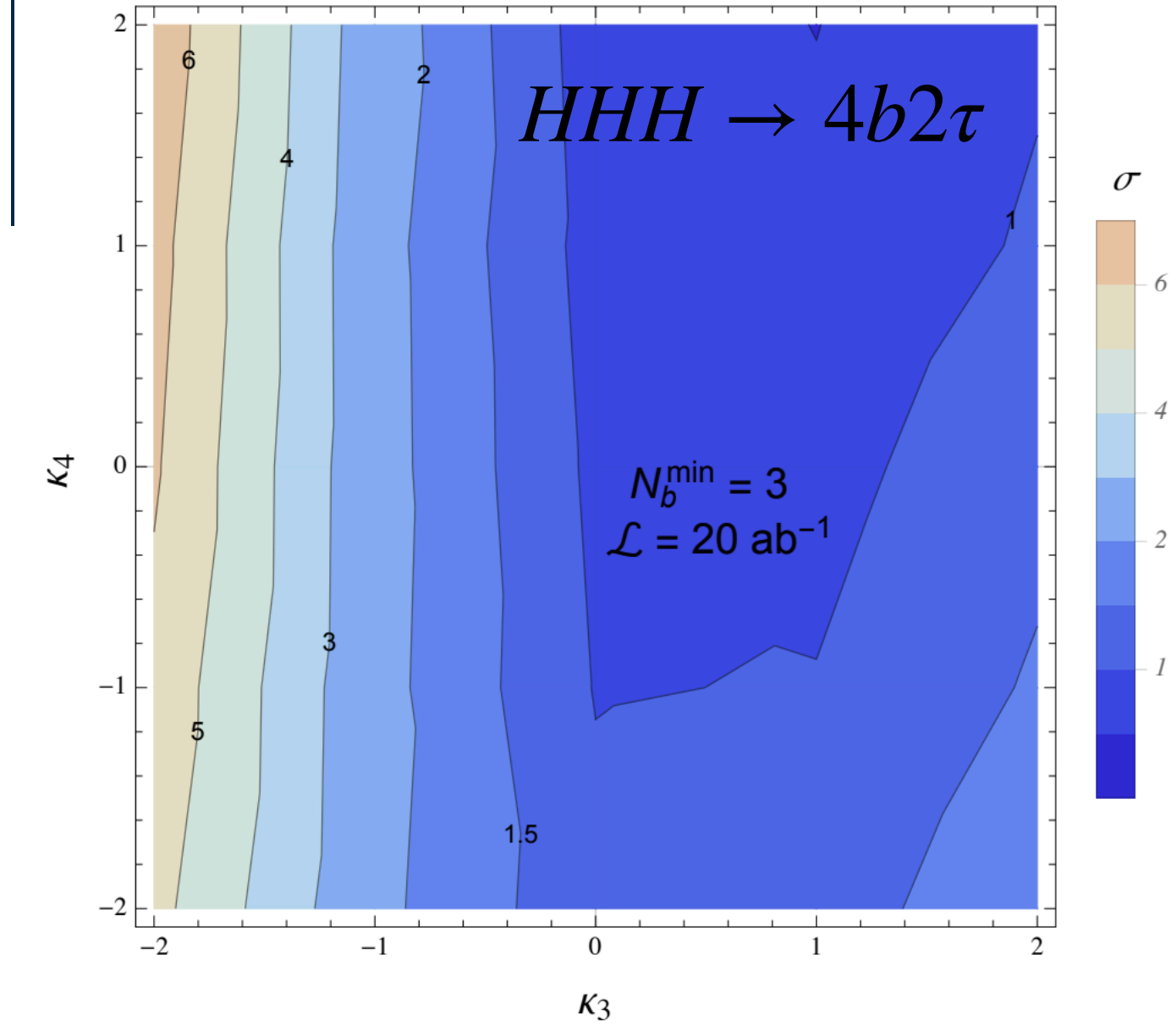
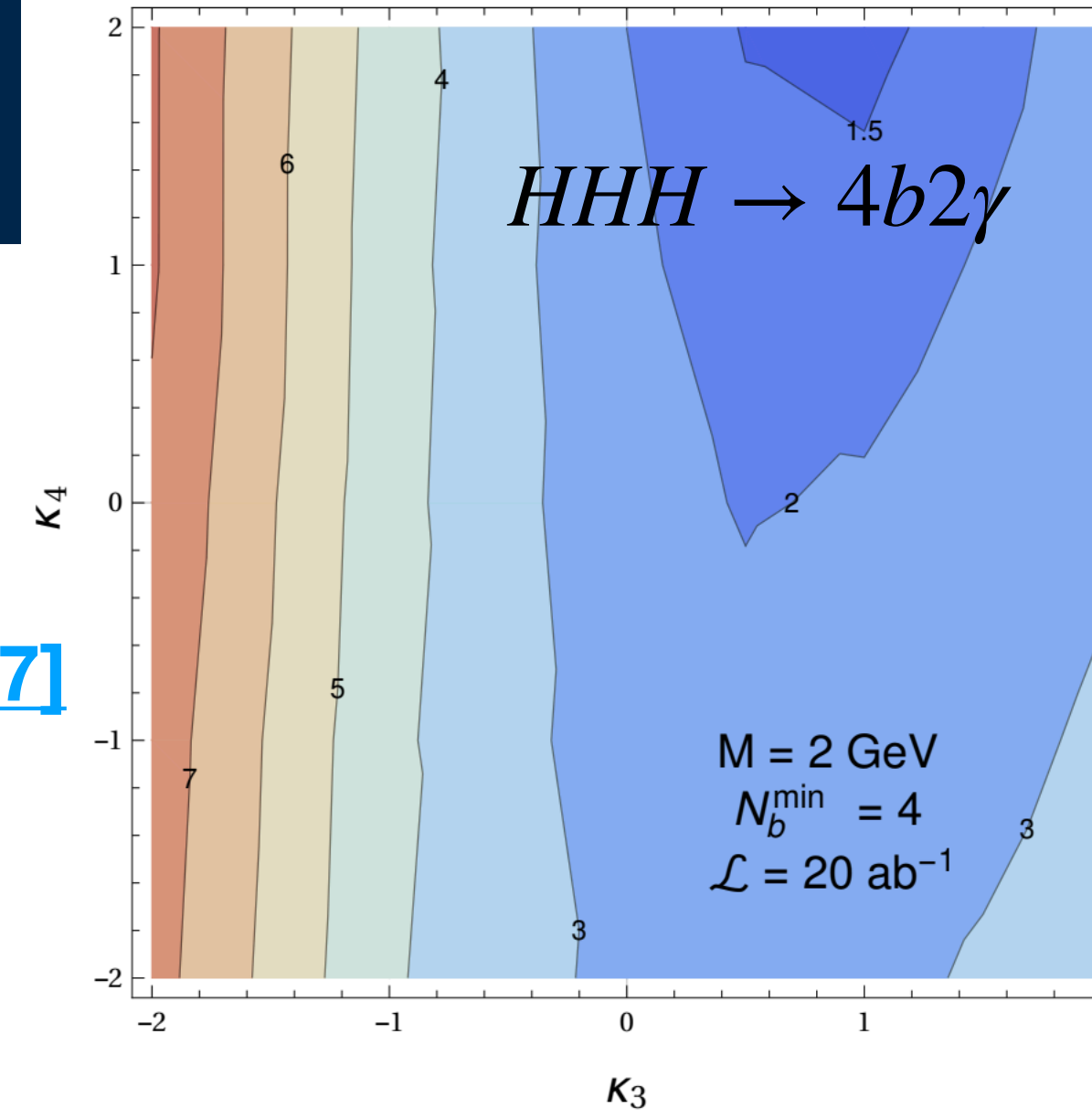


Ordered by likely sensitivity to SM HHH

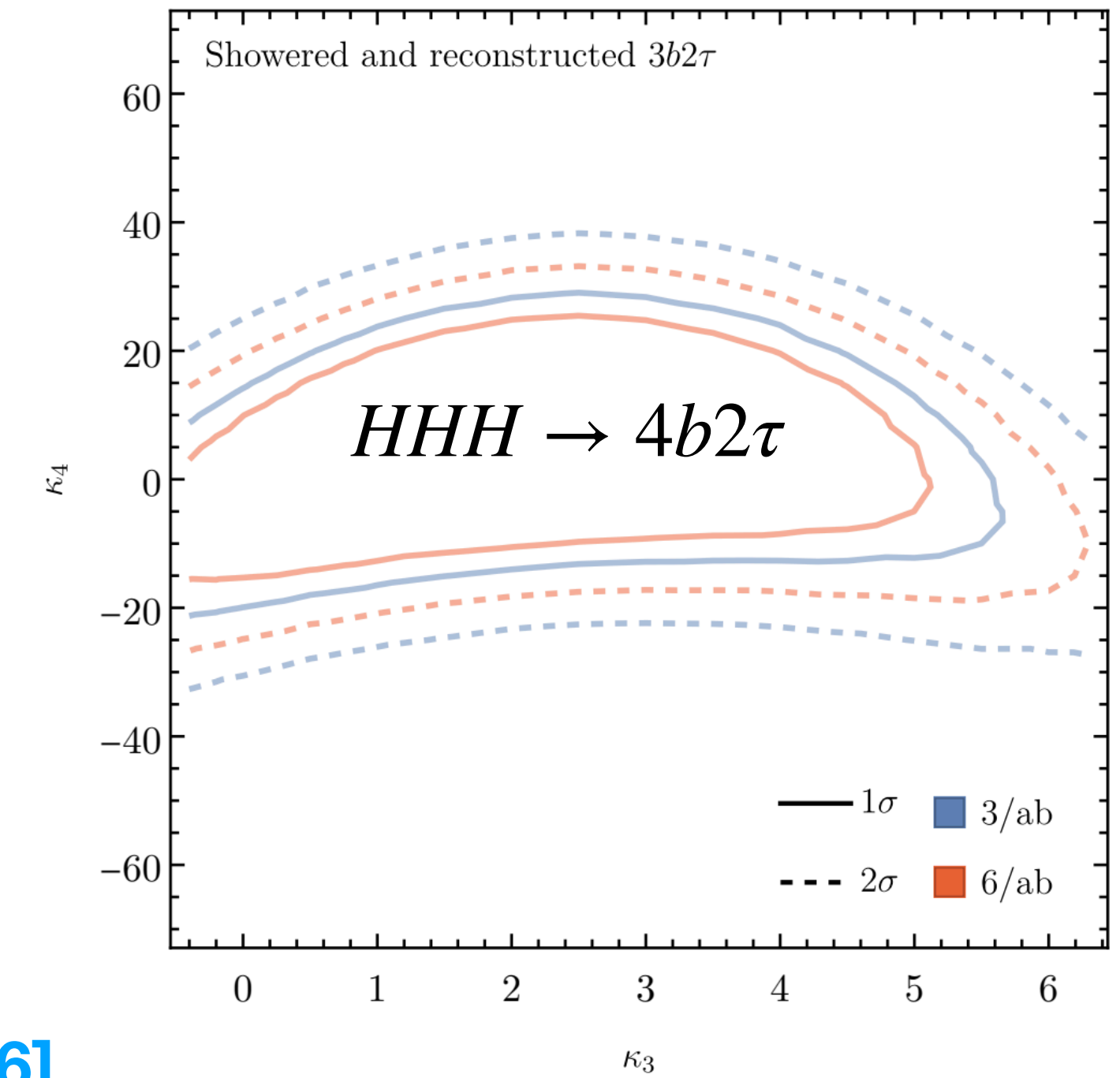


Decay modes

[1510.07697]



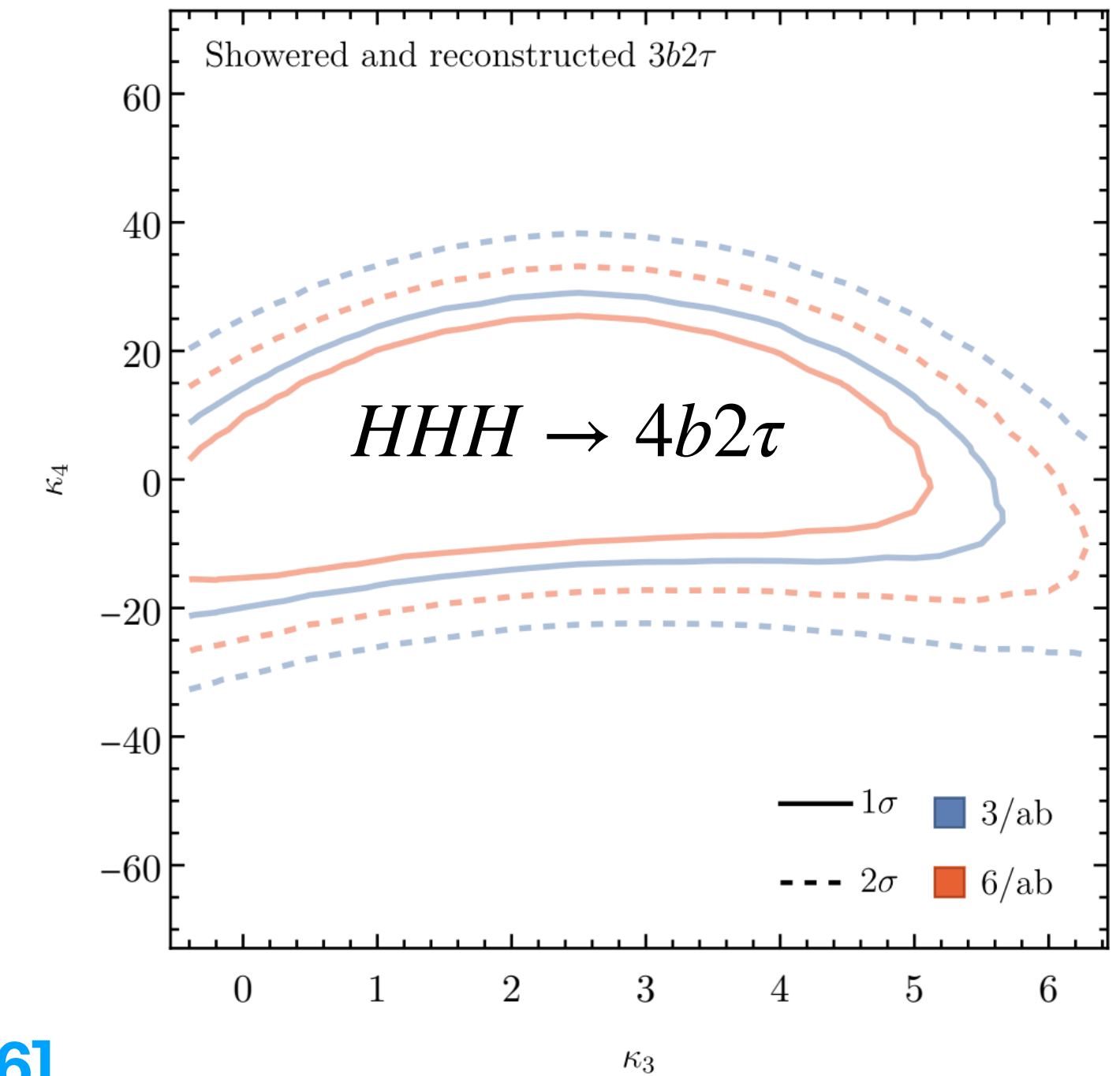
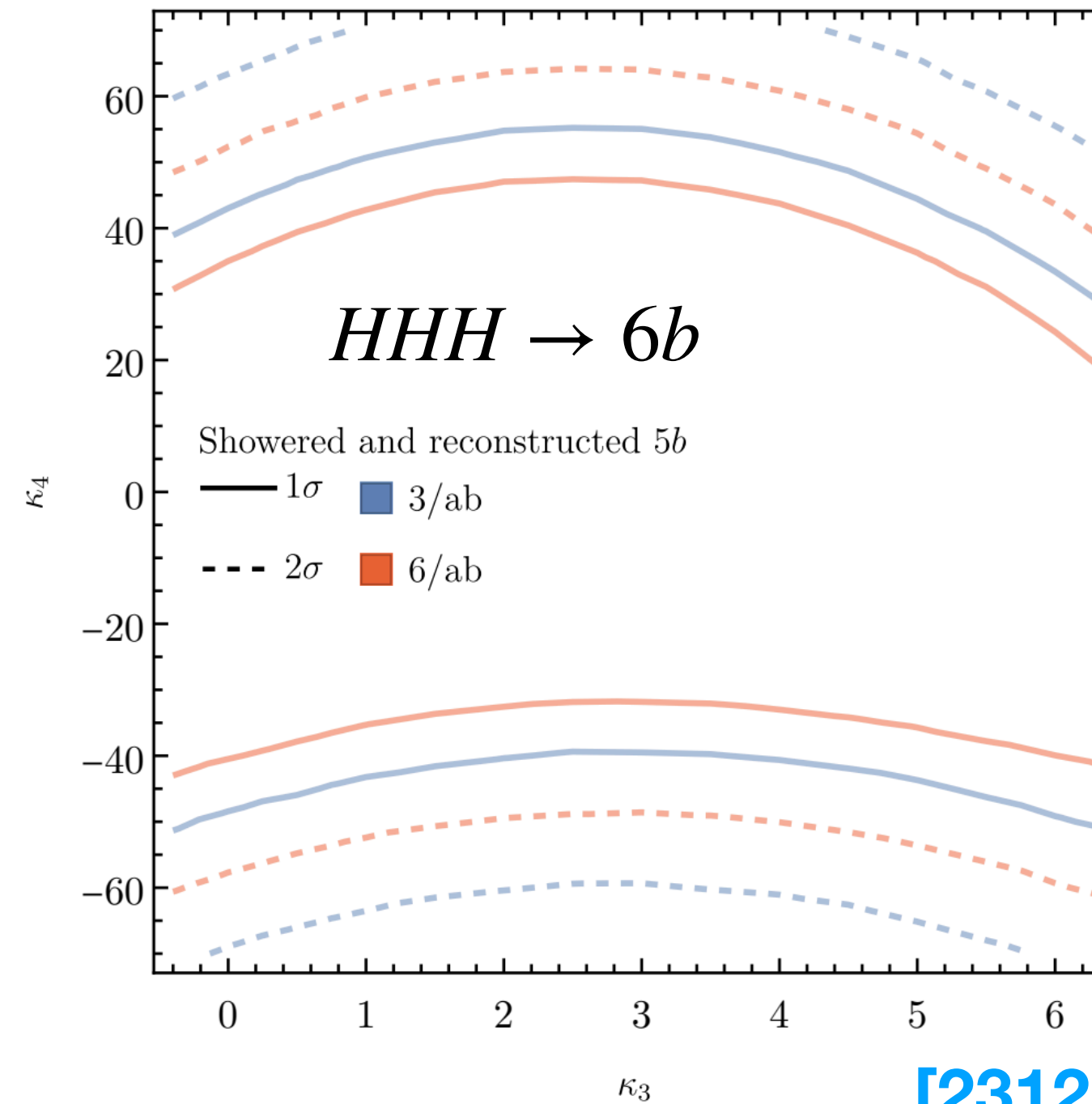
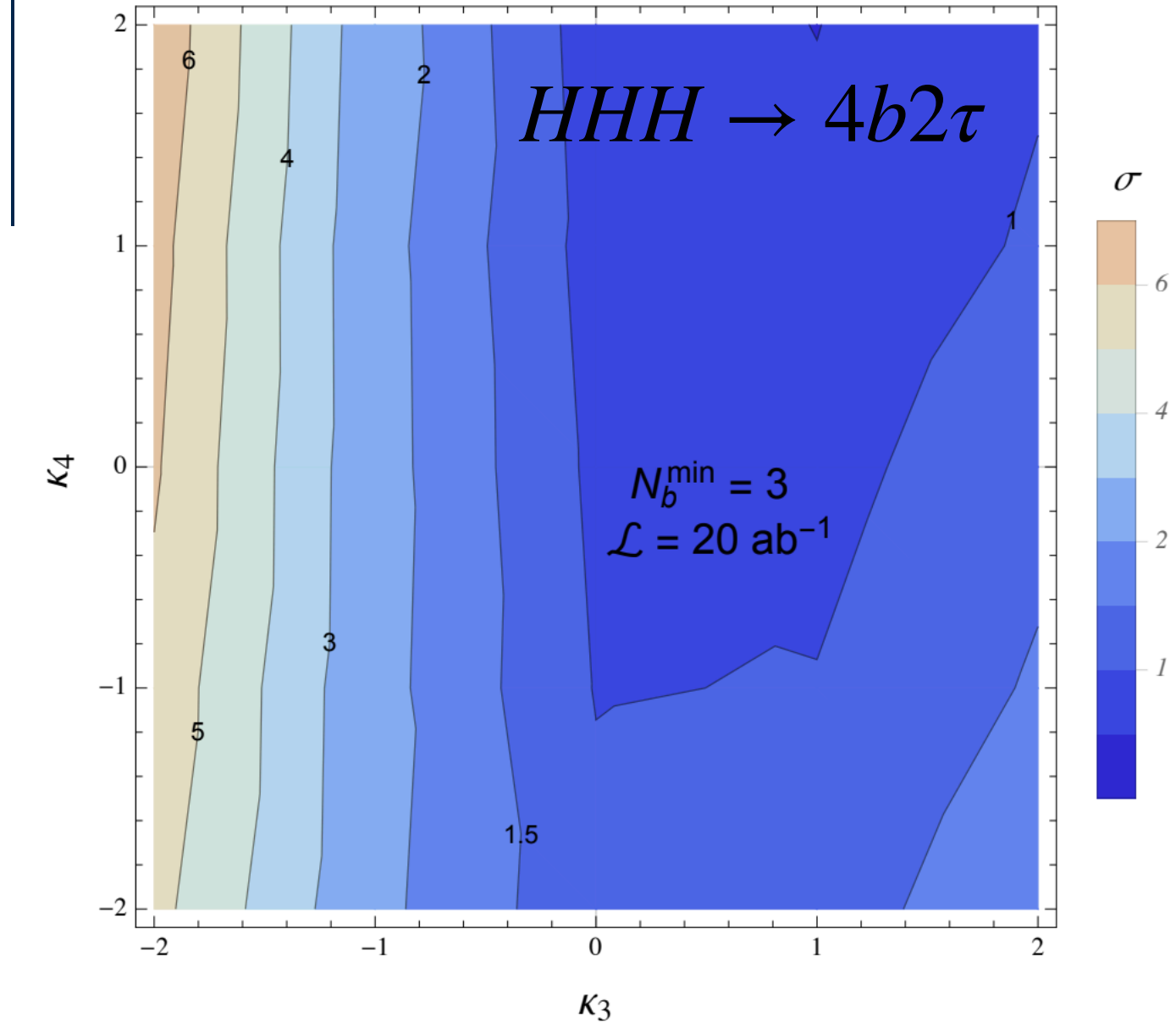
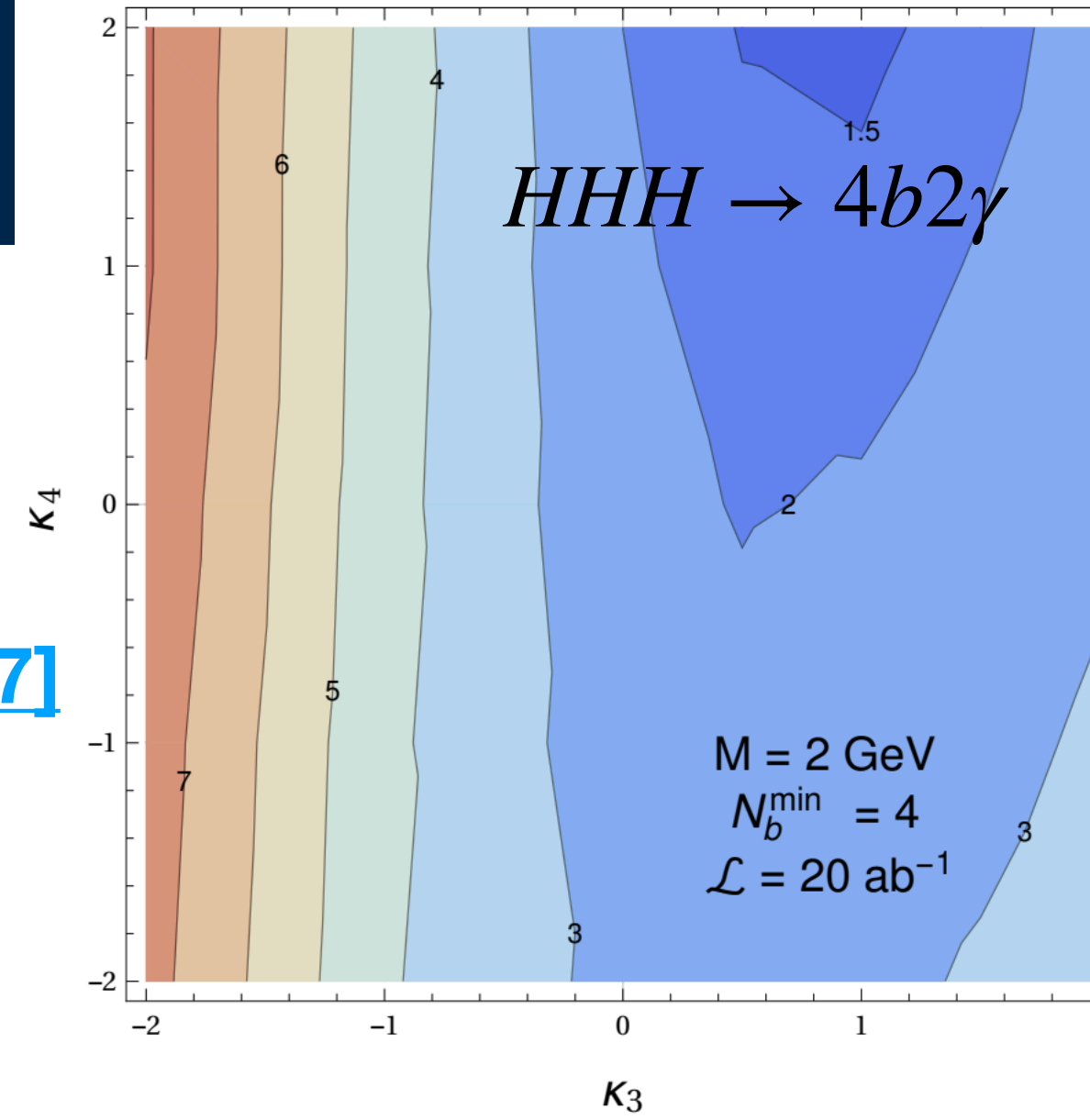
[2312.04646]



Decay modes

- Studies on sensitivity of decay modes are in (slight) conflict

[1510.07697]



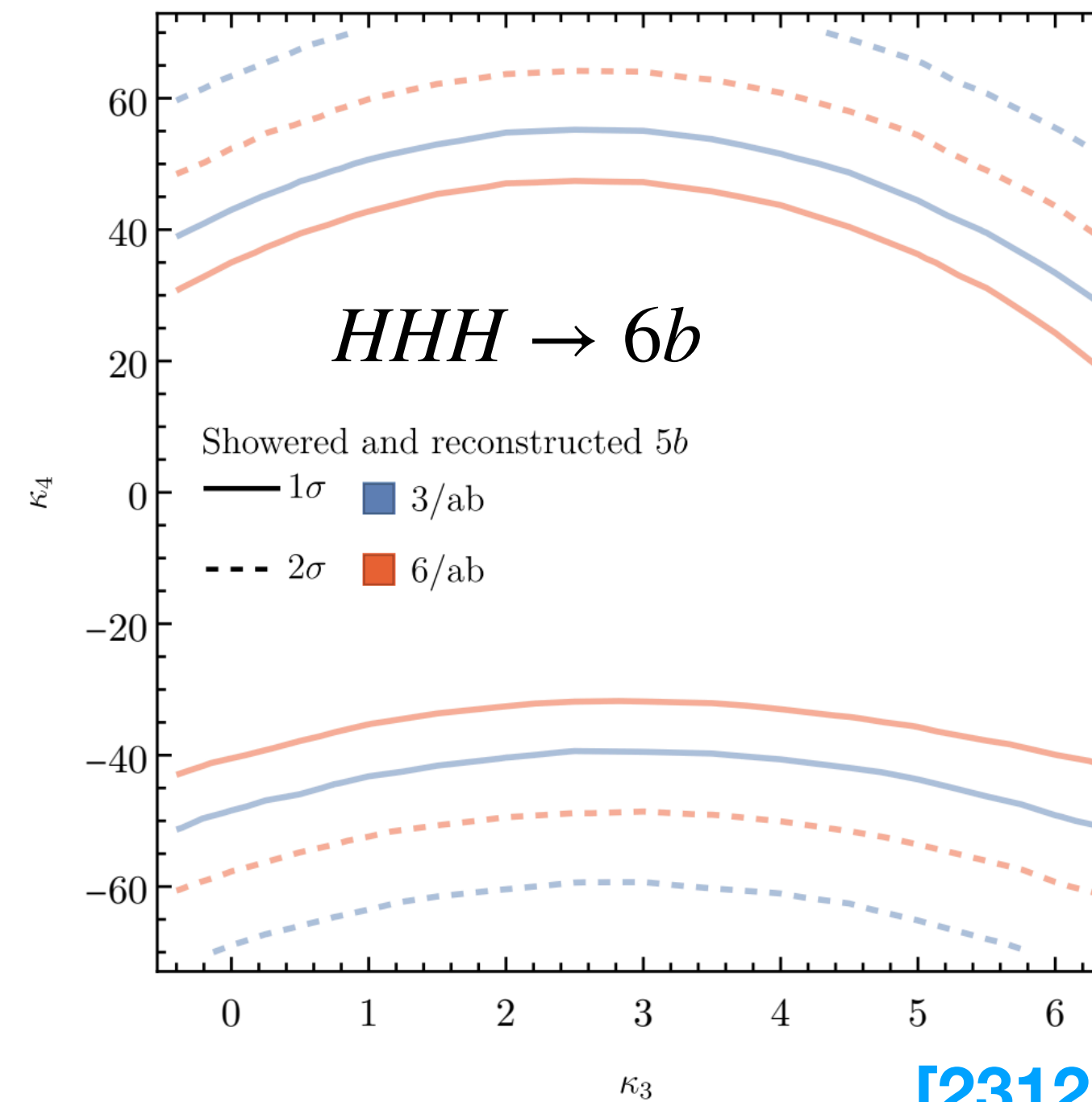
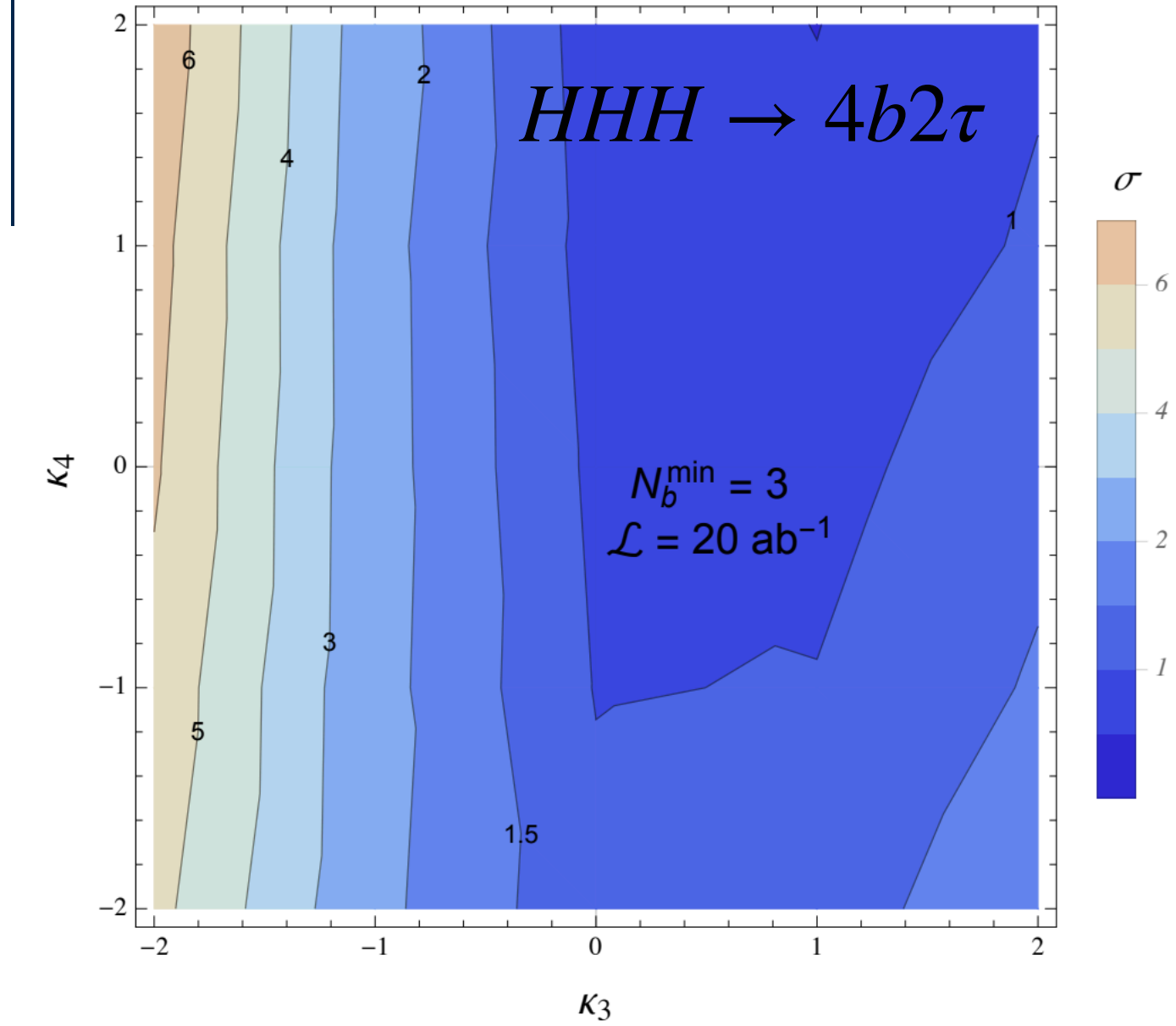
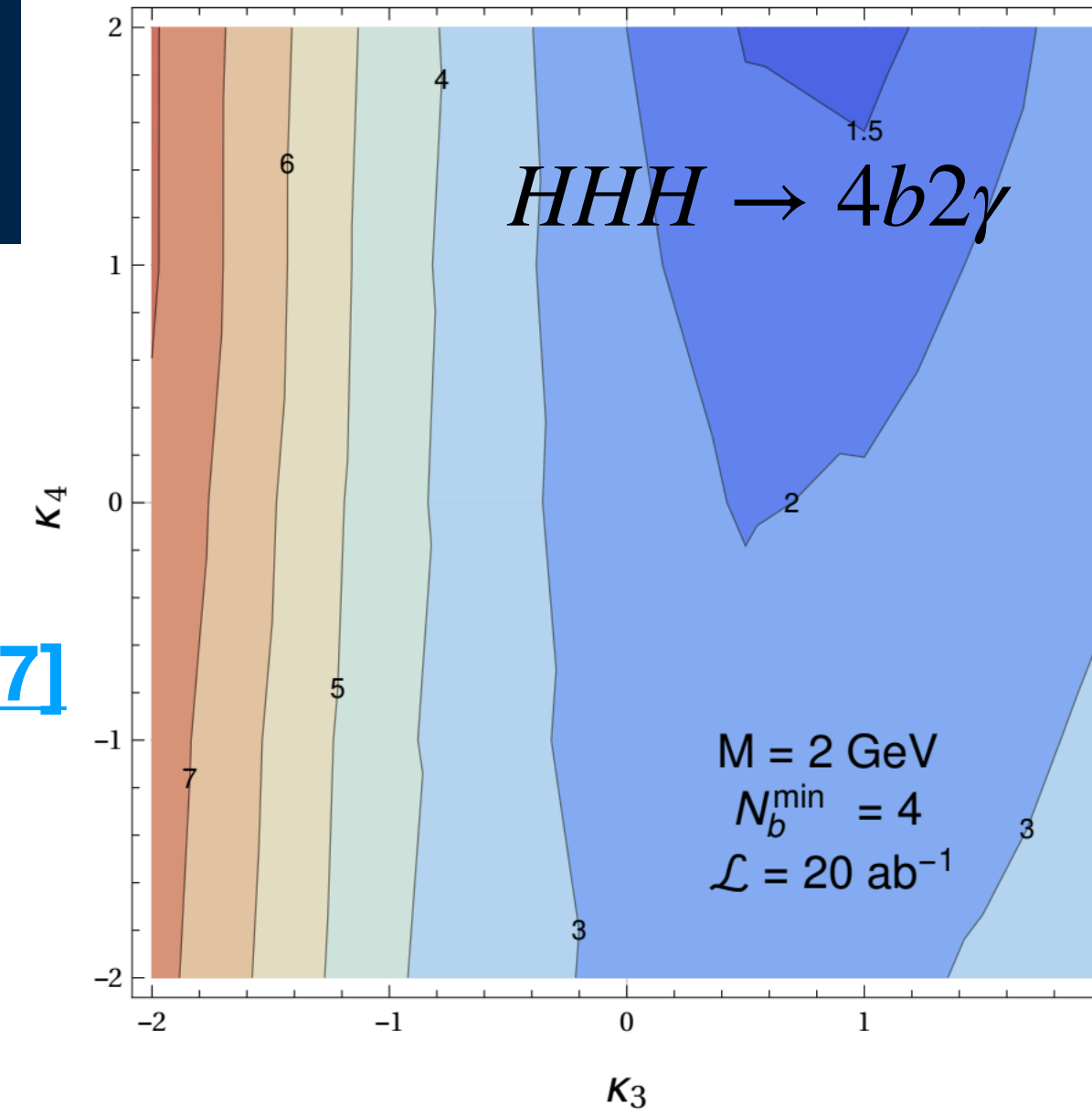
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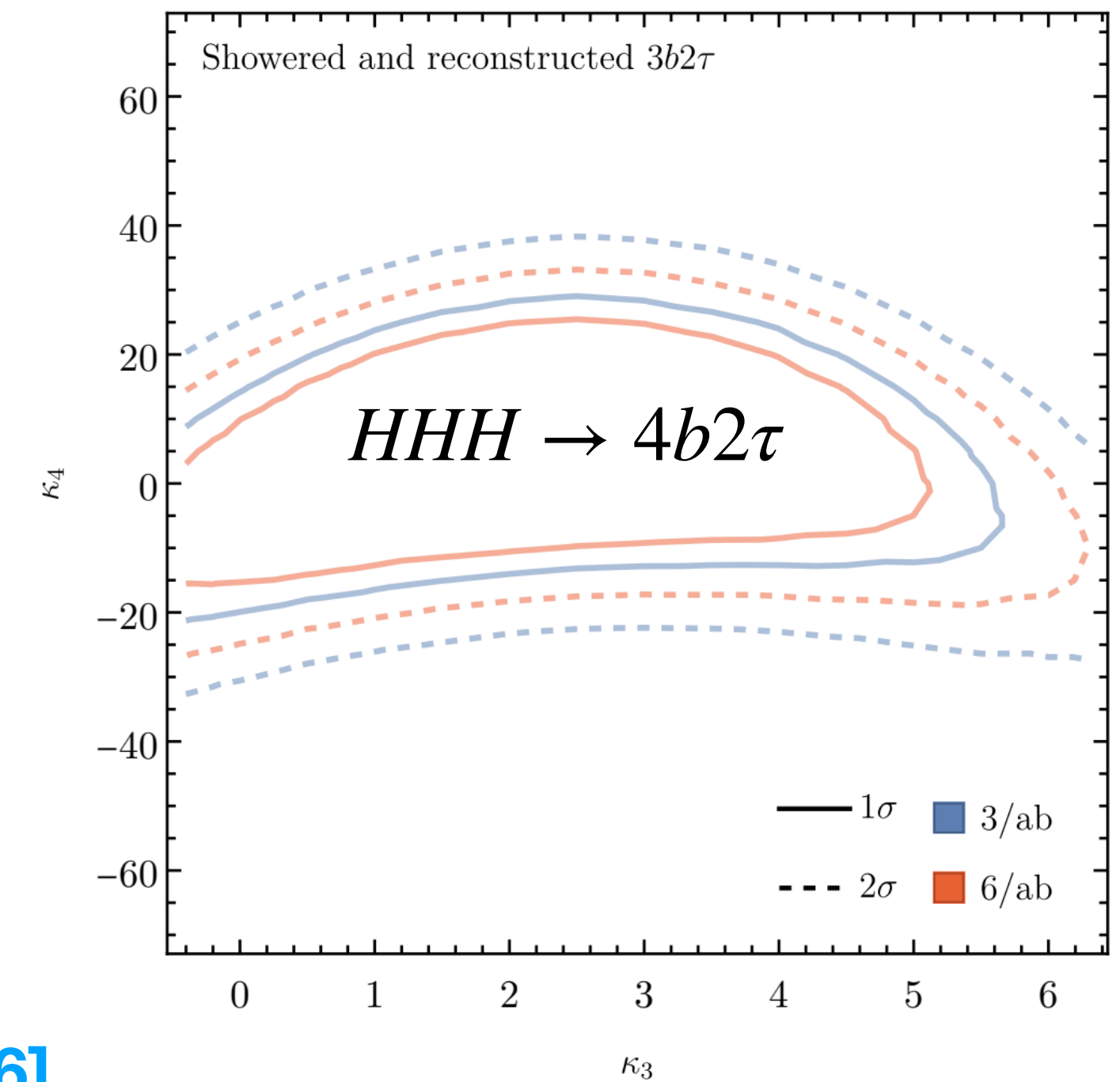
Decay modes

- Studies on sensitivity of decay modes are in (slight) conflict
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[\[1510.07697\]](#)



