

Improving light coupling into instruments with integrated photonic circuits and low temperature detectors:

Or does it make sense to include optical integrating in the RDC work package?

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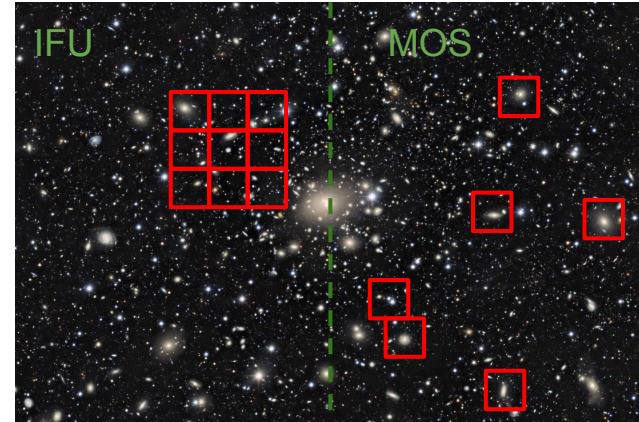
Motivation

Objective: Make spectrometer to advance capability of determining atmospheric composition/habitability on exoplanets

Method: Make a medium resolution spectrograph for use in multi-object or integral field spectroscopy at next generation extremely large telescopes (>30 m) with wavefront control/single mode fiber injection.

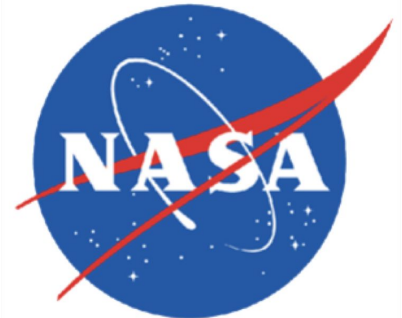
Innovation: Use photonic circuits integrated with low temperature detector (i.e. MKIDs, TES, SNSPDs) instead of bulk optics and conventional semiconductor photodetectors as the spectrometer

Advantage: Significant reduction of size weight cost of instrument, and potential decrease in measurement time.



Got small amounts of funding.

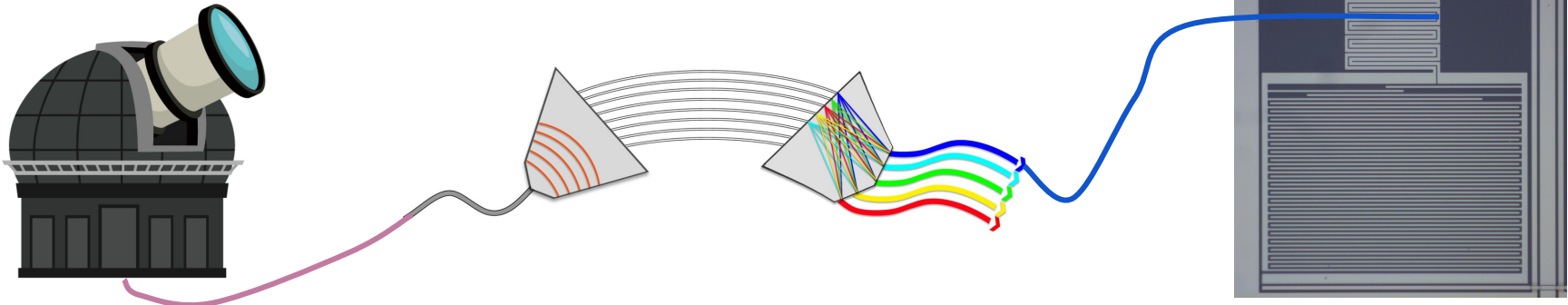
- Found Industry partner (Small Business Technology Transfer) STTR funding from NASA
- Later got a grad student on NSF fellowship
- NSF EARly-concept Grants for Exploratory Research (EAGER) funding.



Specification Goals

Photonic circuit disperse light,
then photon detection by
MKIDs!

- $R = \lambda / \Delta\lambda \sim 5000$
- Bandwidth: 400 nm - 800 nm
- Fiber to MKID throughput: 60%
- Number of detector channels: 1024
- Operation temperature: ~ 100 mK
- Integrated MKID + Waveguide Chip ~ 10 cm²



Current State of the Art

Compare to Dark Energy Spectroscopic Instrument (DESI)

10 spectrographs each illuminated by 500 multimode fibers

CCD detectors at ~140K

Bandwidth 360 nm - 980 nm

$R = \lambda/\Delta\lambda$ 2000 - 5000

Throughput 60% - 78%