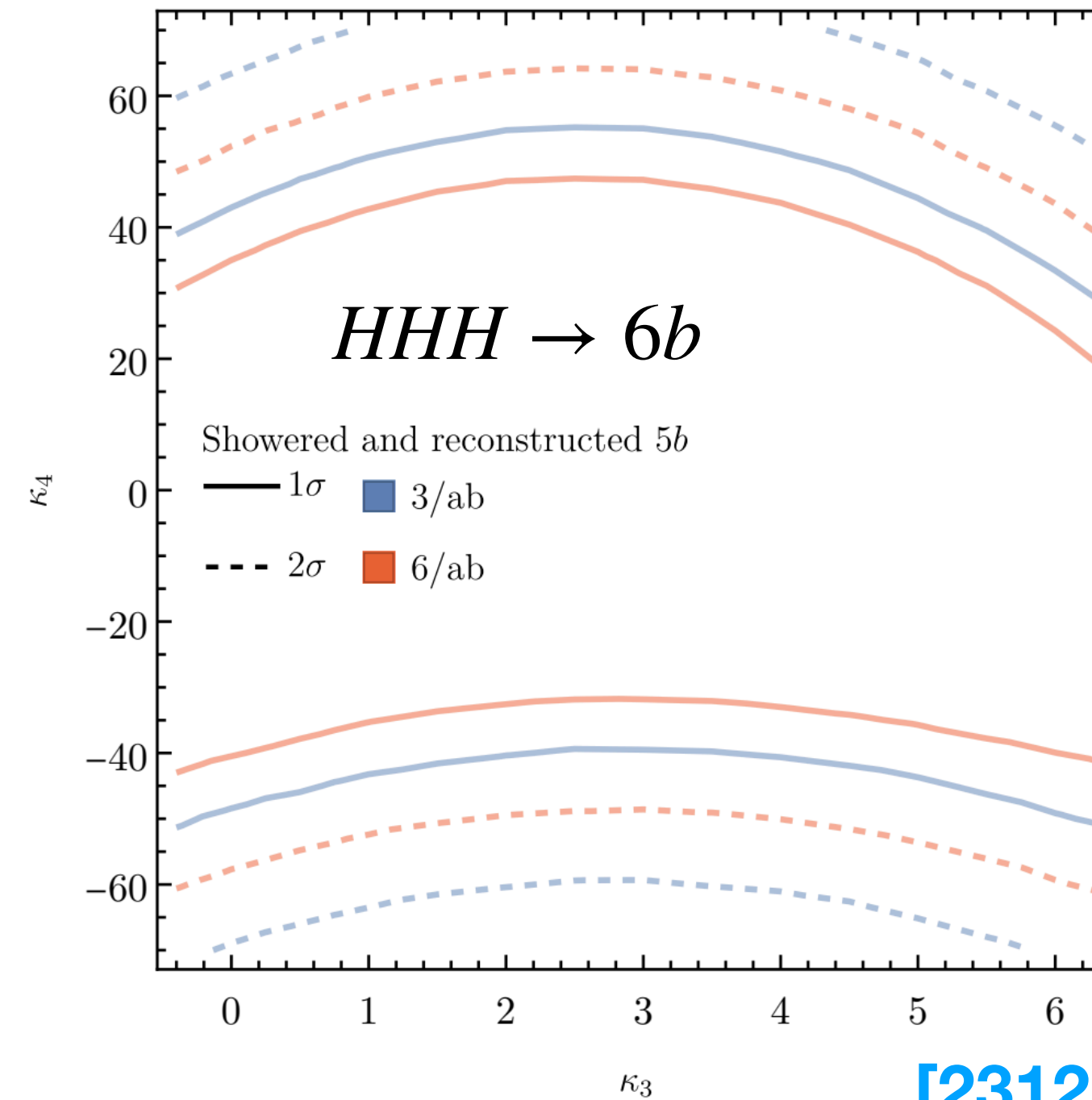
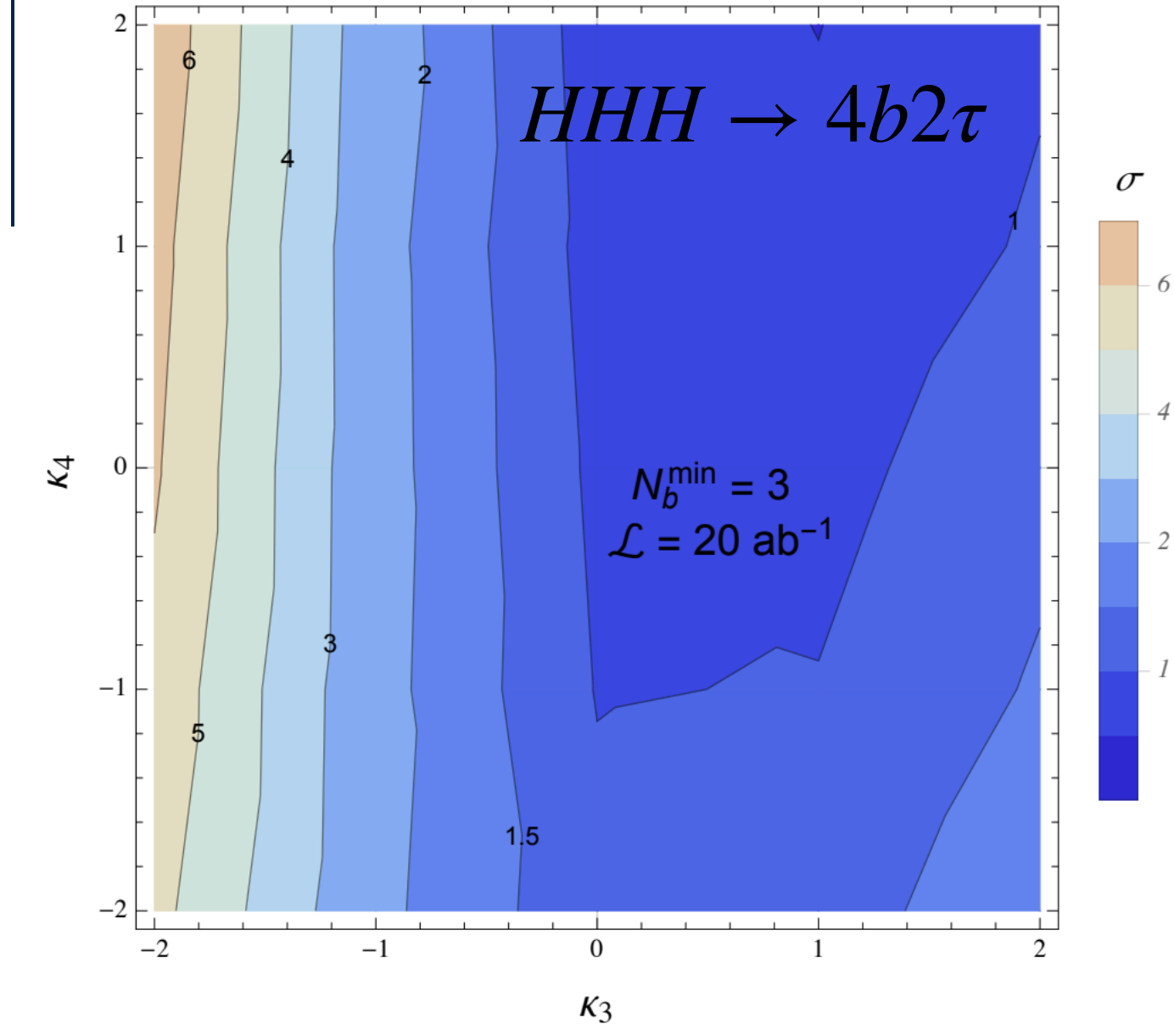
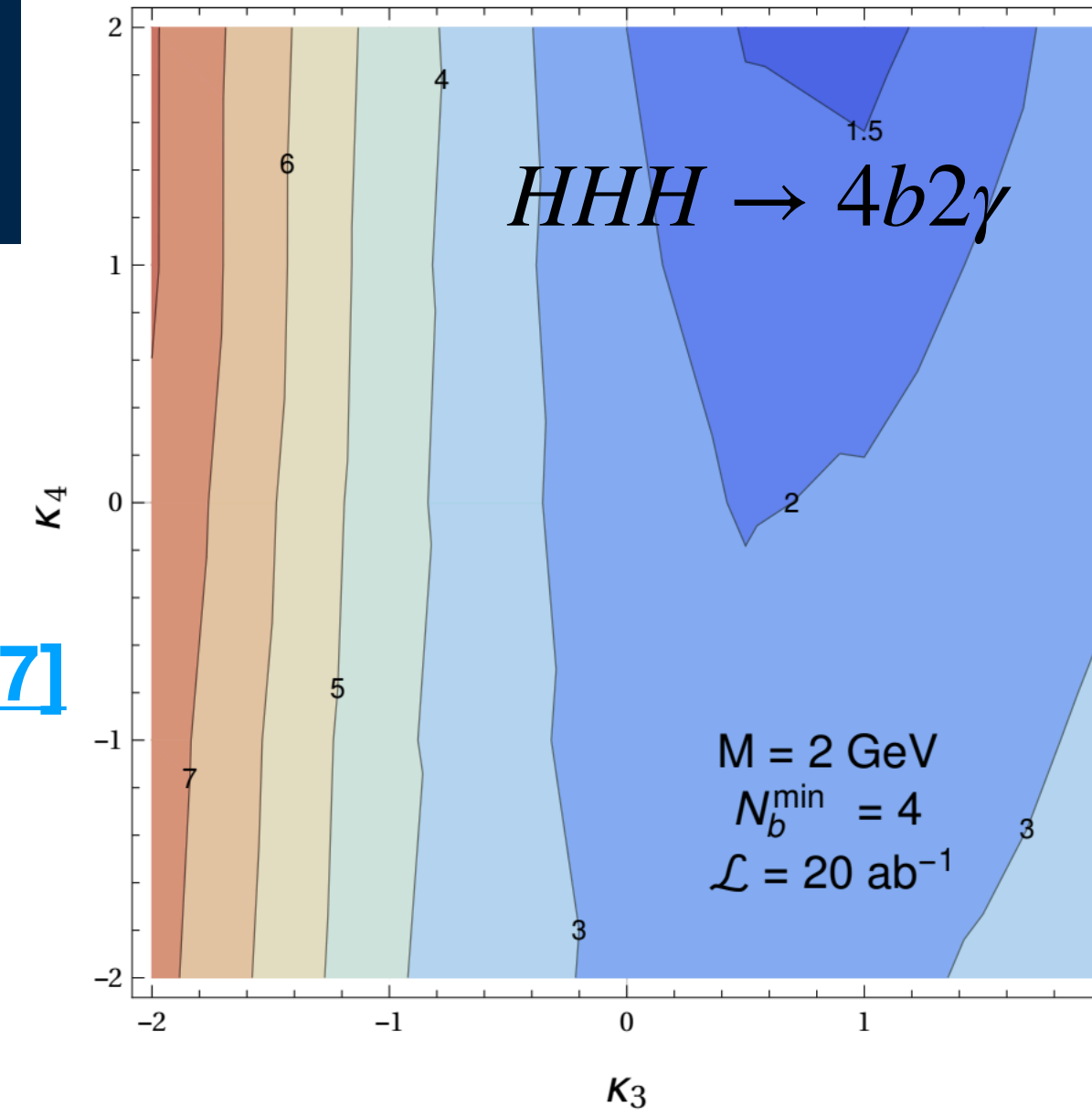
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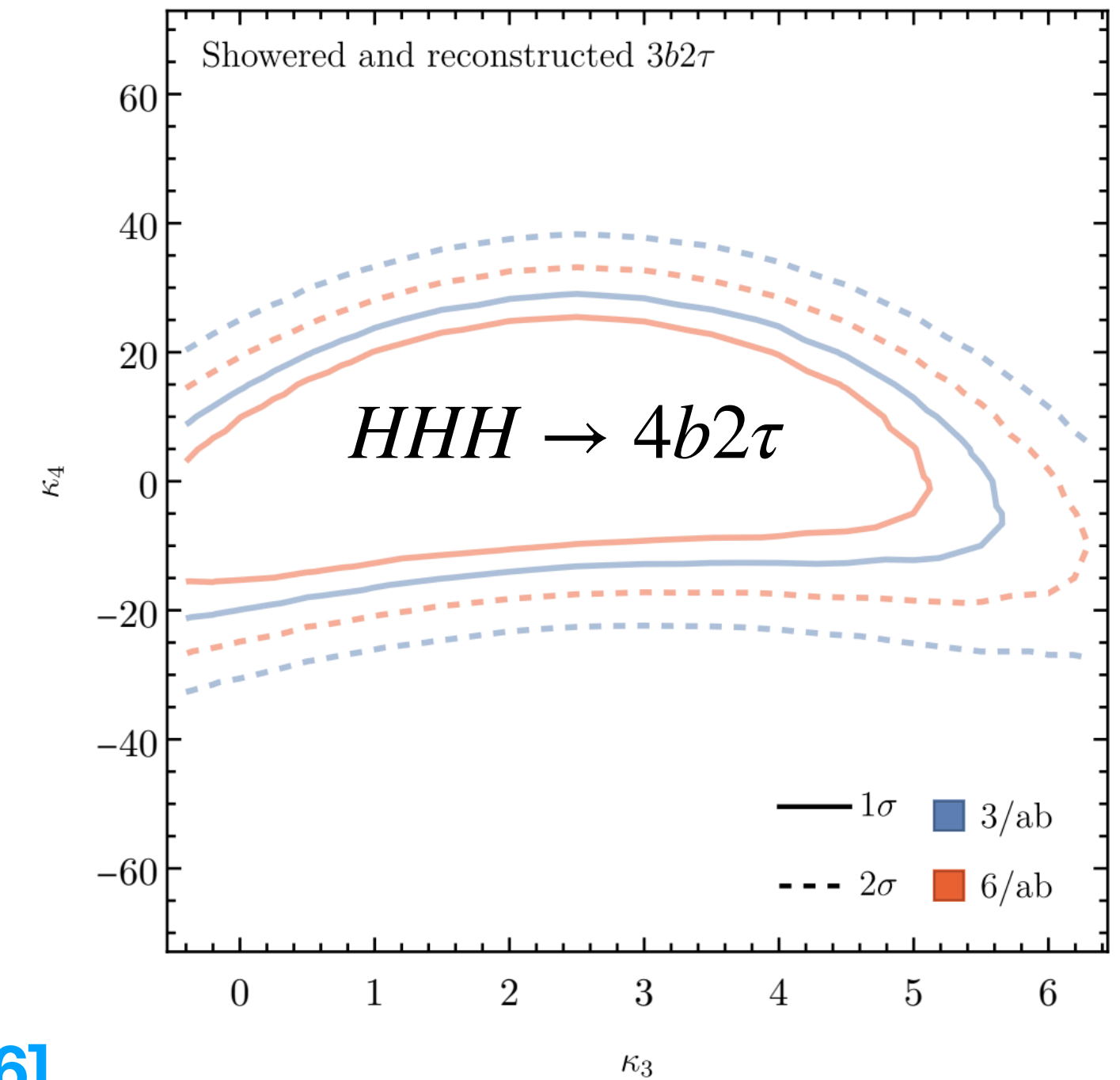
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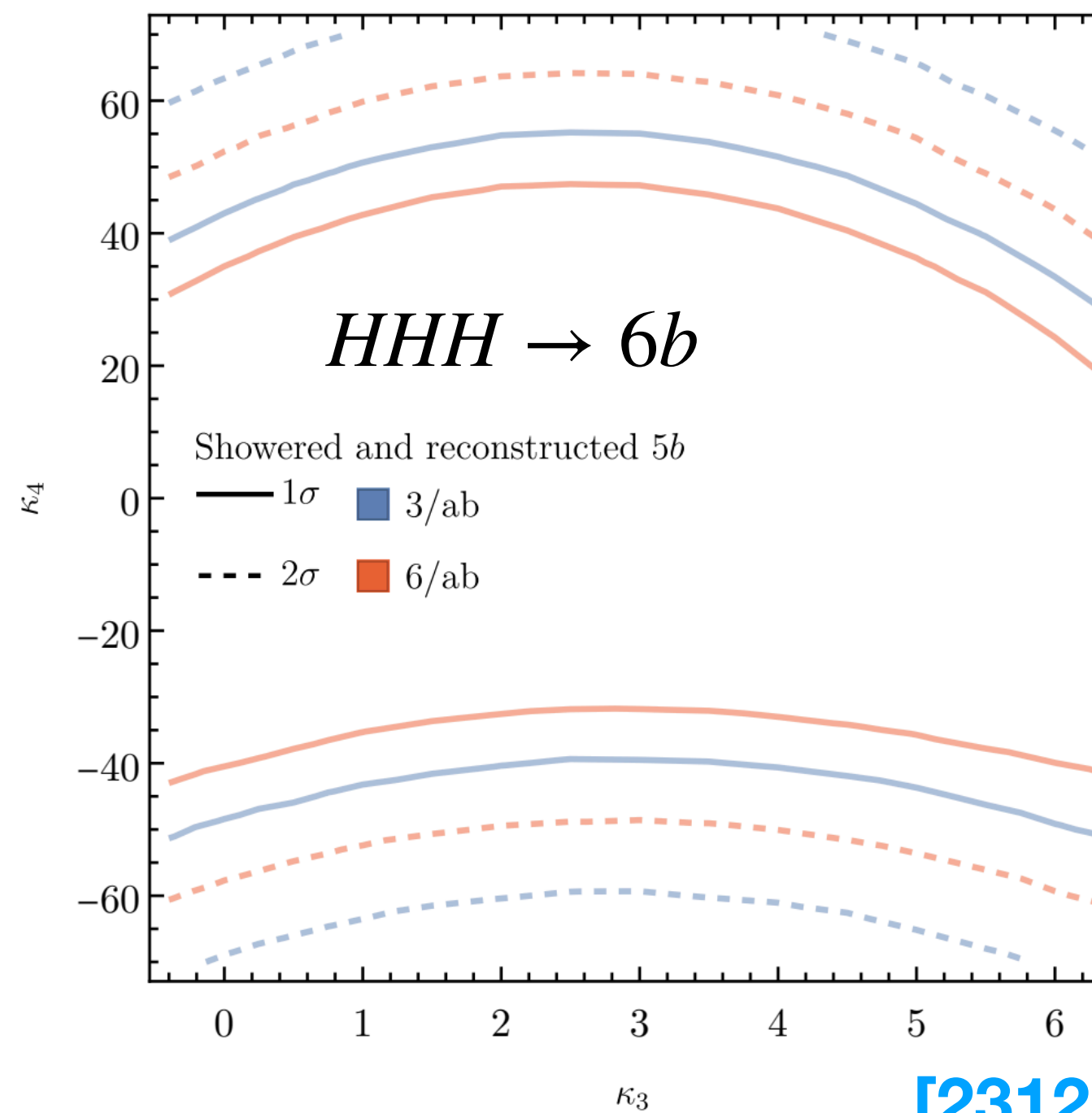
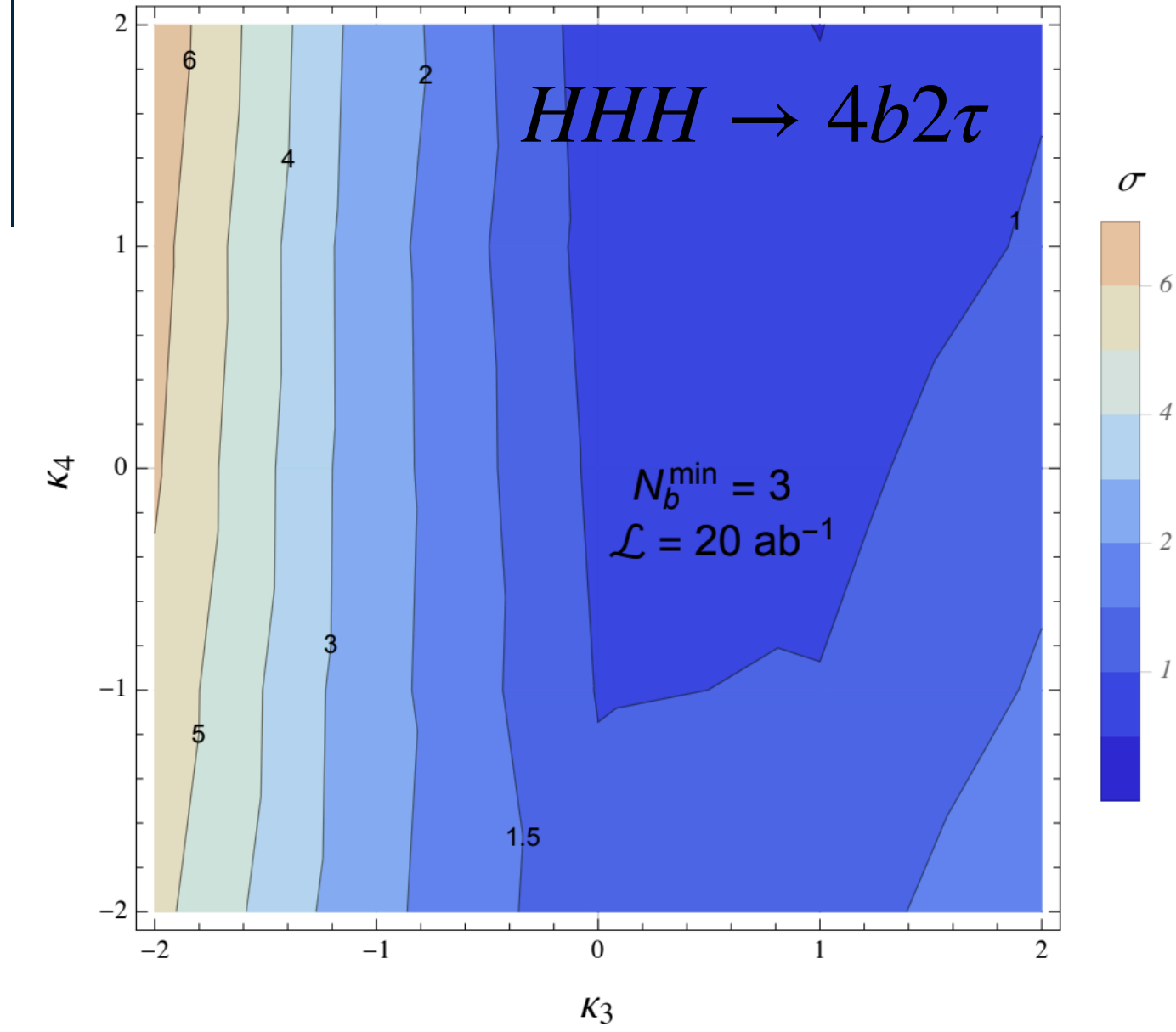
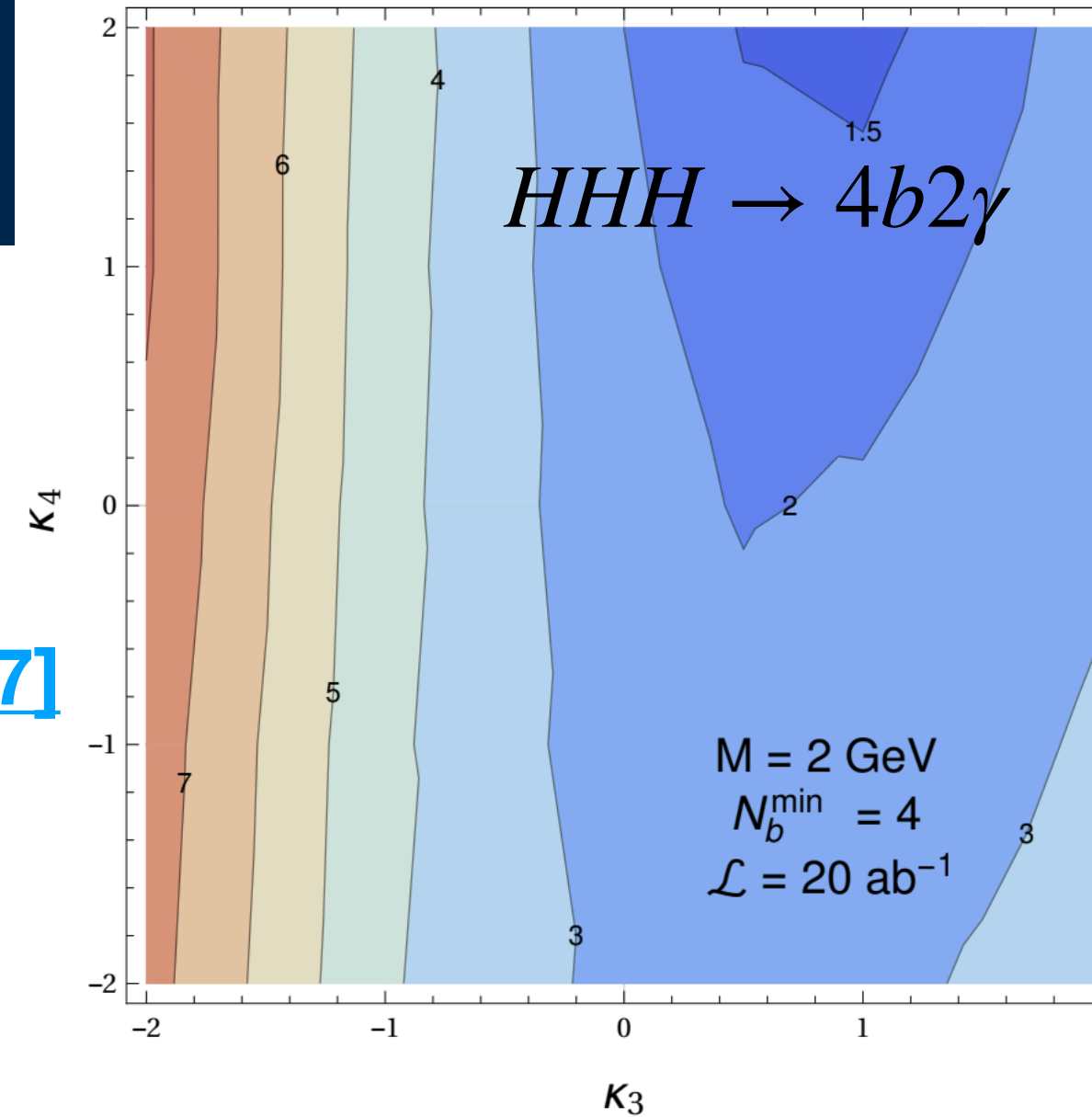
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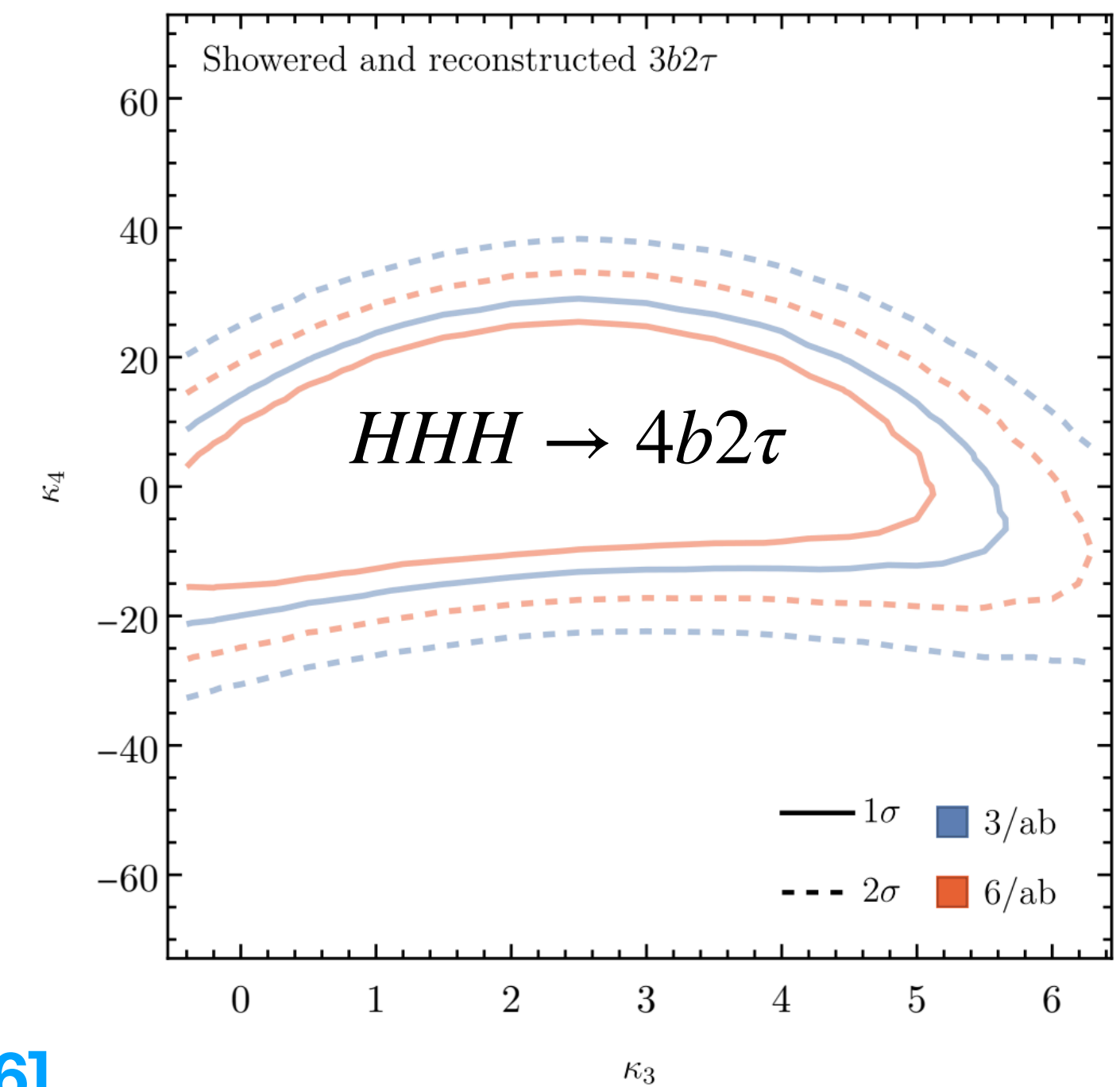
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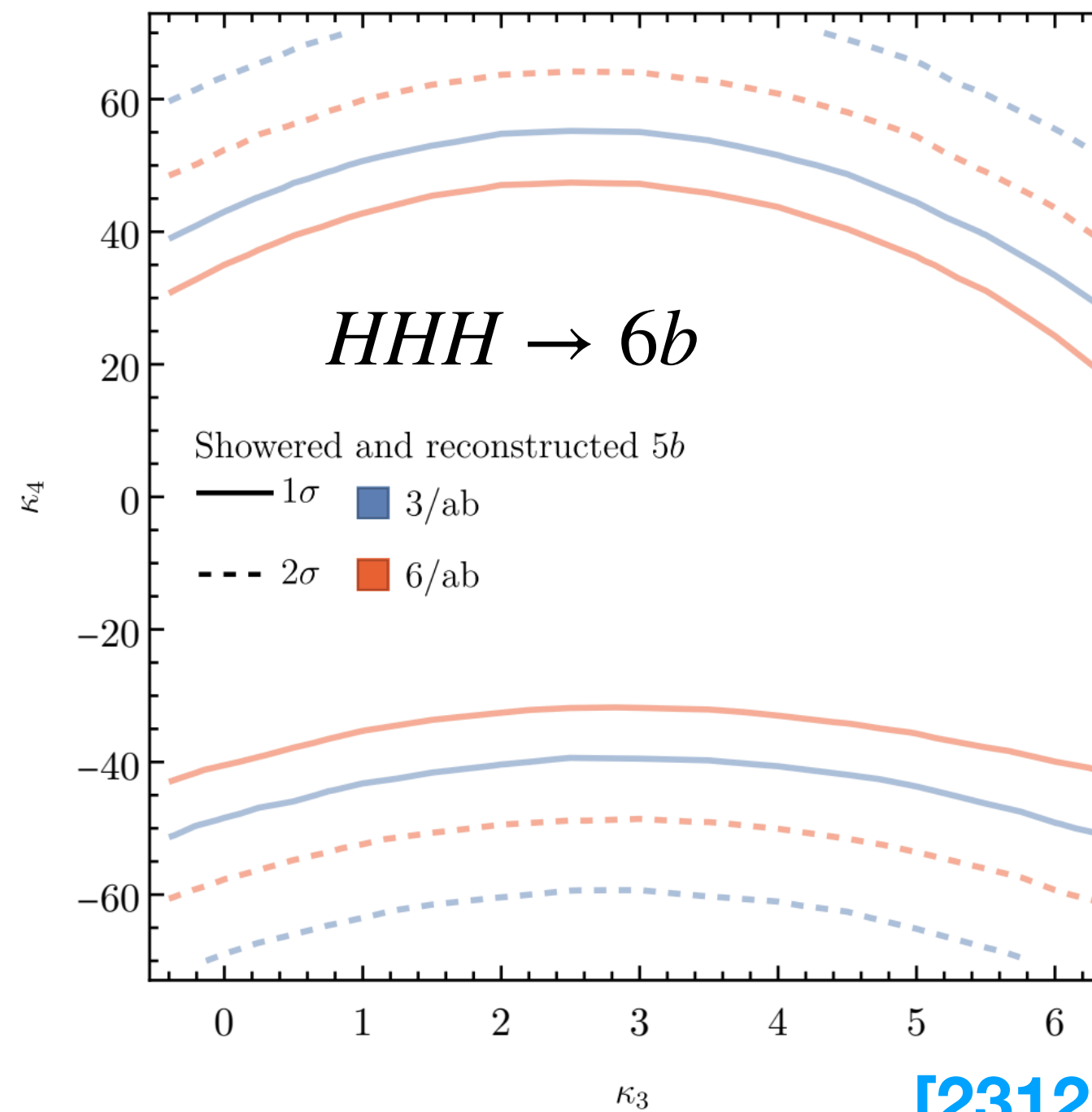
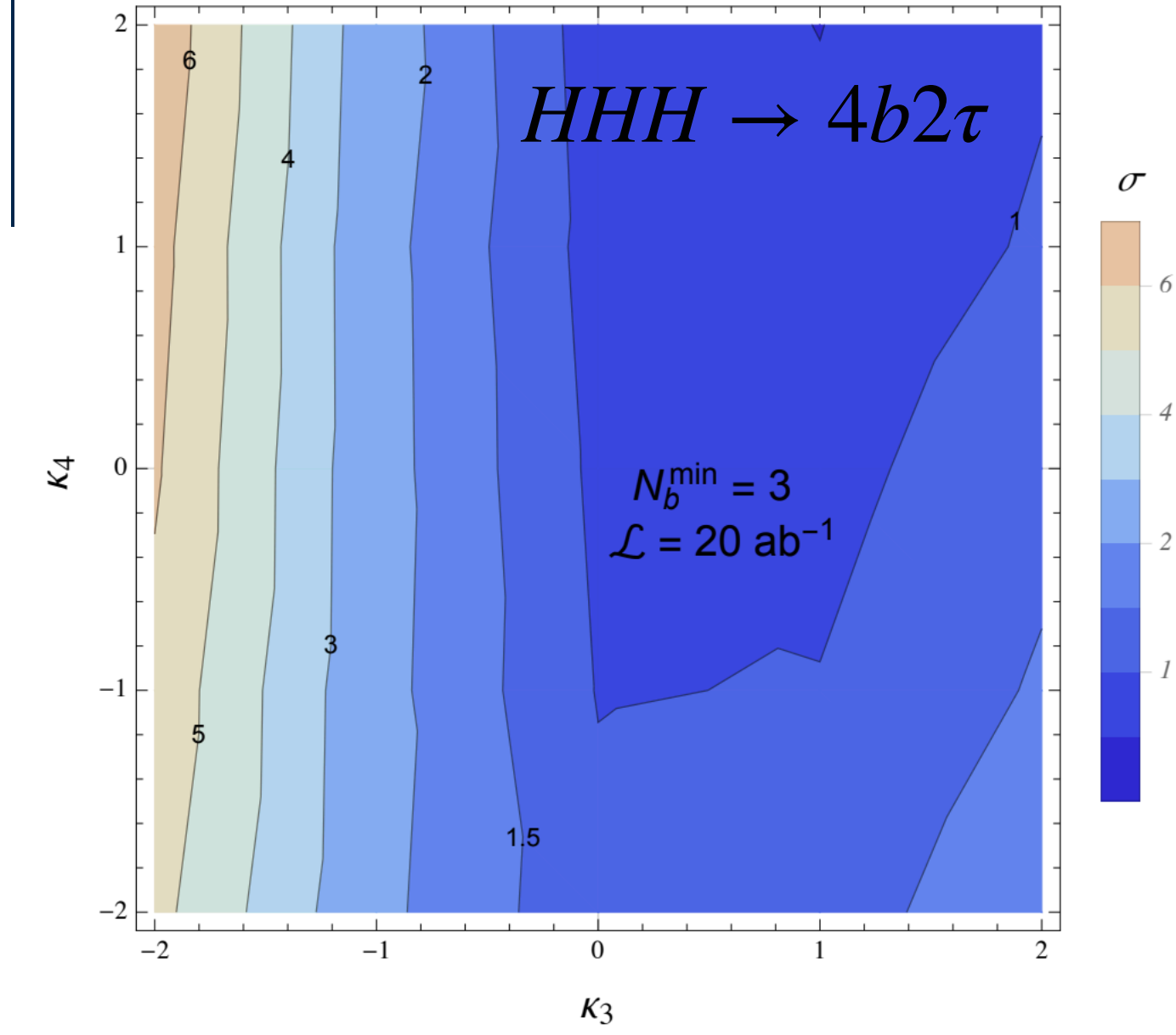
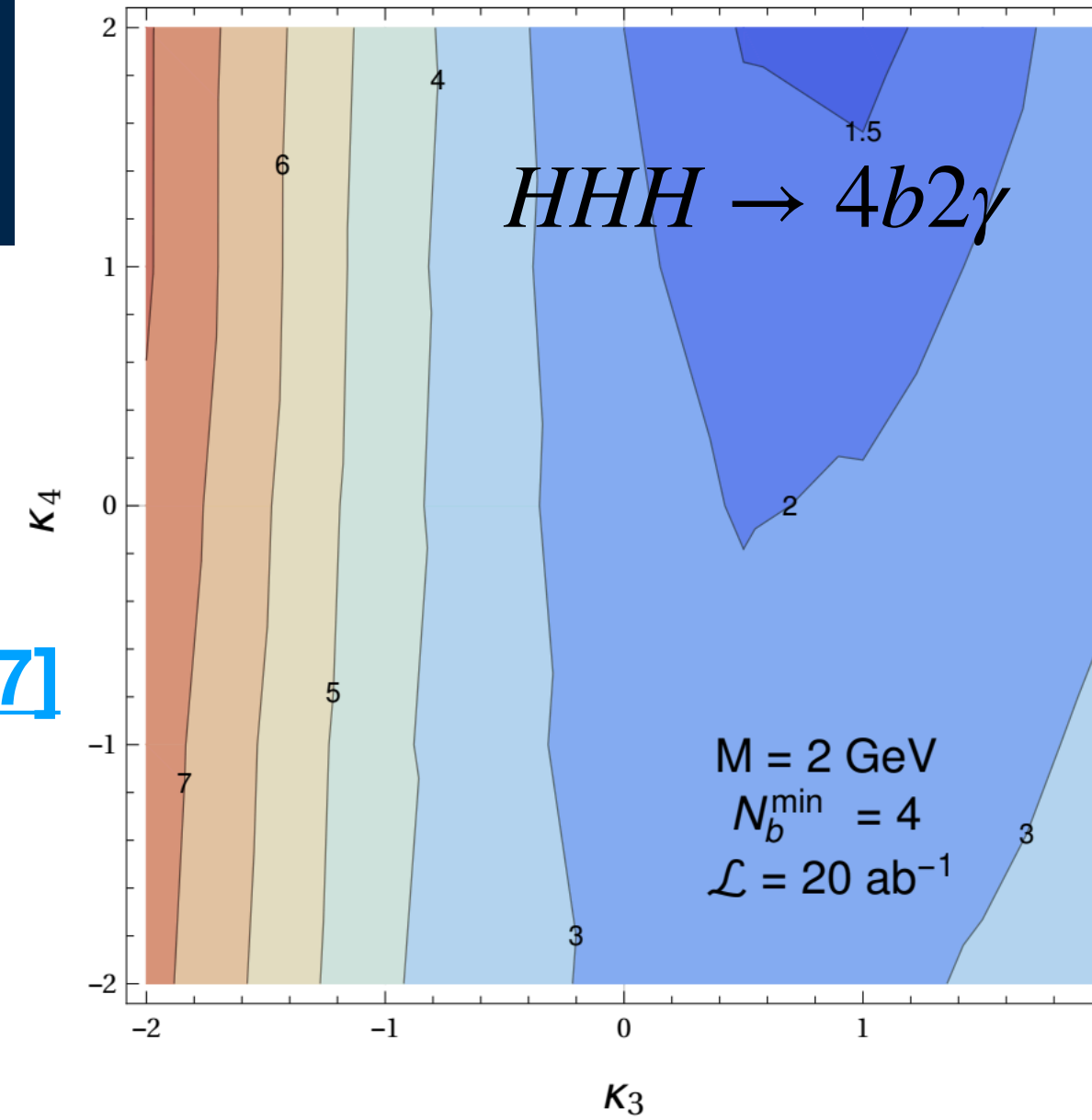
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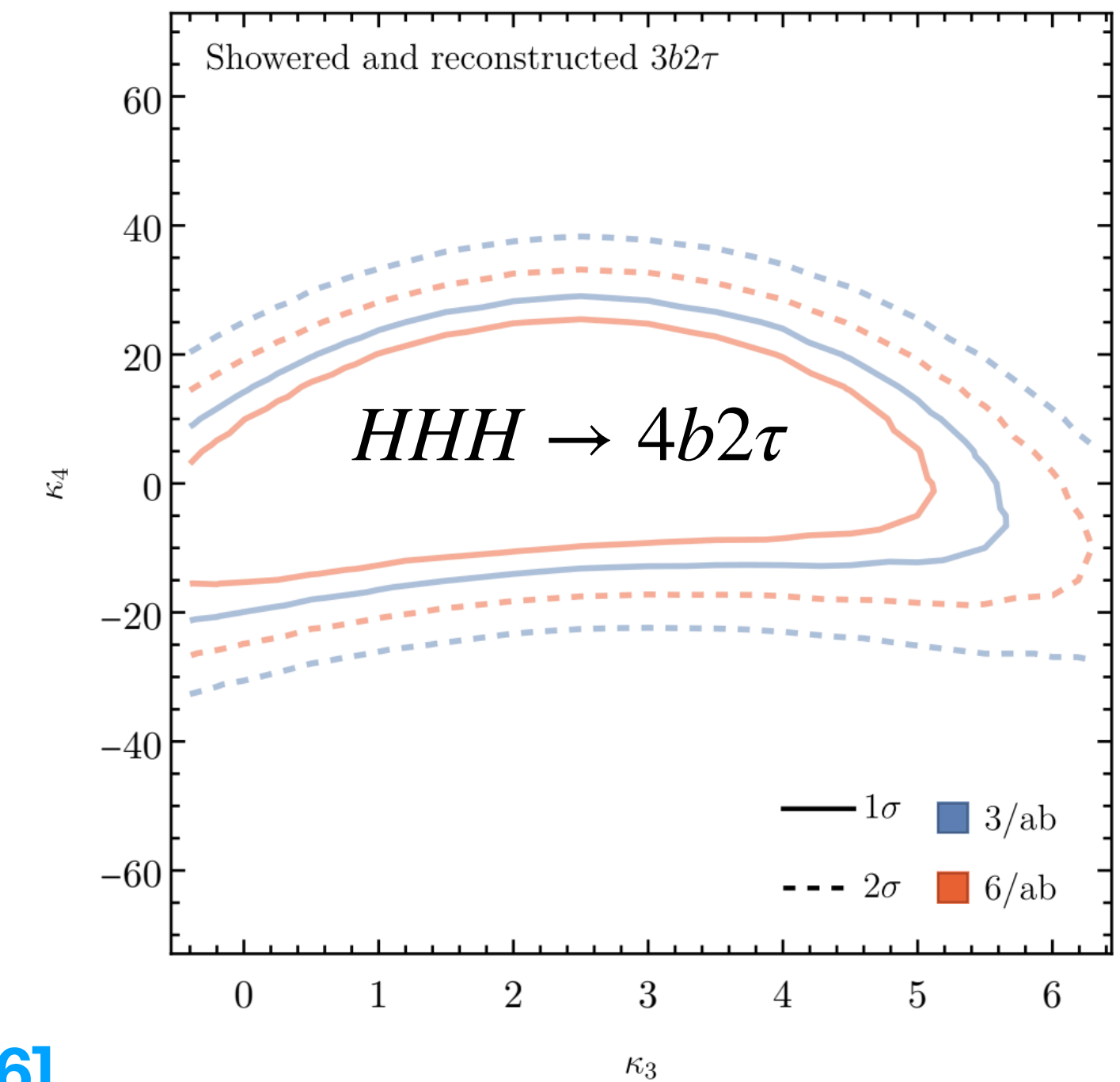
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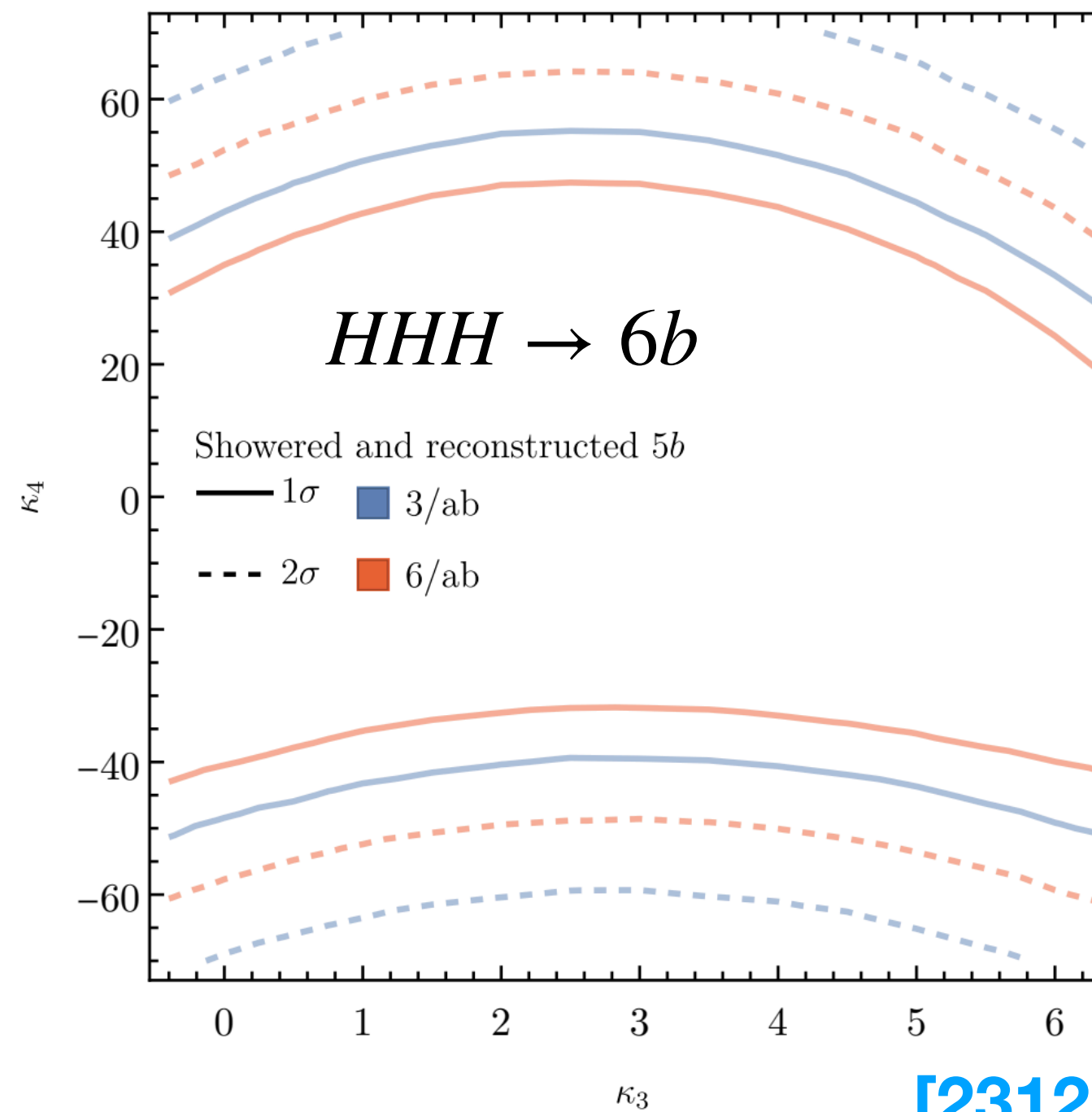
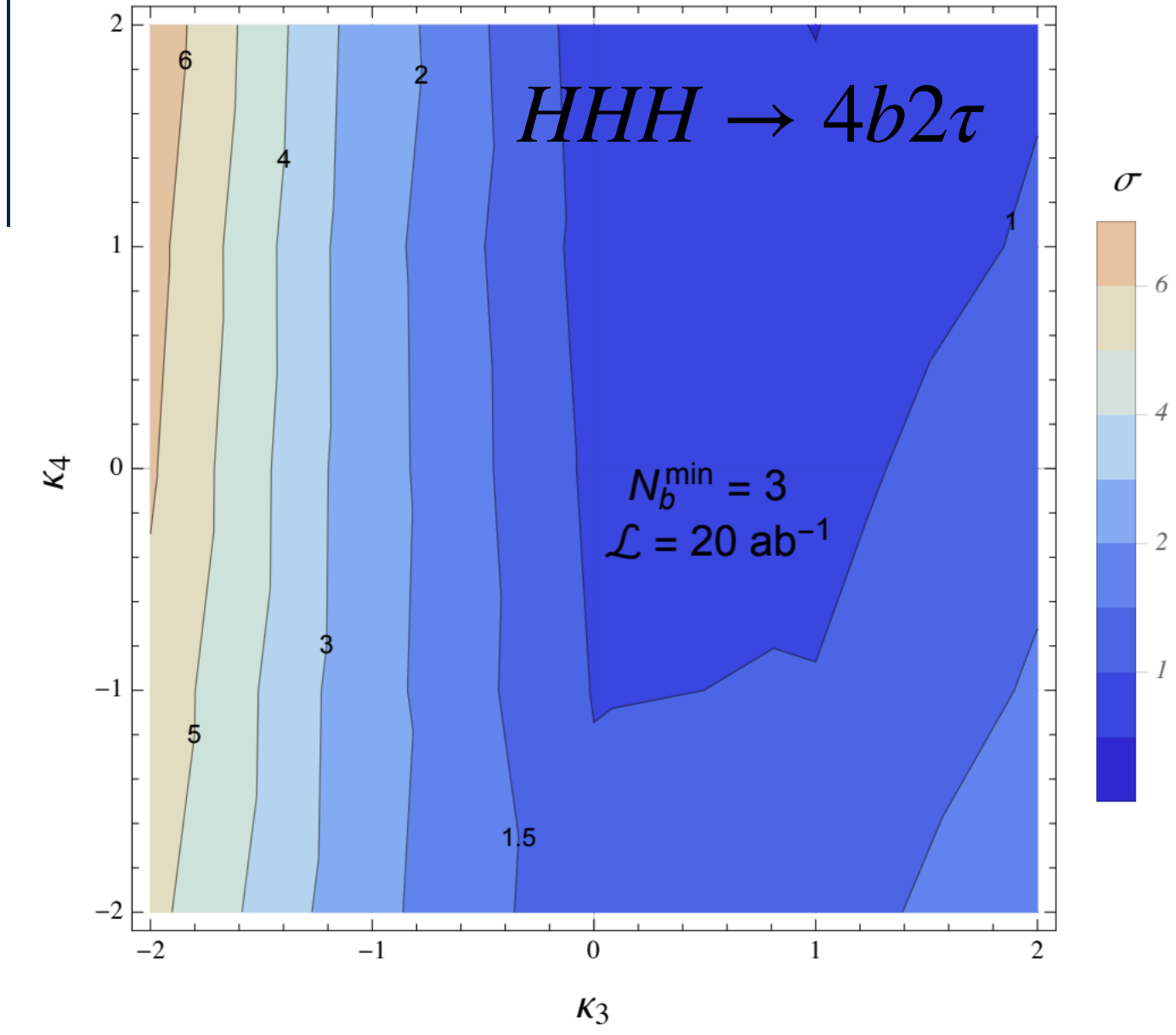
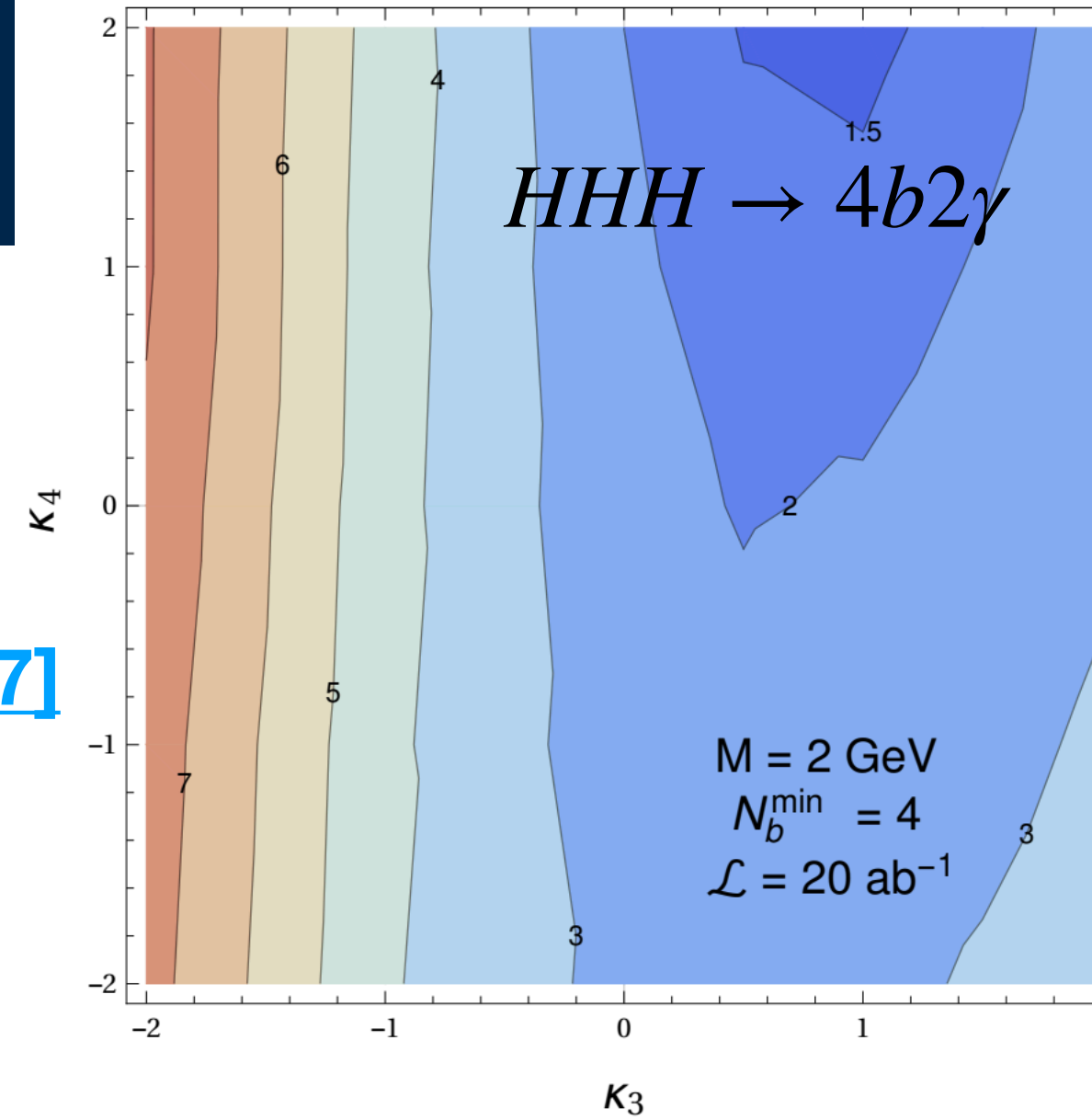
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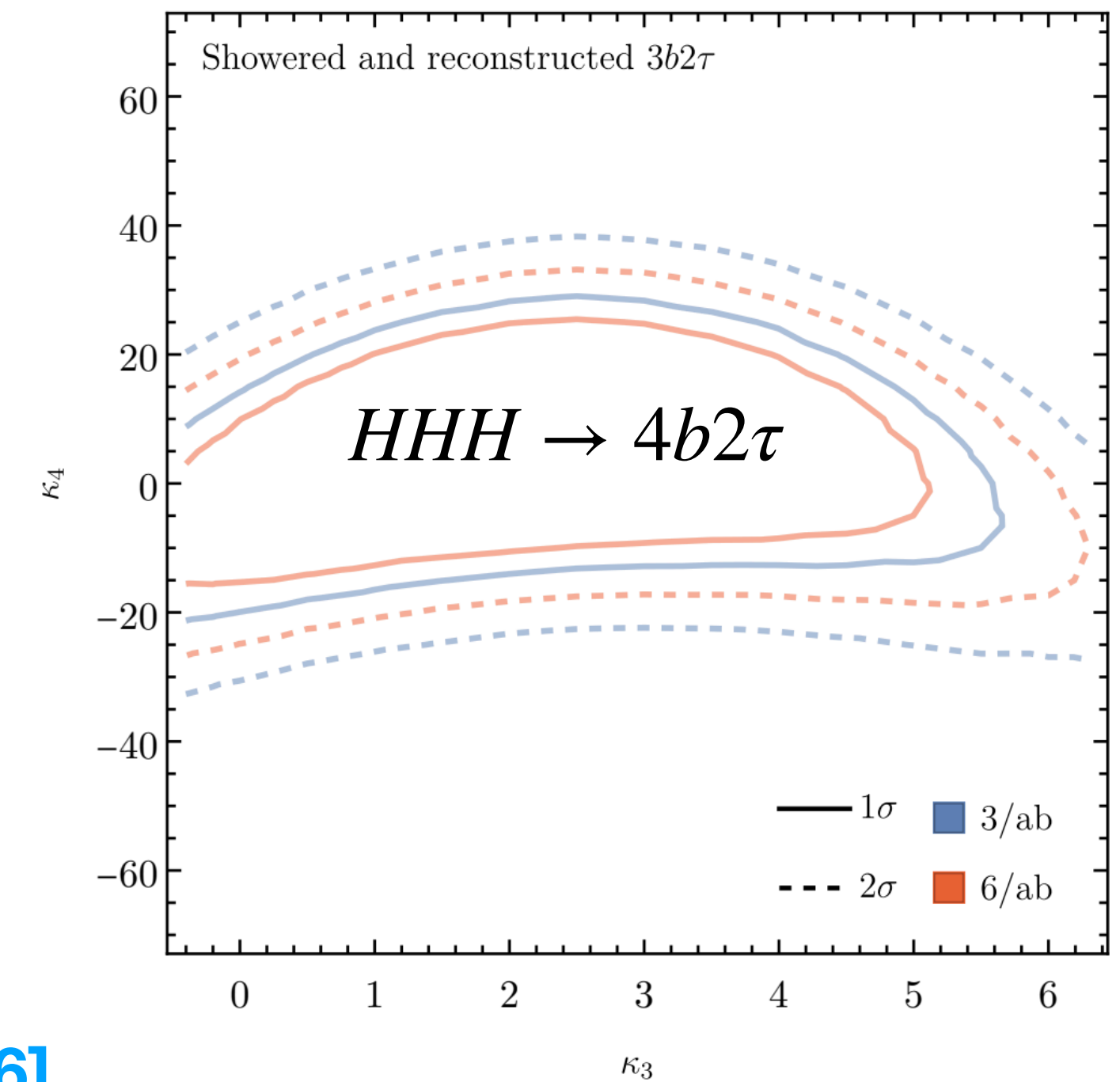
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 - Benefit from at least one lepton from many final states: WW^* , $\tau\tau$, ZZ^*
- ATLAS chose $6b$ for 2 reasons:
 1. Our projections showed this was best
 2. The team had expertise from $HH \rightarrow 4b$

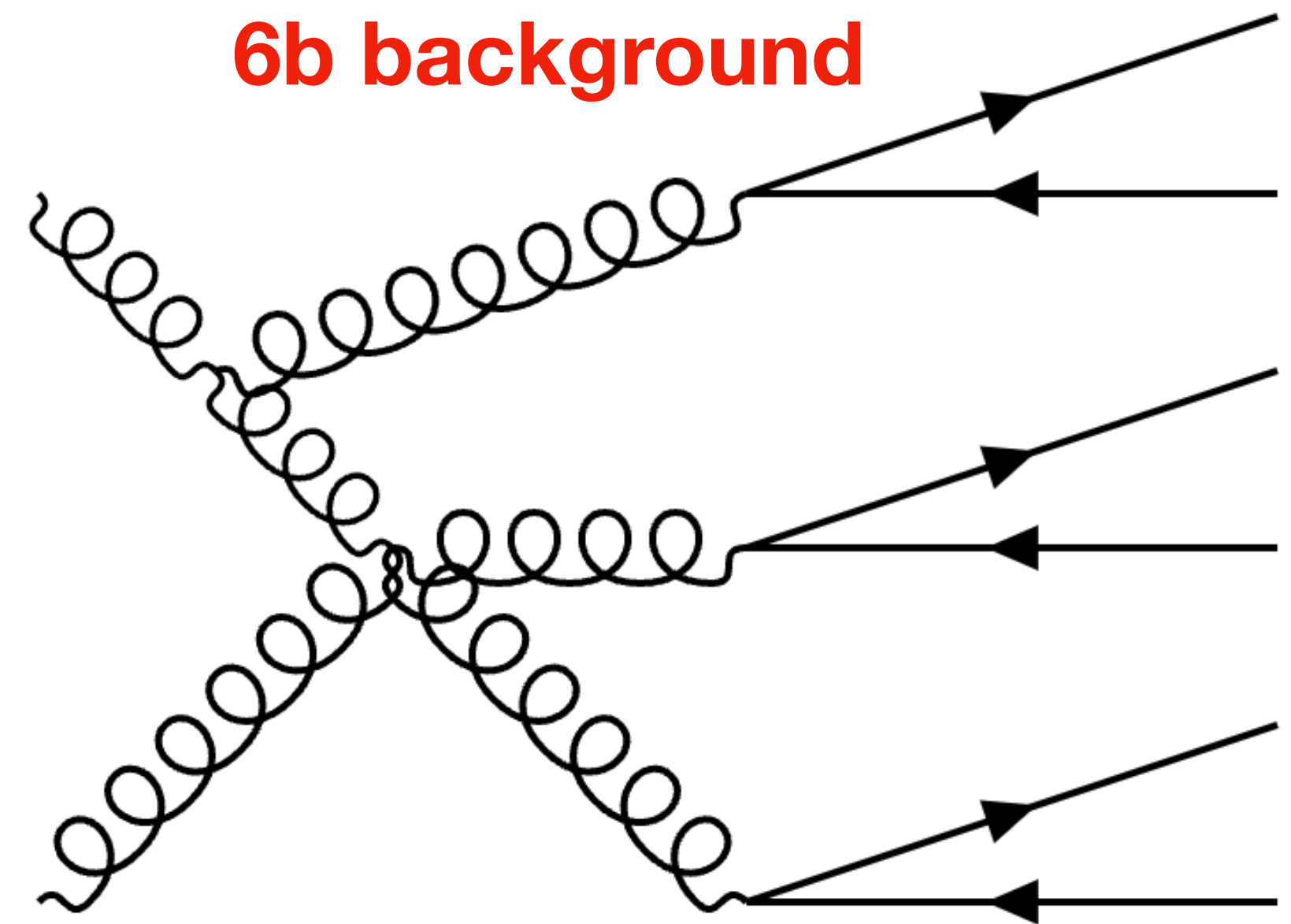
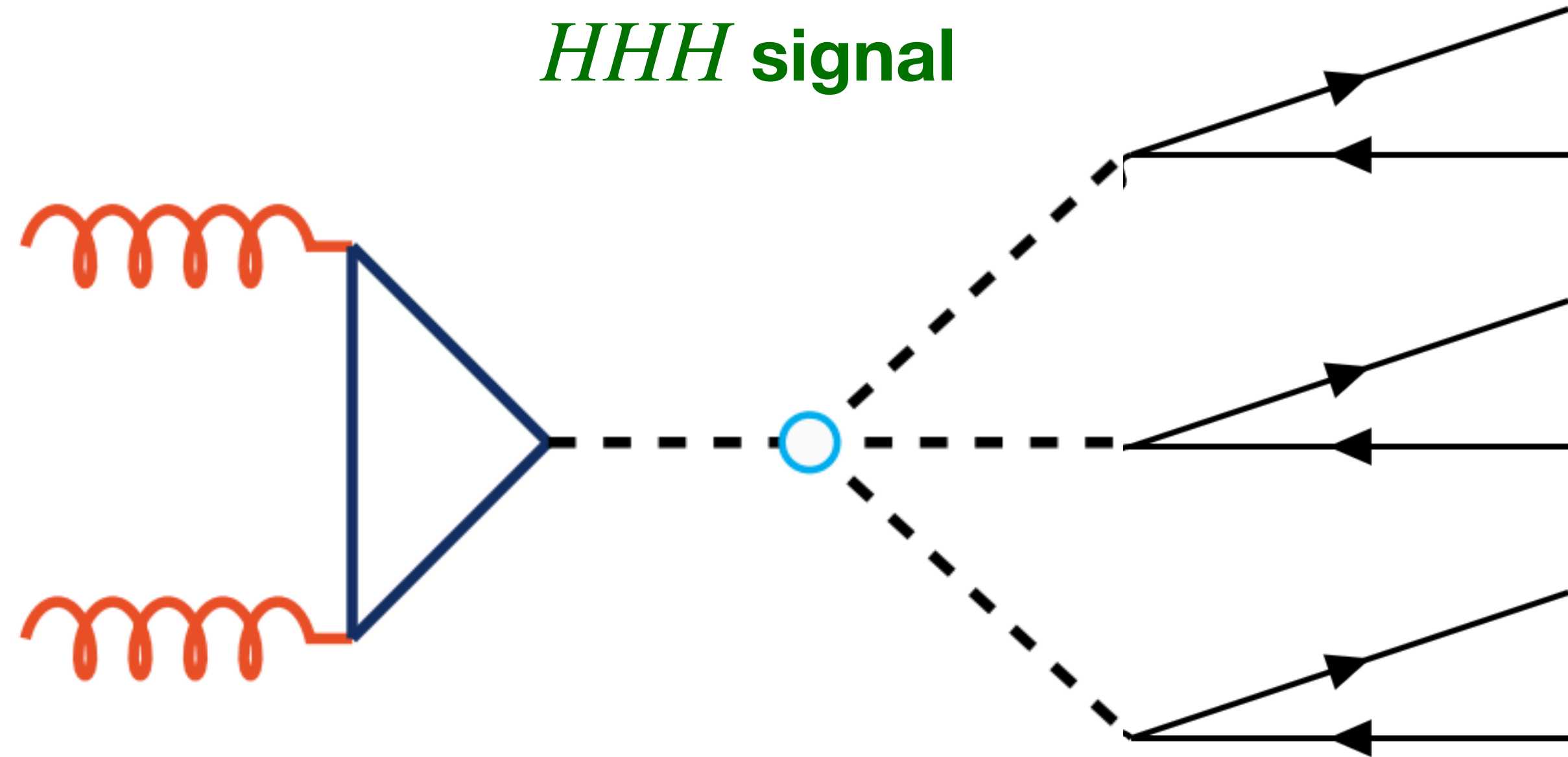
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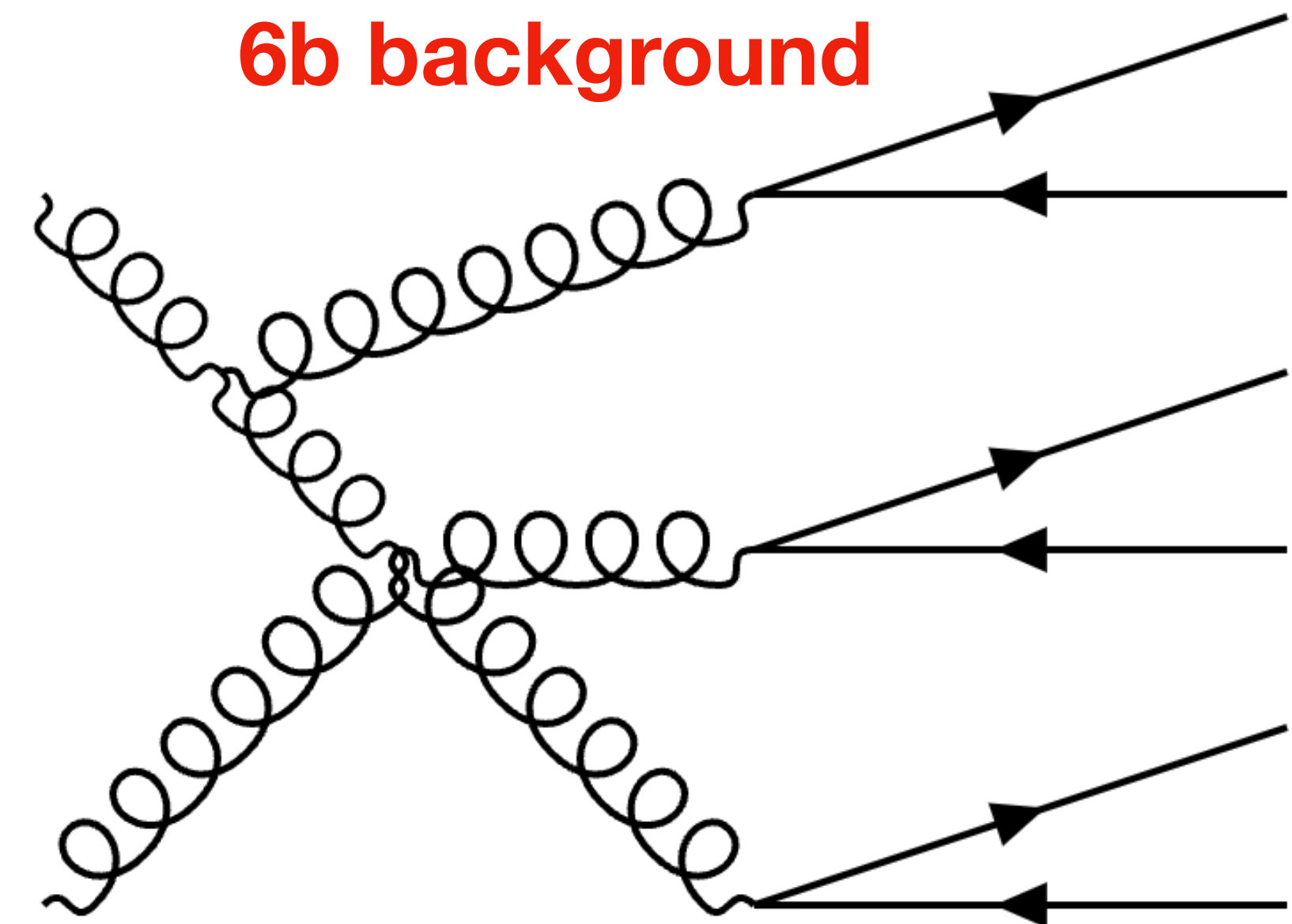
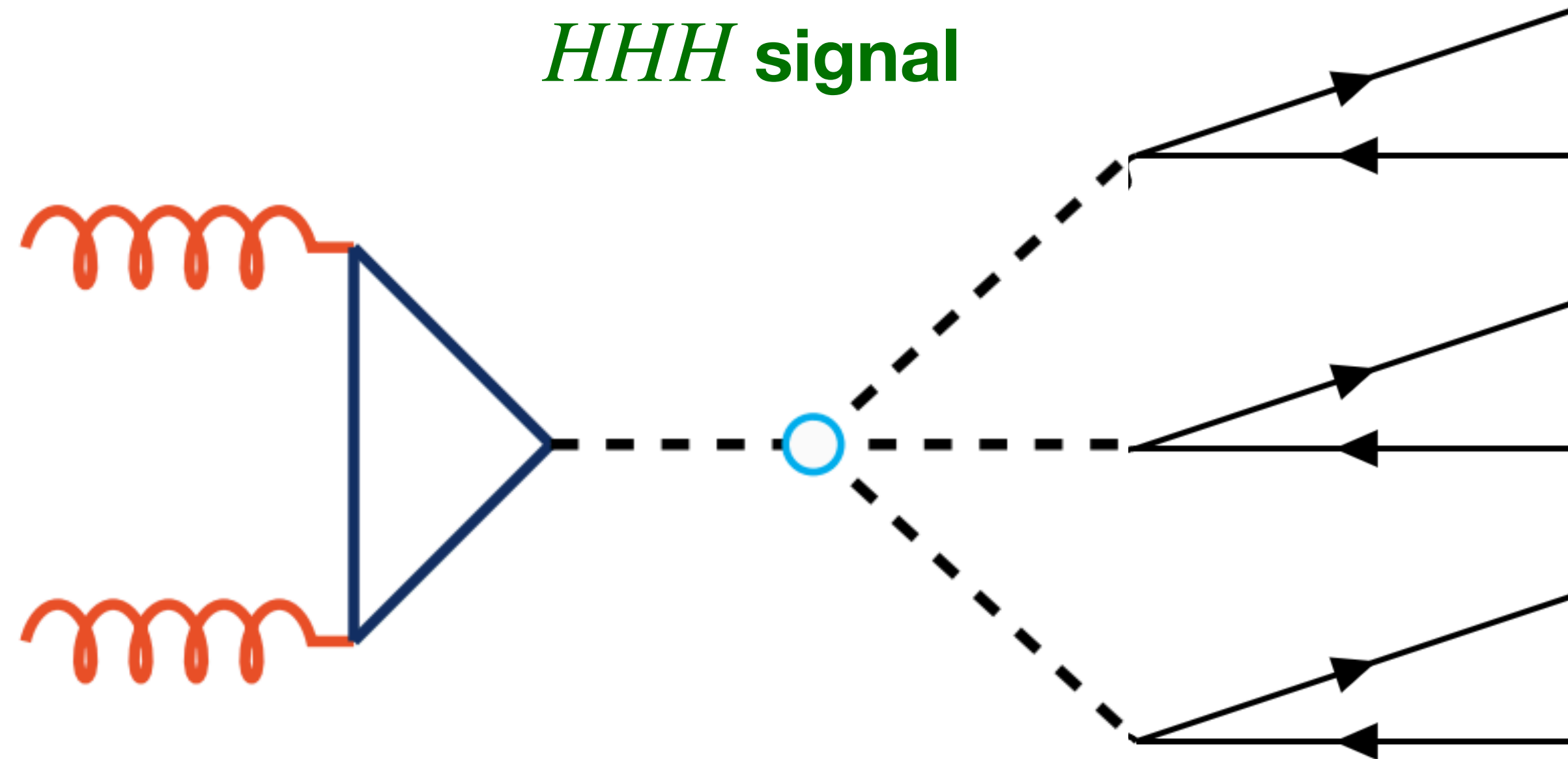


Signal and background



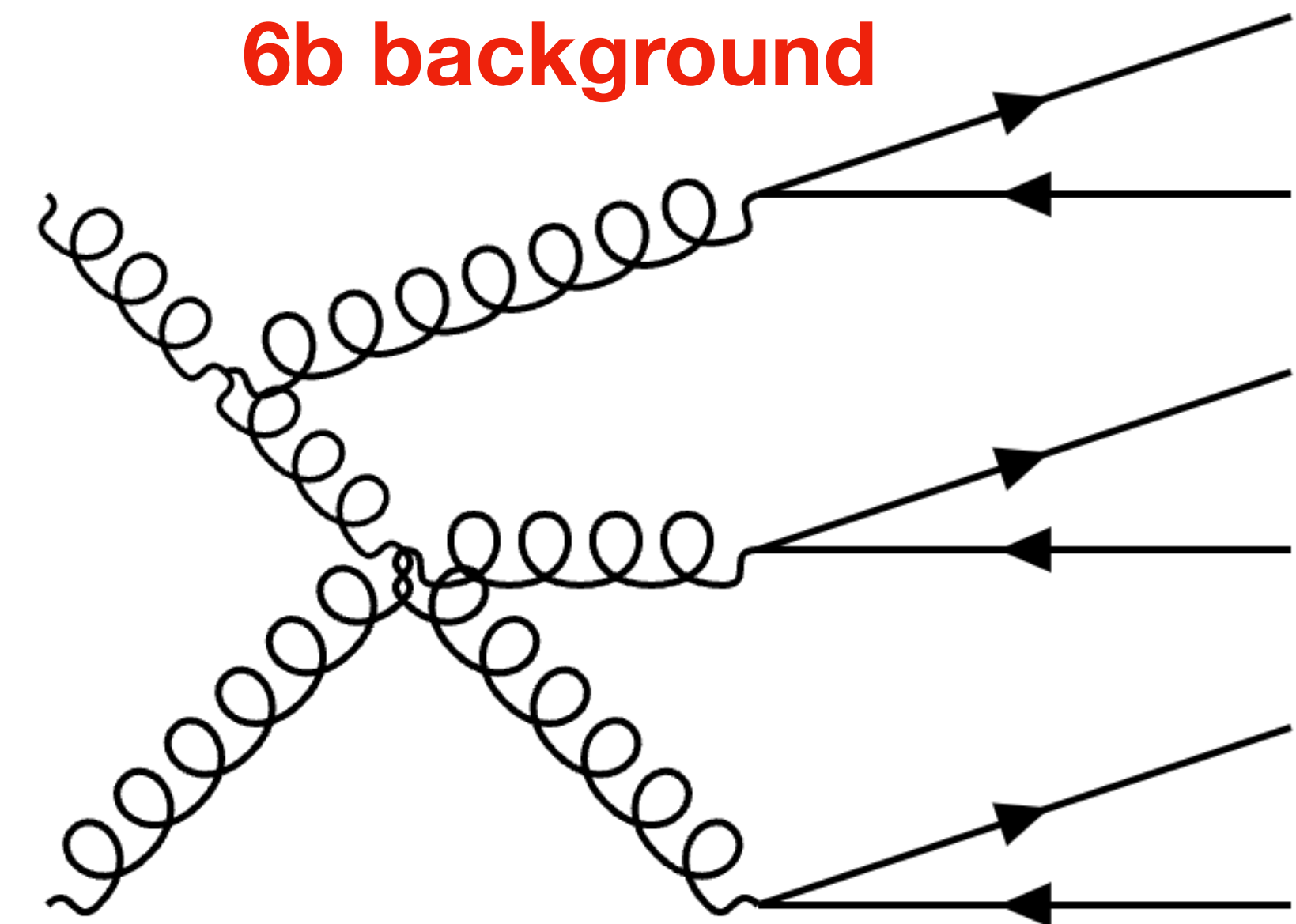
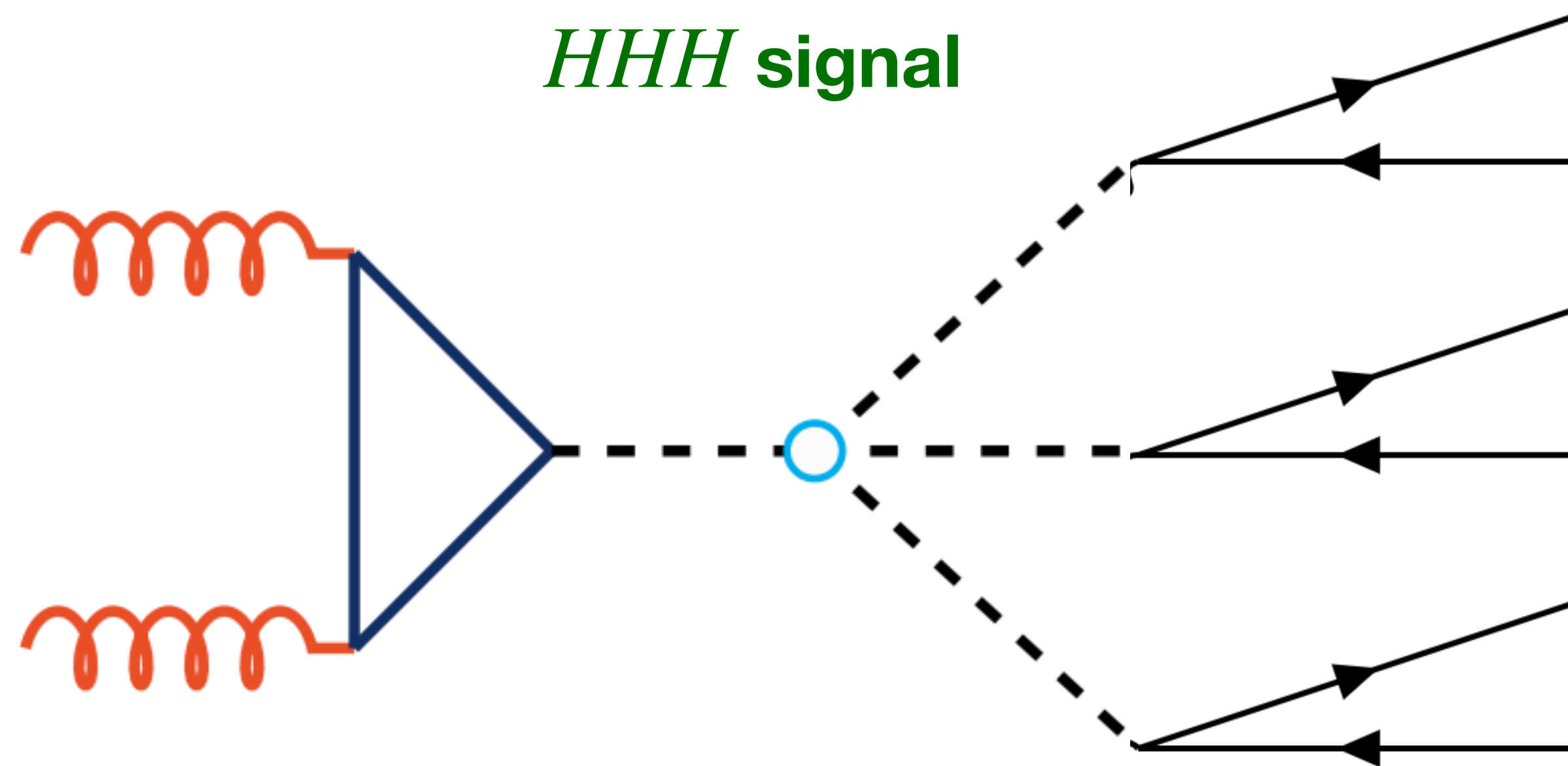
Signal and background

- We search in the 6b final state
- Primary background is from QCD multi-jet production



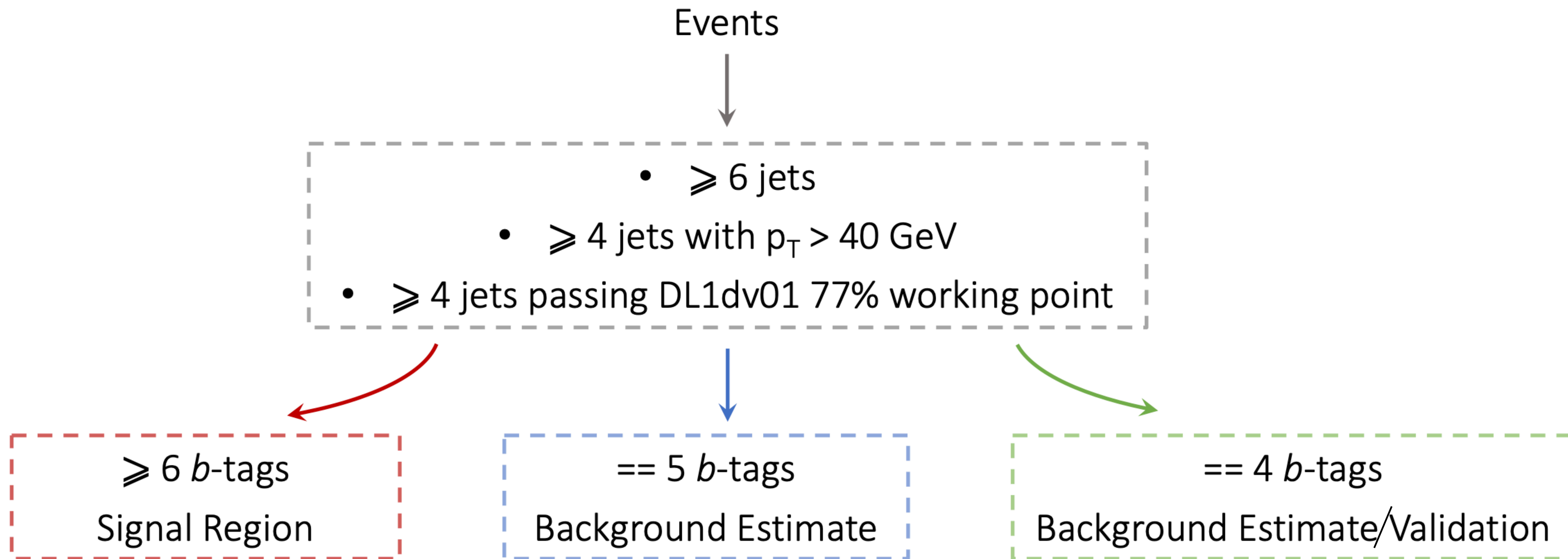
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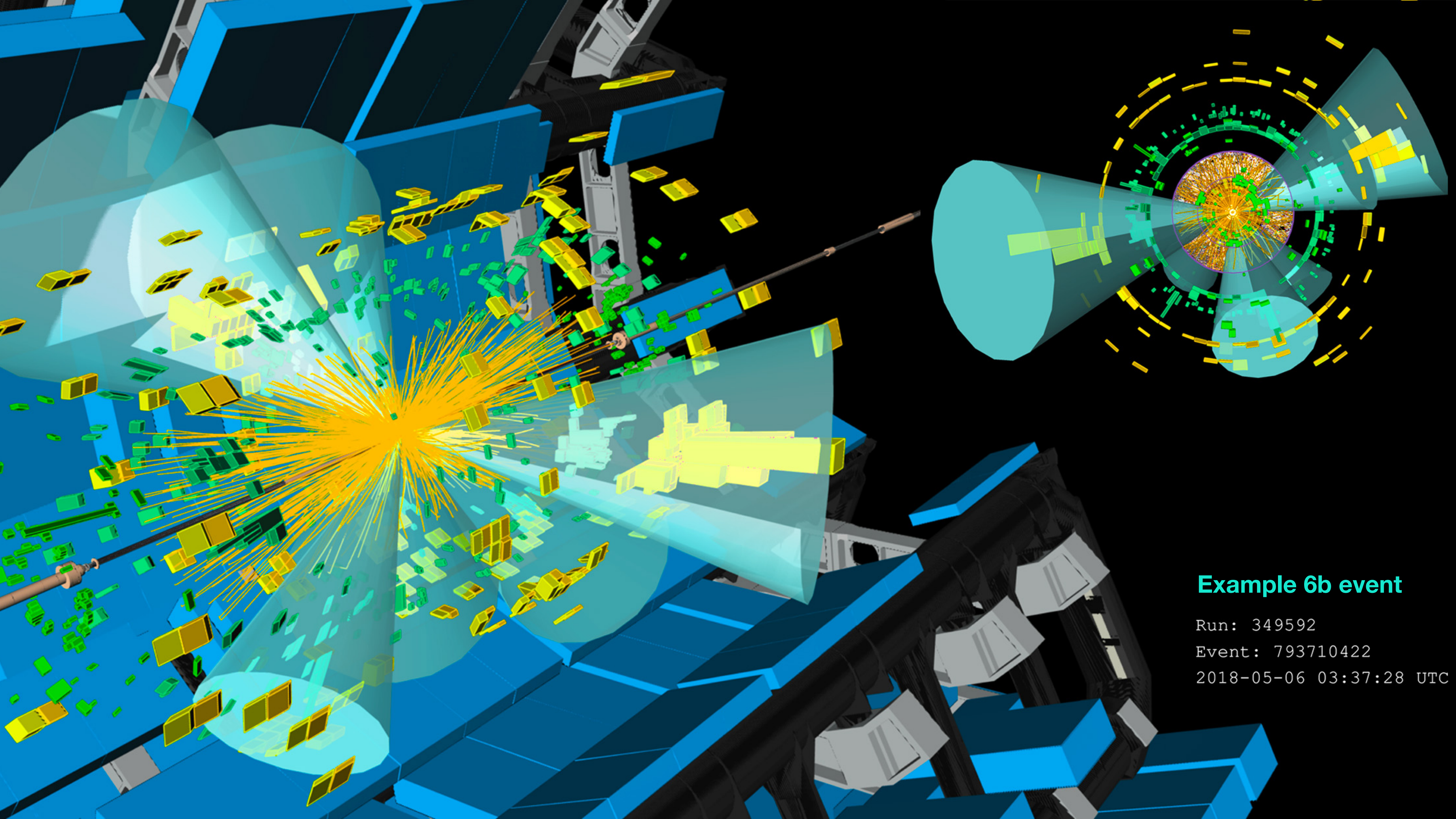
- We search in the 6b final state
- Primary background is from QCD multi-jet production
- Other backgrounds like $t\bar{t} + HF$ still have many additional b-jets and are difficult to model with MC
- Motivates a data-driven background estimate



Event selection

- Trigger on the jets/b-jets (some online b-tagging is used in the trigger)
- Select all events to have at least 6 jets, categorize into the 4b, 5b, 6b regions





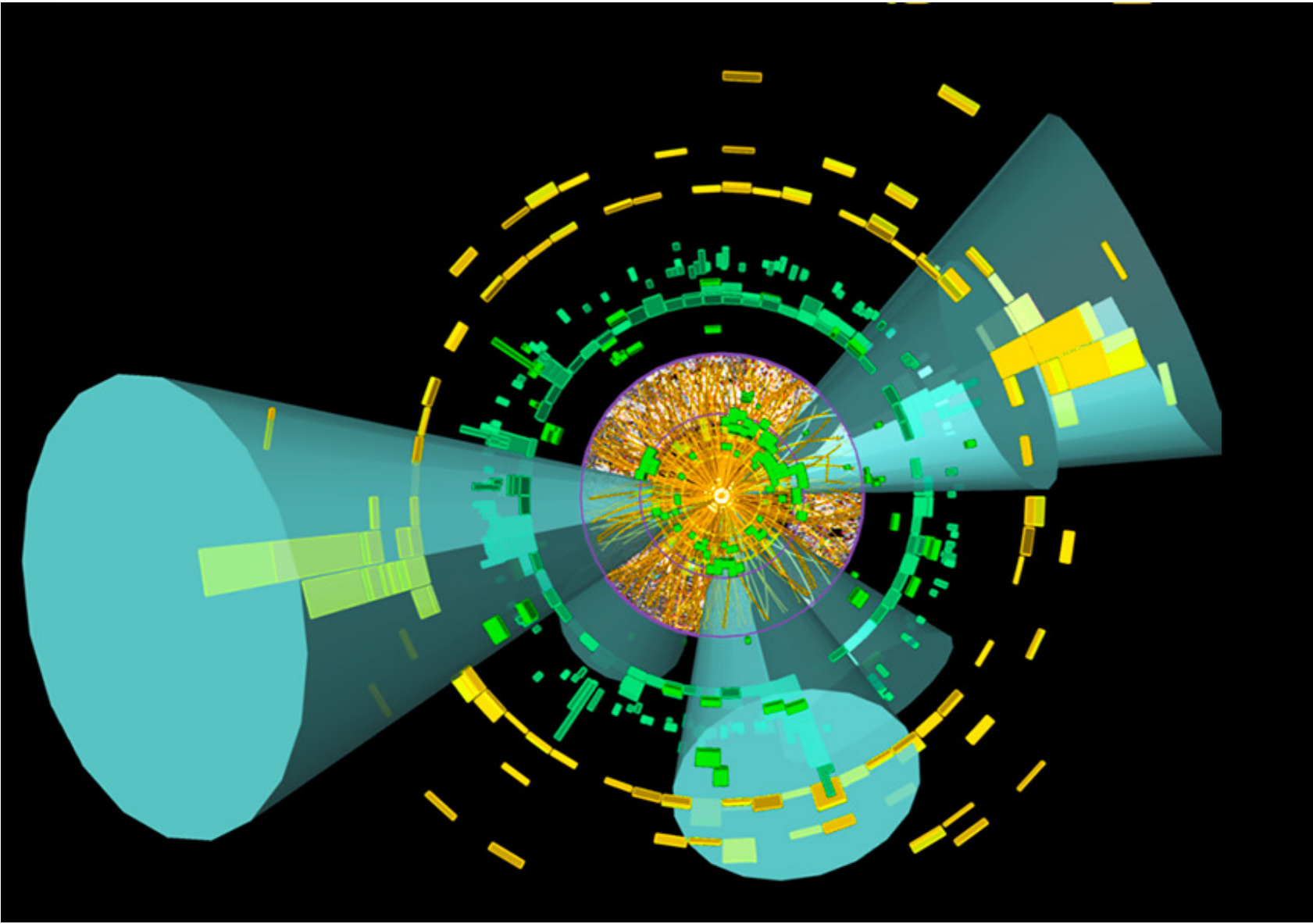
Example 6b event

Run: 349592

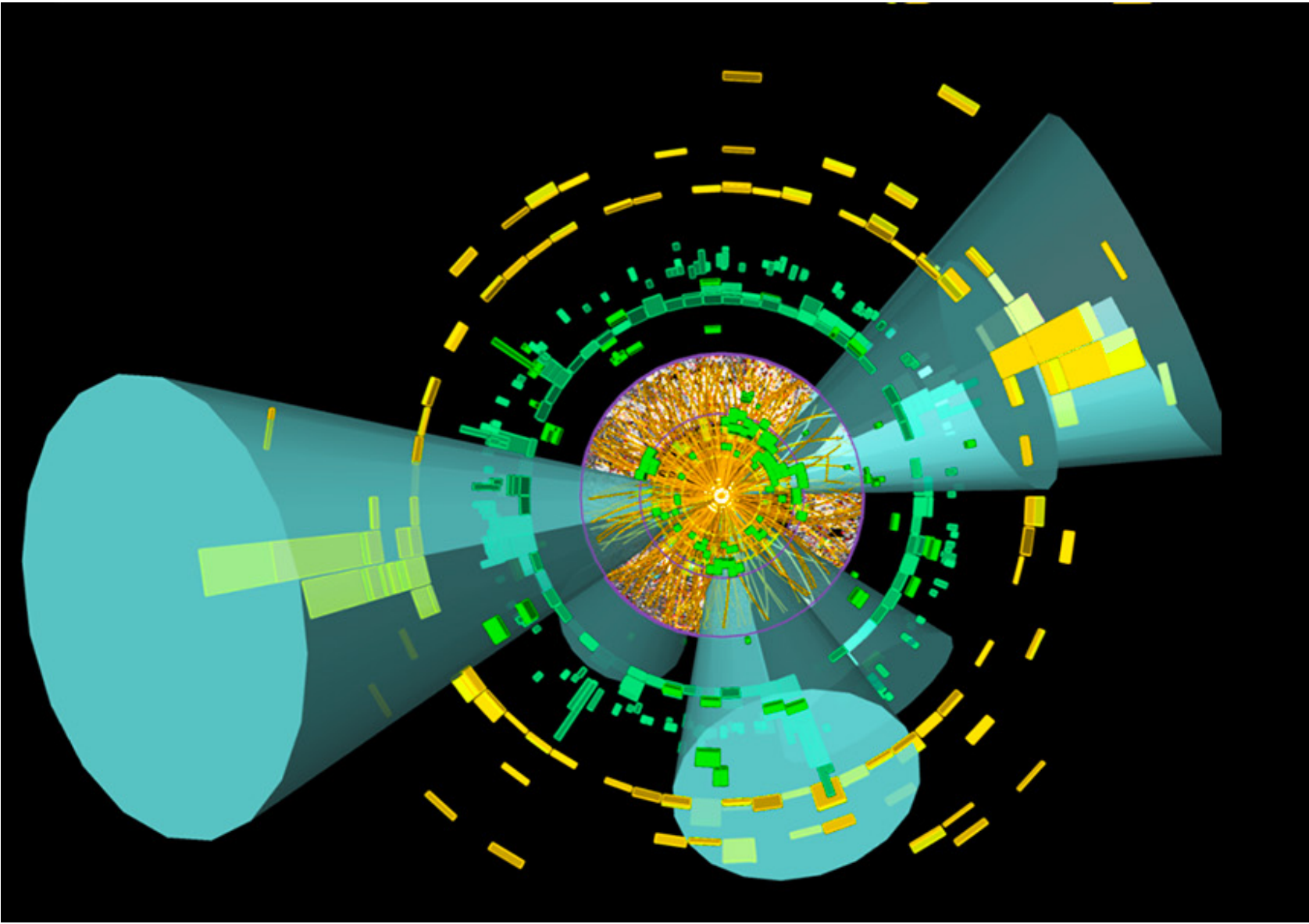
Event: 793710422

2018-05-06 03:37:28 UTC

Analysis strategy

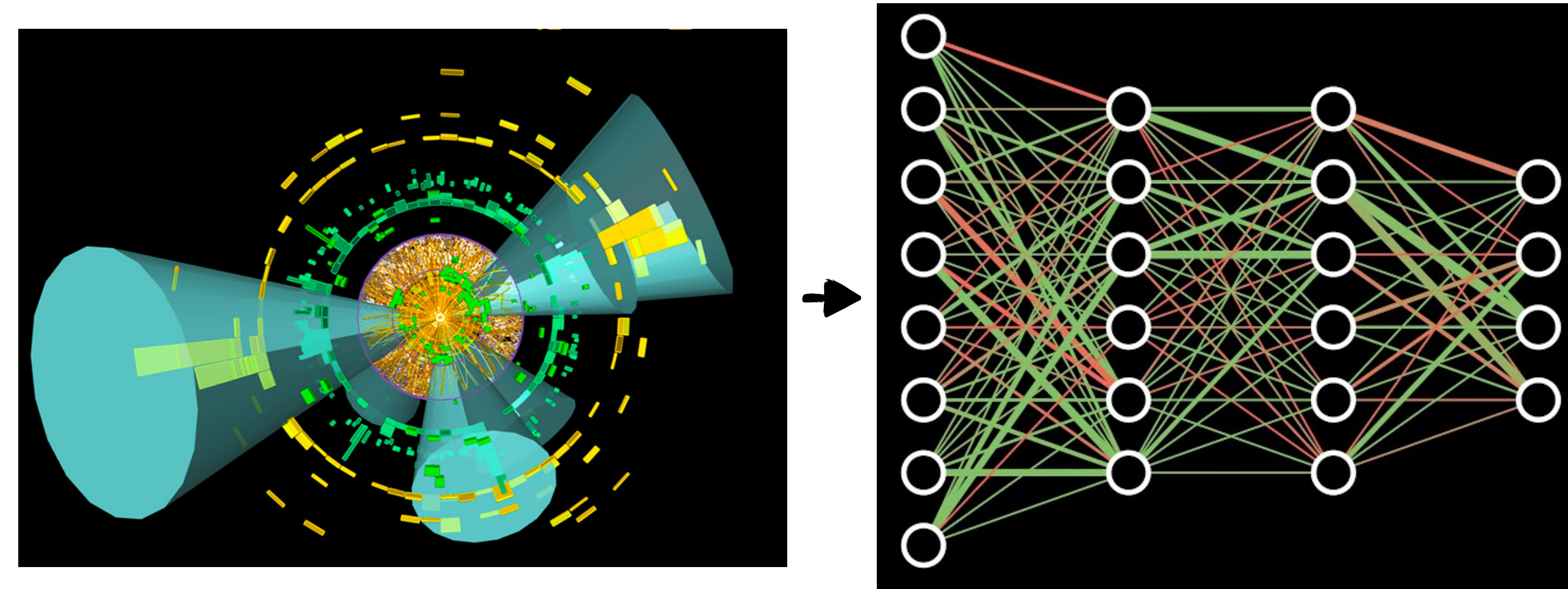


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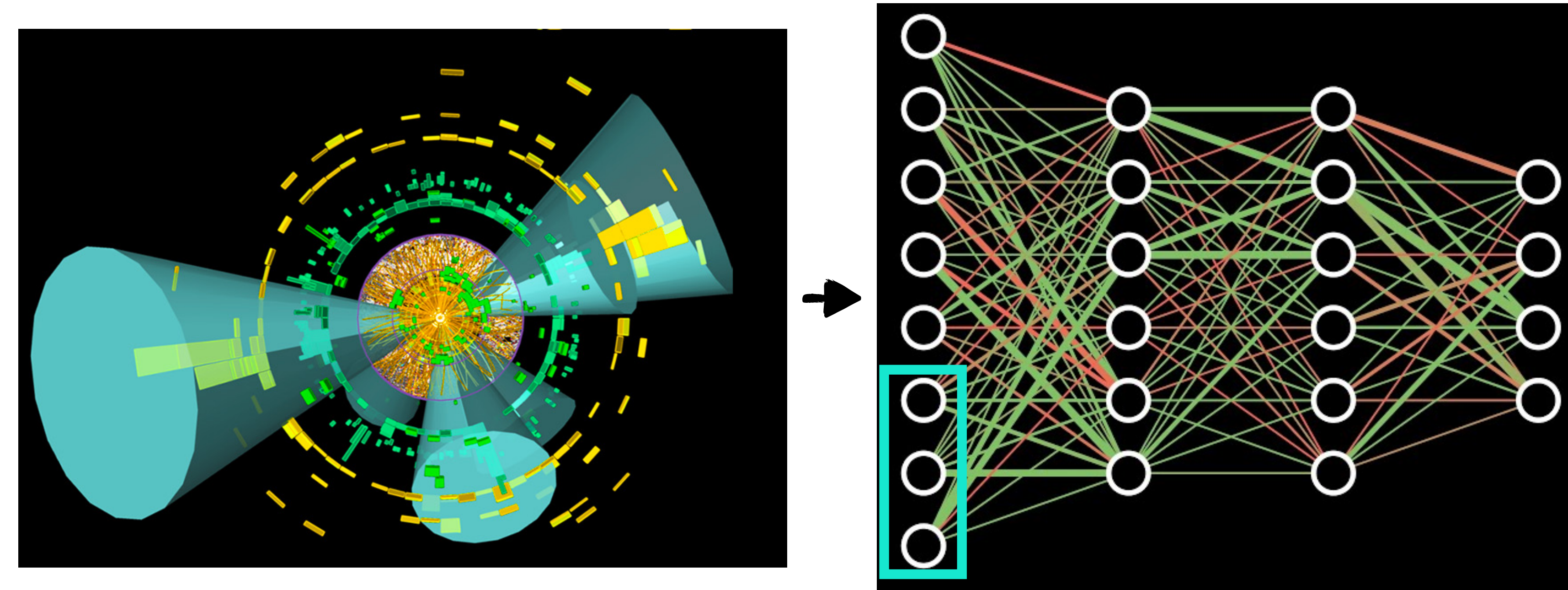
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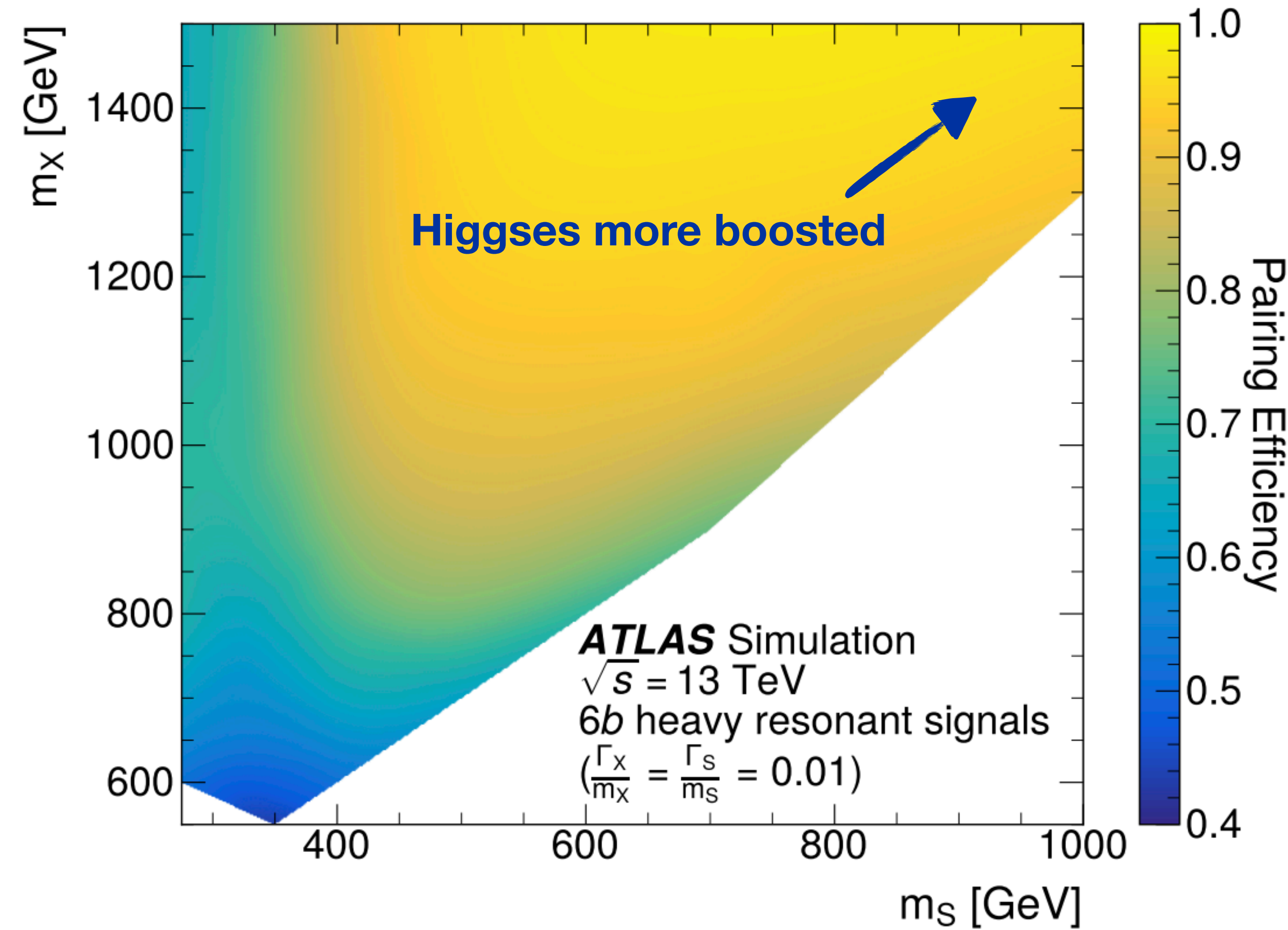
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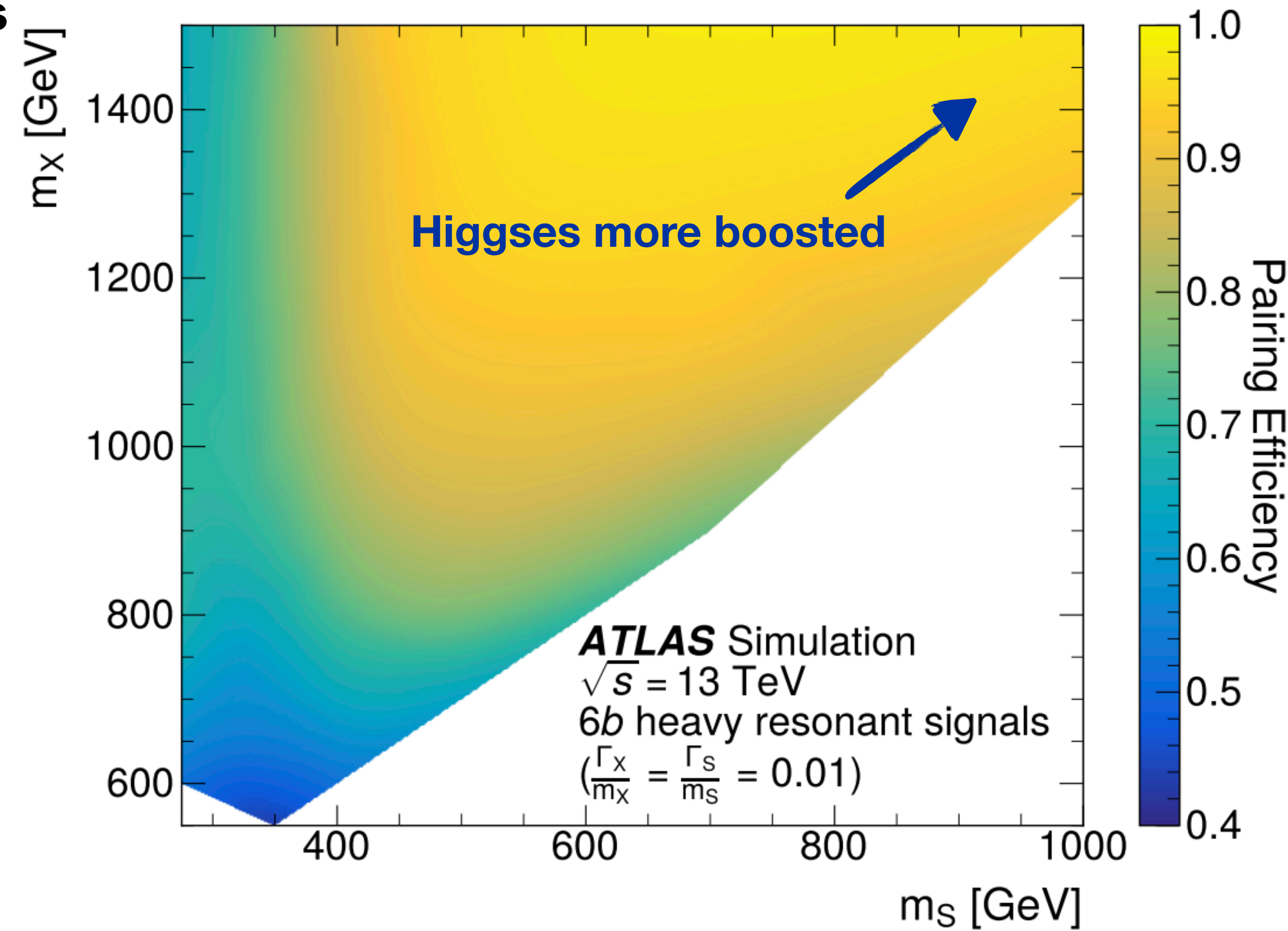
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- Train a DNN to discriminate between HHH signal and 5b background.
 - Some of the DNN inputs are reliant on paired jets, for example to compute $m(bb)$ for the three pairs of jets.

b-jet pairing



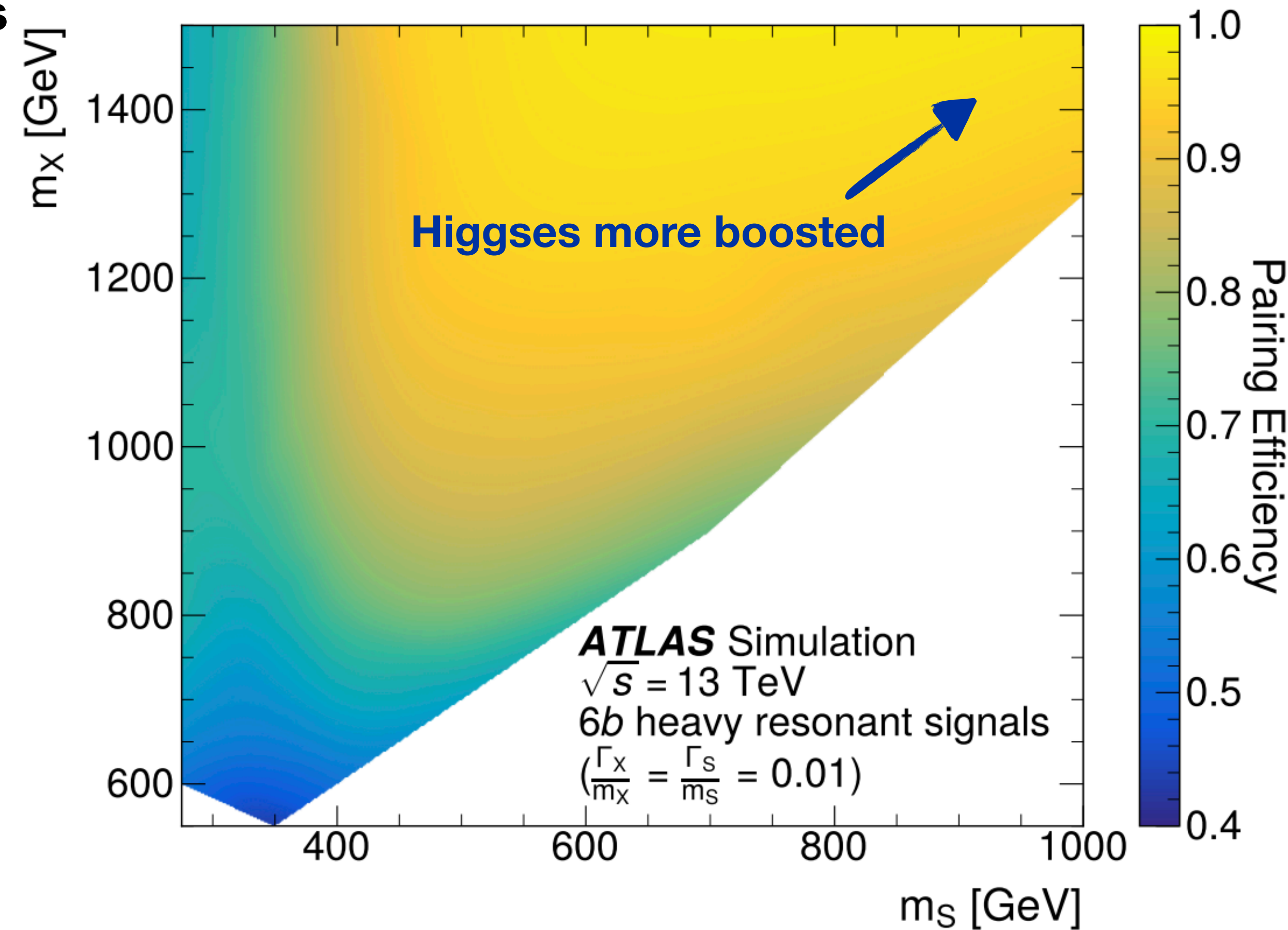
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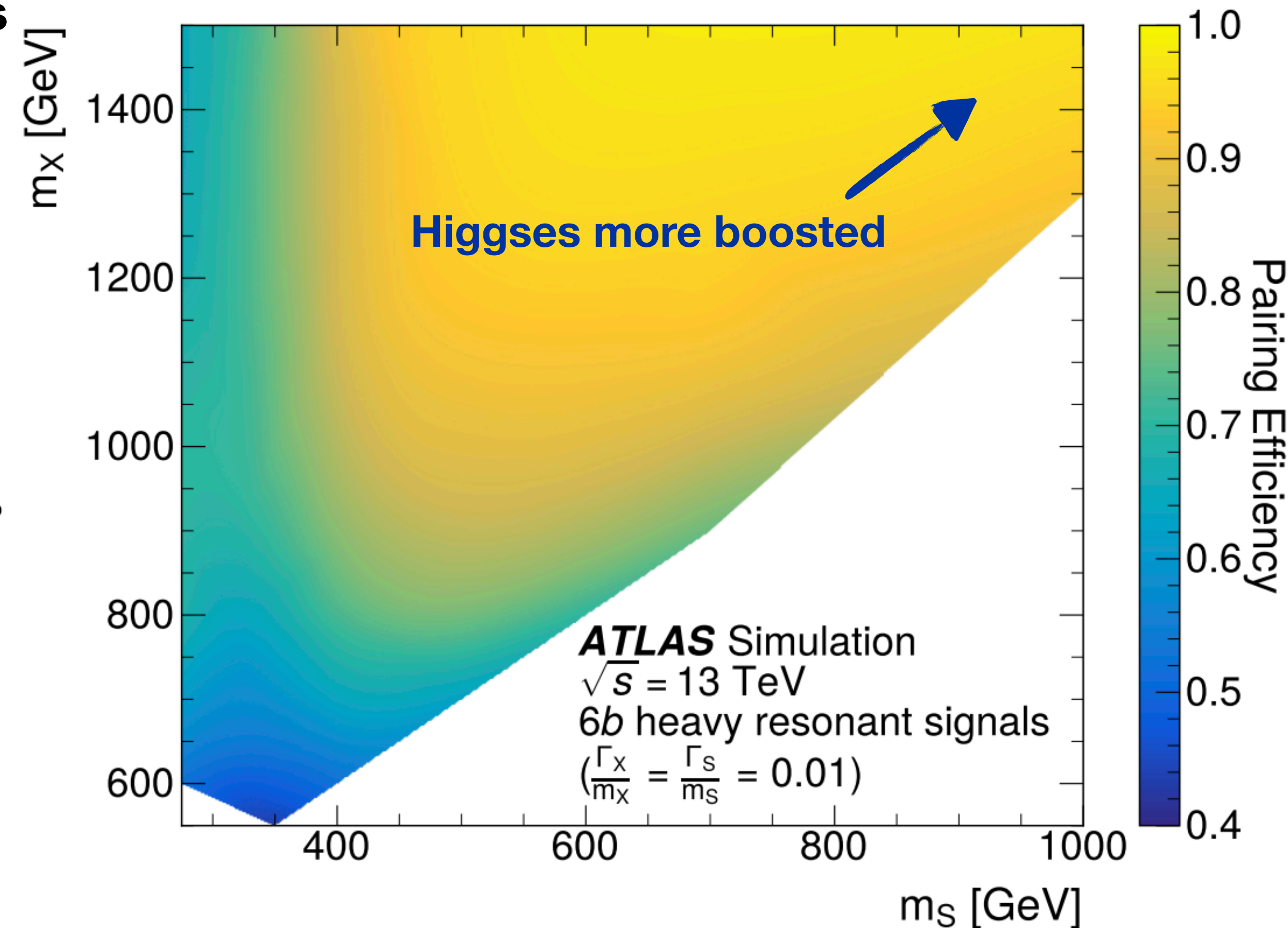


b-jet pairing

- Some variables in the DNN will be based on paired jets
- There are 15 ways to pair 6 jets
- We use a mass-based pairing with additional constraint on p_T -ordering, reducing from 15 possibilities

$$|m_{H1} - 120 \text{ GeV}| + |m_{H2} - 115 \text{ GeV}| + |m_{H3} - 110 \text{ GeV}|,$$

Leading **Subleading** **Third**



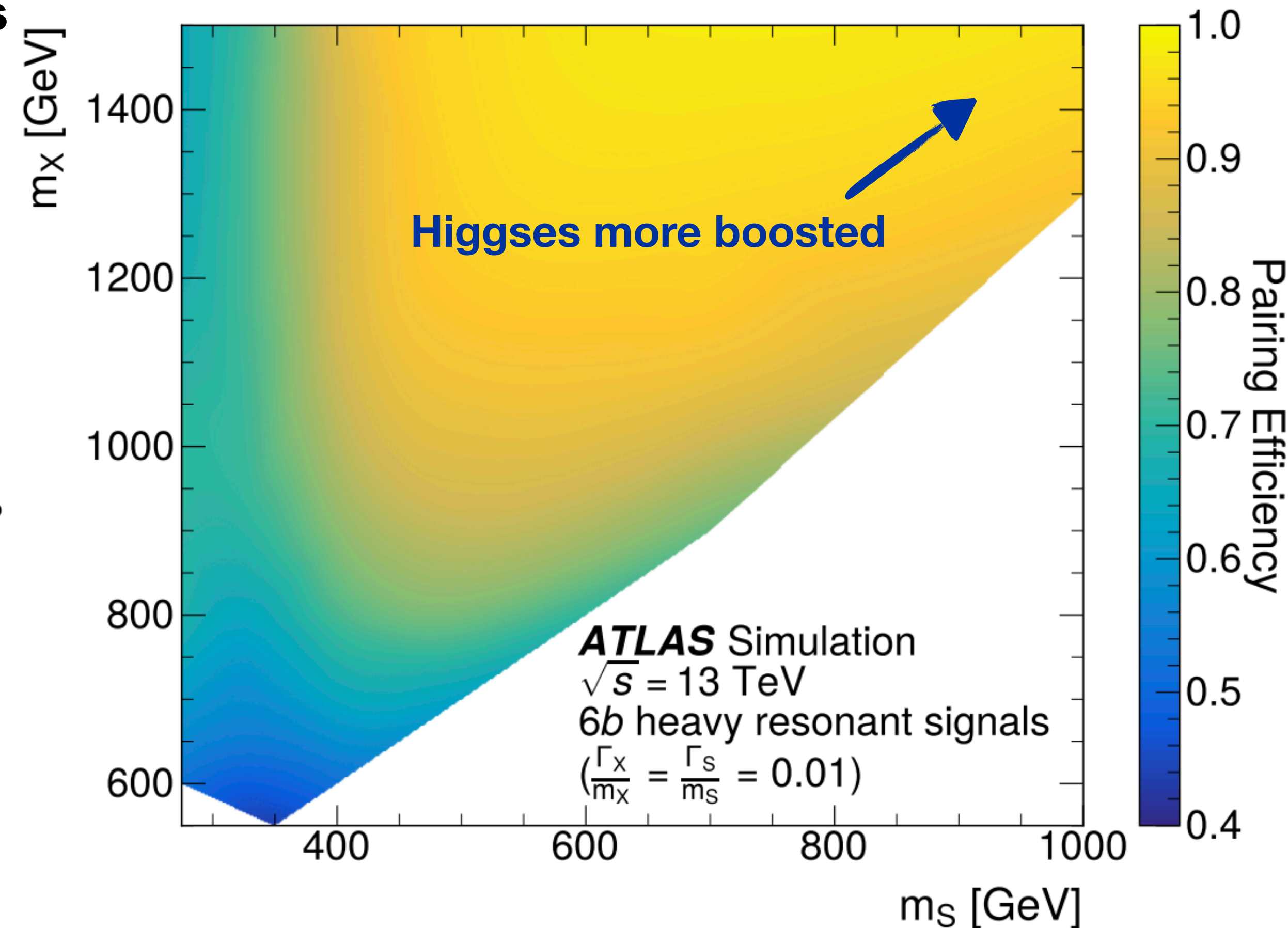
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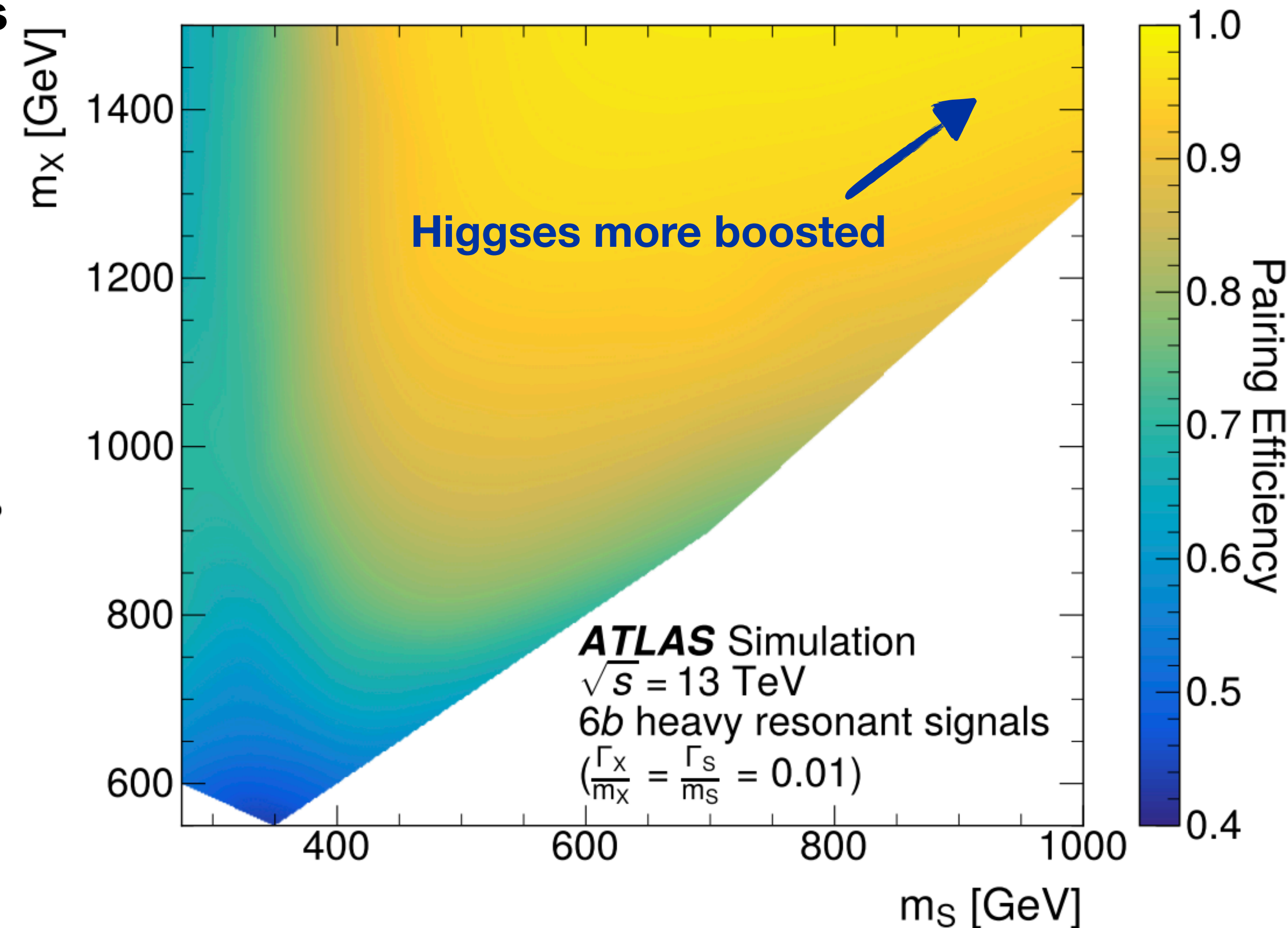
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- **Performs very well in the boosted regime.**
 - Still far better than random chance when not boosted, but that is a low bar.



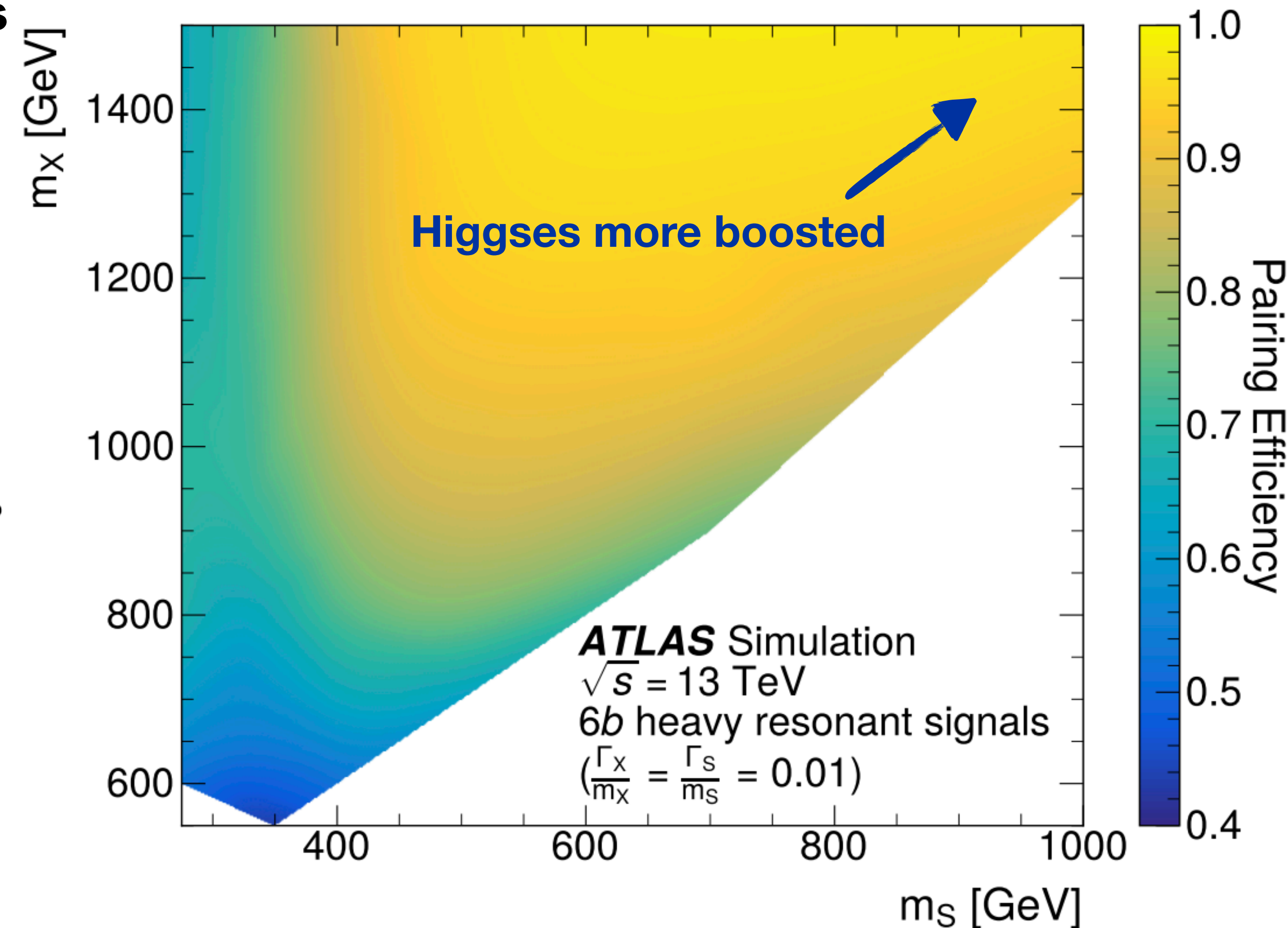
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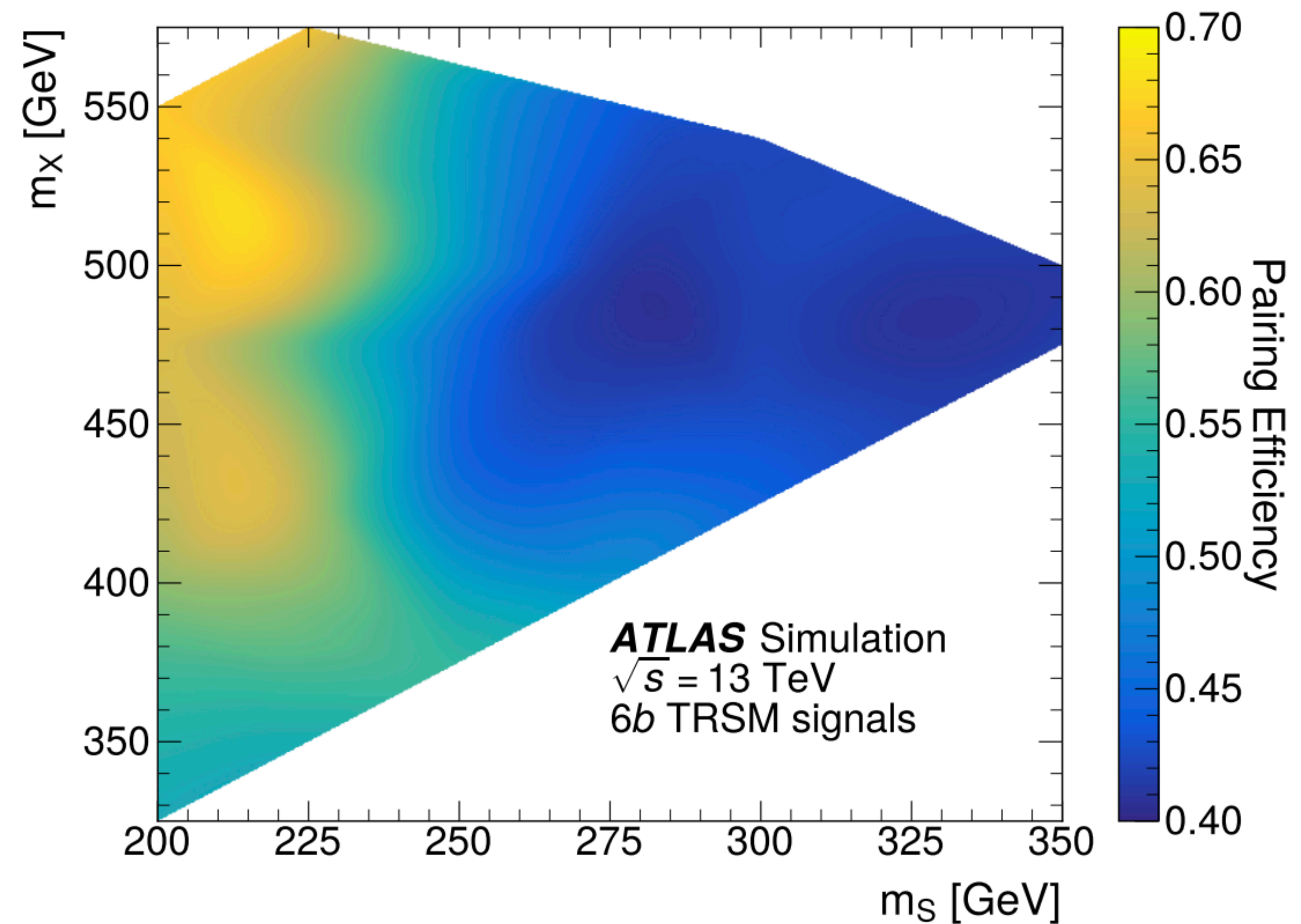
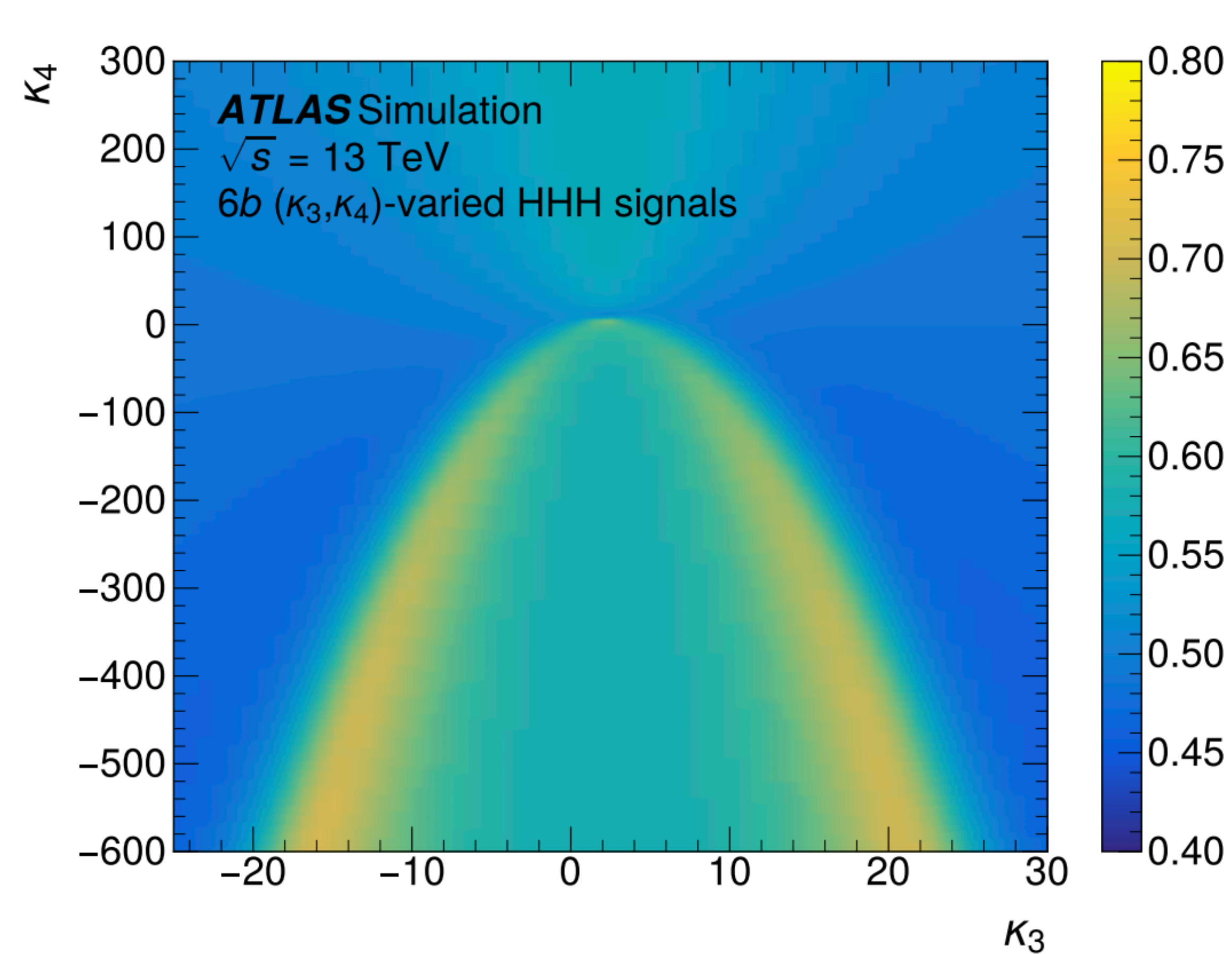
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Leading **Subleading** **Third**

- **This is empirical!**
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- **Pairing algorithms are one area for future R&D on HHH**

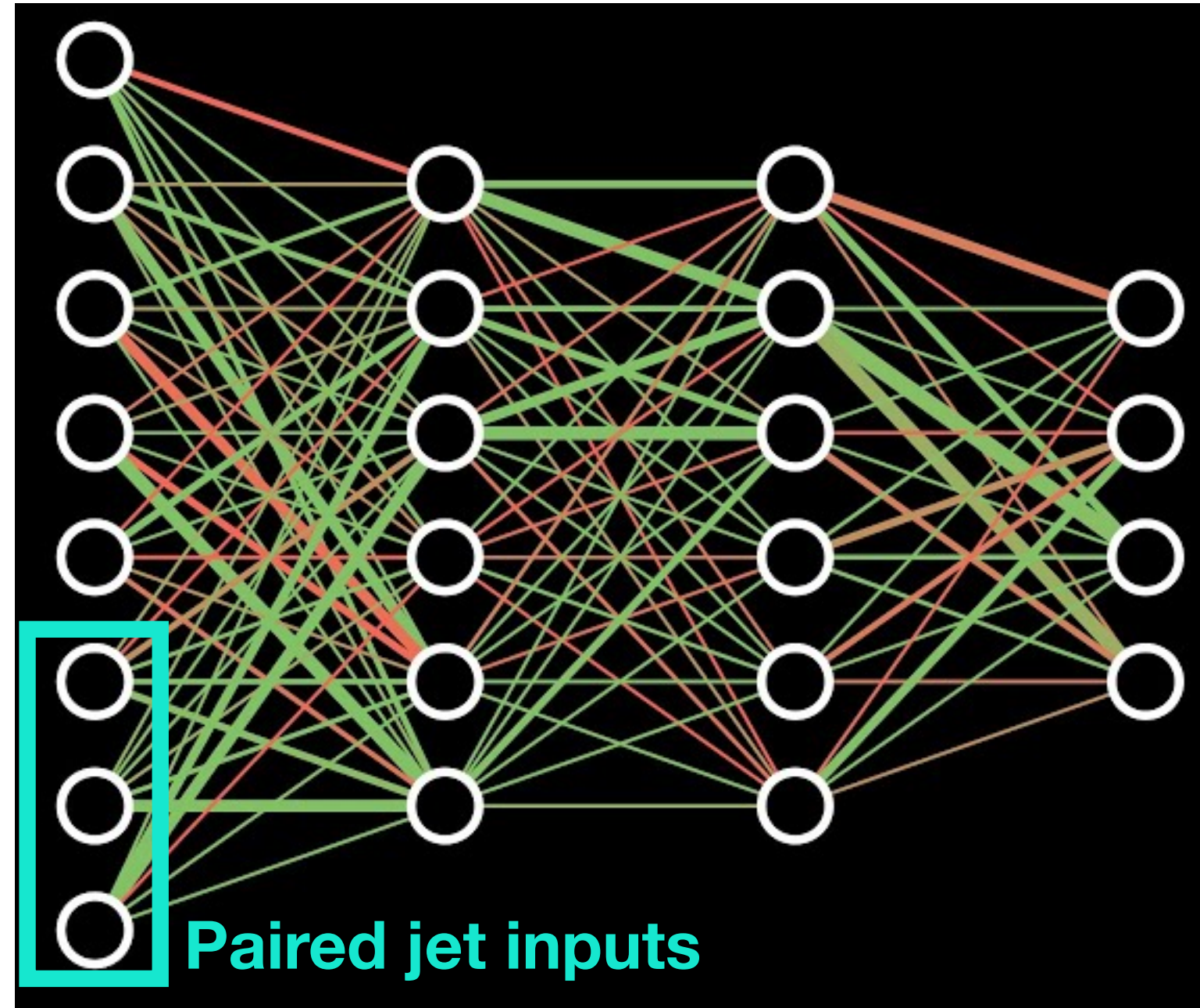
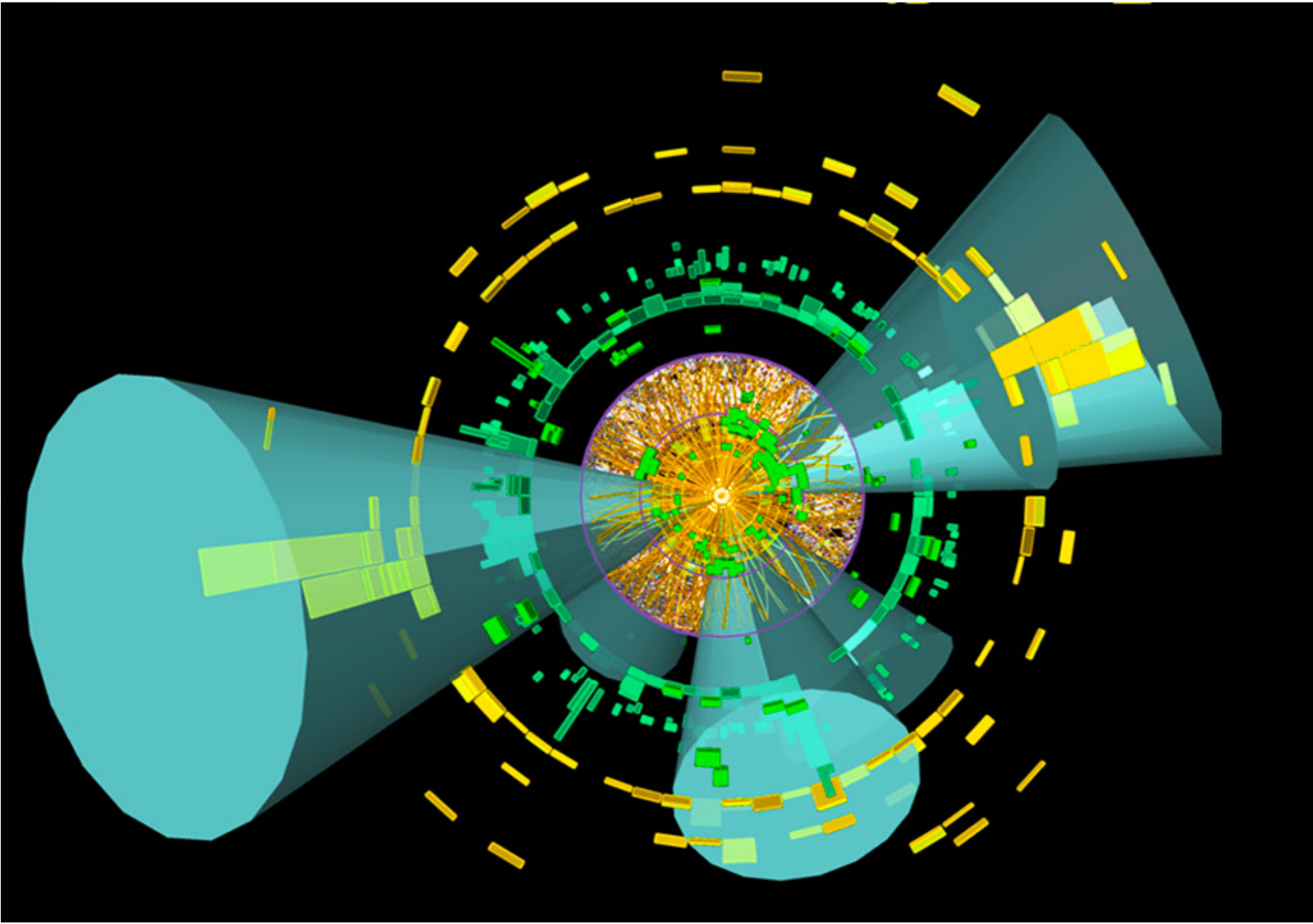


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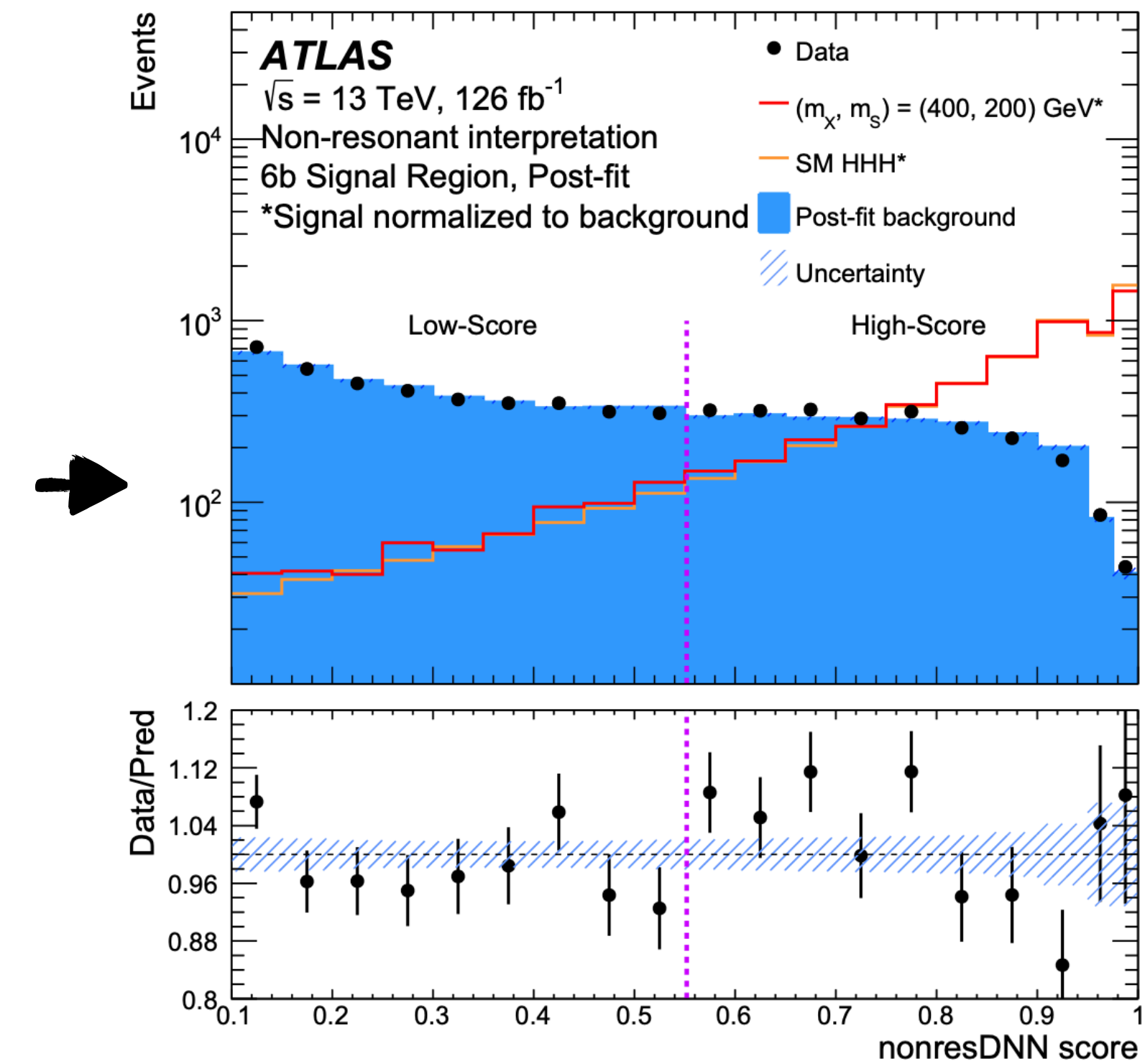
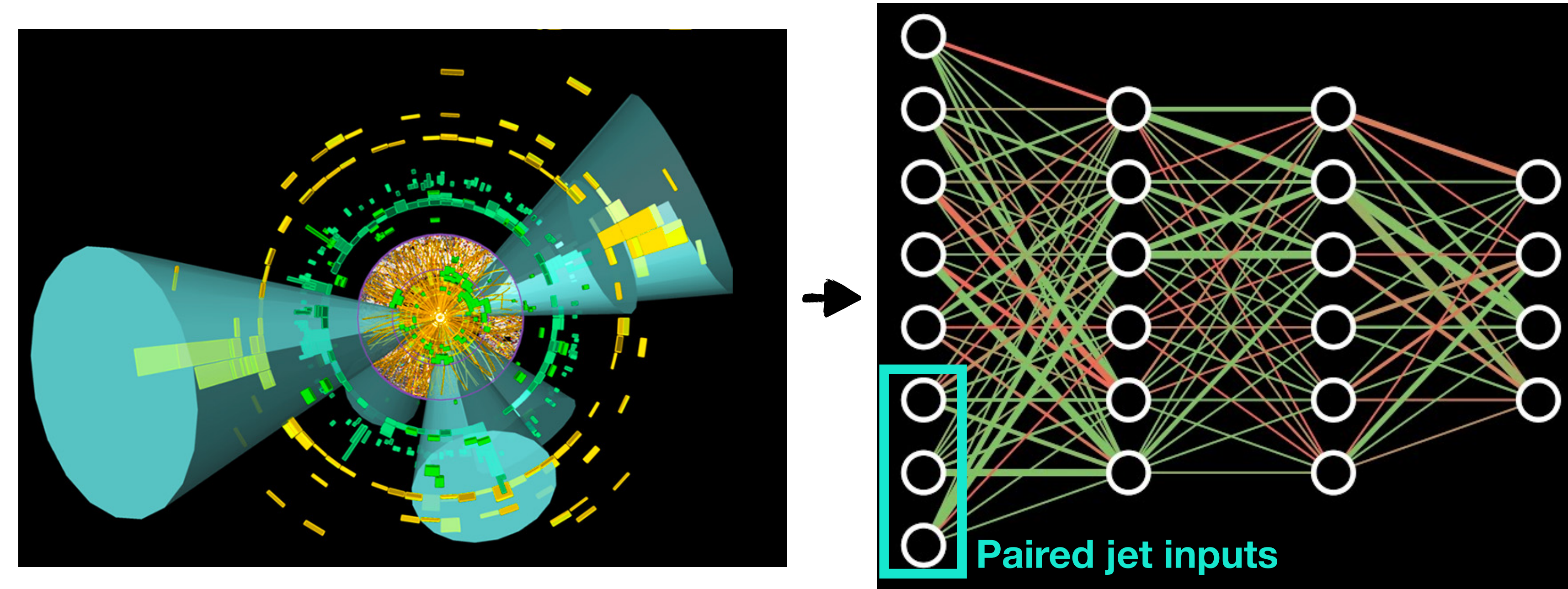


- Without large boost, pairing efficiency is as low as 40%
- SM pairing efficiency: 51%

Analysis strategy

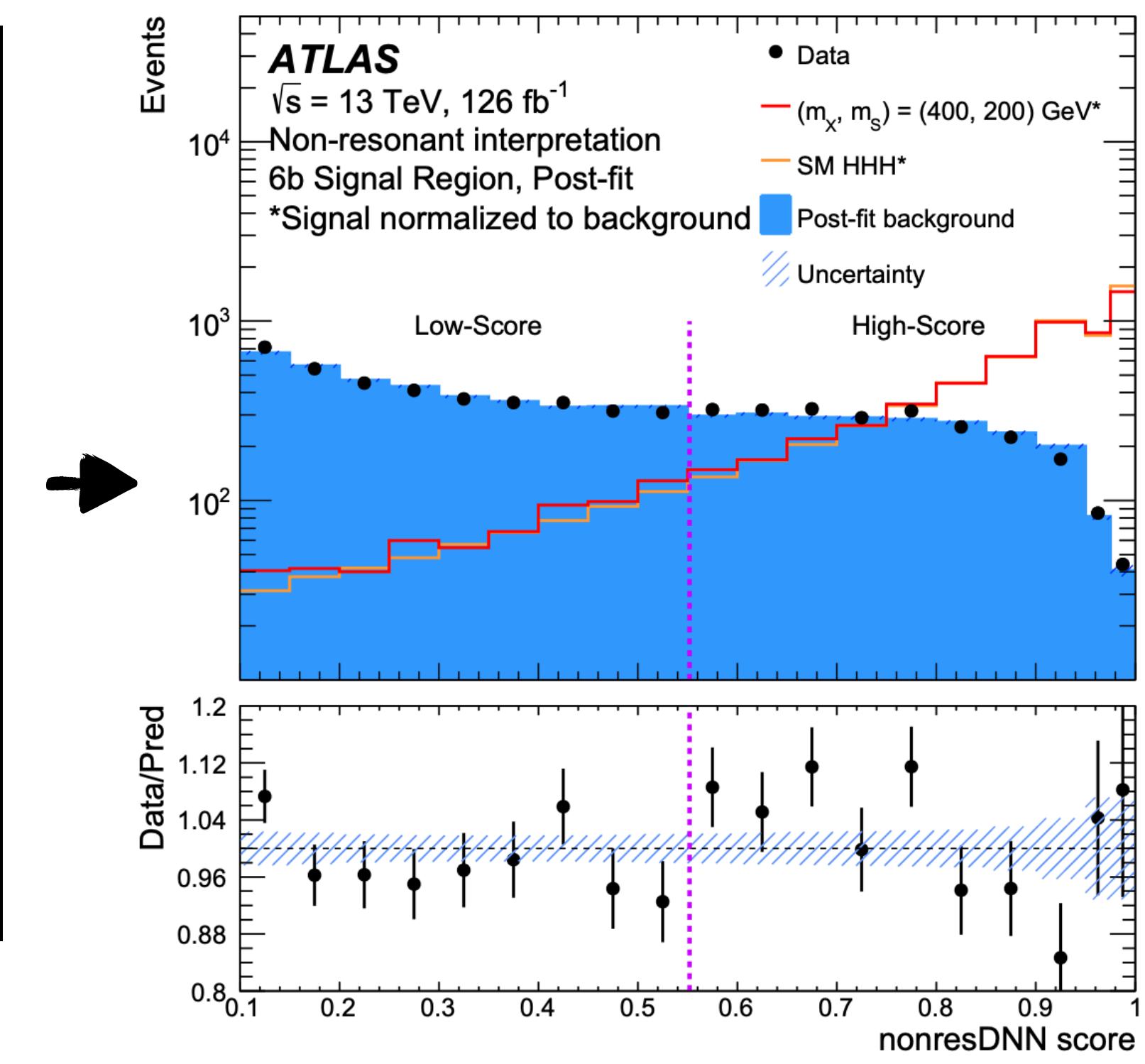
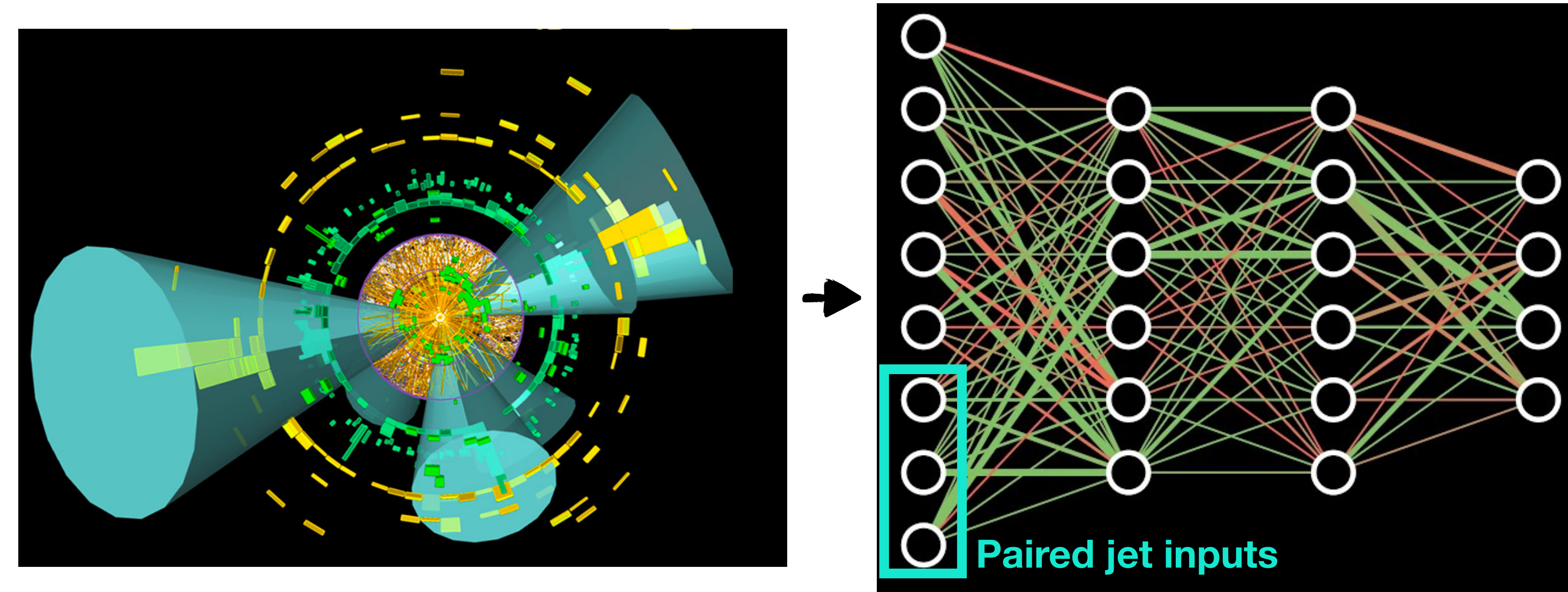


Analysis strategy



- The DNN output score is a discriminant we fit and compute p-values/limits.

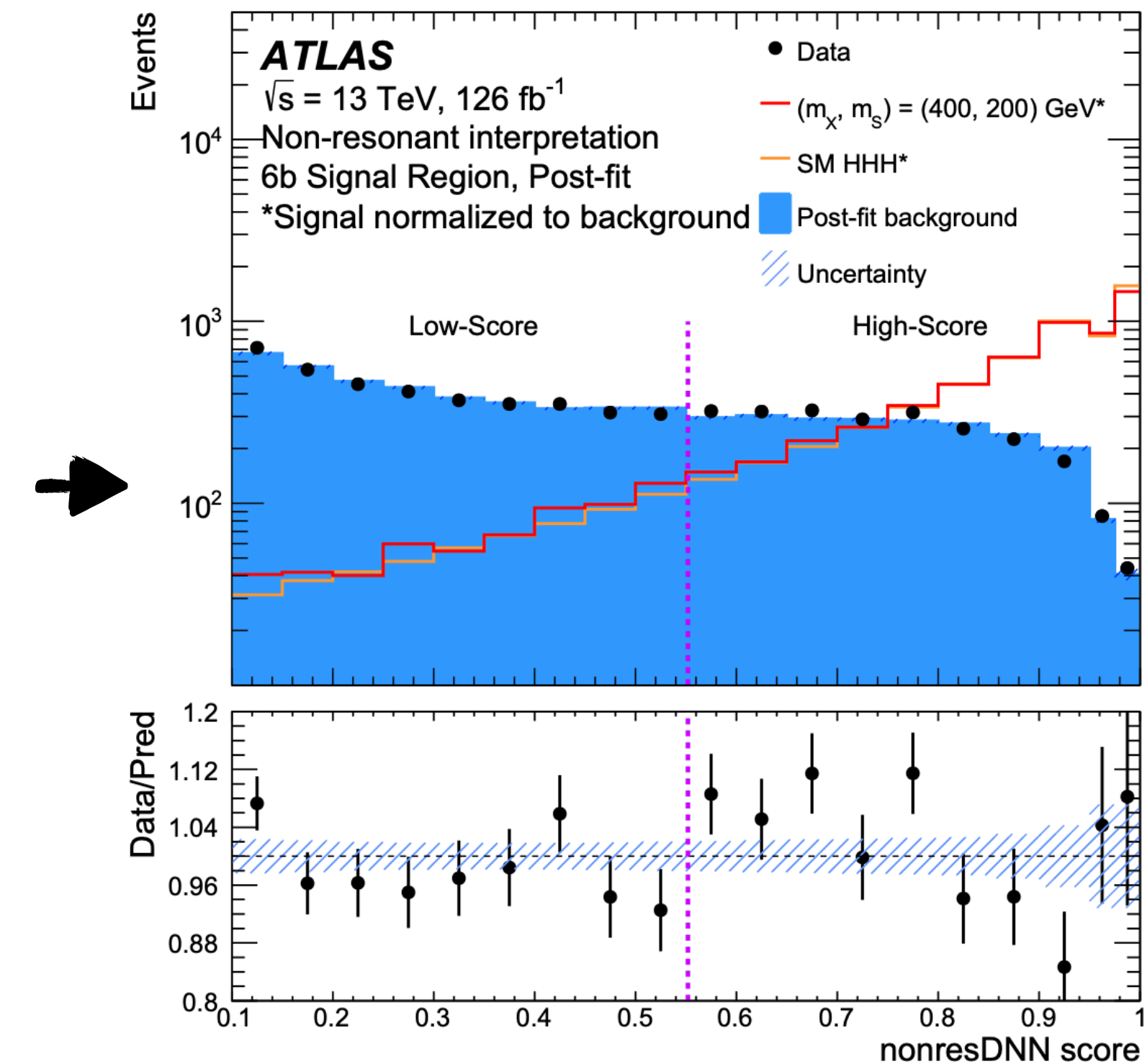
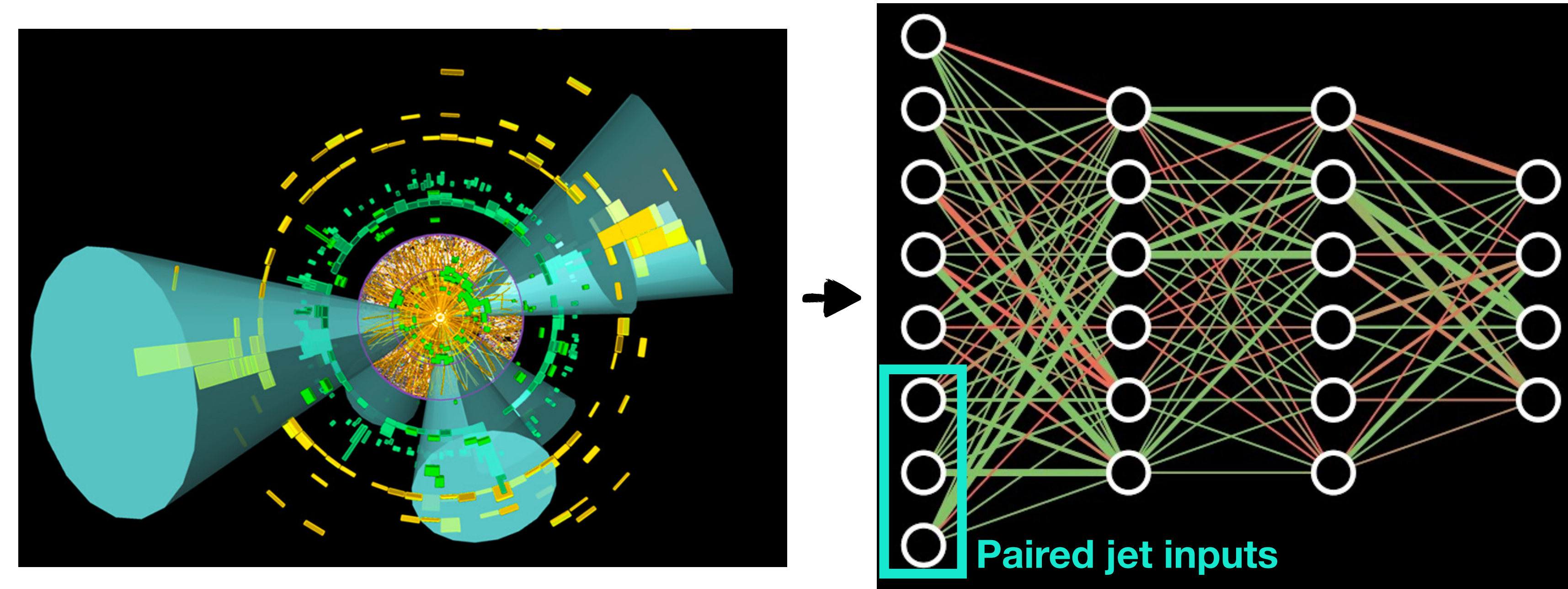
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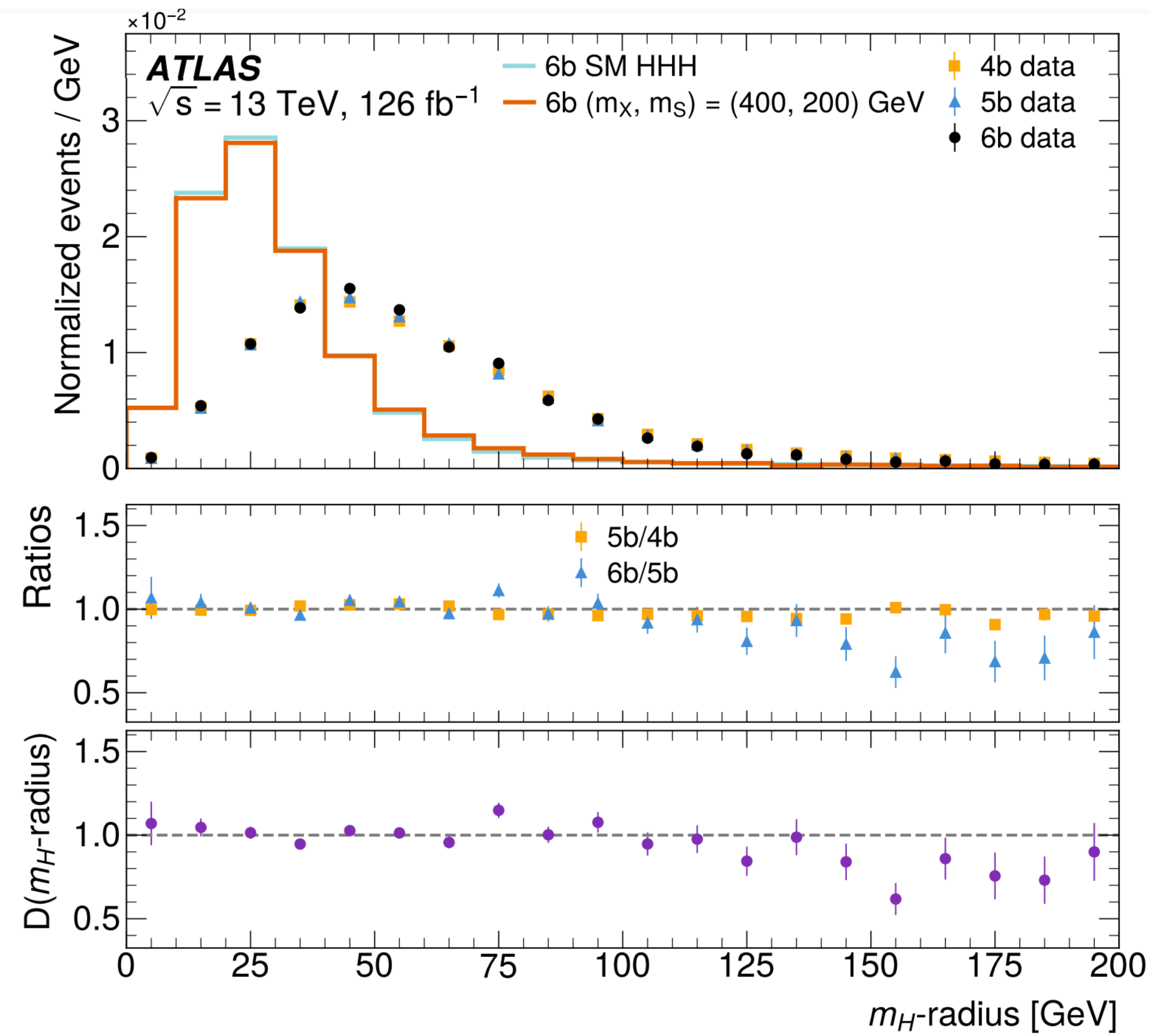
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 - Larger difference between 4b, 5b, 6b means larger systematic uncertainties
 - Need to choose DNN inputs carefully!

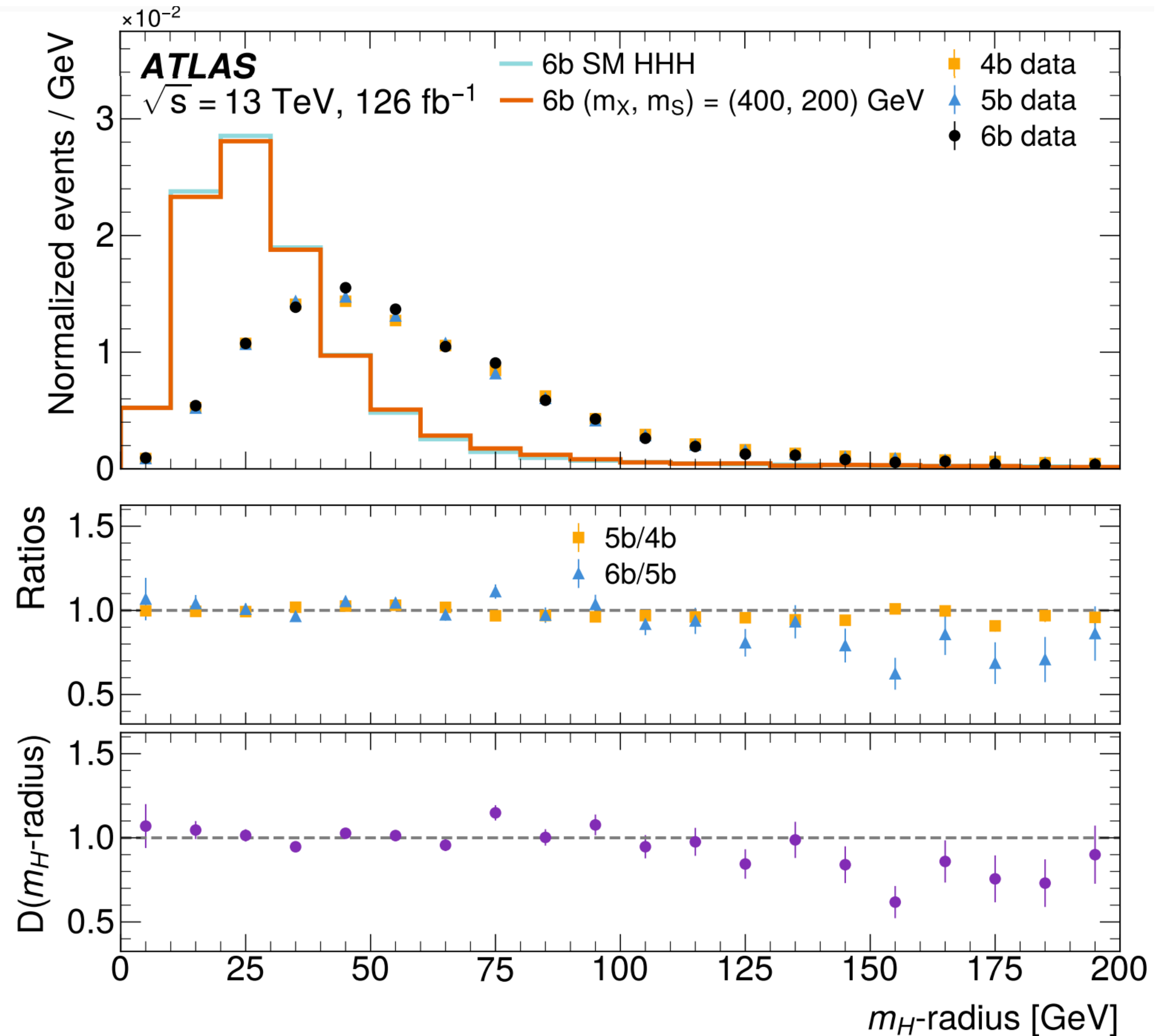
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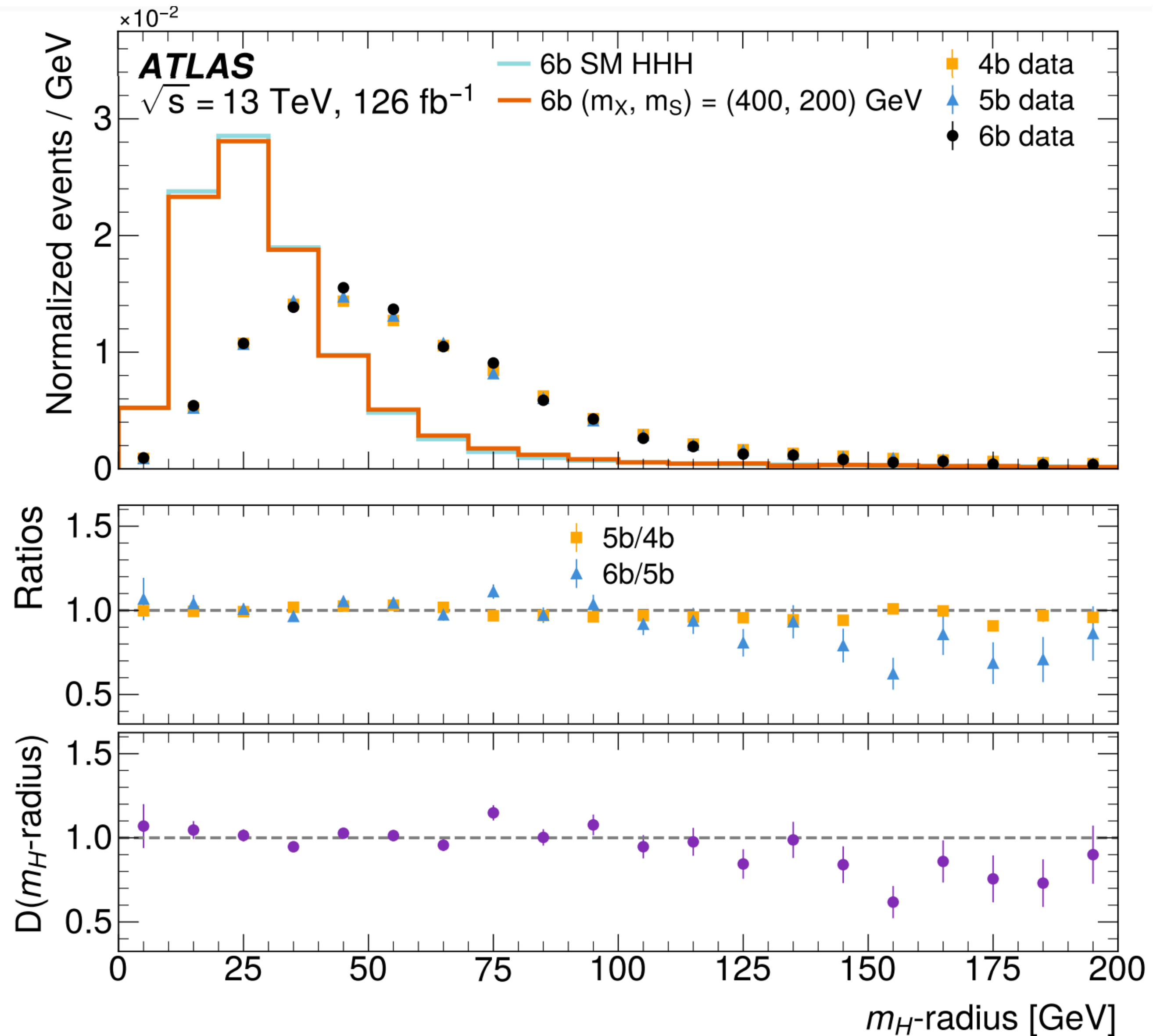
DNN inputs

- **Example of one DNN input shown.**
 - Distance from pairing center: (120, 115, 110) GeV
 - This is an example of a variable for which pairing matters

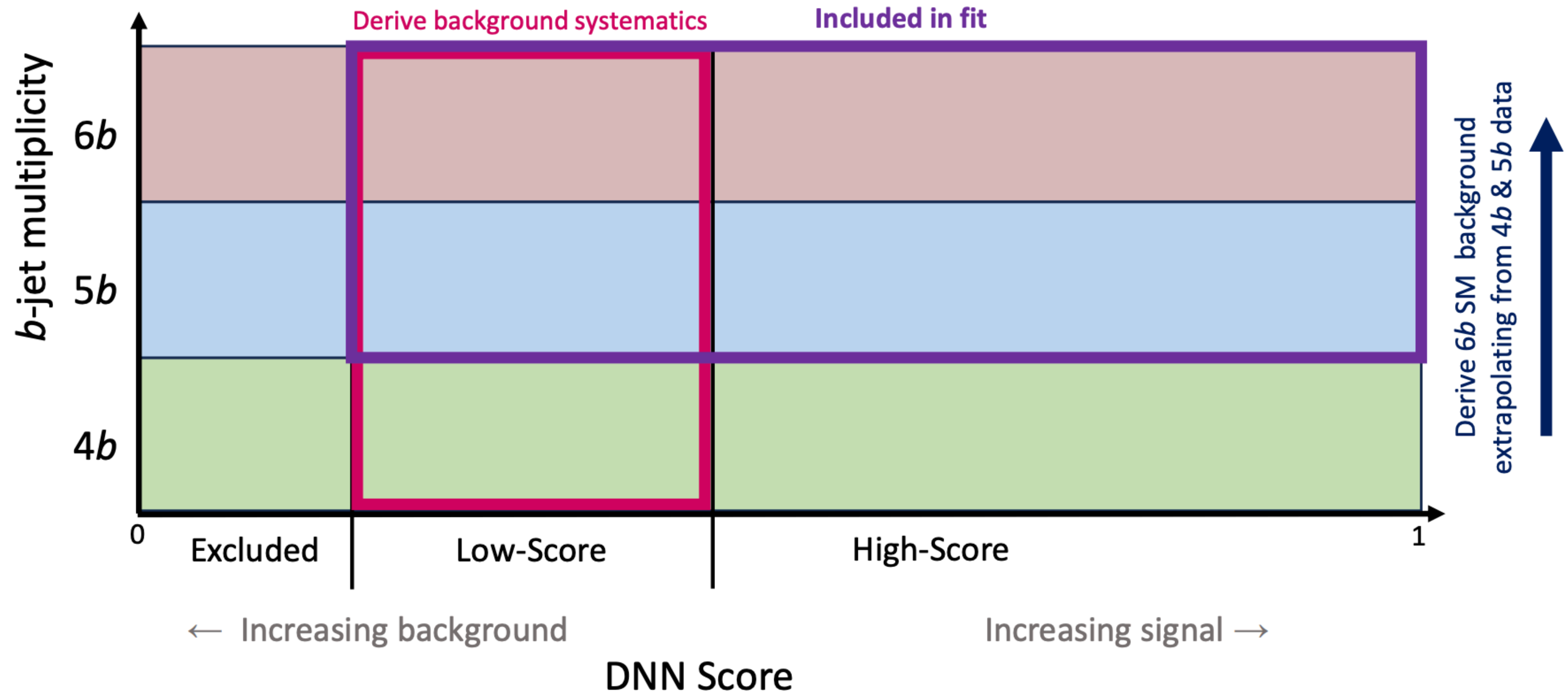


DNN inputs

- **Example of one DNN input shown.**
 - Distance from pairing center: (120, 115, 110) GeV
 - This is an example of a variable for which pairing matters
- **Also shown are 4b, 5b, 6b data and their ratios**
 - We limit ourselves to variables which have small shape difference vs. b-jet multiplicity.
 - This is a trade-off between using more discriminating variables vs. having larger systematic uncertainty on the b-jet extrapolation.

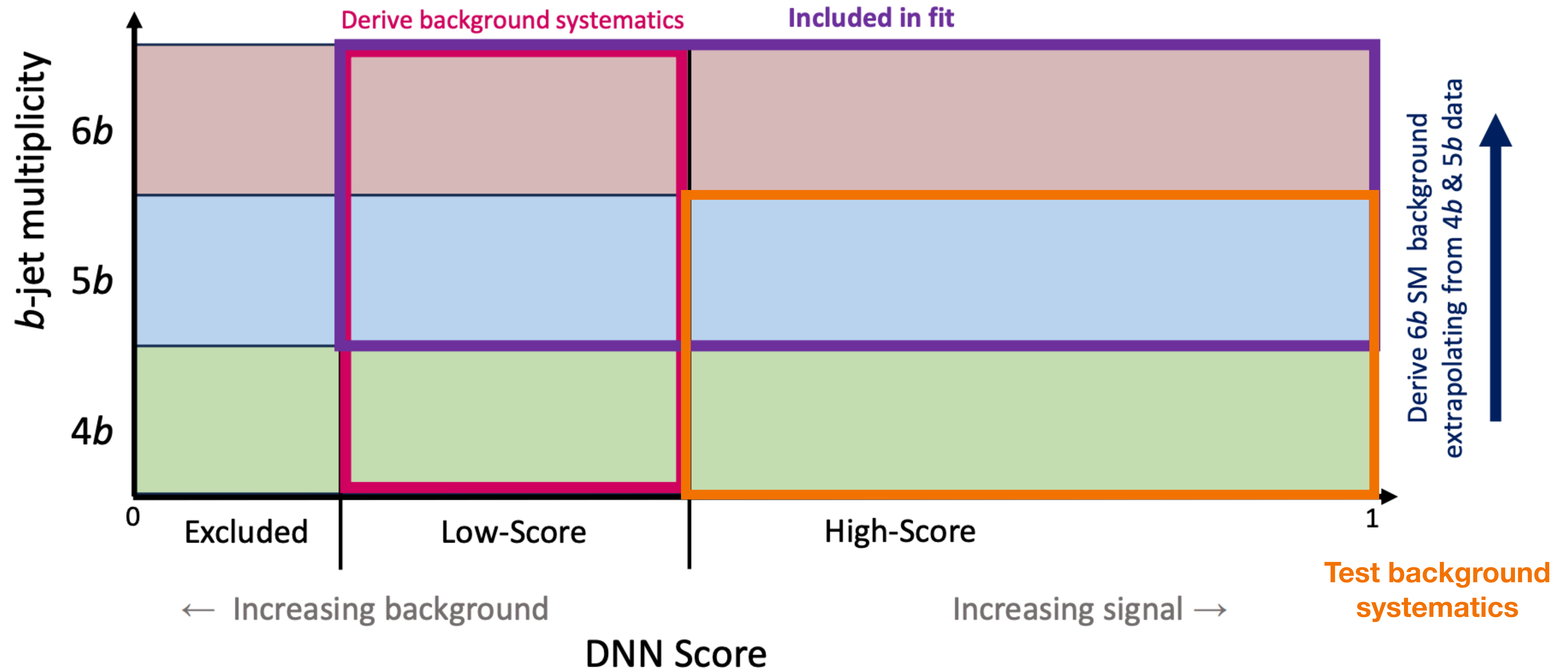


Analysis strategy



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Background shape uncertainties



Background shape uncertainties

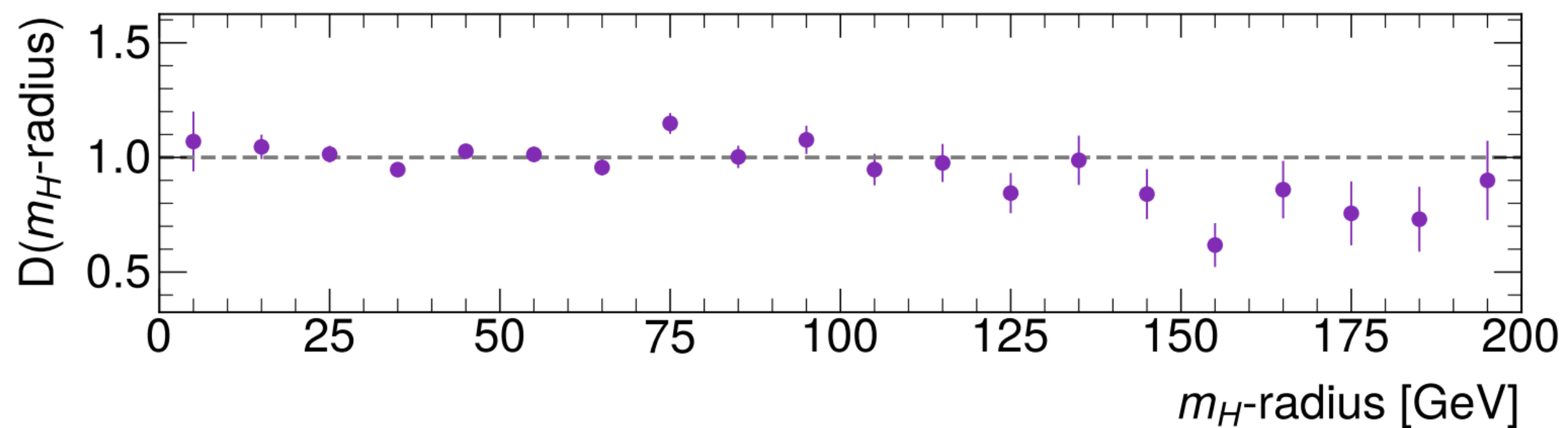
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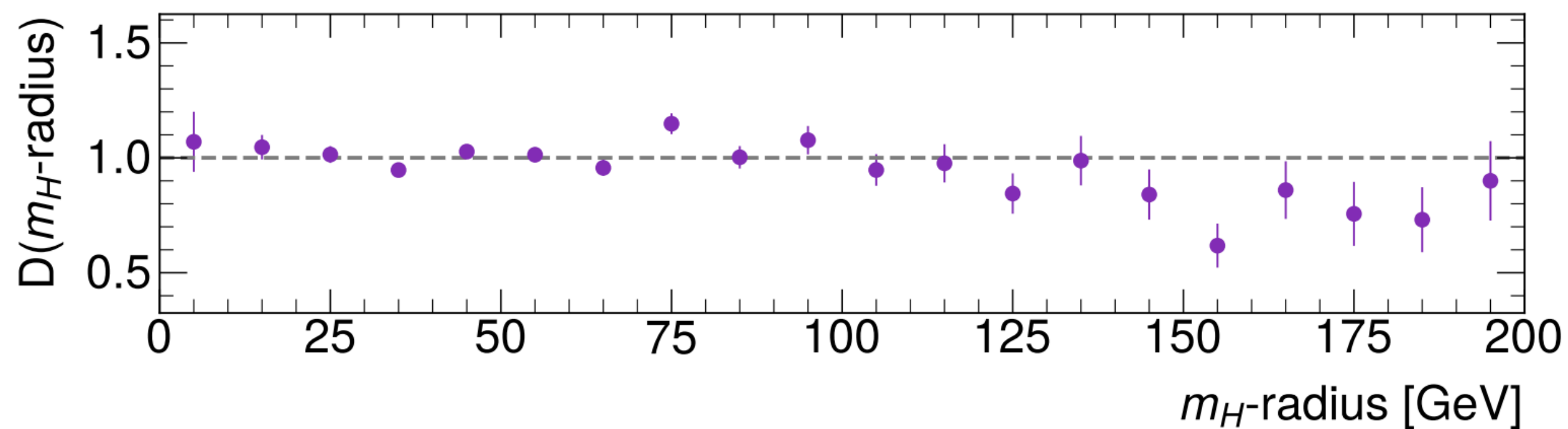
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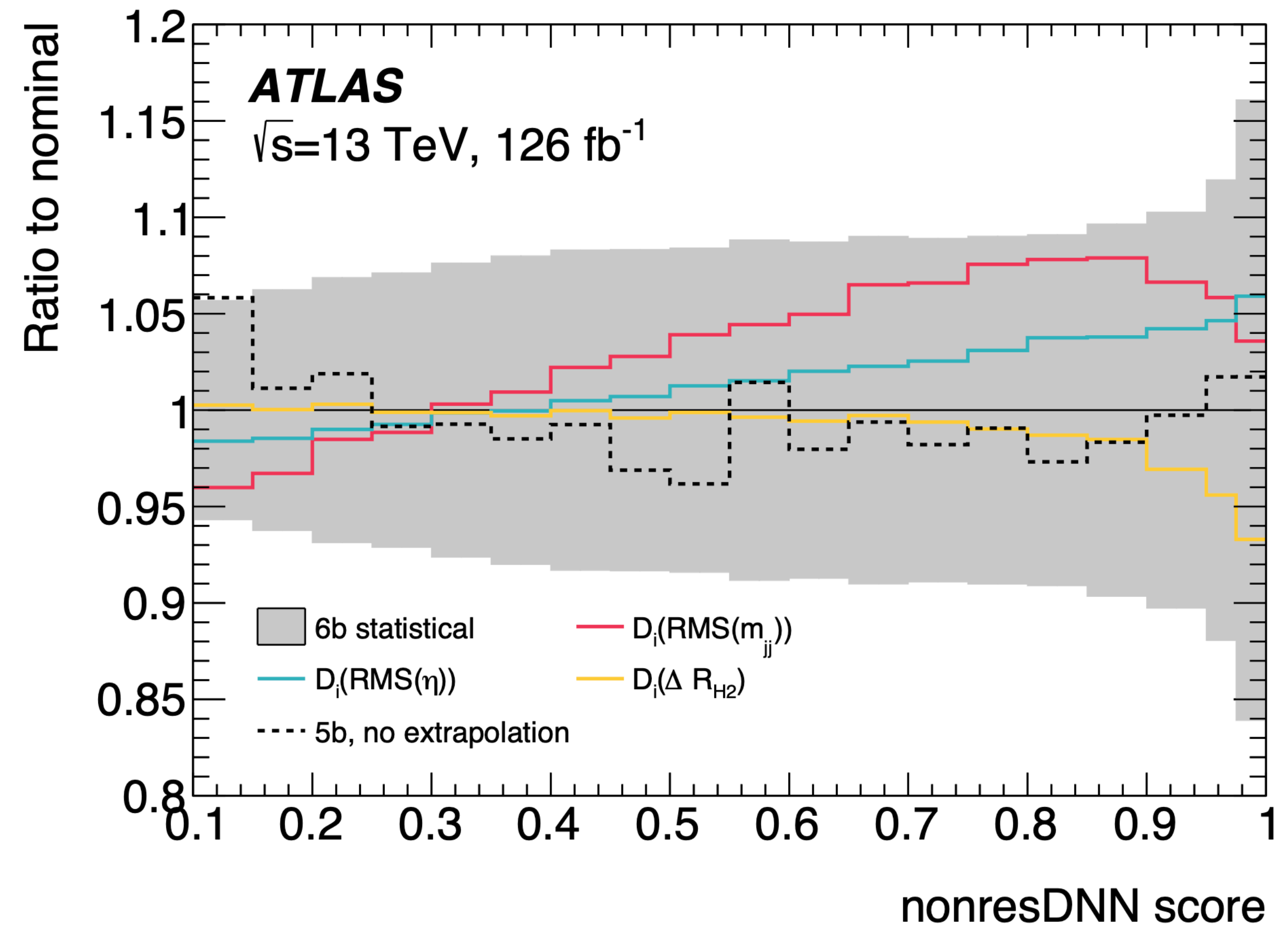


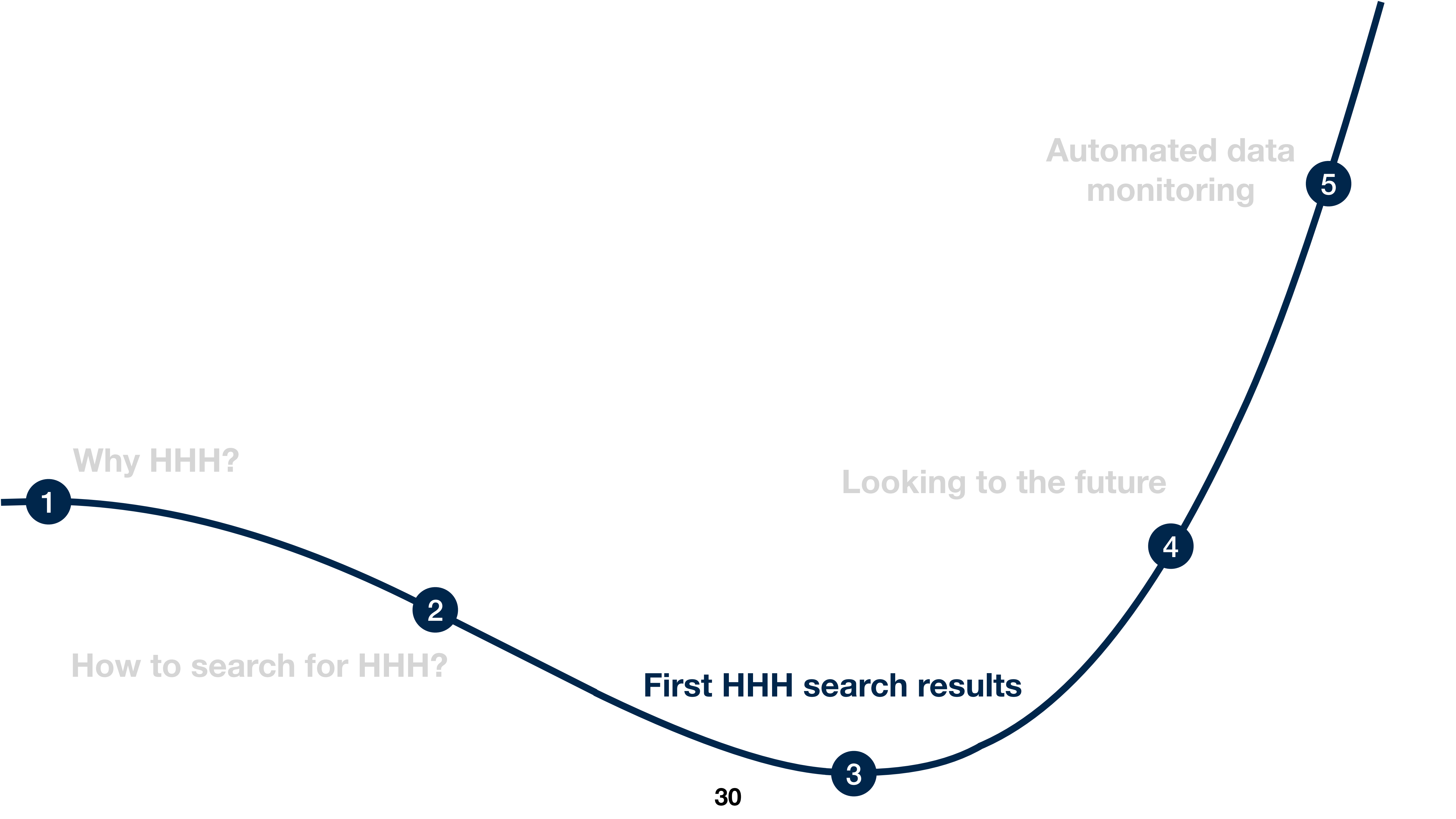
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- Reweight each event by this histogram. Select a linearly independent subset of alternate shapes.





1

Why HHH?

2

How to search for HHH?

3

First HHH search results

30

4

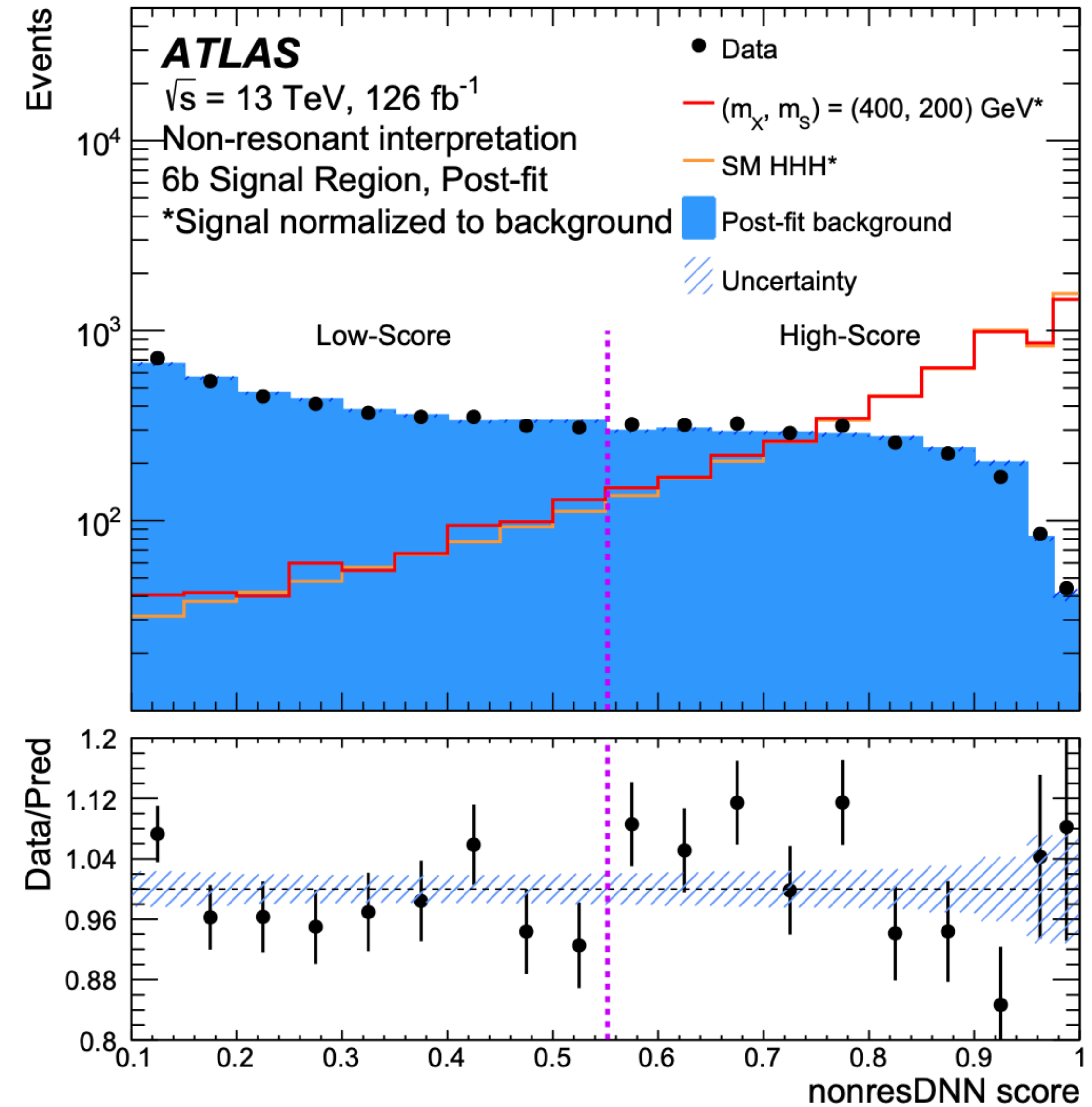
Looking to the future

5

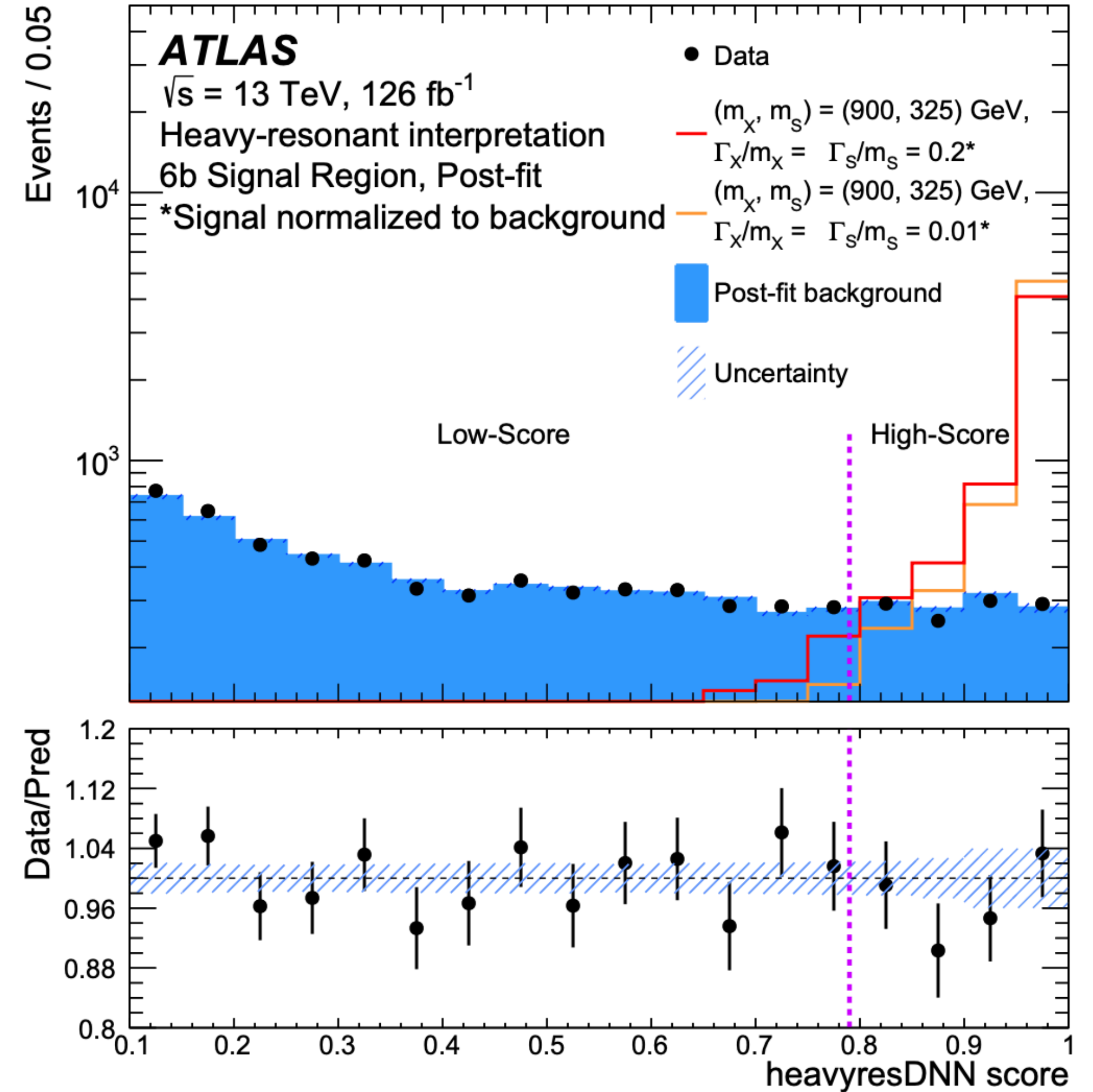
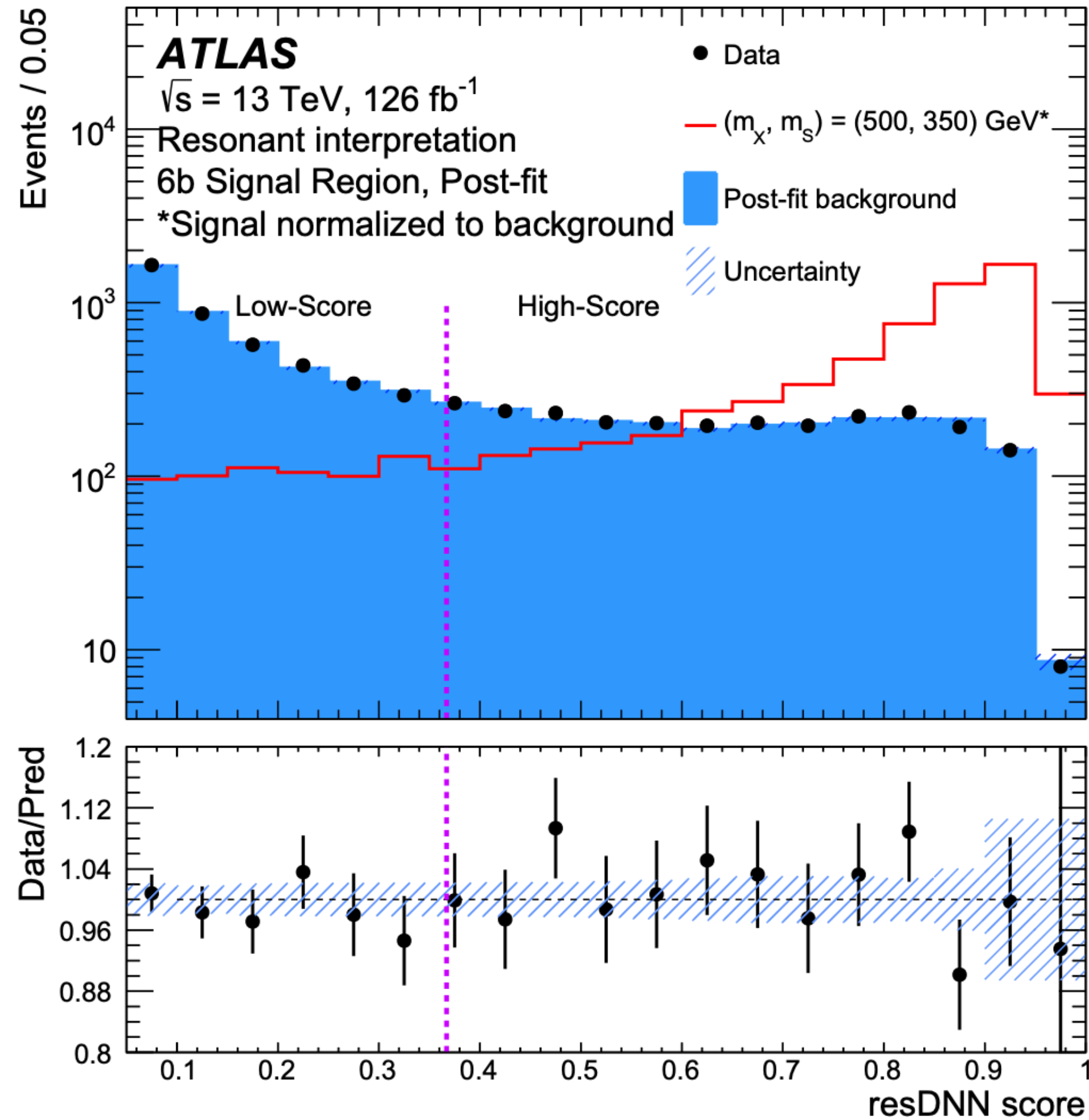
Automated data monitoring

Results

- No evidence for non-resonant HHH production
- Set limit at $\mu = \sigma/\sigma_{SM} < 750$, or 59 fb
- Non-resonant BSM modes can have nearly identical signature to SM HHH.



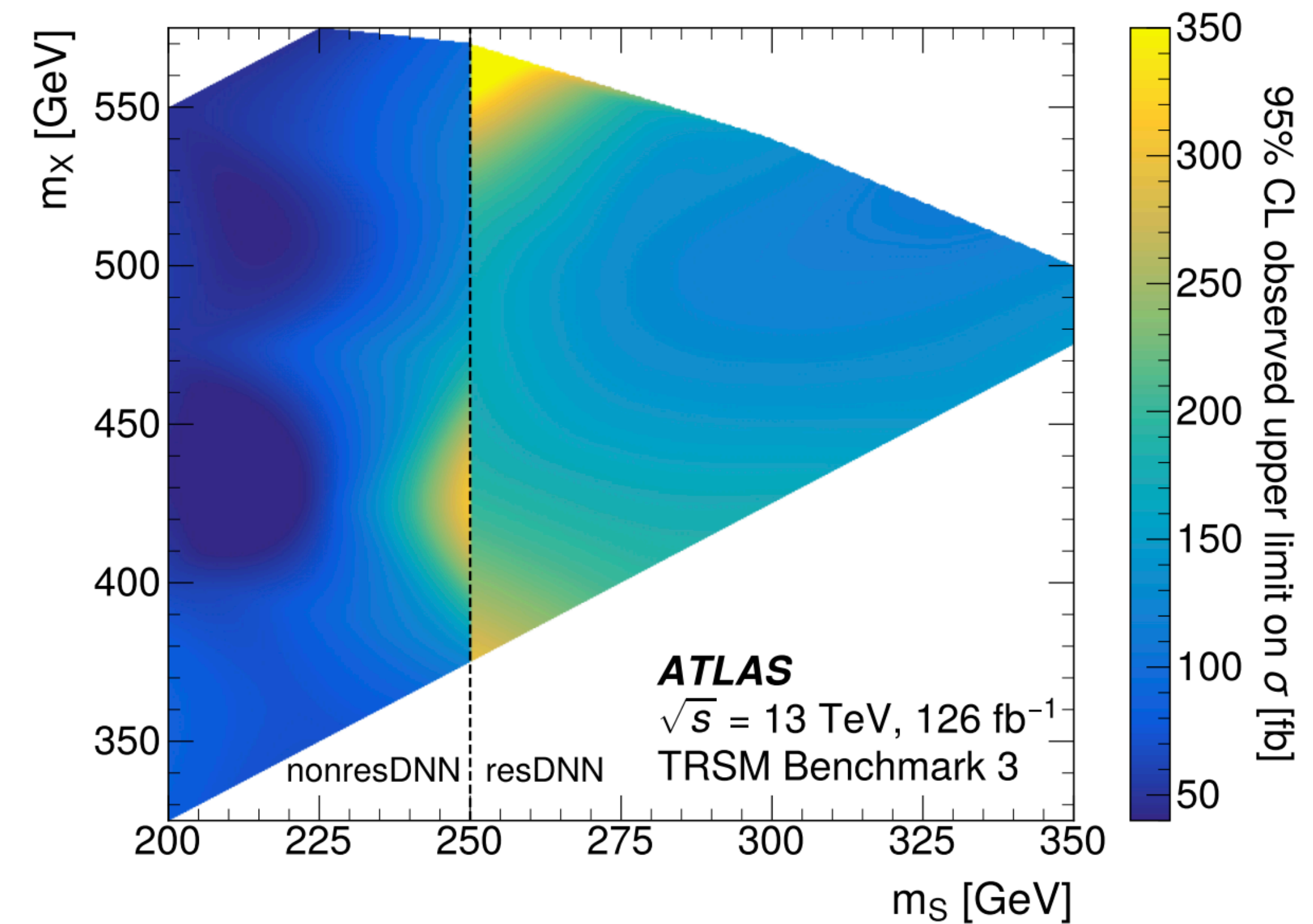
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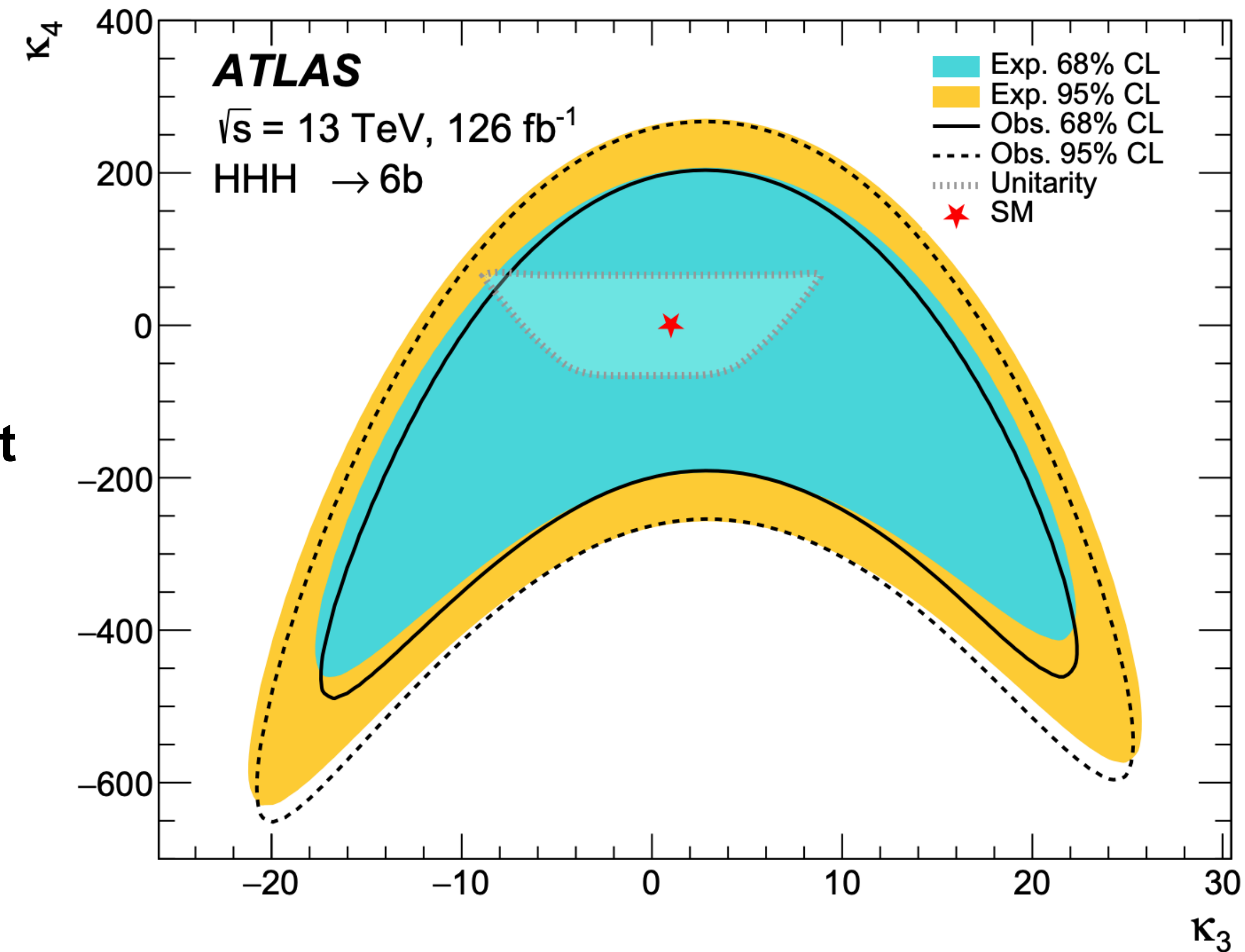
Conclusions: BSM searches

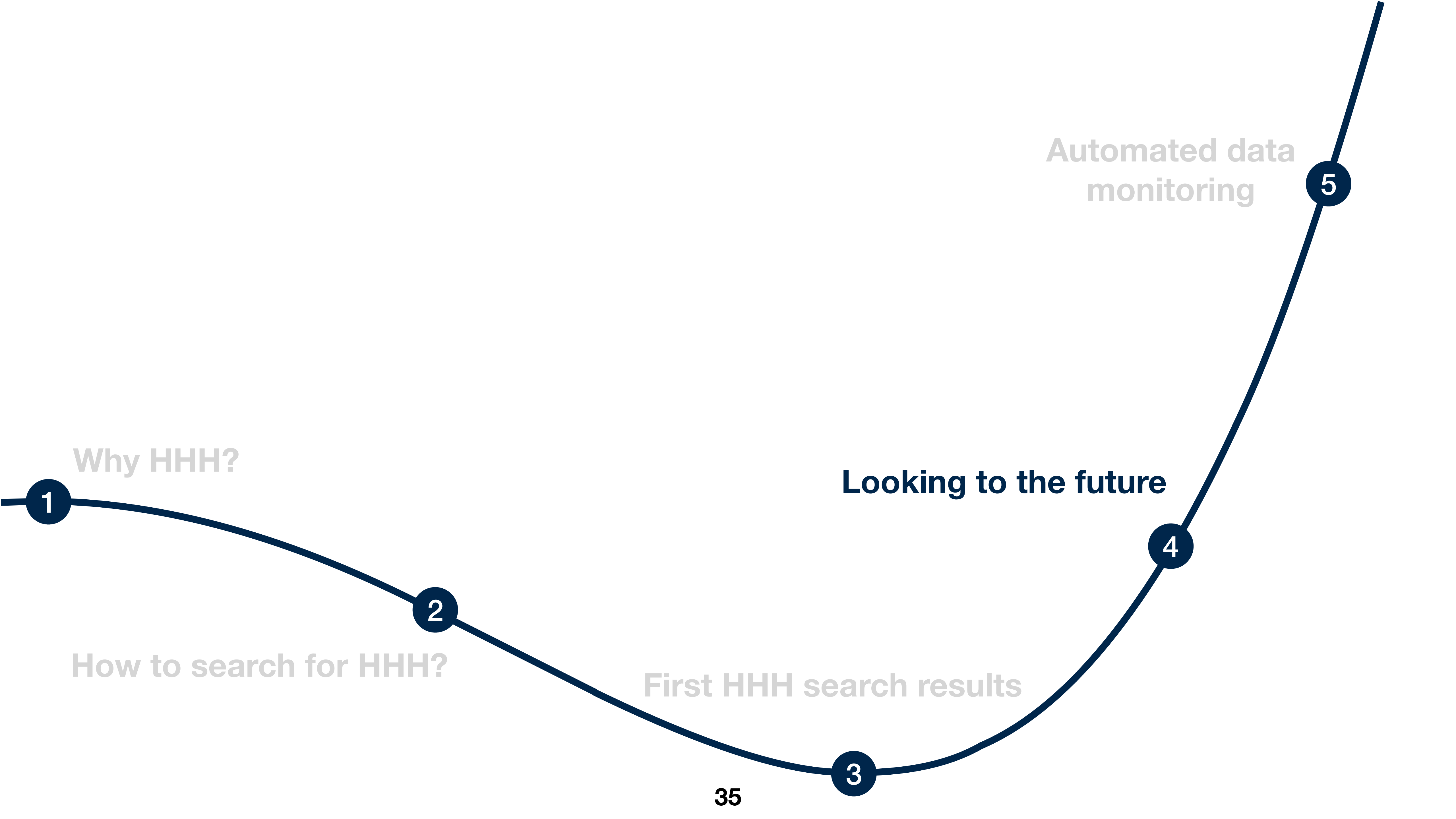
- The first search for HHH production was recently [published](#)
- [TRSM model](#) predicts HHH production cross sections up to ~ 150 fb
- HHH production limits are on the same order as predicted cross section (50-350 fb as a function of m_X, m_S)
- Searches for BSM physics in HHH is relevant TODAY!



Conclusions: SM-like interpretation

- The first search for HHH production was recently [published](#)
- Results include first ever limits on κ_4 , the quartic Higgs self-coupling
- Current results suggest $\kappa_4 = 1 \pm 30$ is an ambitious target at the HL-LHC.
 - **ATLAS projects in progress. This is back of envelope math**
 - Need 10x improvement on top of lumi (6 ab^{-1}).
 - With no improvement other than lumi, more like $\kappa_4 = 1 \pm 100$
 - Some improvement expected from:
 - Multiple channels
 - Reduction of theoretical uncertainties
- HHH may be especially suited to AI/ML due to the complicated final state.





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Low hanging fruit



Low hanging fruit

- **Theoretical uncertainties**

- A HHH cross-section at 13 TeV was not available in the literature.
- We computed a quadratic fit to the 14 TeV, 27 TeV, 100 TeV cross-sections and added an additional 20% uncertainty.
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- **Exploration of more channels**

- There is slight tension in phenomenological studies with respect to which channel is best
- This is suggestive that in a HHH combination many channels would contribute

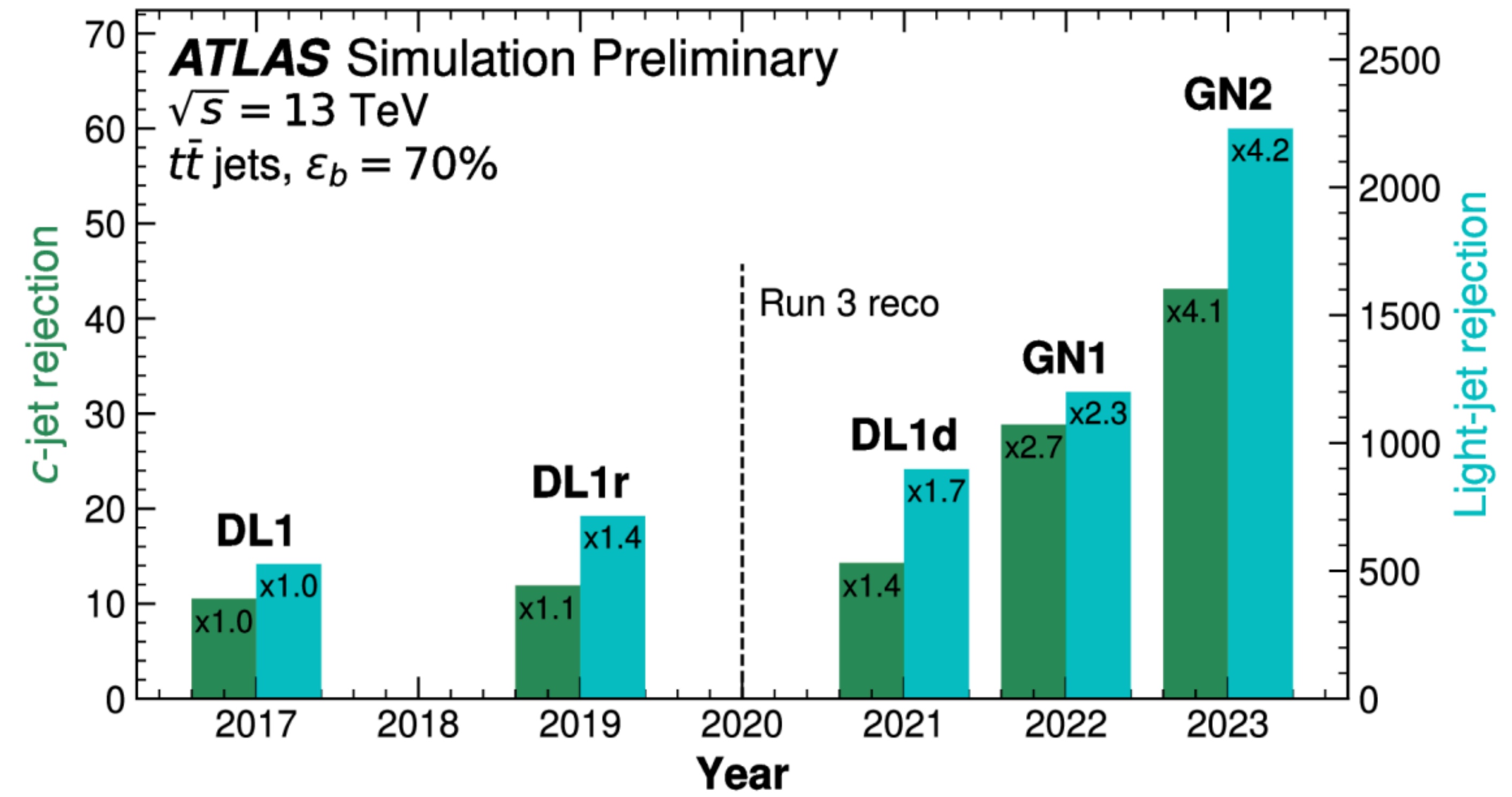
Decay Mode	Branching Fraction
bbbbbb	19.5%
bbbb+1 lepton*	15.2%
bbbb+yy	0.2%
bbbb+taus(had)	2.5%
bb+SS leptons	1.9%

HHH as an AI laboratory



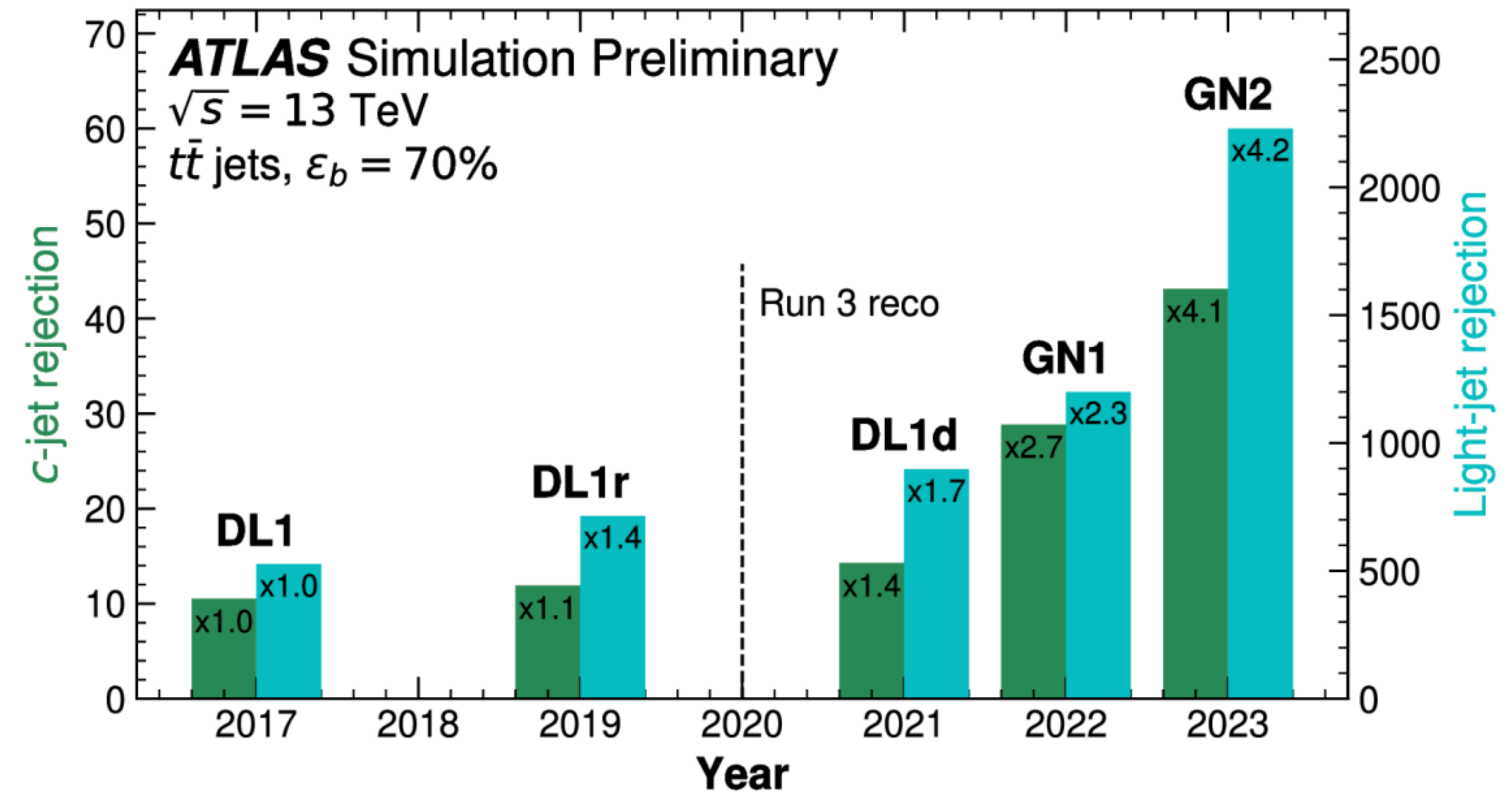
HHH as an AI laboratory

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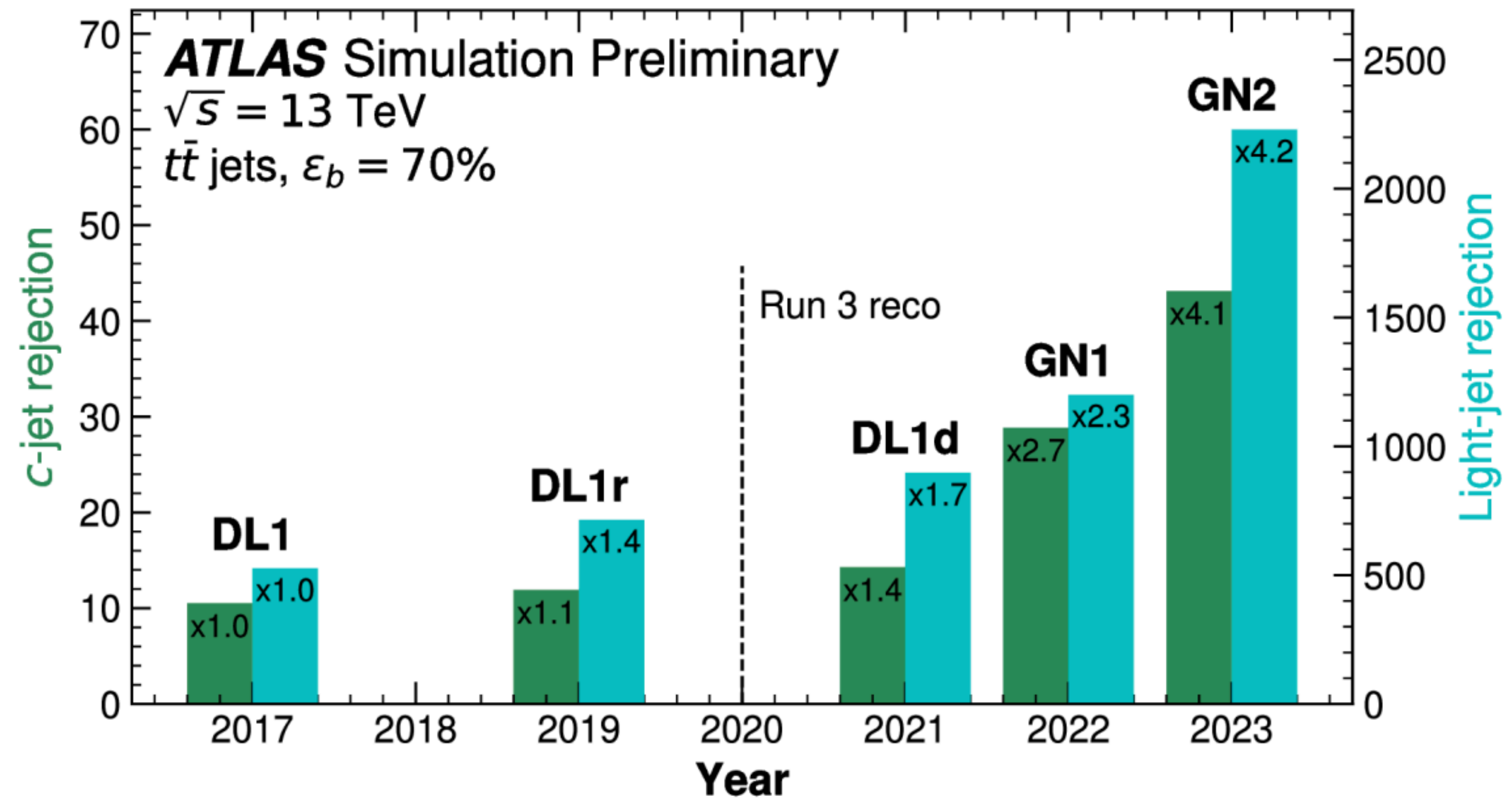
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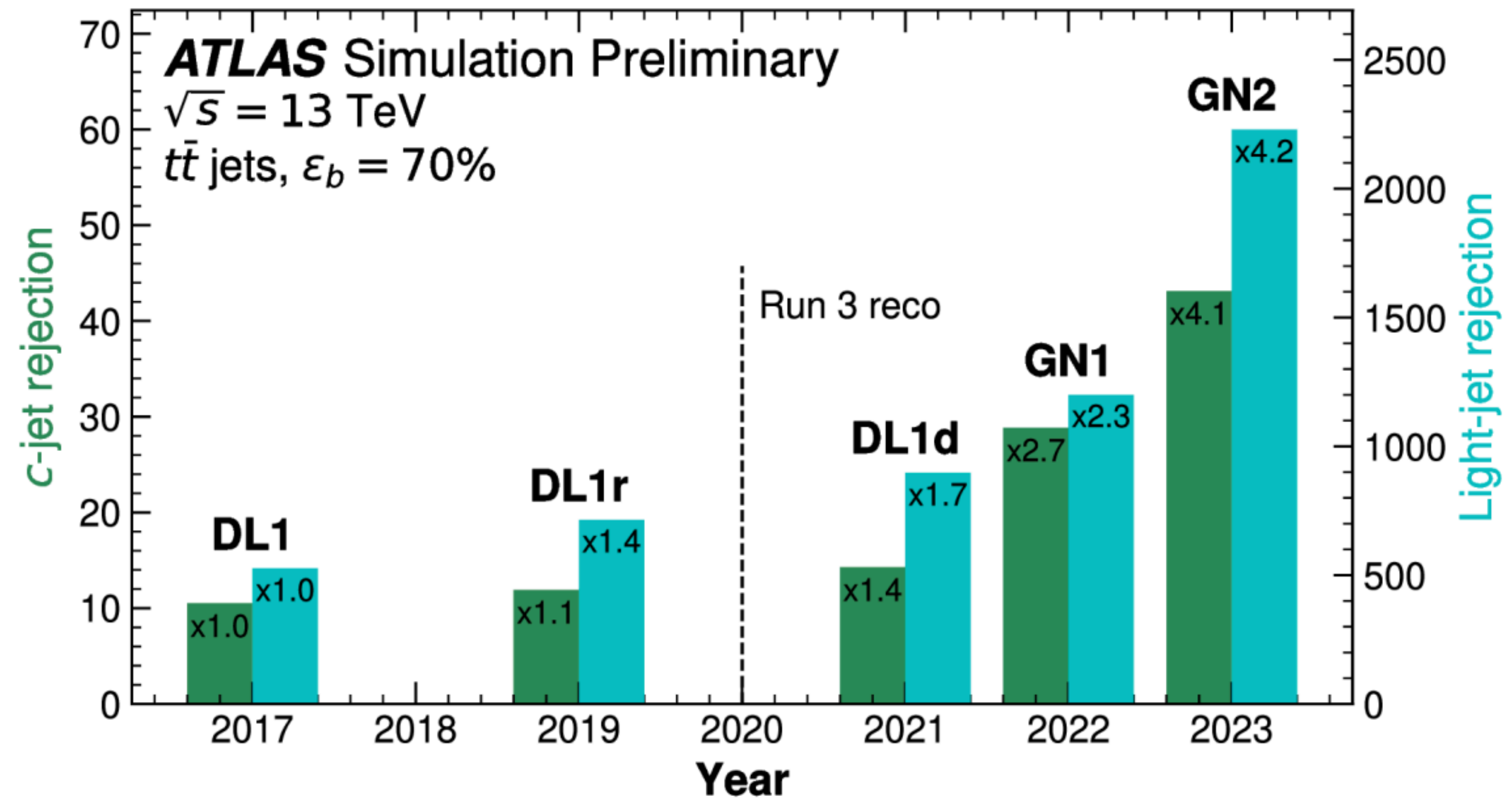
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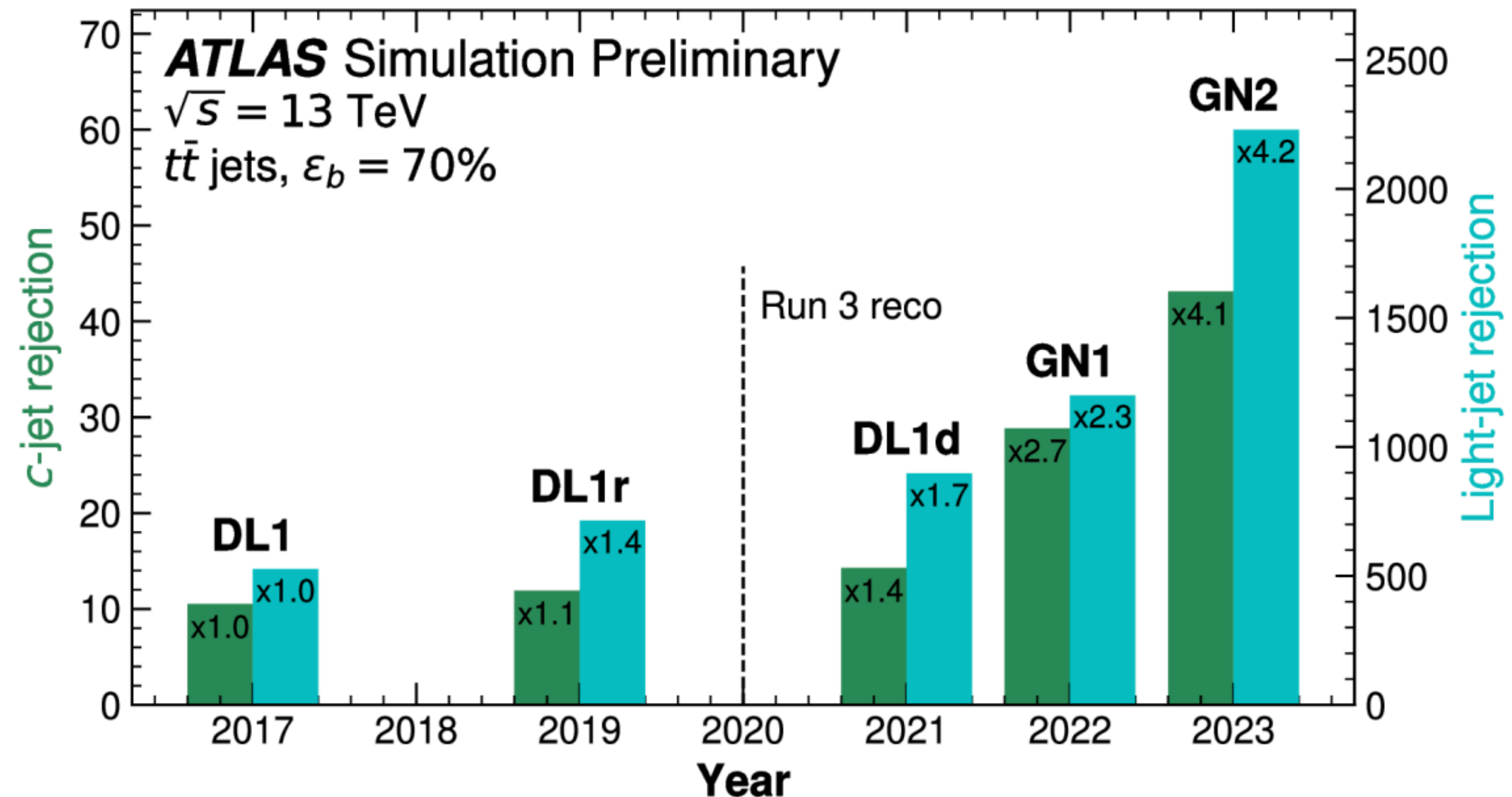
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- Need some classic datasets “MNIST of ATLAS/LHC” for ML challenges



Pairing b-jets



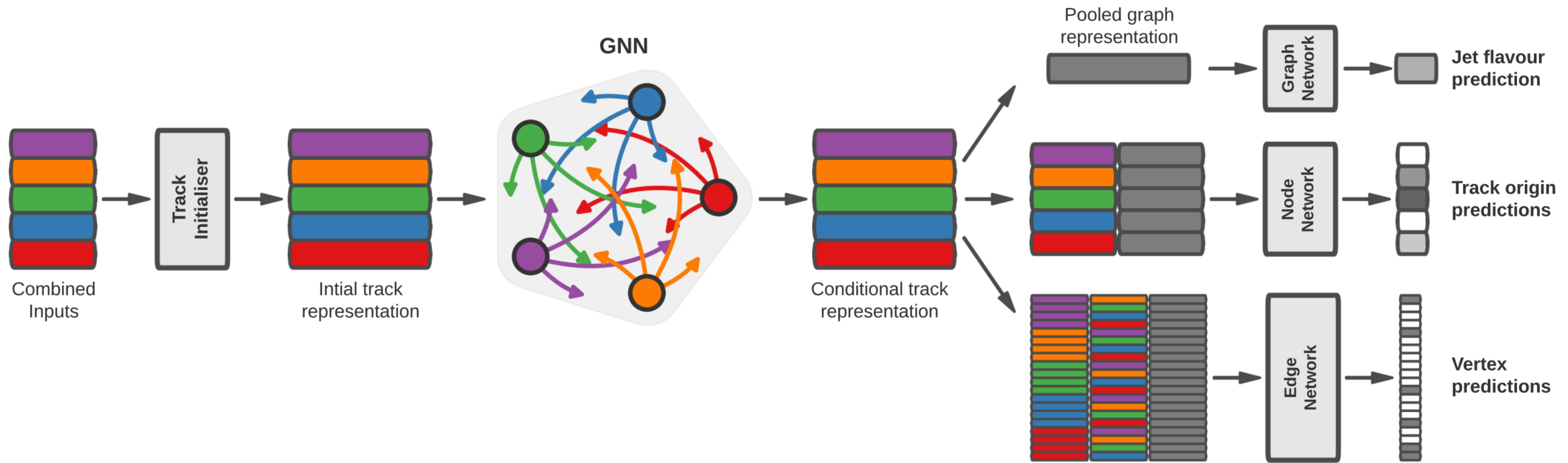
Pairing b-jets

- Pairing currently uses simplest-possible mass-based method.



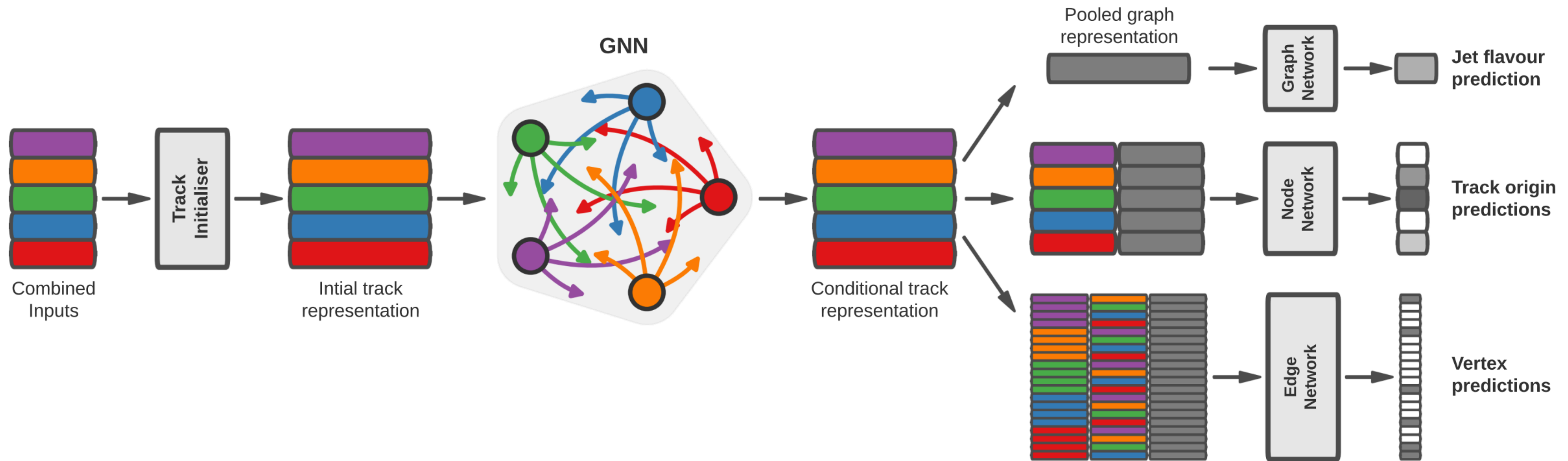
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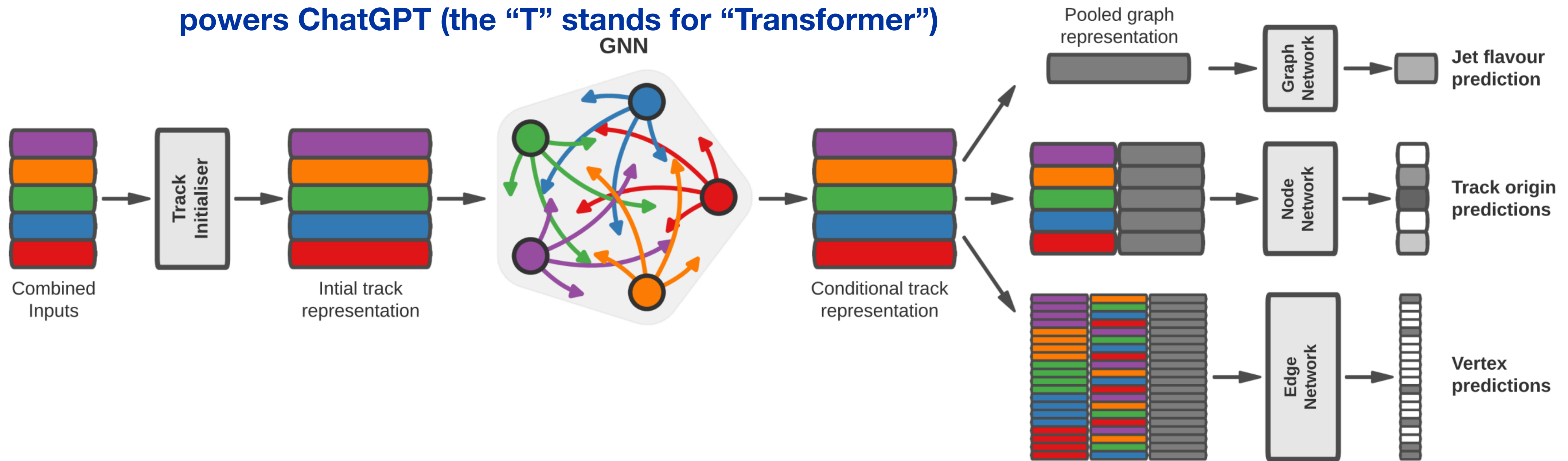


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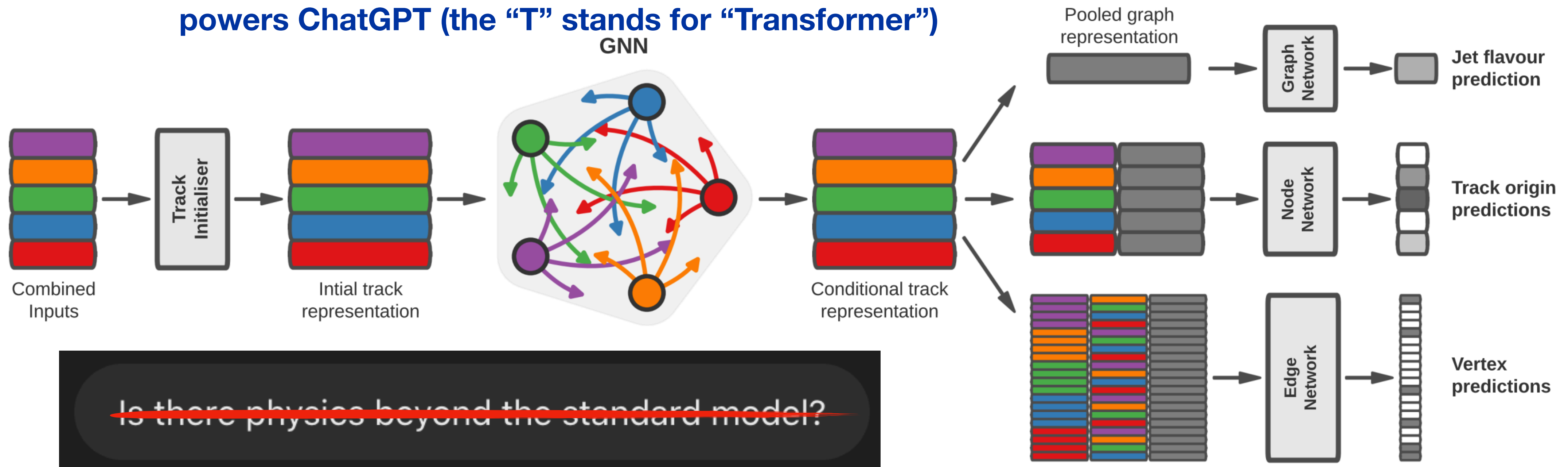


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~~Is there physics beyond the standard model?~~

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Why would transformers help pairing?

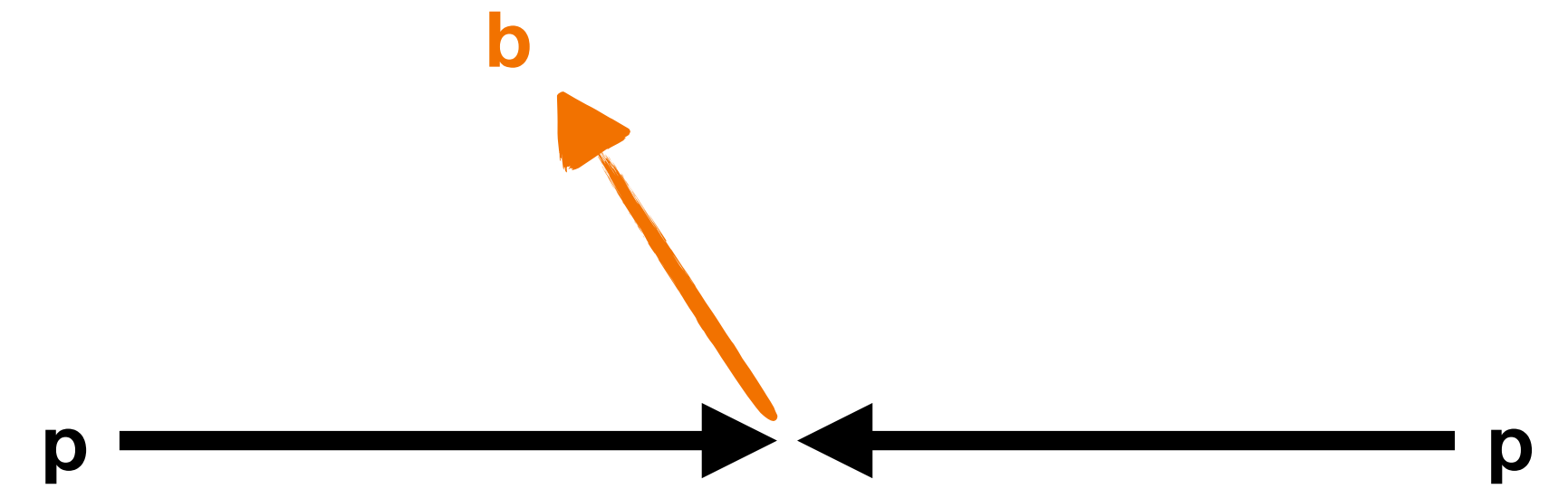


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- **Transformers allows for context-dependent encoding of particles**
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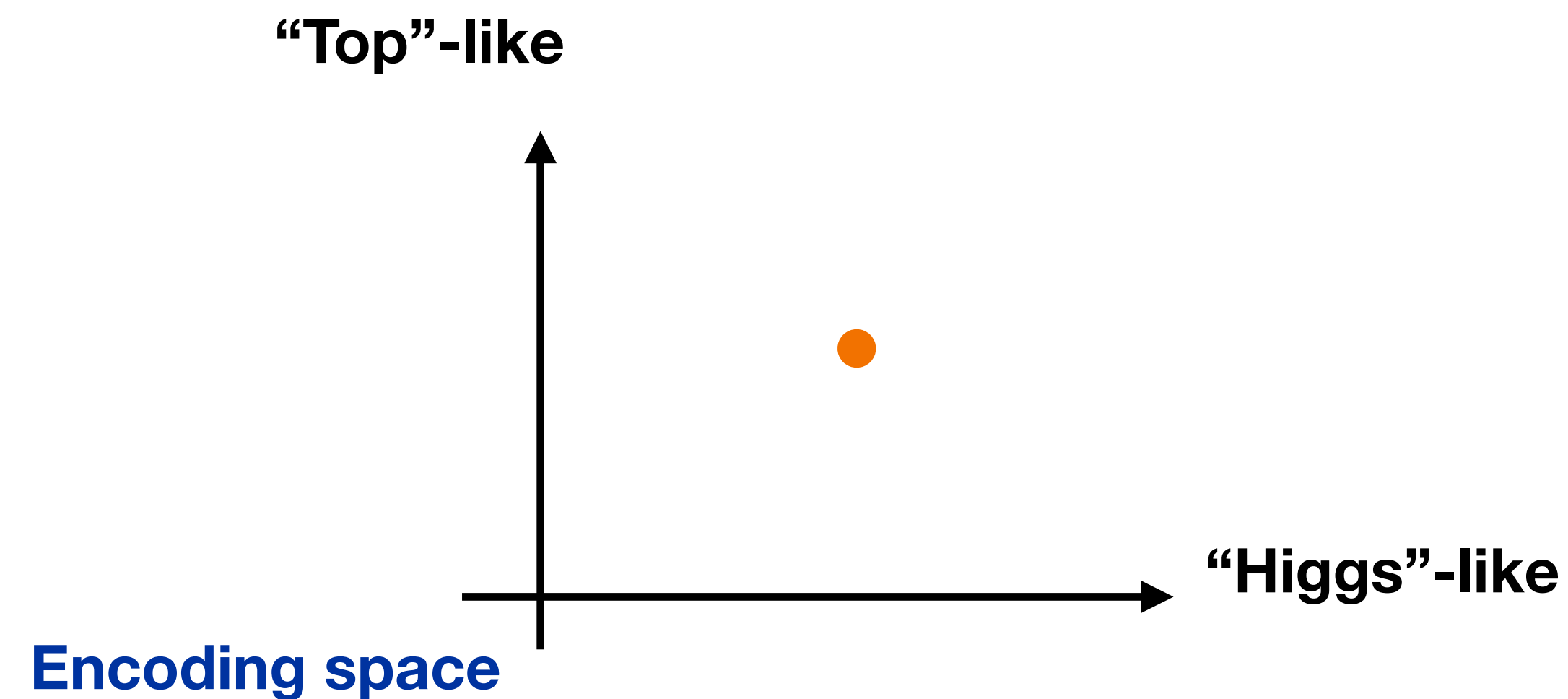
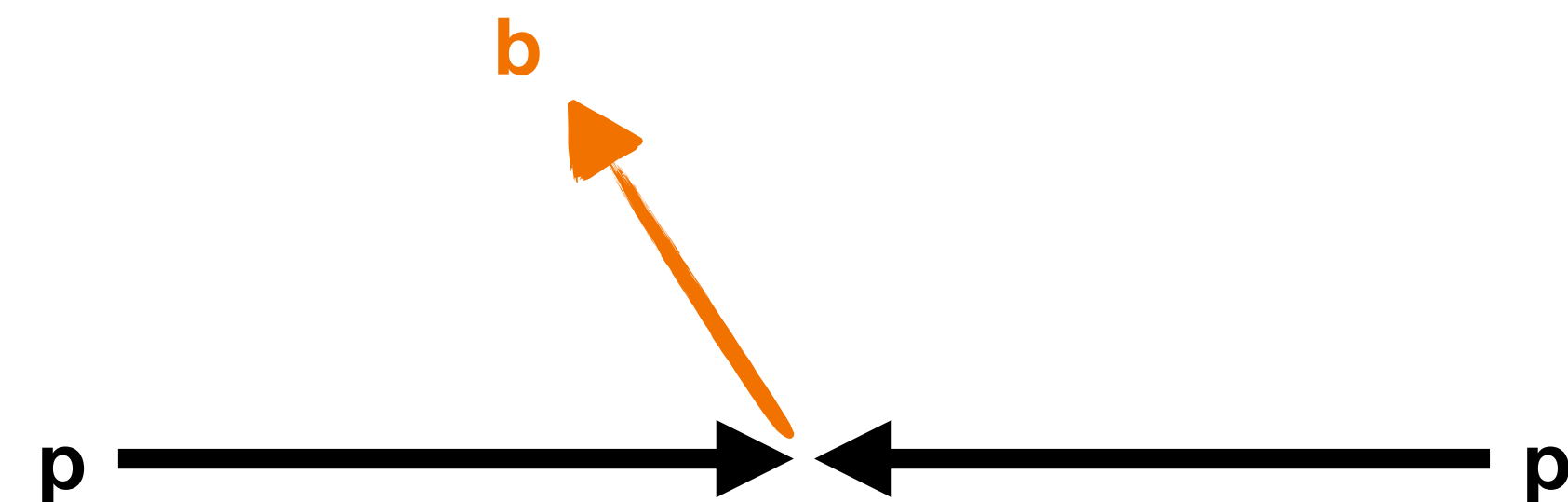
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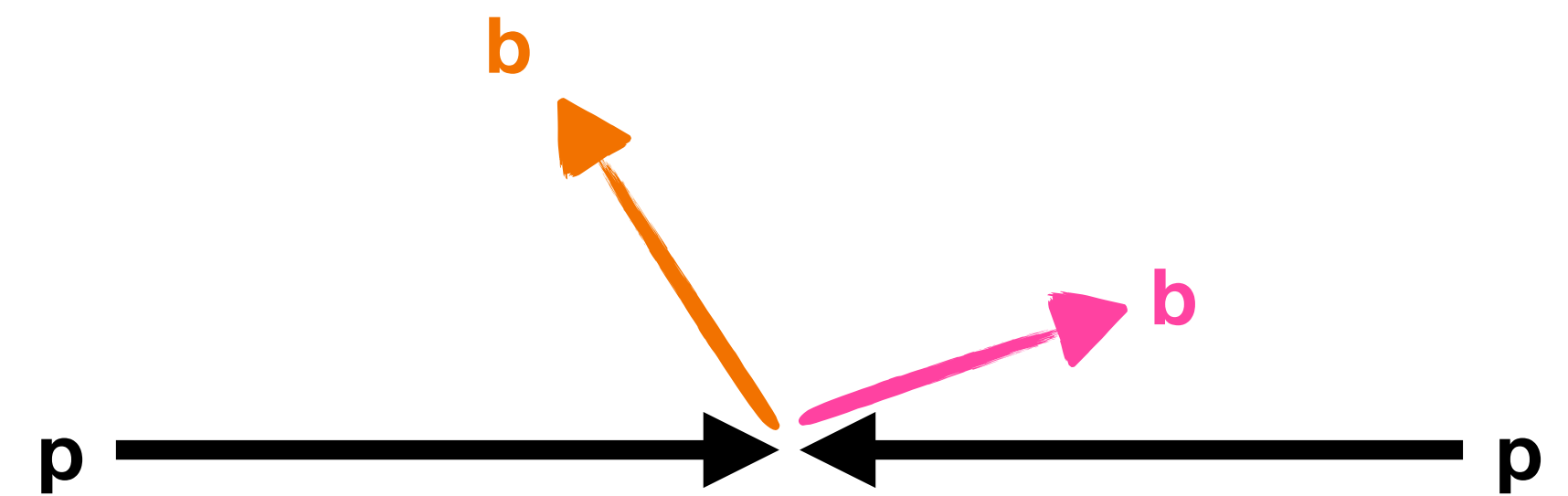
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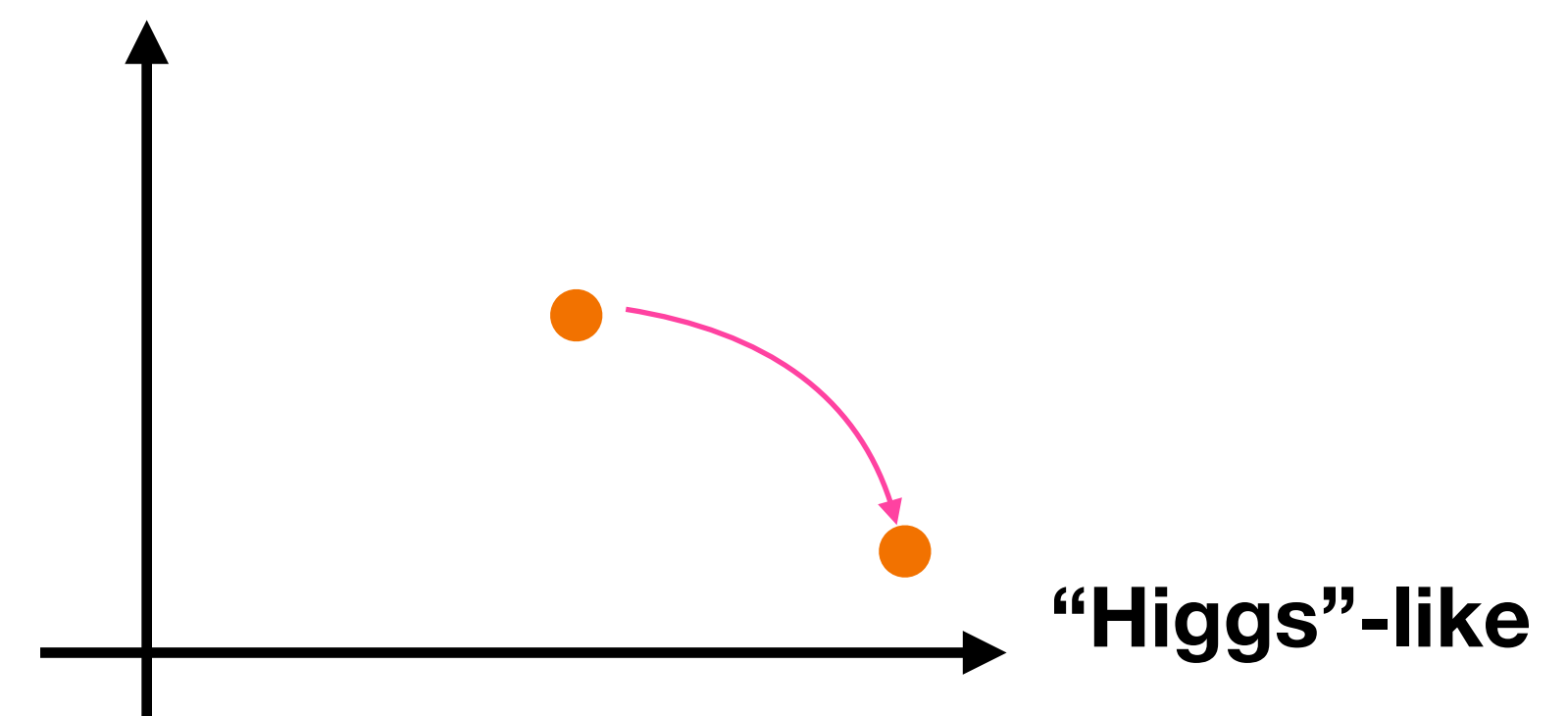


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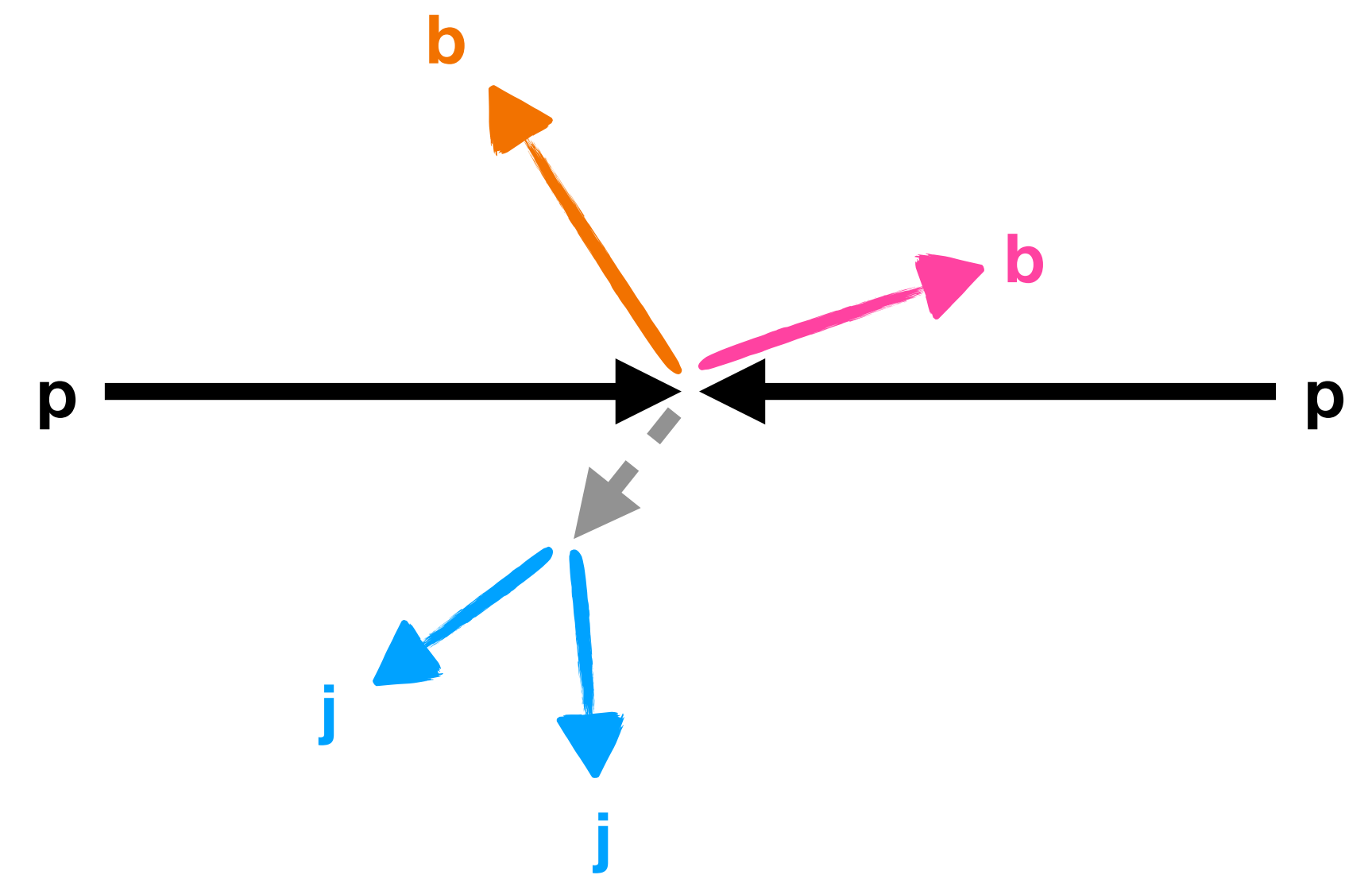


Encoding space

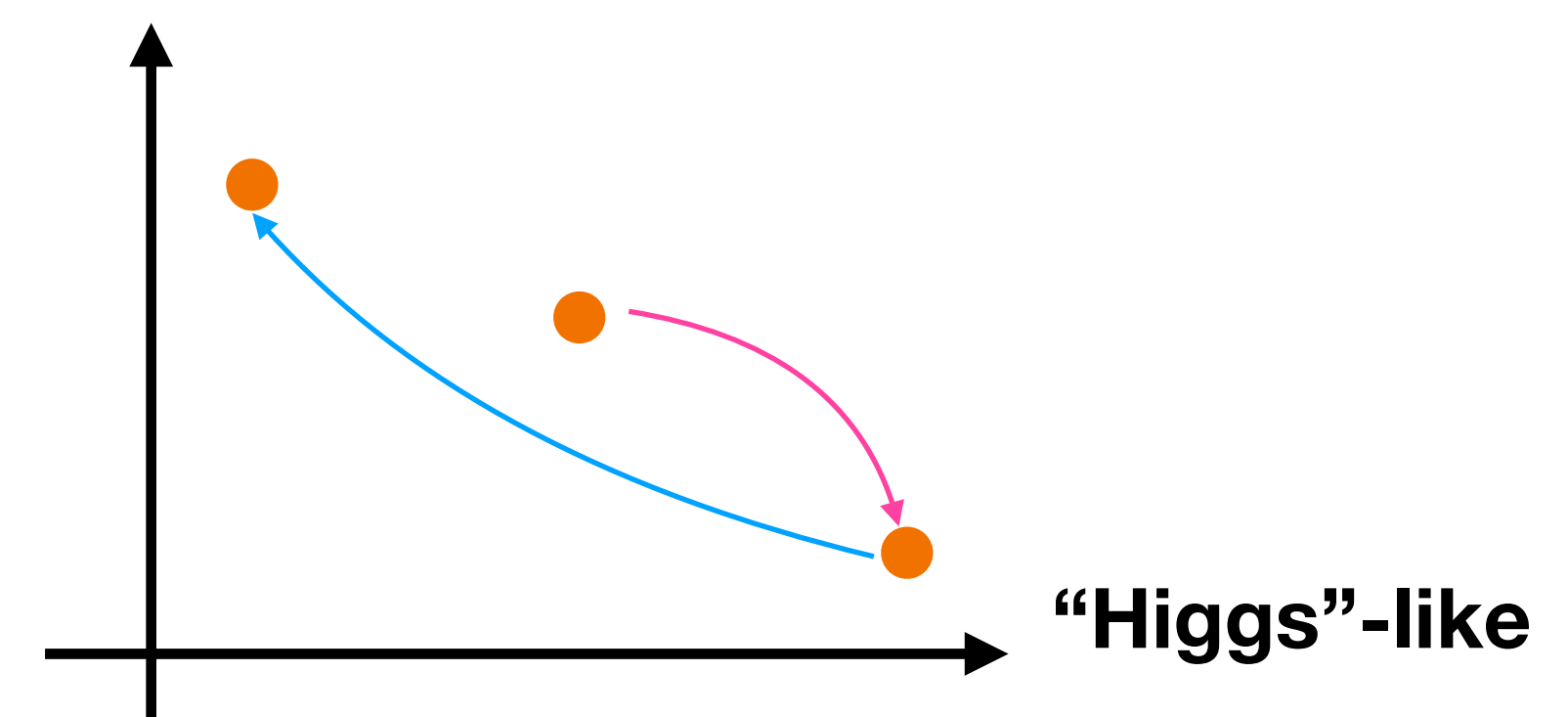


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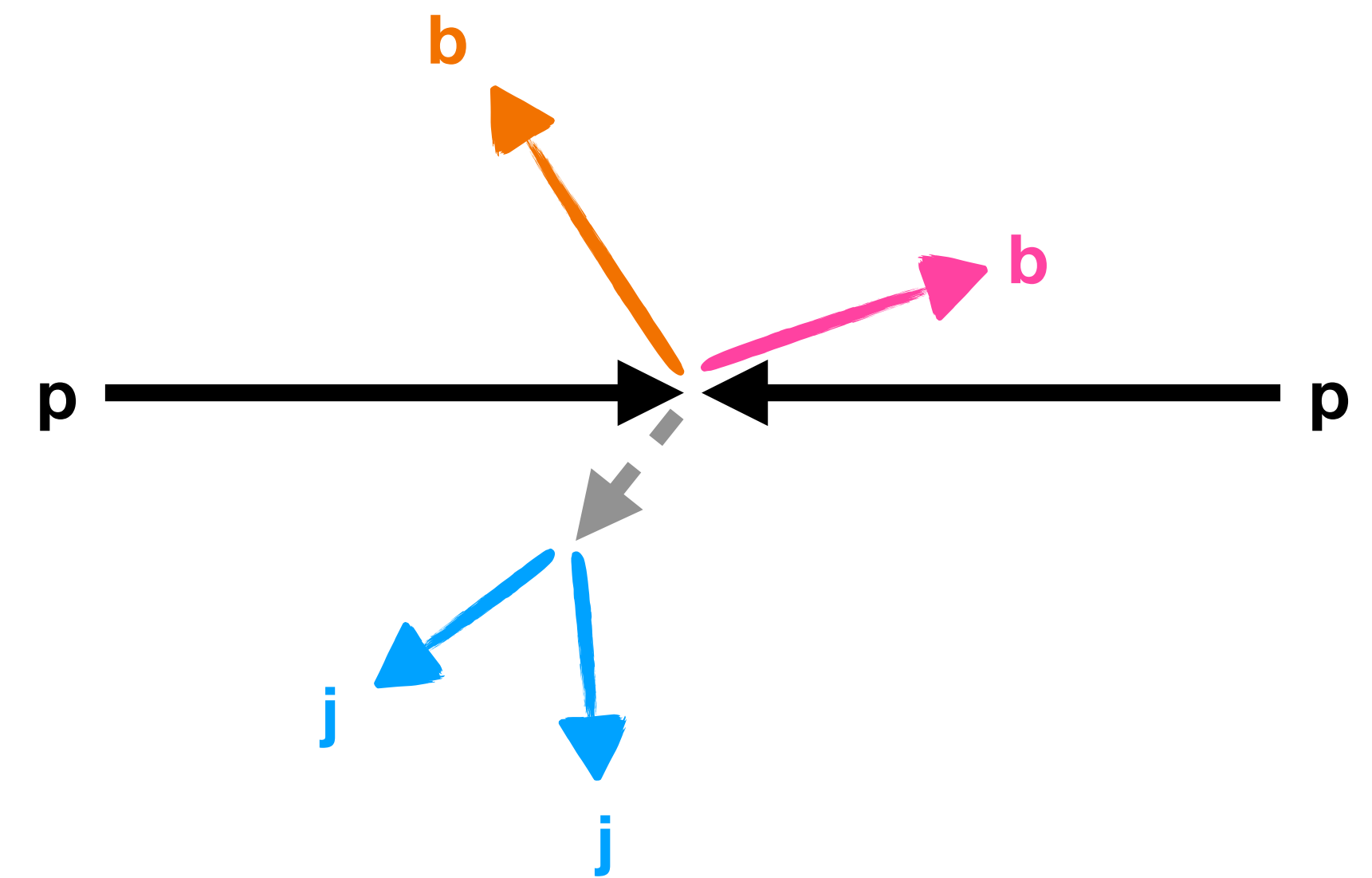


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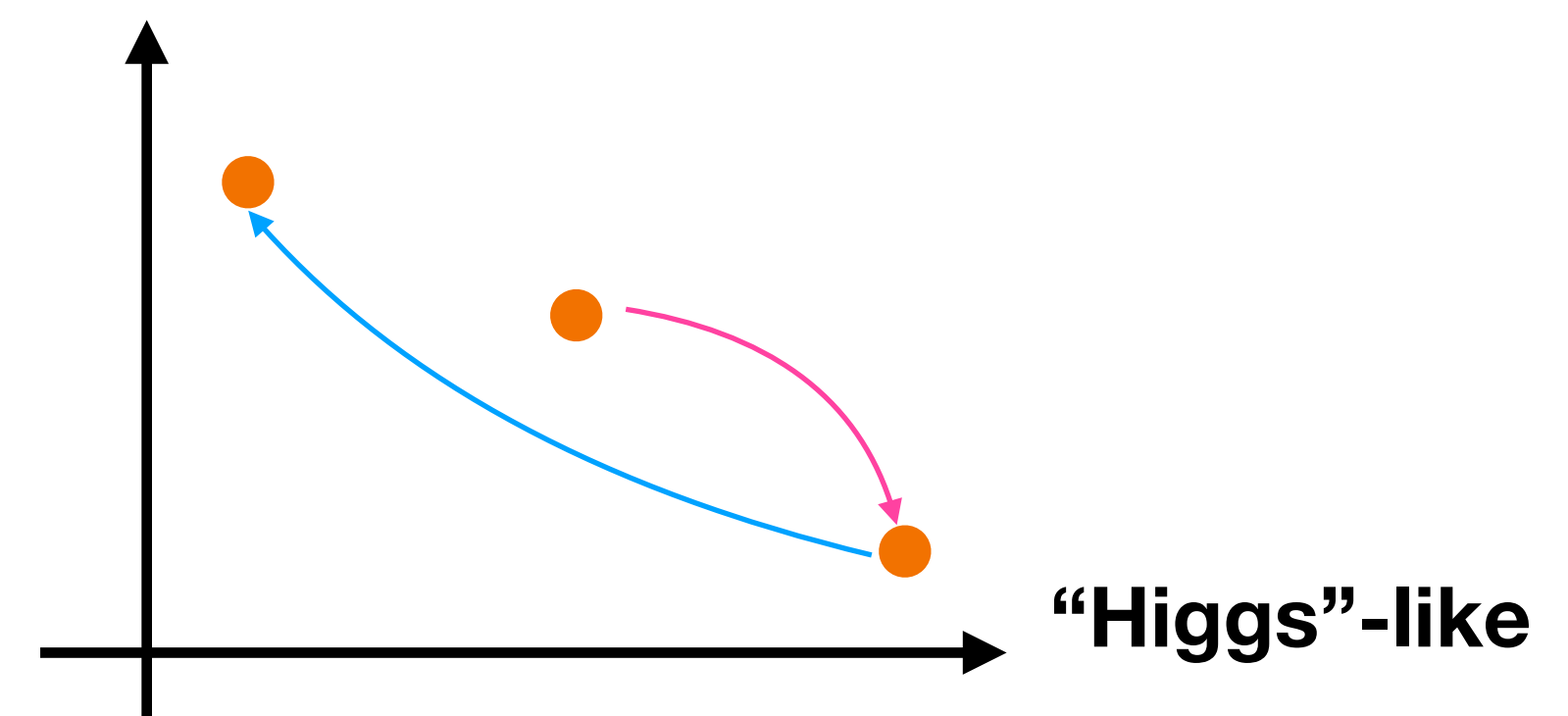


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- Transformers can add these conflicting contexts in the embedding space and pair in auxiliary tasks



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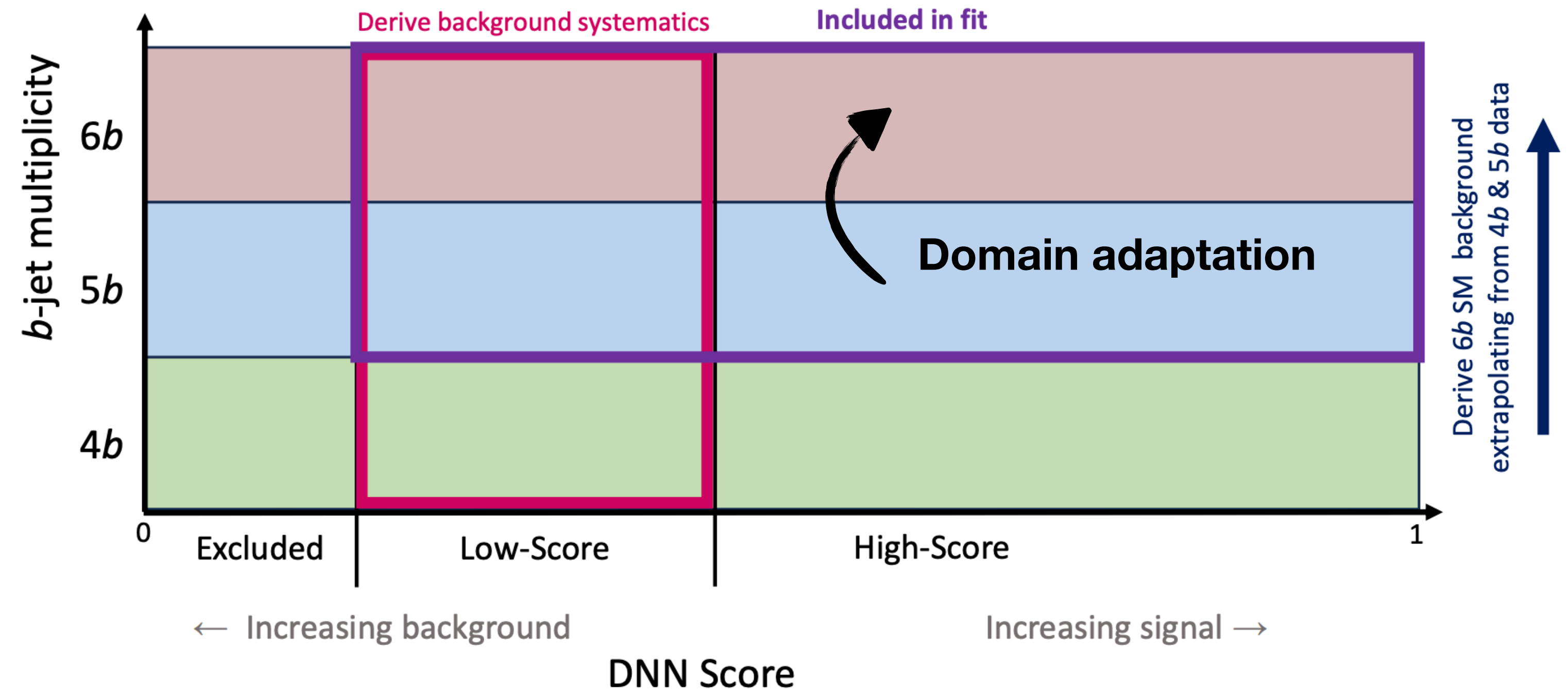


Domain adaptation



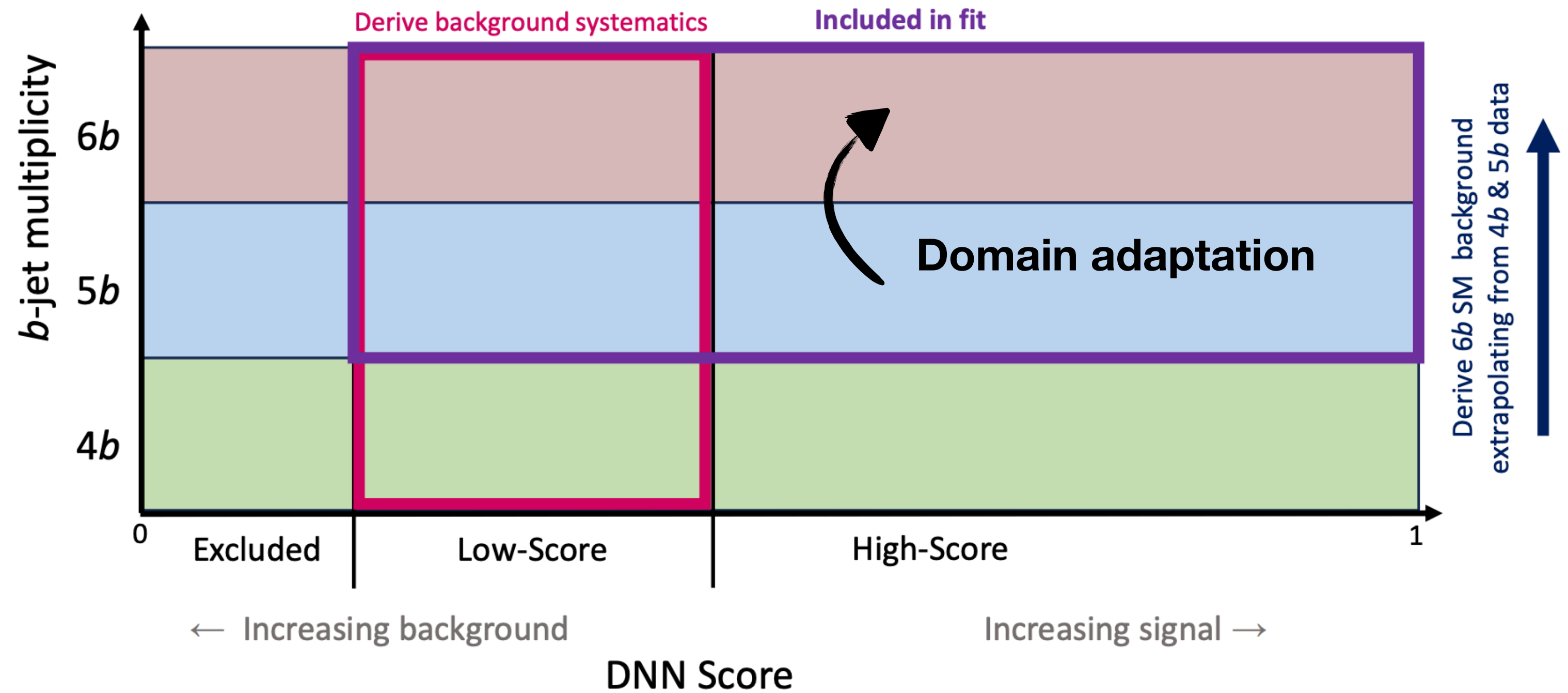
Domain adaptation

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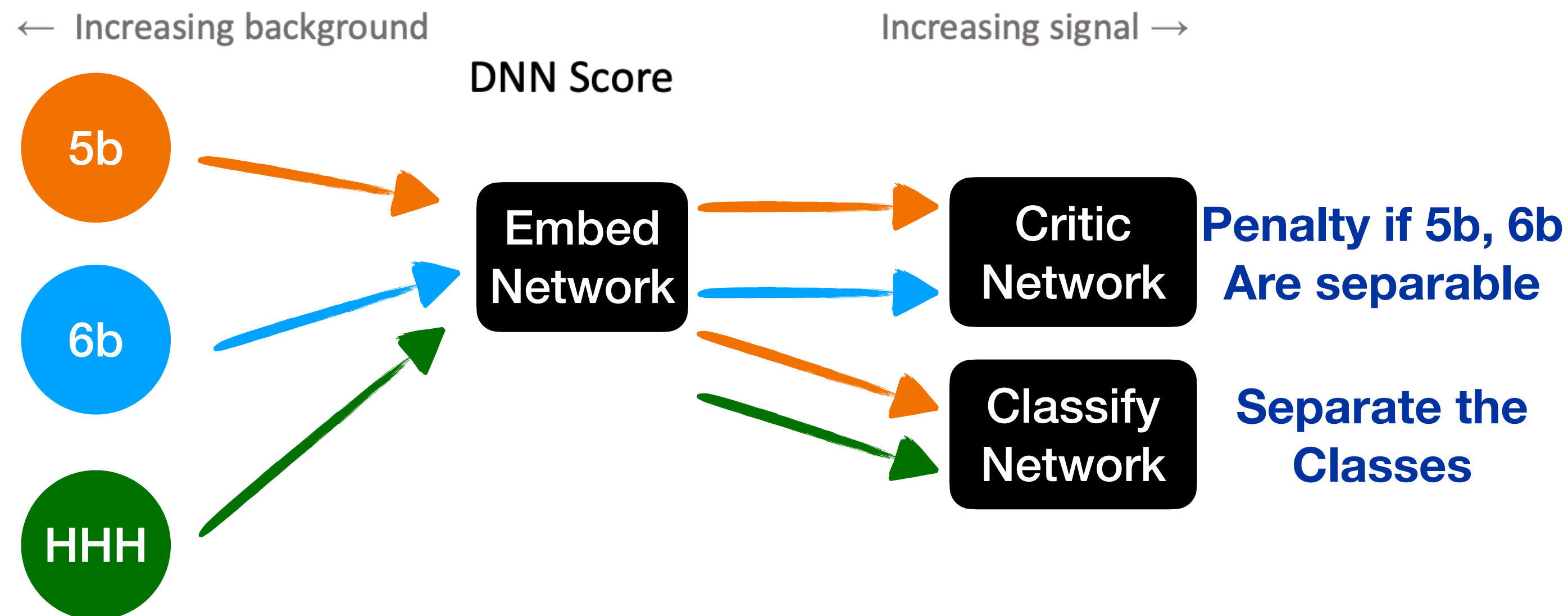
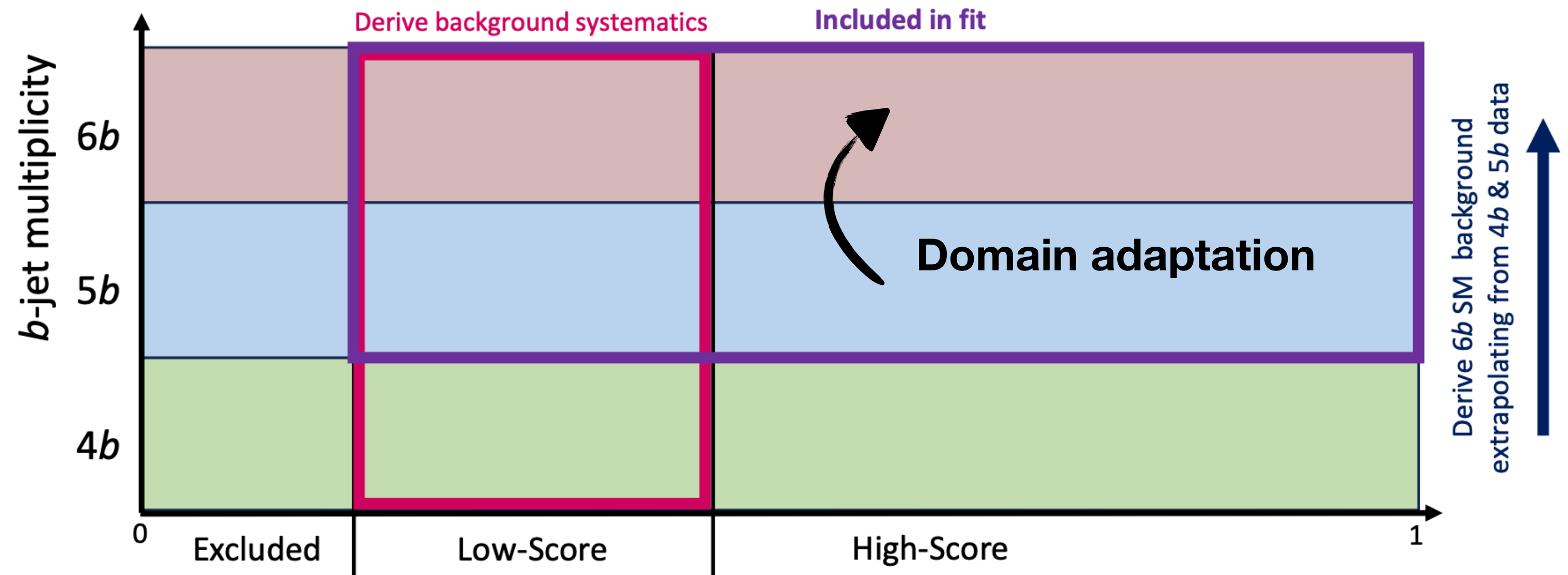
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Domain adaptation

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- We selected input features which minimized difference between 5b and 6b data
 - But this may have left us with a non-optimal DNN
- One possible solution is to use an adversarial training, in which a critic network penalizes the training if it can learn any difference between 5b and 6b.
- The loss function has some “pressure” to learn difference between 5b and HHH signal, but also a competing “pressure” to learn no difference between 5b and 6b.



Functional decomposition



Functional decomposition

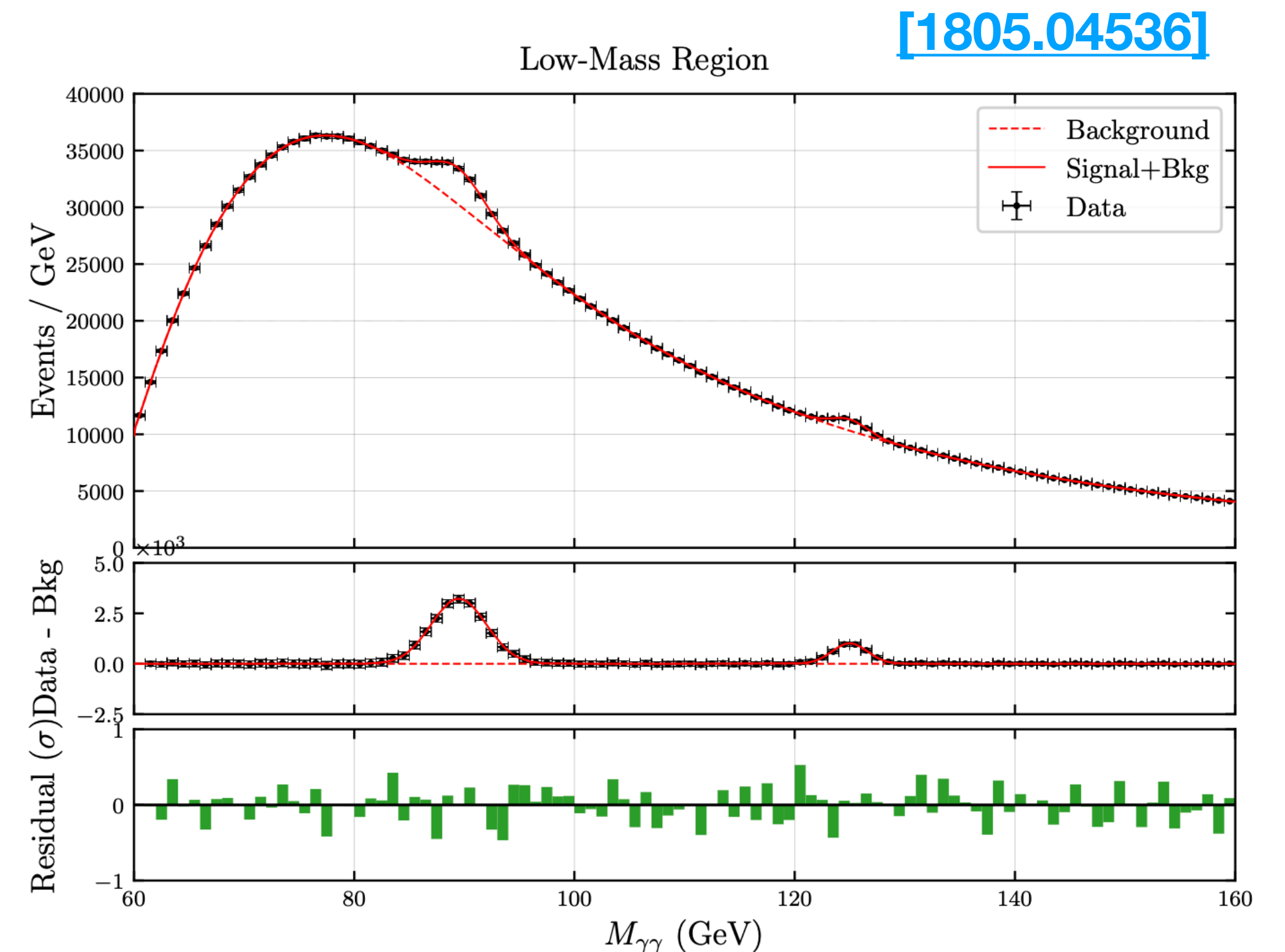
- **Background modeling uncertainties are currently dominant systematic uncertainties.**
 - Generally 2x bigger than FTAG, JET/MET, etc *combined*.

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All uncertainties	24
Experimental	22
Detector response	7.4
Background modeling	16

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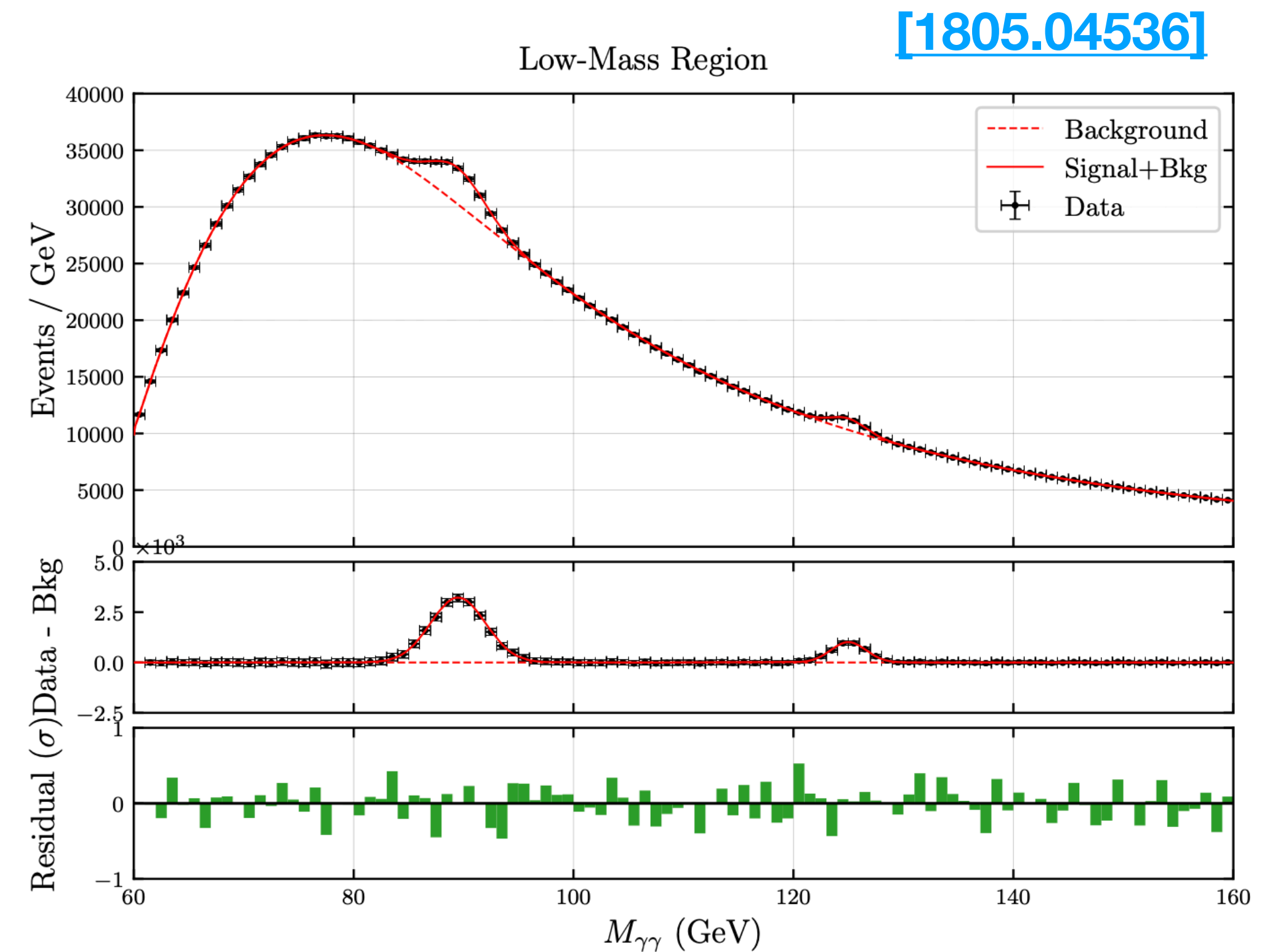
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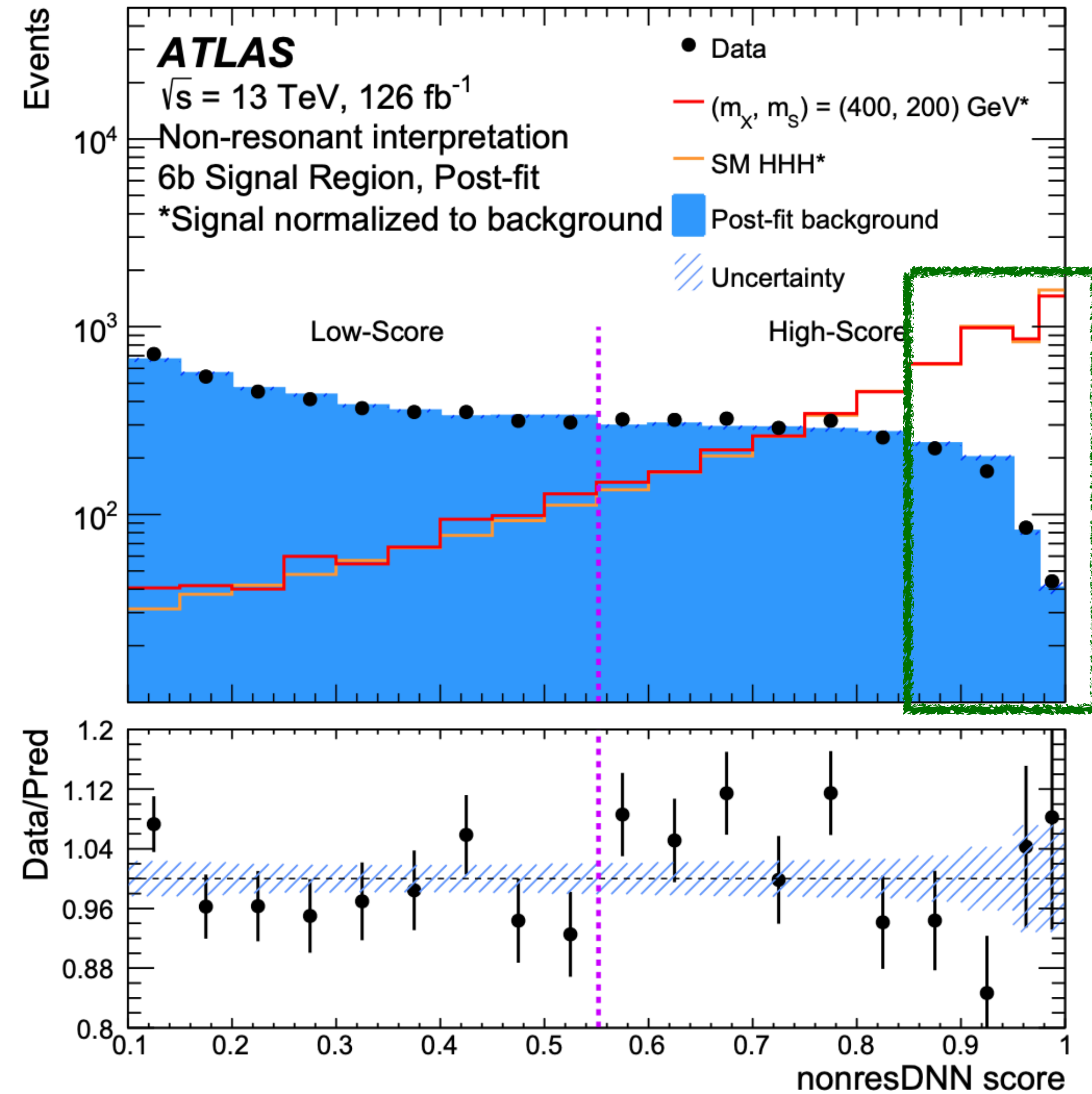
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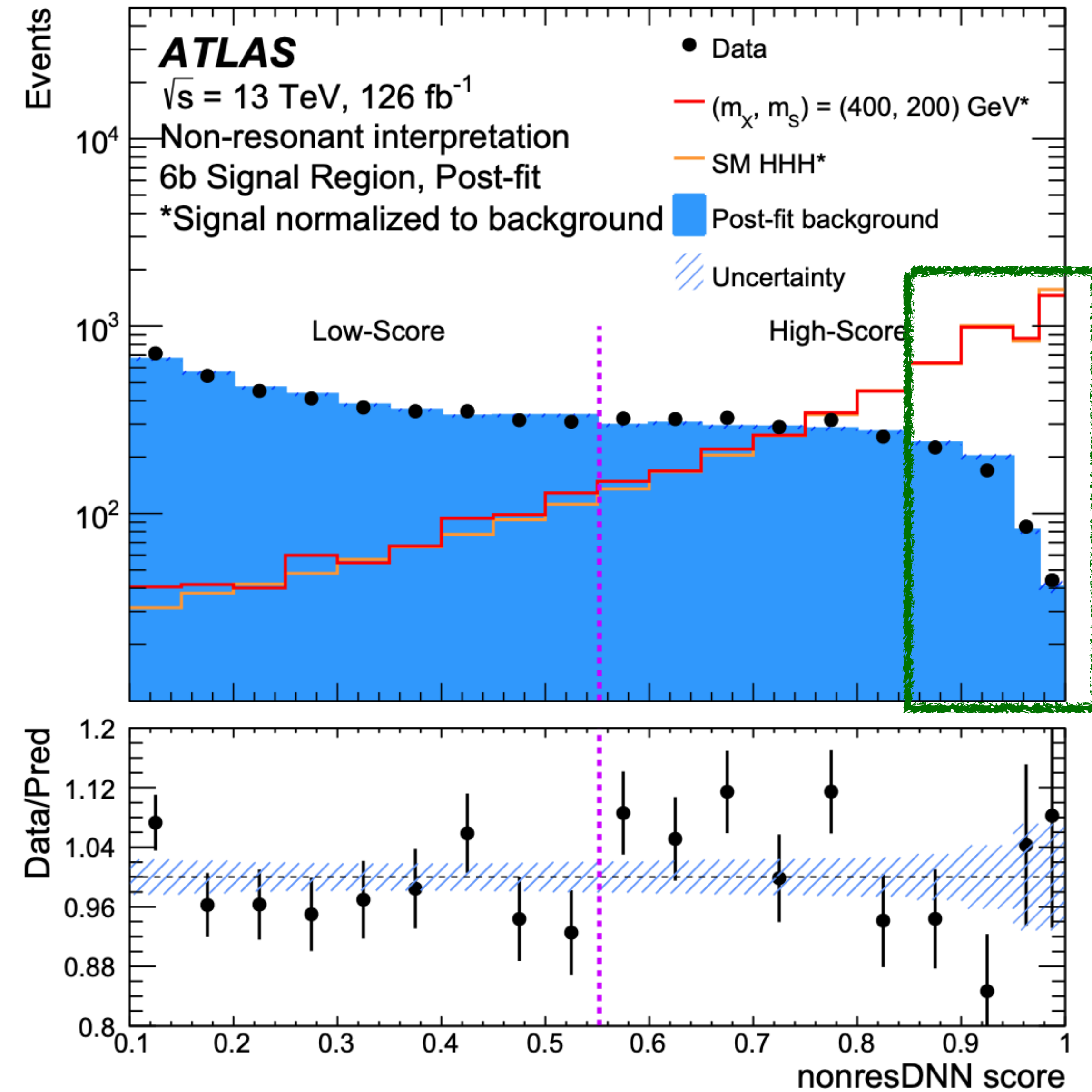


Differentiable p-values with FD



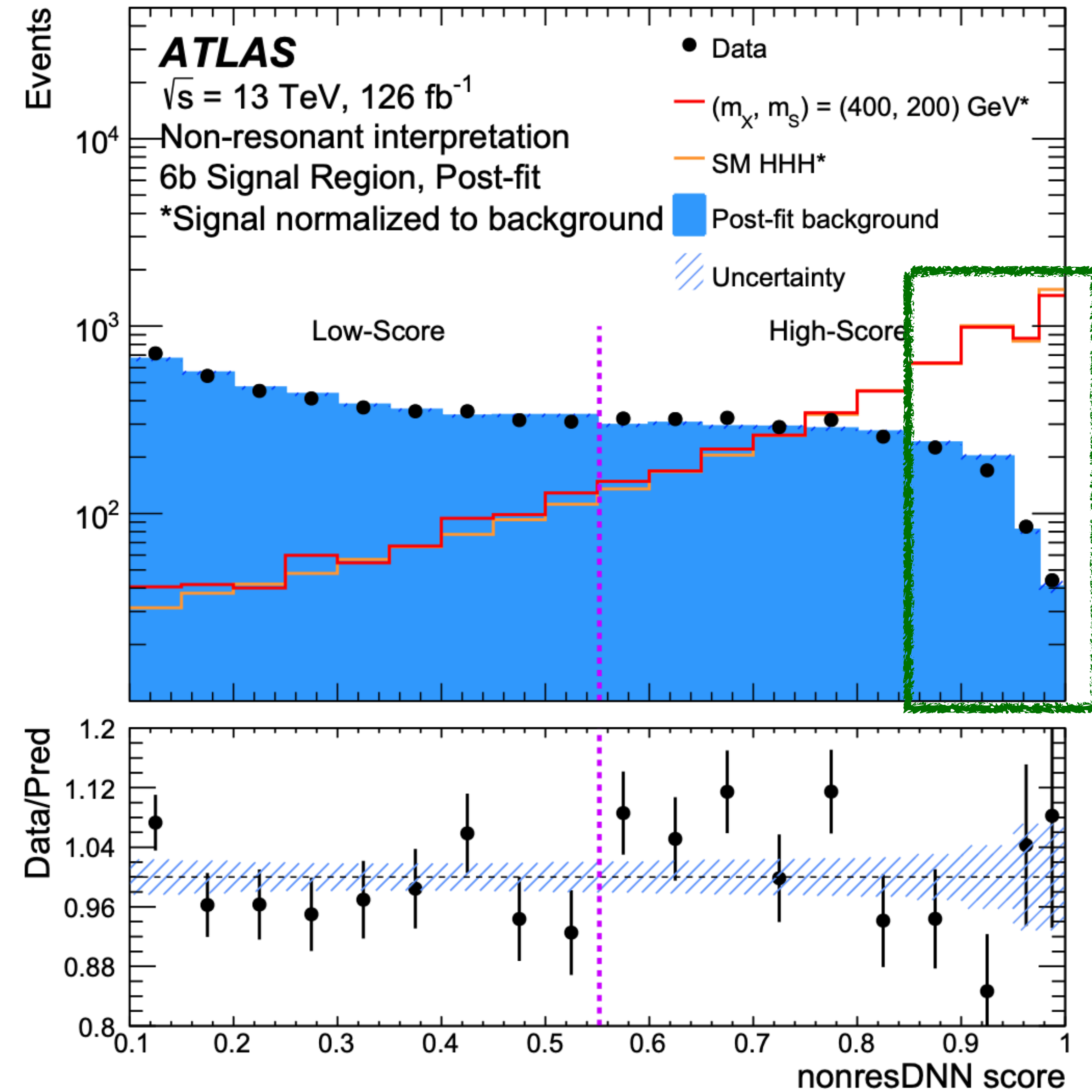
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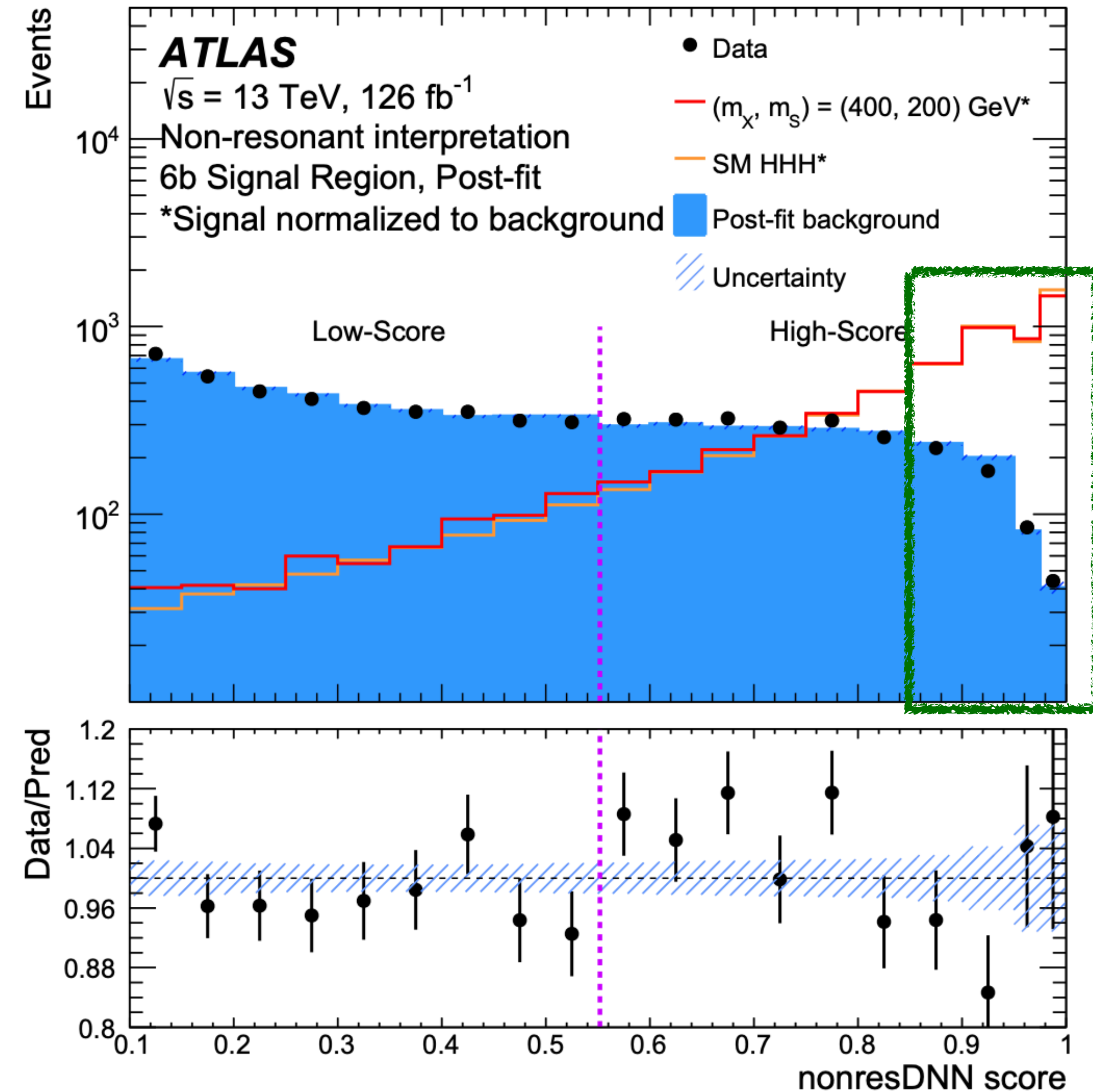
Differentiable p-values with FD

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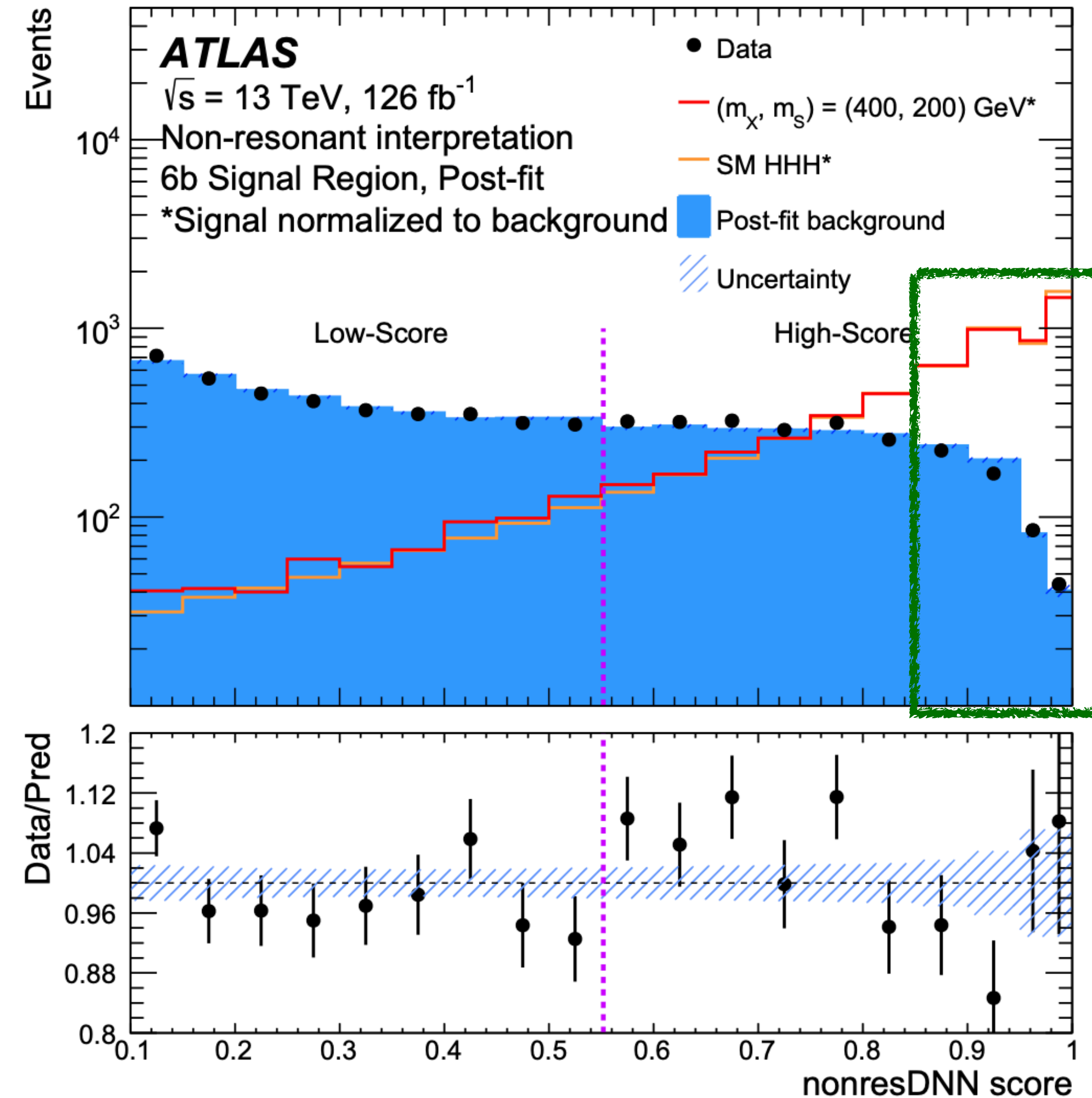
Differentiable p-values with FD

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- **Major caveat: of course we need to be careful not to over-fit! This is active research...**



Differentiable p-values with FD



Differentiable p-values with FD

- **FD is a computationally inexpensive way to compute p-values directly from unbinned data**
 - The covariance of the Hilbert space coefficients is directly calculable from the coefficients themselves
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 - Once we know the covariance, we can compute a p-value in the frequency domain

$$\begin{aligned}\hat{\Sigma}_{\tilde{f}_{nm}} &= \frac{1}{M} \sum_{i=1}^M E_n(z_i) E_m(z_i) - \tilde{\mathbf{f}}_n \tilde{\mathbf{f}}_m \\ &= \tilde{\mathbf{f}}^i \hat{\mathbf{I}}_{inm} - \tilde{\mathbf{f}}_n \tilde{\mathbf{f}}_m\end{aligned}$$

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Unbinned data
NN output score

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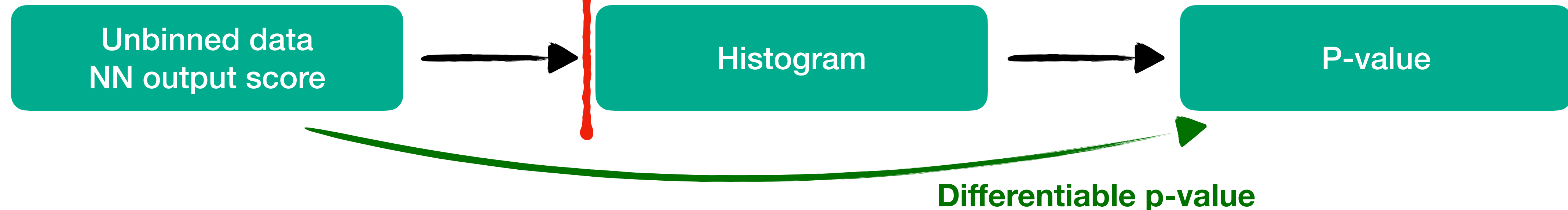


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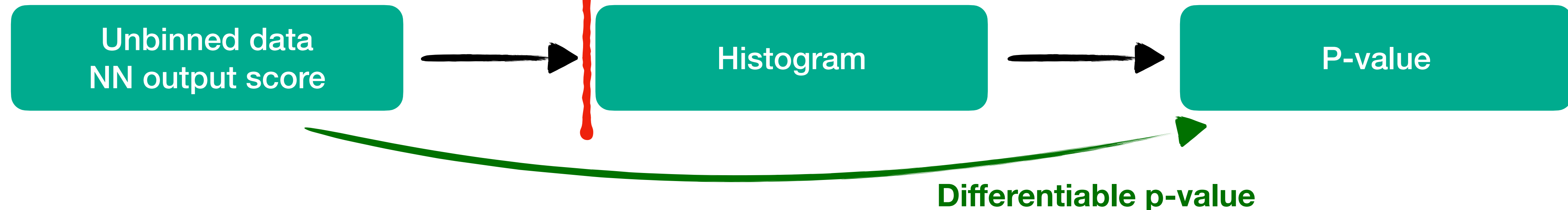
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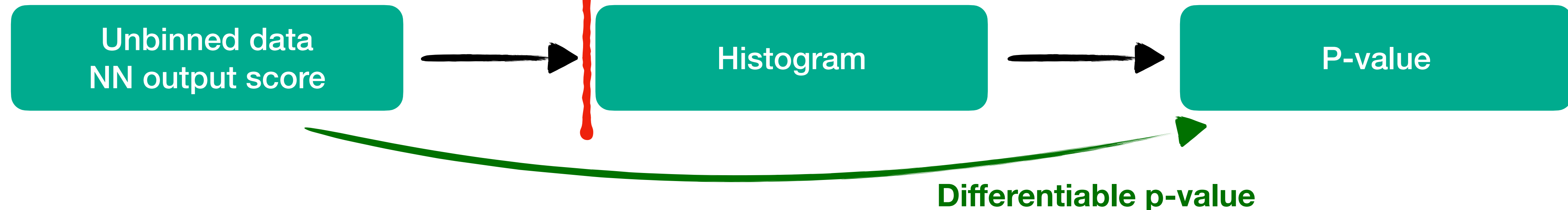
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- Another benefit: with FD background model is learned in a semi-supervised manner

Summary



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- **Yes!**
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 - Novel machine learning ideas: domain adaptation, functional decomposition, pairing tasks (GNN/transformer), etc.

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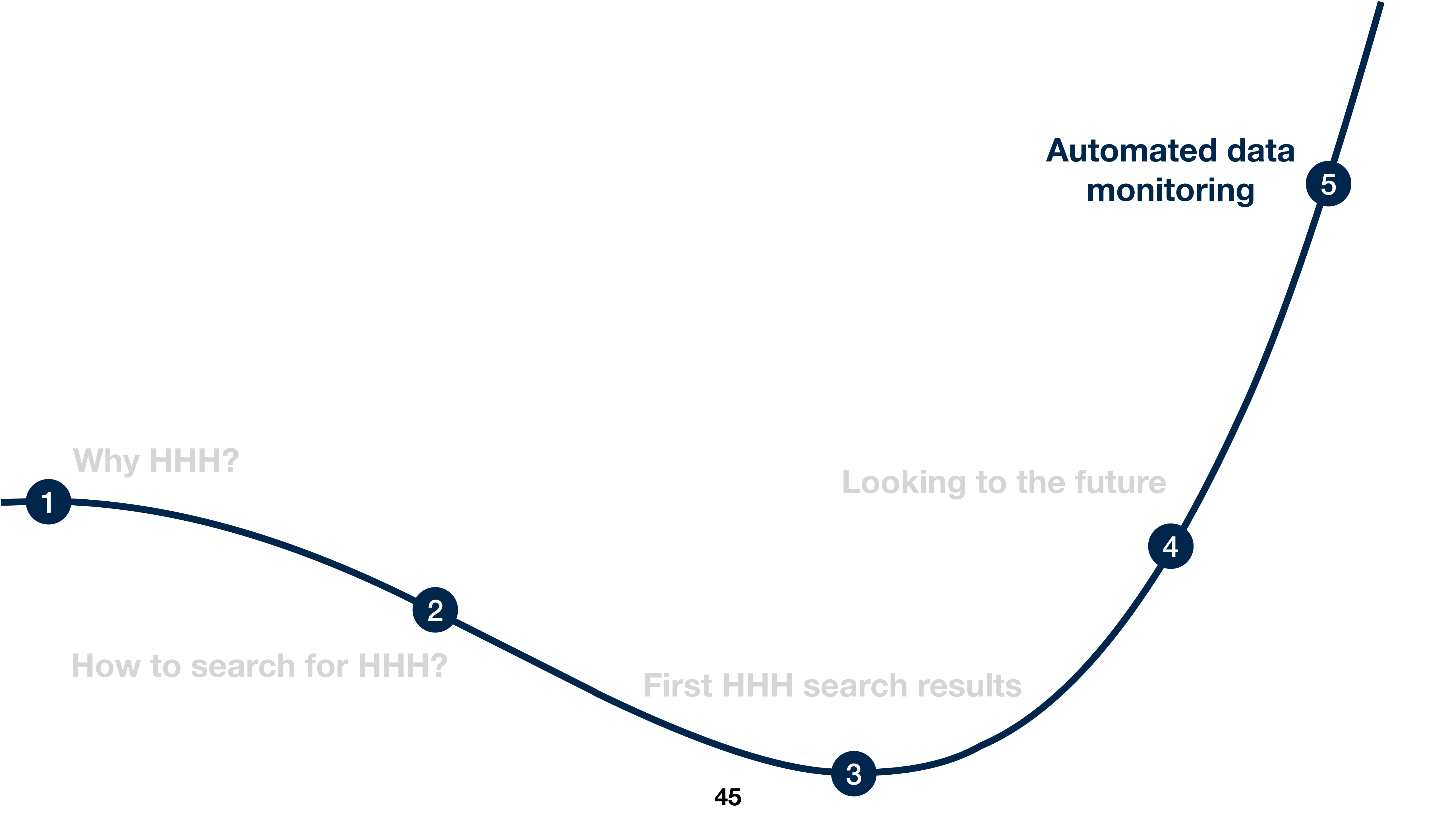
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- **Even if you aren't excited about HHH, is there reason to be excited about this talk?**
- **Yes!**
 - Many of the ML/AI ideas discussed at the end are entirely general!





1

Why HHH?

2

How to search for HHH?

3

First HHH search results

45

4

Looking to the future

5

Automated data monitoring

Muon performance studies

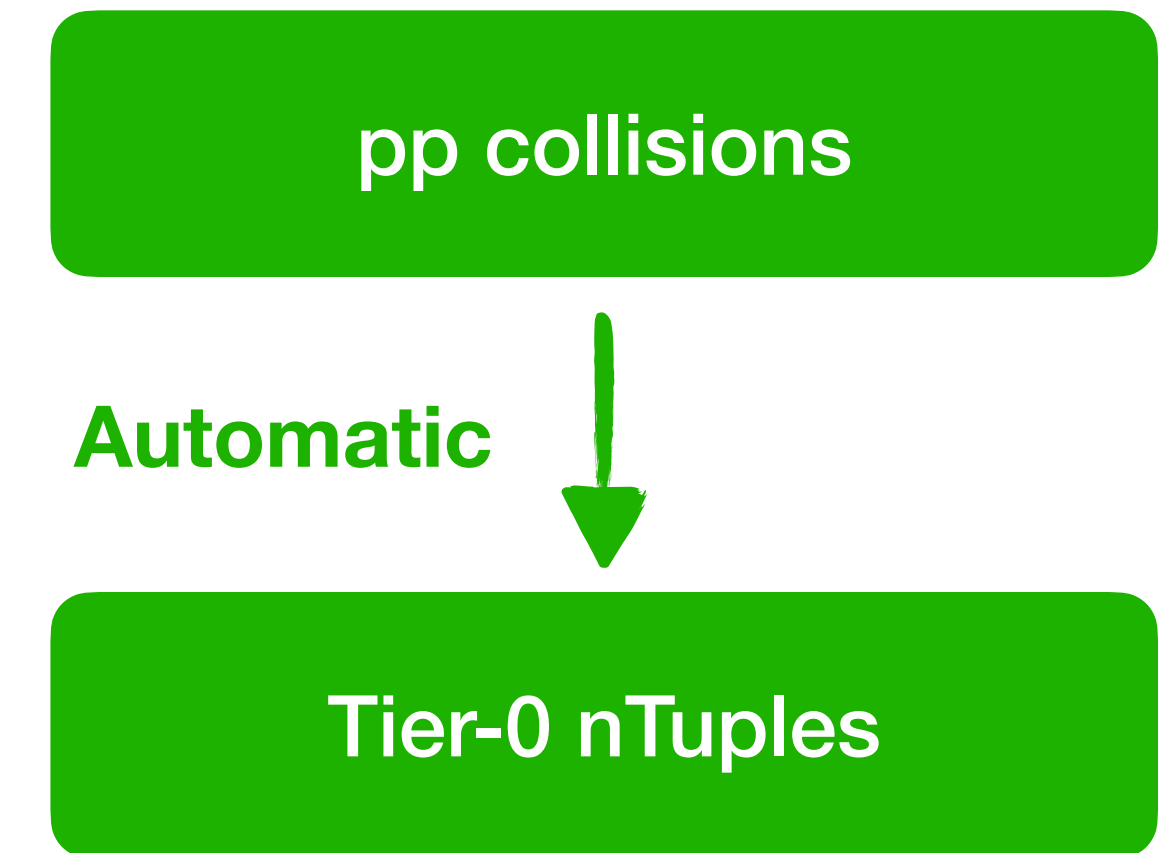


Muon performance studies

- In the muon performance group we analyze data and MC to:
 - Measure SFs
 - Calibrate muon p_T
 - Apply corrections to data

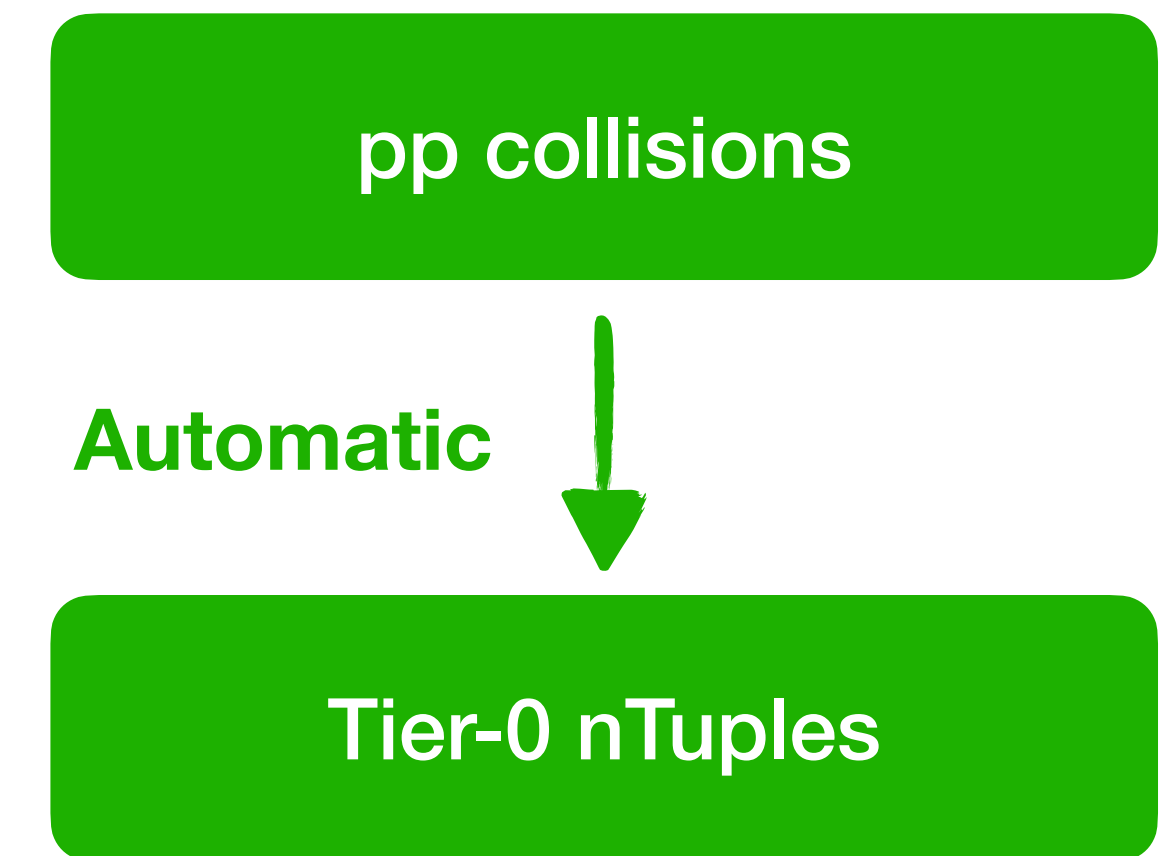
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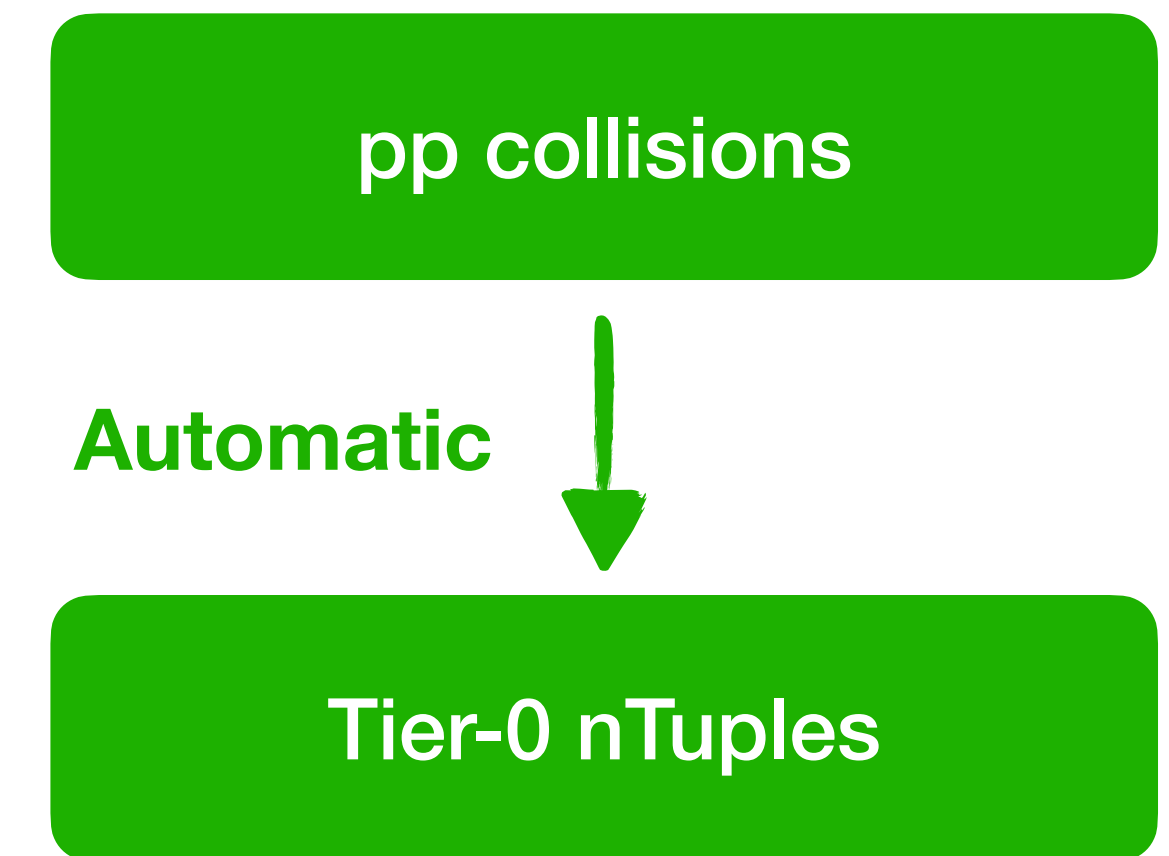


Muon performance analysis toolkit



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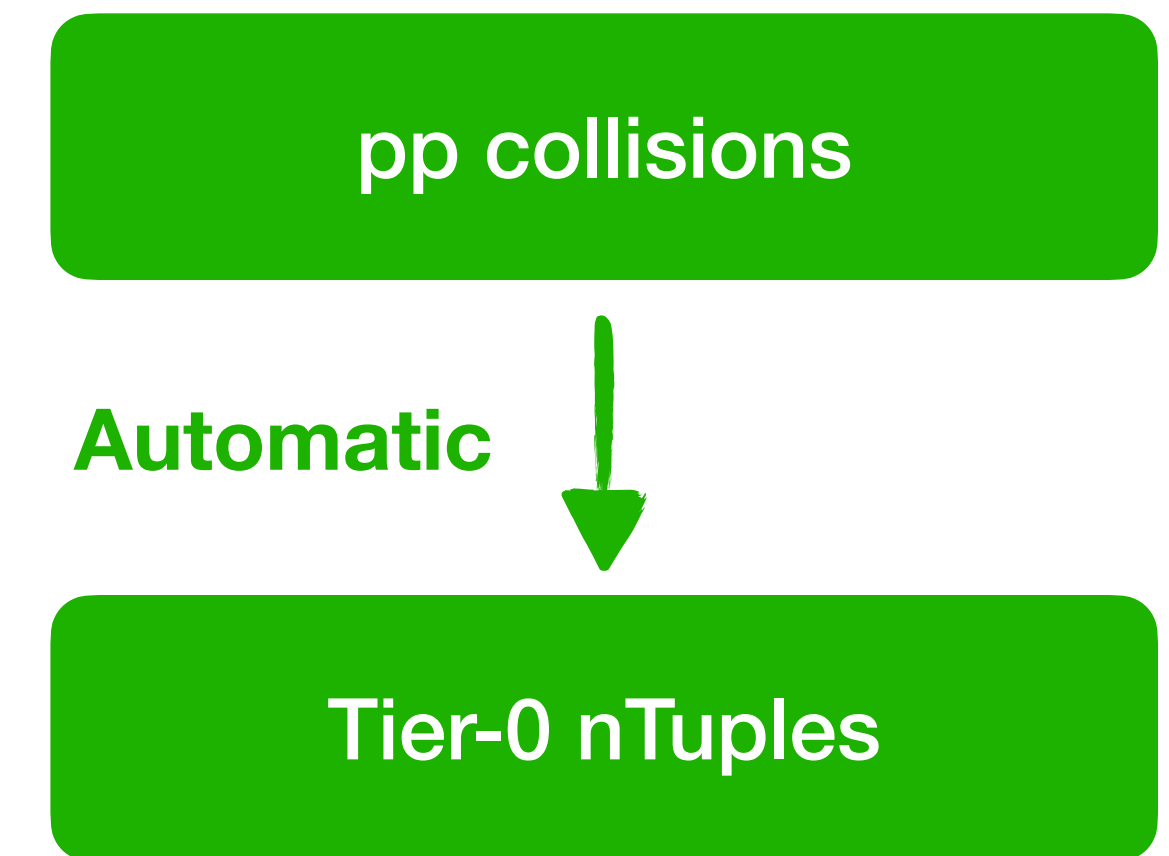


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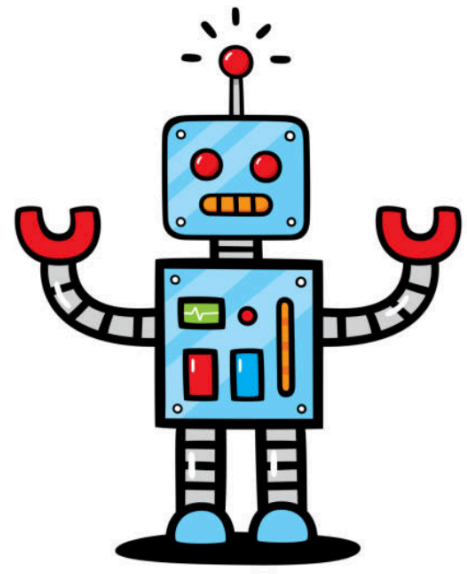
New pipeline framework

- The goal is to **accelerate muon performance checks** via **preservation** and **automation**



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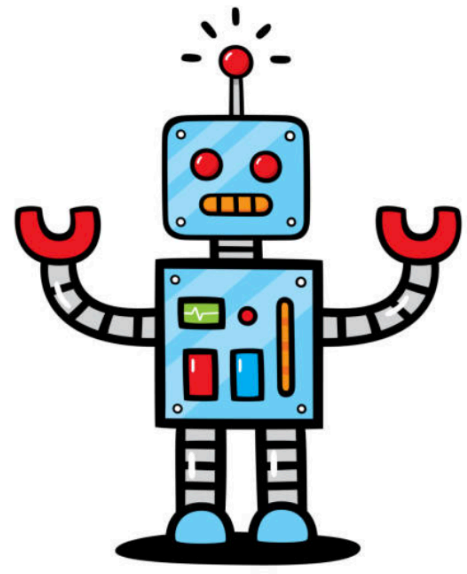
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Example when a new nTuple
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On demand user trigger:
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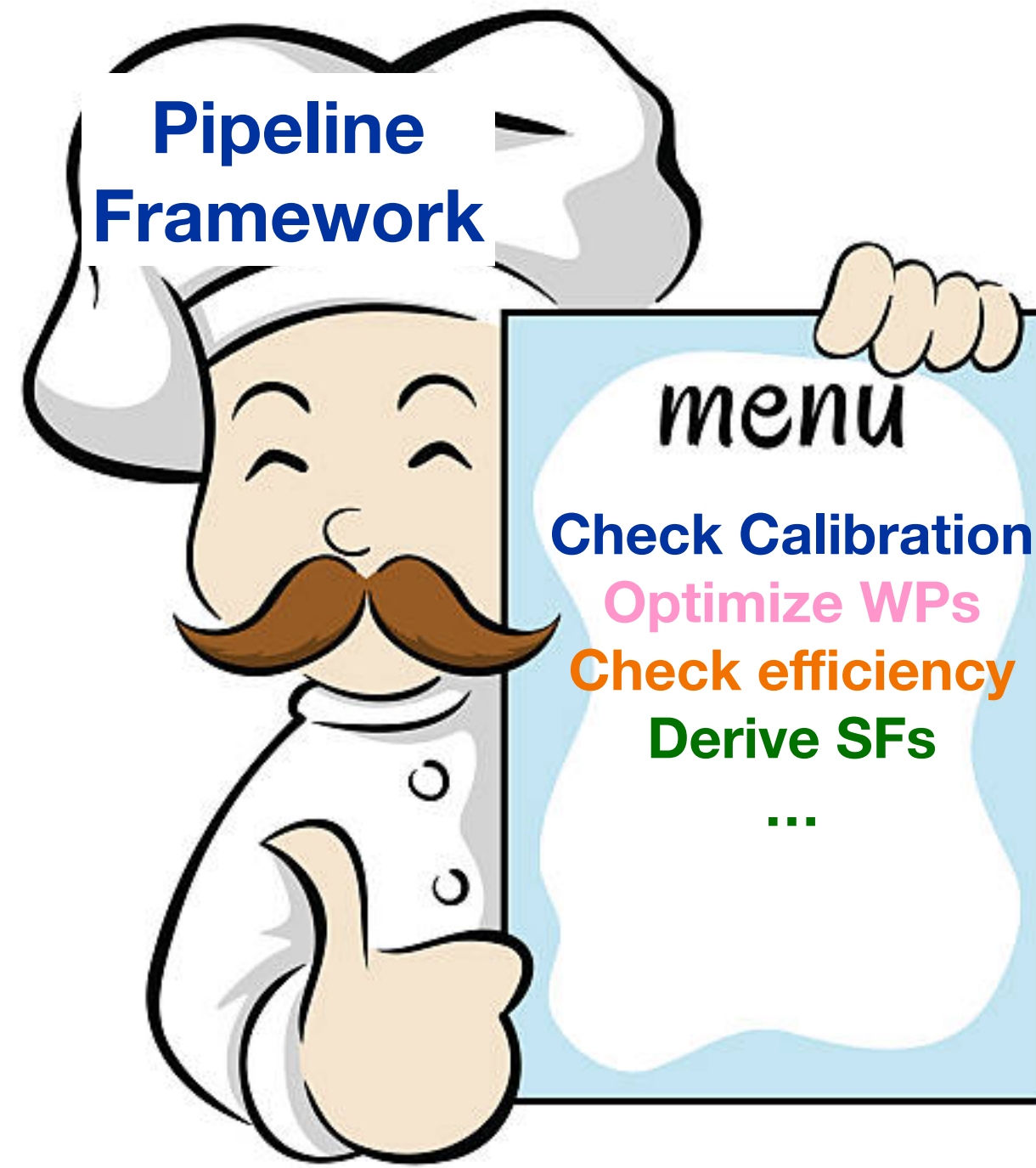
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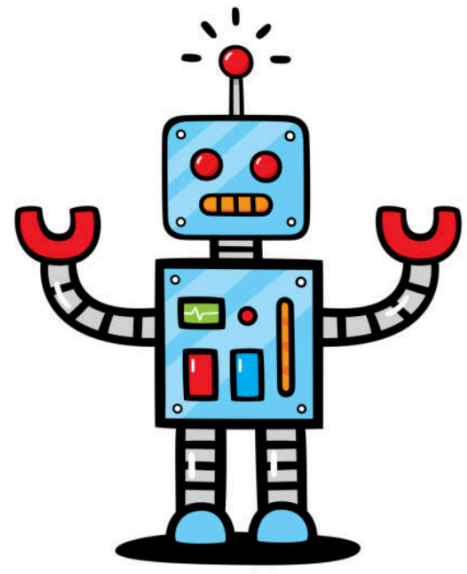


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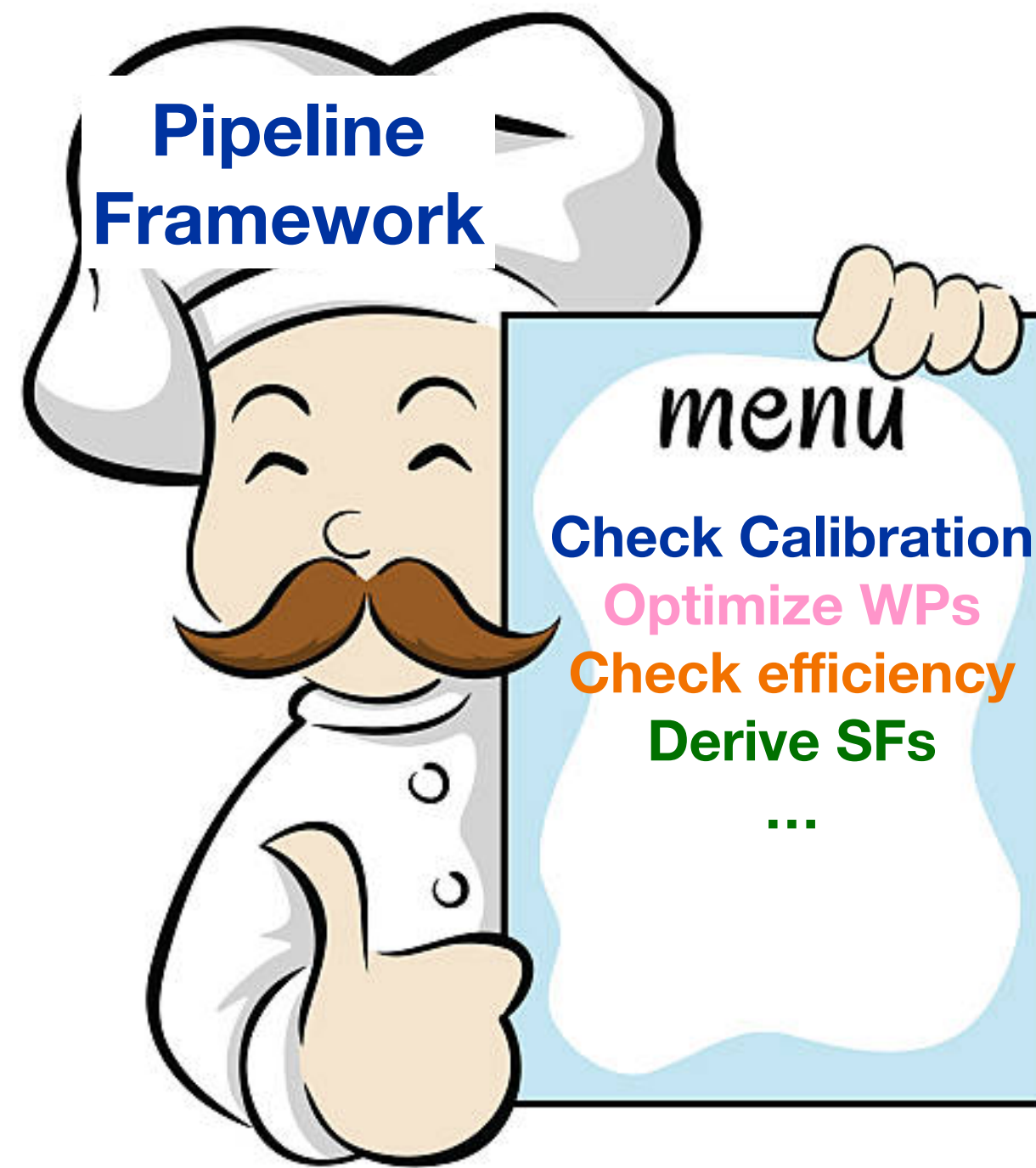
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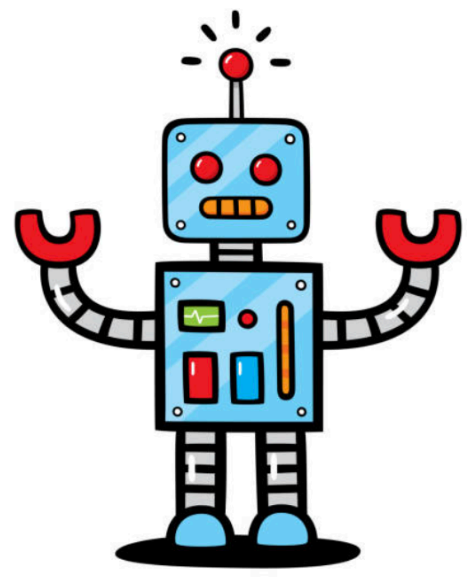
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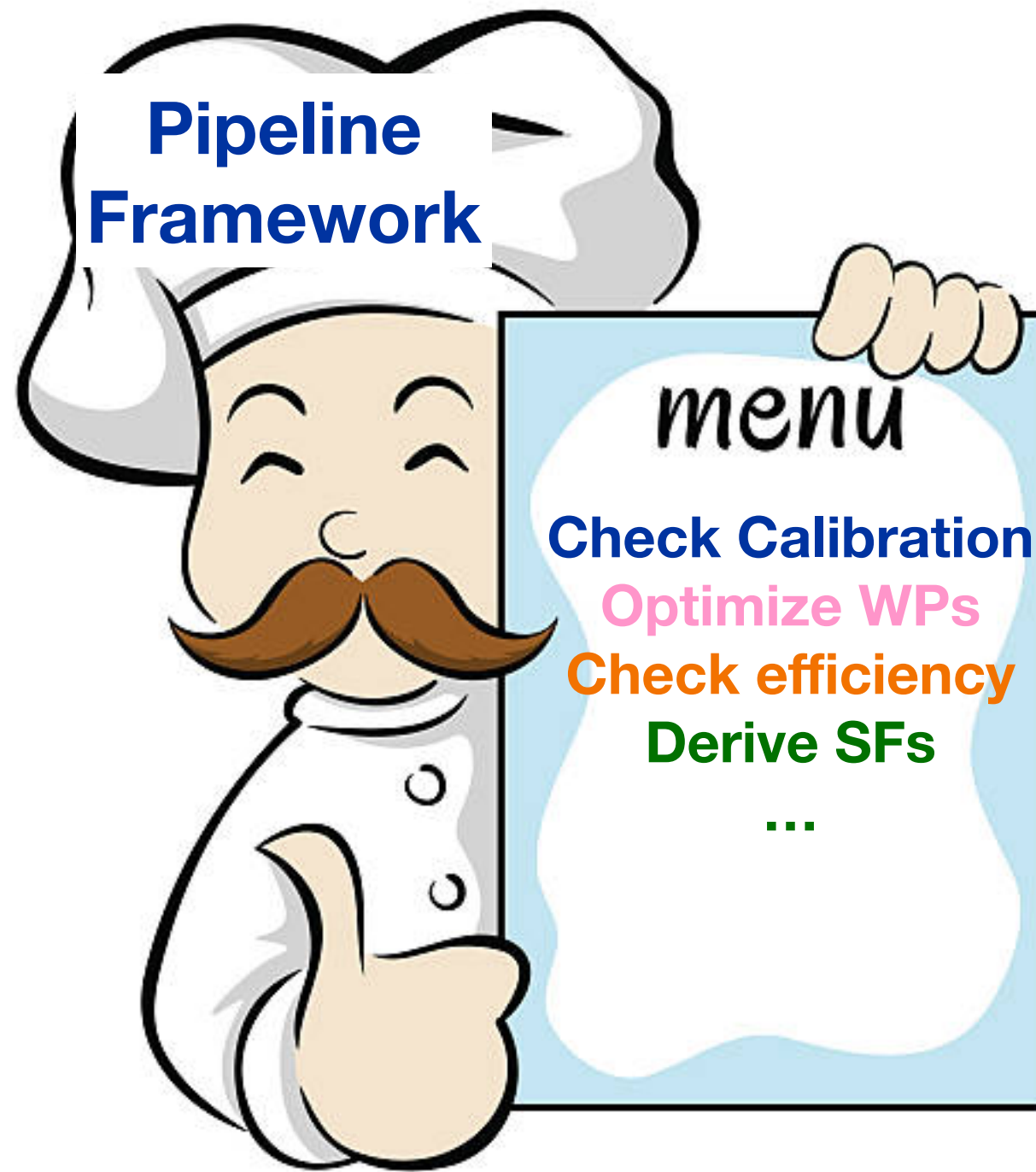
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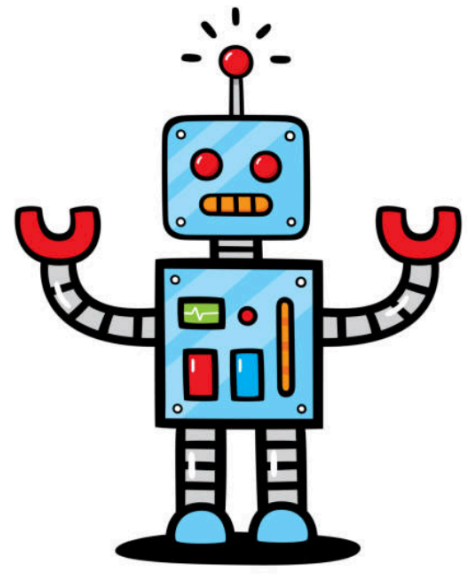
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Packaged into docker images



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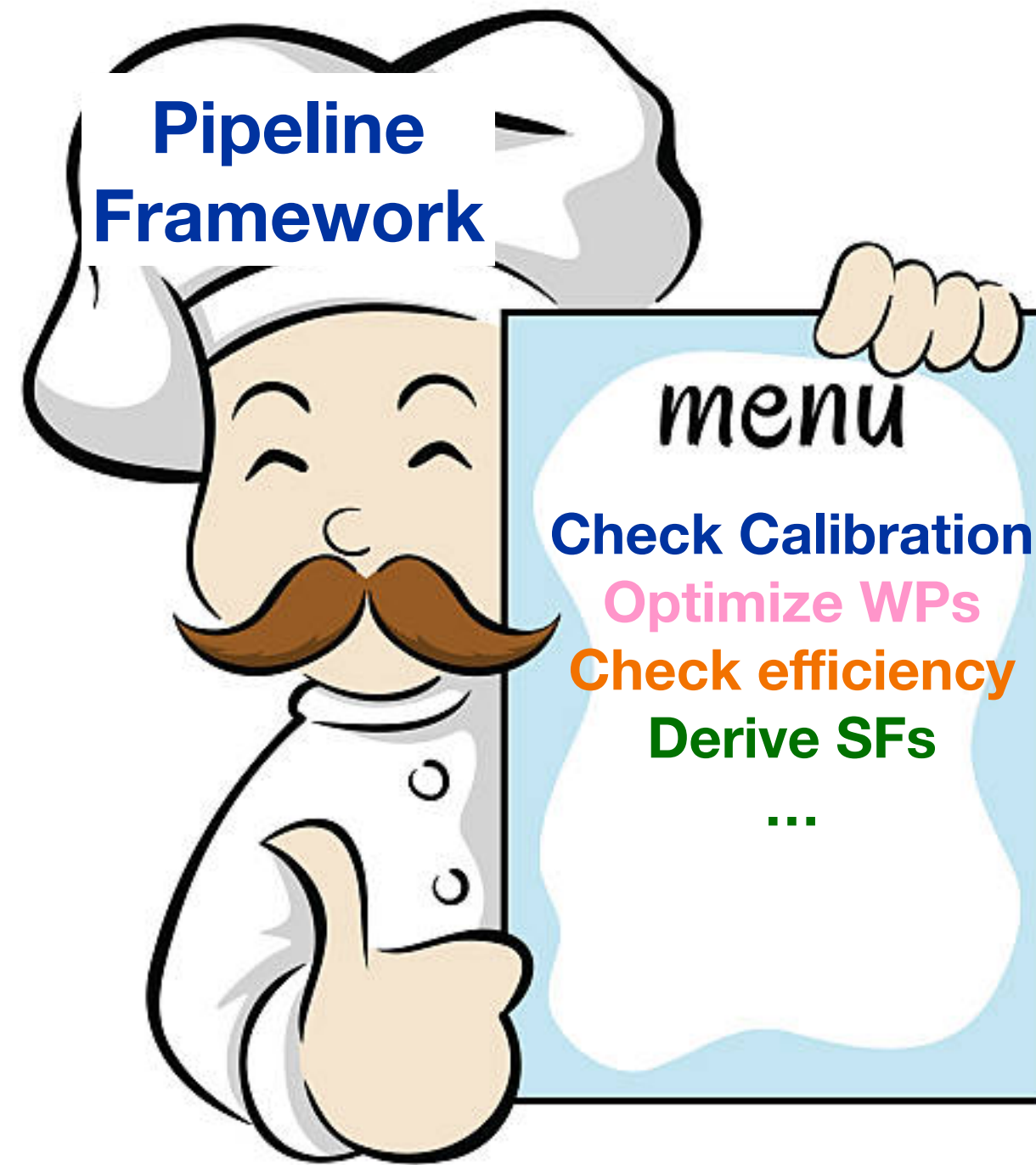
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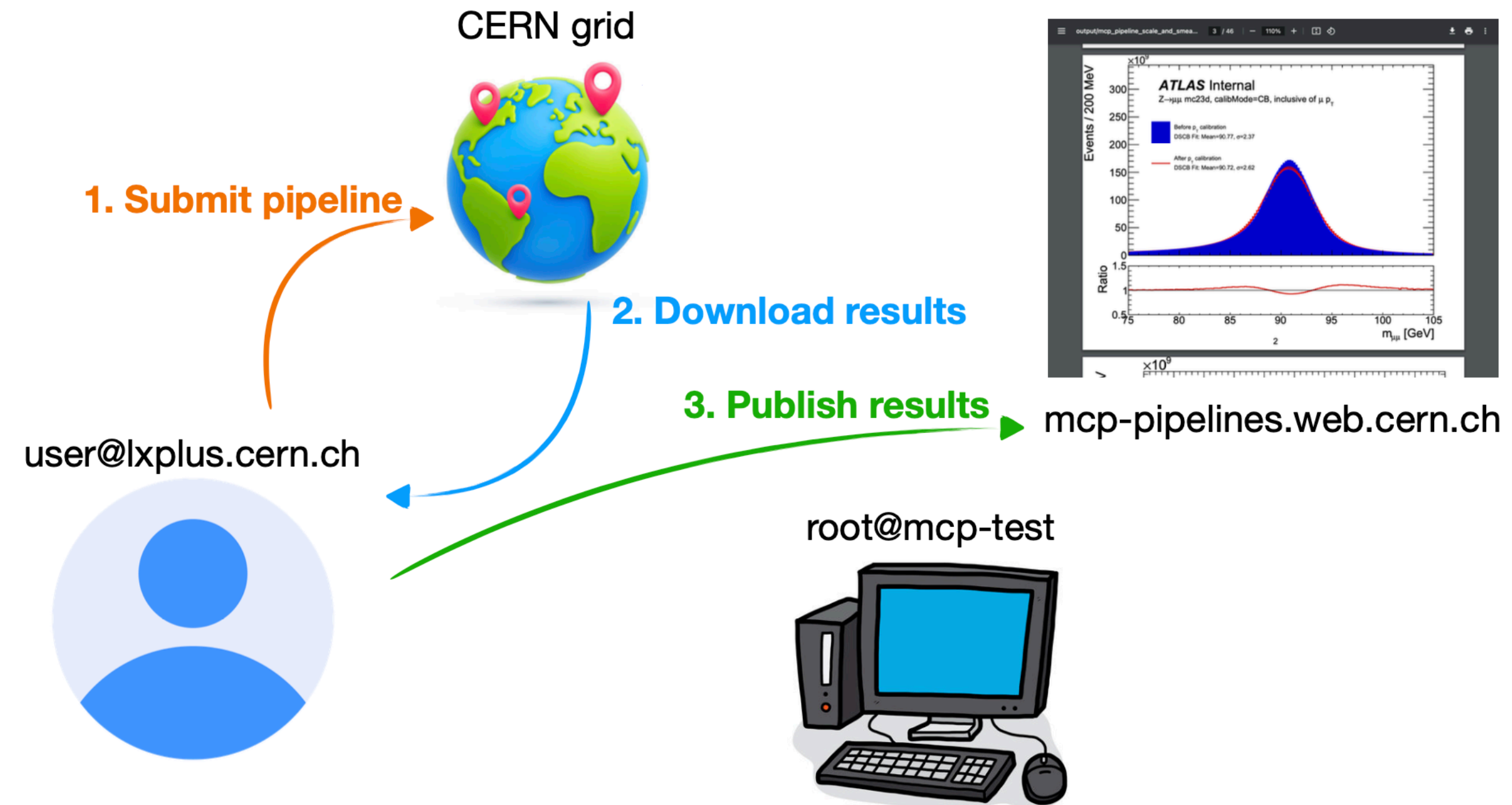
**Muon performance analysis toolkit
Packaged into docker images**



.pdf, .root output files
Published to the web
Can compare previous runs
of the same workflow and
check consistency

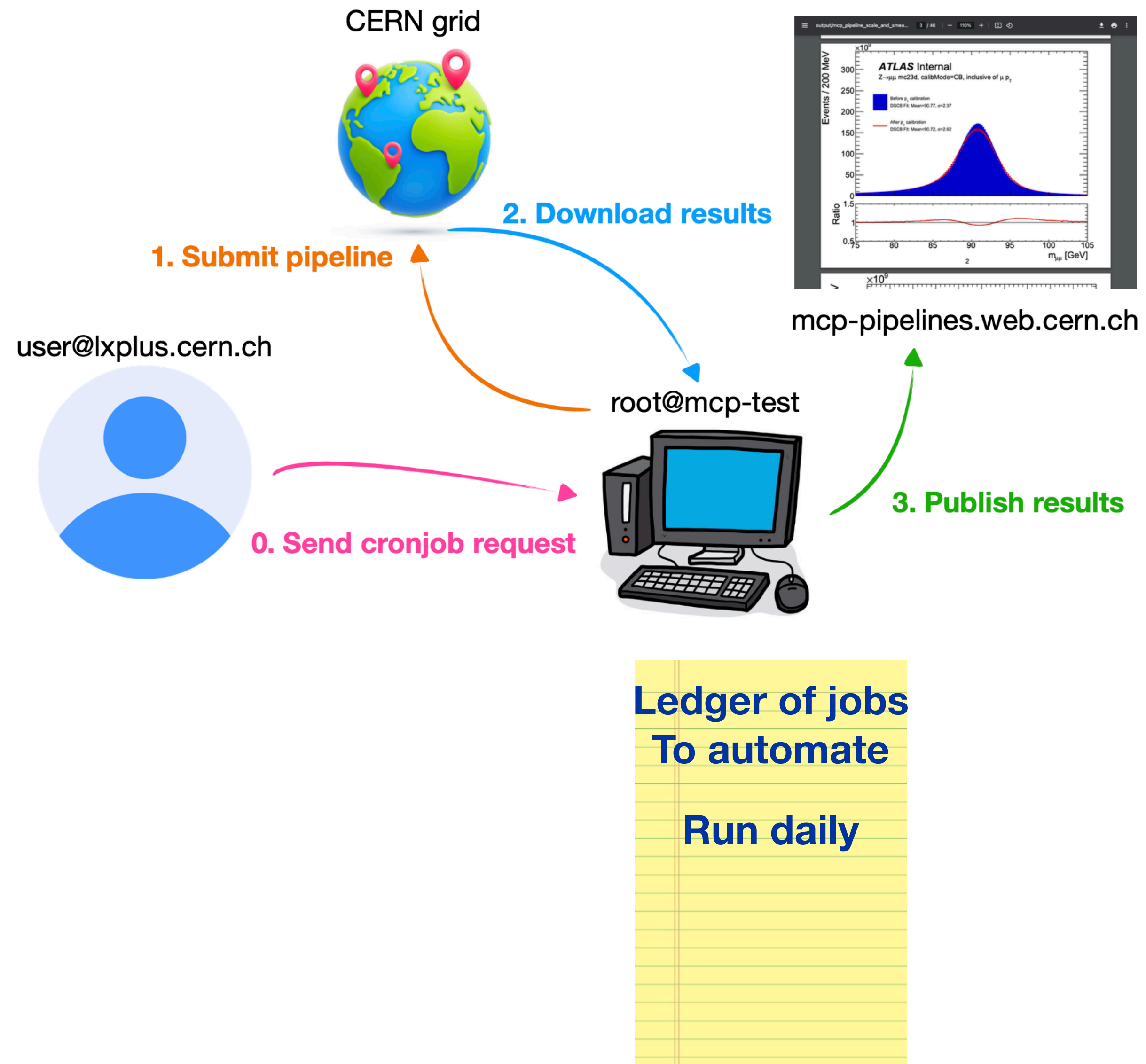
Automation

- The typical workflow in the pipeline framework
- A lot is already automatic, as a full workflow is encapsulated into a single interaction with the grid
 - Ntuple creation, histogram creation, plot creation all happens in a single “chain” of grid jobs.
- But, what if we could remove the user entirely??



Automation

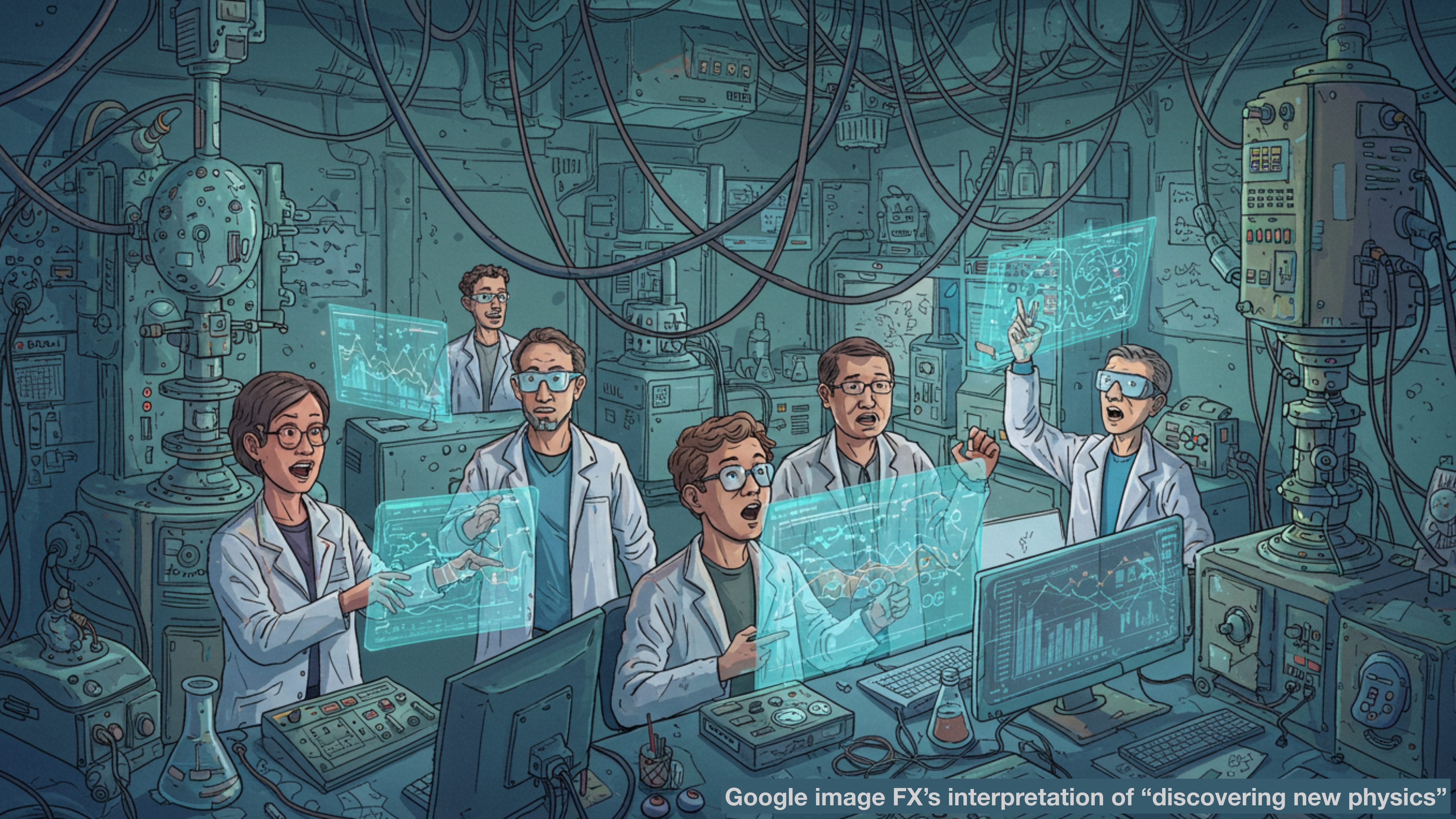
- The workflow of an automated job is similar
- The user can modify the jobs.txt file on the remote server at any time with a command line interface
- Whenever new data is taken, nTuples are automatically created at Tier0 computing center
- Previously, we did very little with these nTuples!
- Now, we analyze them automatically and check for changes in muon performance on a regular basis.
- This automated method will be deployed in 2025!



Conclusion

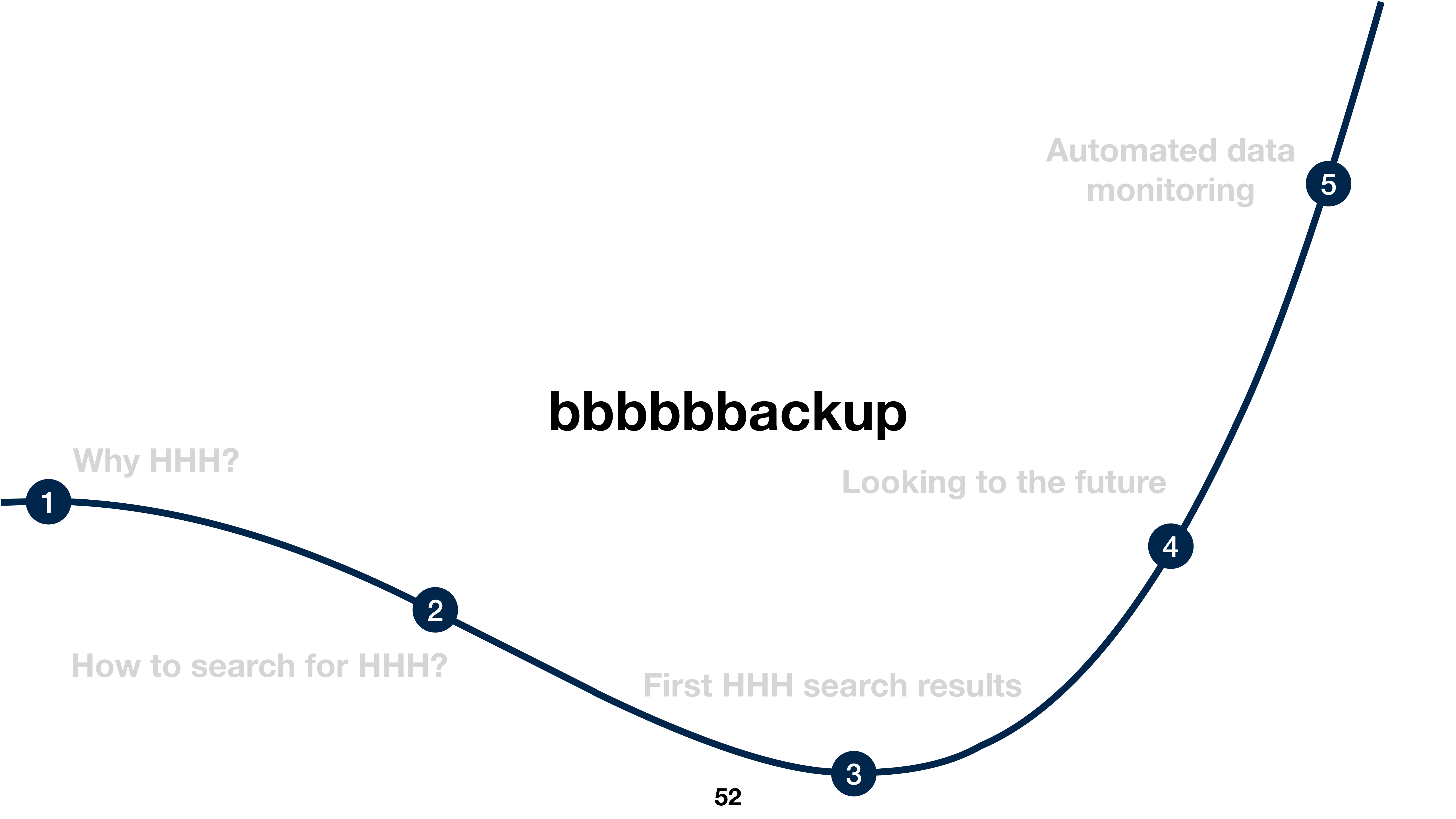
- **The first search for HHH has been performed**
 - Unique sensitivity to κ_4 , which is the limiting factor on our knowledge of the Higgs potential shape for large deviations from TeV
 - Relevant *today* because of possible HHH enhancement in models with extra scalars (TRSM, DM-CPV)
- **HHH is a suitable laboratory for AI studies**
 - Statistics limited, simple data-driven method.
 - Applications of modern AI methods can be studied:
 - Domain adaptation
 - Functional decomposition
 - Transformers and object pairing
- **Muon performance studies are *accelerated* via *preservation* and *automation* in a new framework**





Google image FX's interpretation of "discovering new physics"

bbbbbbbackup



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52

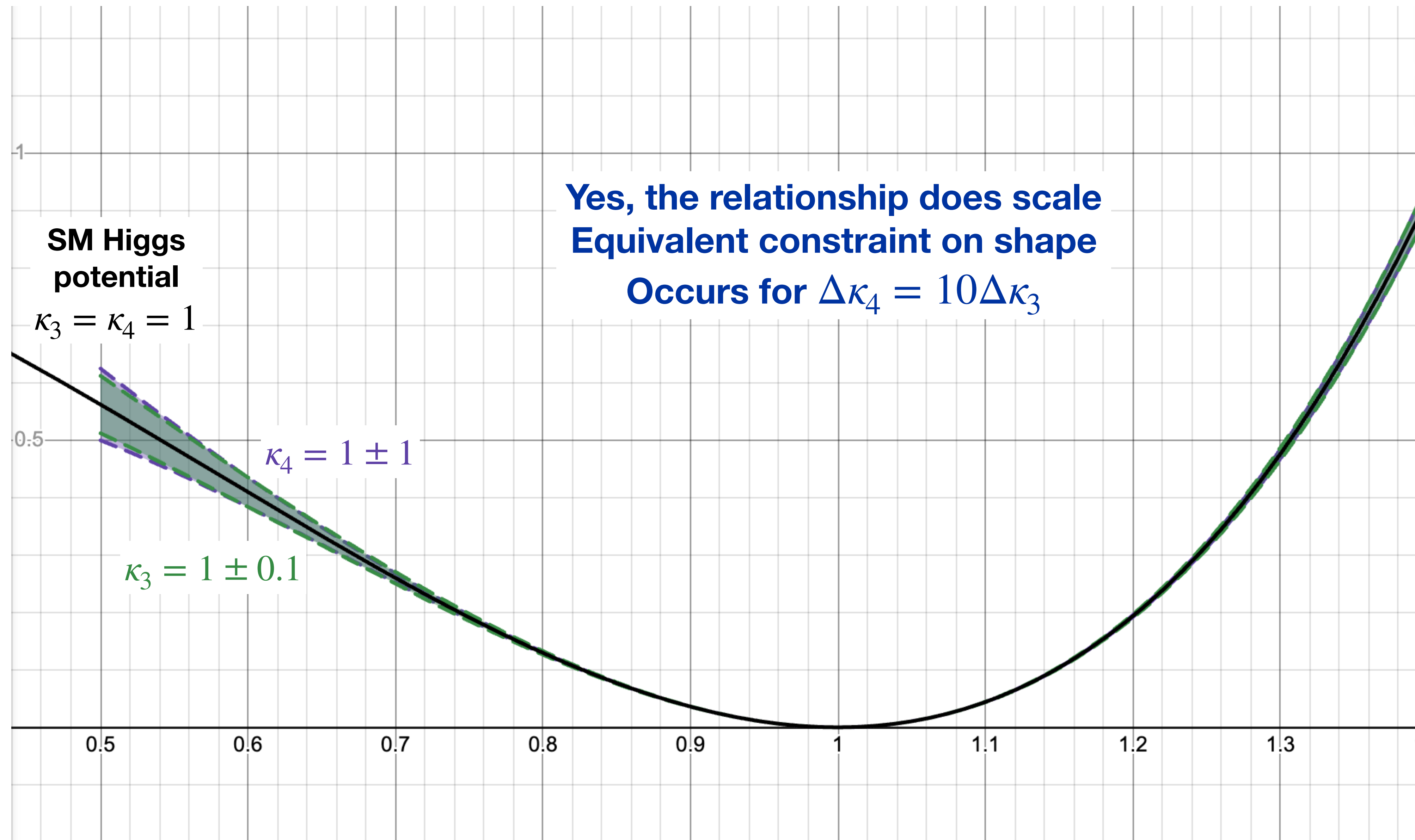
Looking to the future

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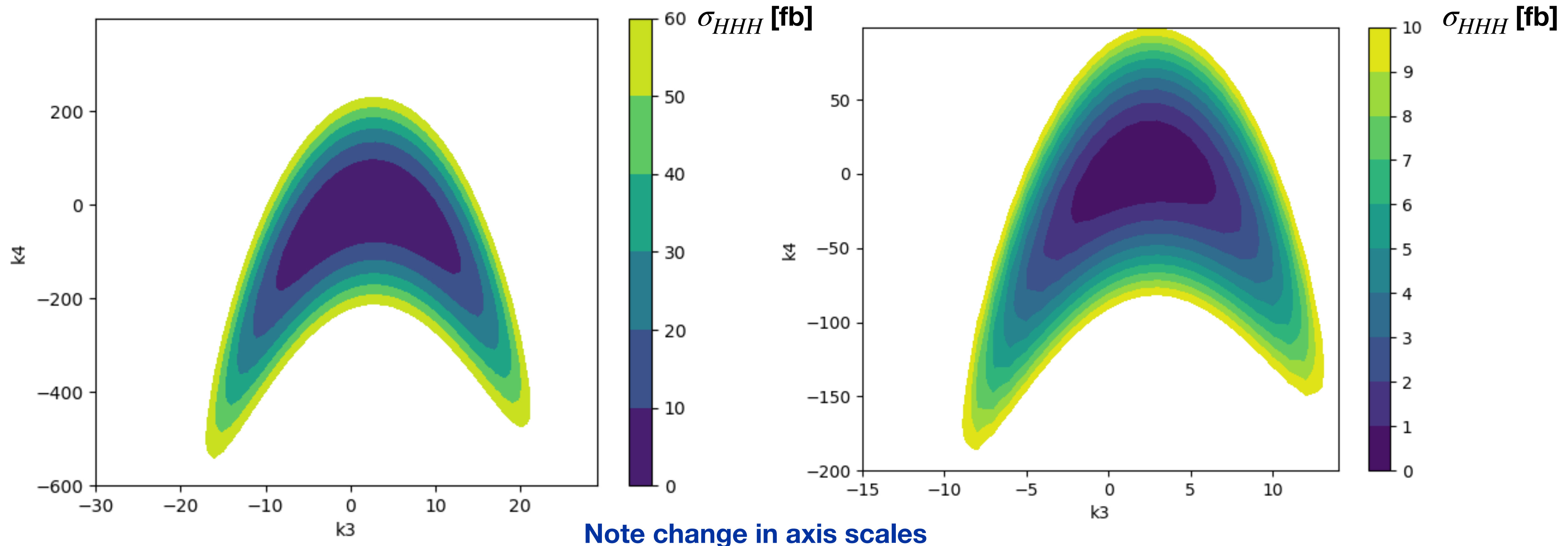
Automated data monitoring

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Scaling of κ_3 vs κ_4

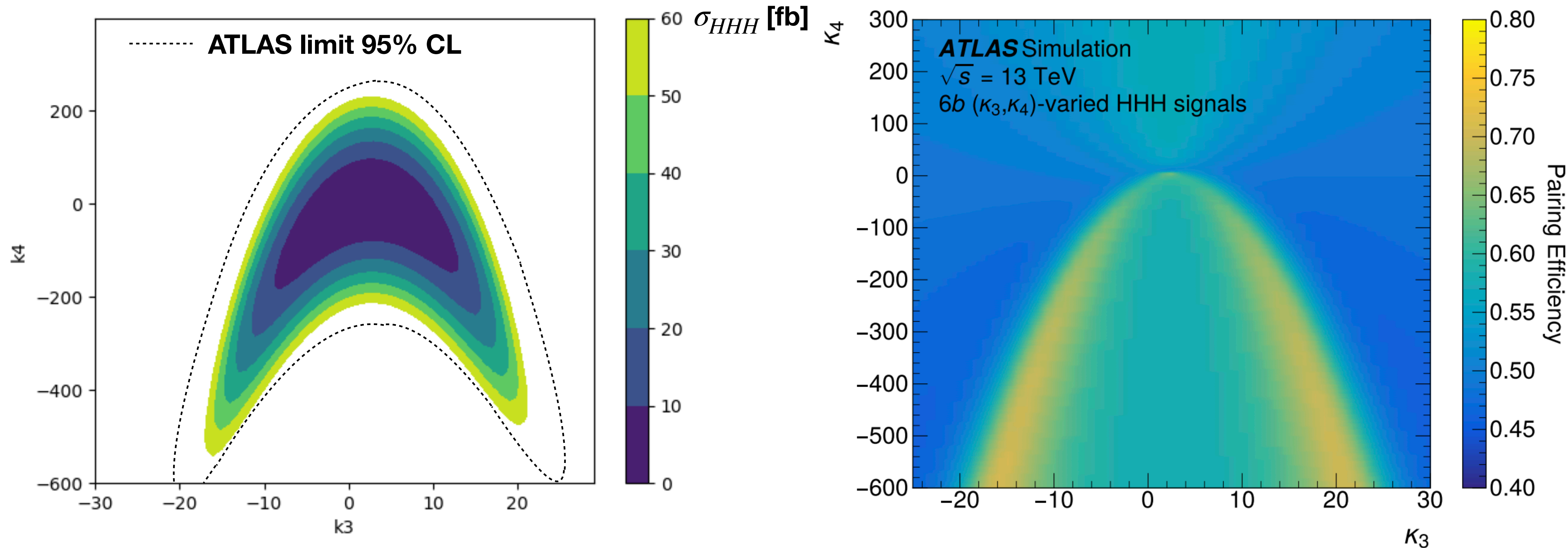


HHH production cross section



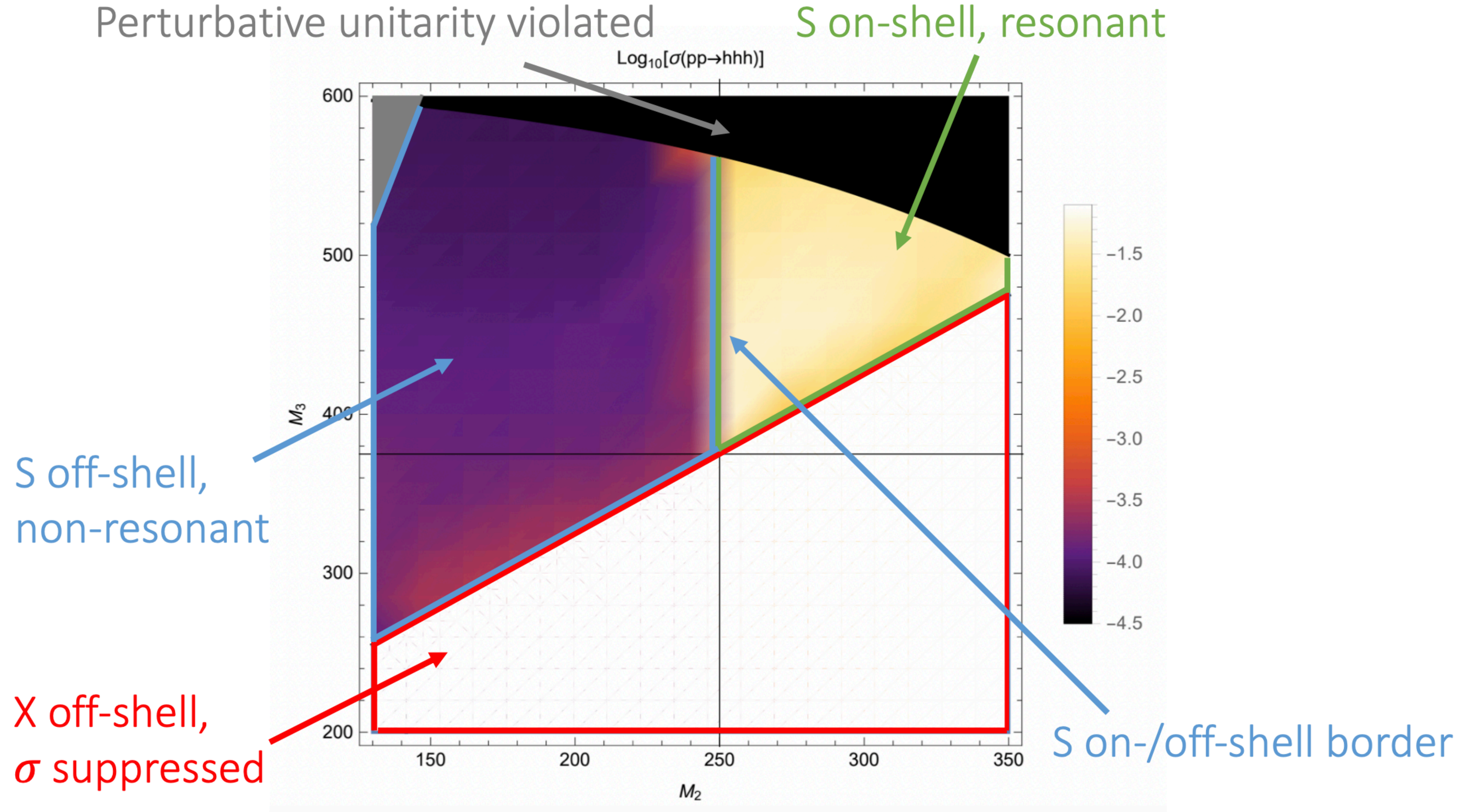
- Constraint of $\kappa_4 = 1 \pm 30$ roughly translates to a constraint $\sigma_{HHH} \leq 1$ fb
- Luminosity-only scaling for HL-LHC should get us to around $\sigma_{HHH} \leq 10$ fb

Shape of ATLAS limits

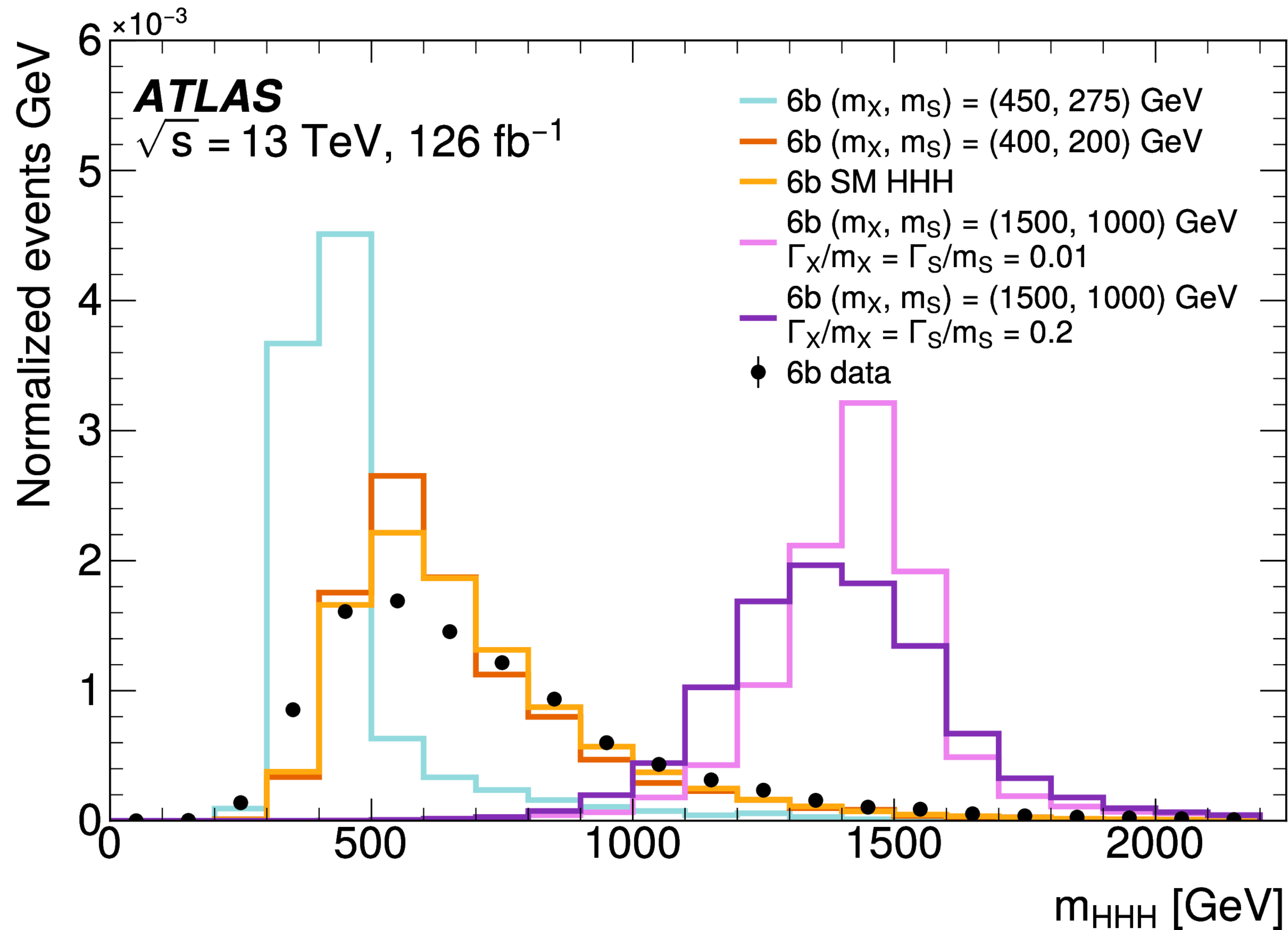


- ATLAS limit inherits its shape from the contours of constant σ_{HHH}
- Slight modulations in shape come from how boosted the Higgses are, and therefore how well we pair the jets.
- In regions of higher pairing efficiency, we set a more strict limit

TRSM perturbative unitarity



Signal kinematics



Trigger strategy

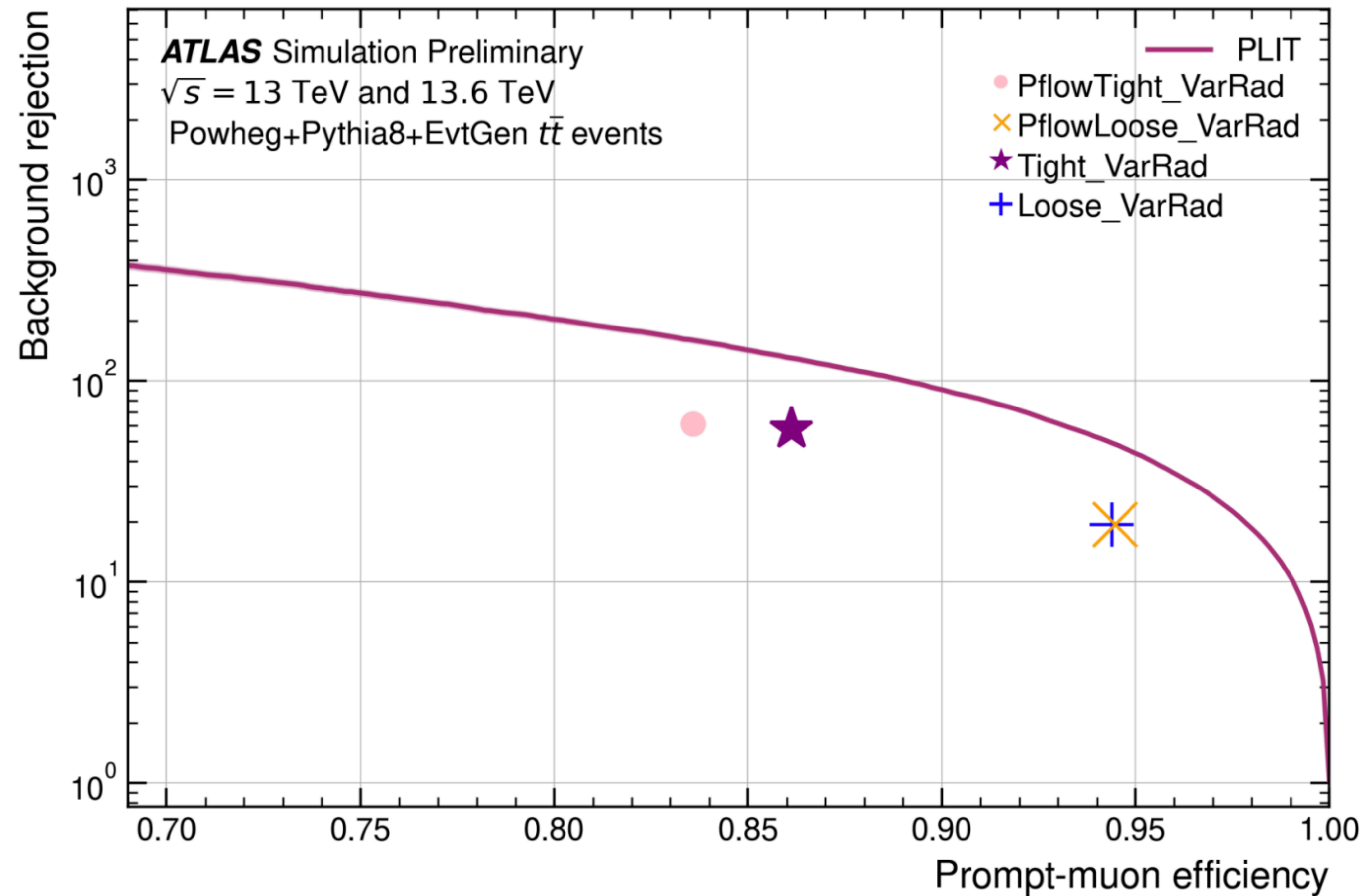
- Using multi- b -jet triggers for a 6 b final state – similar to the resolved HH4 b analysis
 - Omitting 2015 data due to lack of jet trigger-matching information
- 2 b 2 j triggers in 2016, 2018
- 3 b 1 j triggers in 2017
 - Due to very tight online b -tagging working point in 2017 2 b 2 j triggers

Table 5.1: List of b -jet triggers selected for efficiency study.

Year	Type	Trigger Name
2016	2 b +2 j	HLT_2j35_bmv2c2060_split_2j35_L14J15.0ETA25
2017	3 b +1 j	HLT_3j15_gsc35_bmv2c1070_split_j15_gsc35_boffperf_split_L14J15.0ETA25
2018	2 b +2 j	HLT_2j35_bmv2c1060_split_2j35_L14J15.0ETA25

Other applications of transformers

- Factor of 2 improvement in lepton isolation from transformer architecture



Differentiable p-values with FD

- Transformation from physical to frequency domain is fast:
 - Recursion relationships for the orthonormal functions
 - Use unbinned data: inner product with Dirac delta functions is simple to compute
- In the frequency domain background and signal are separable by a cut

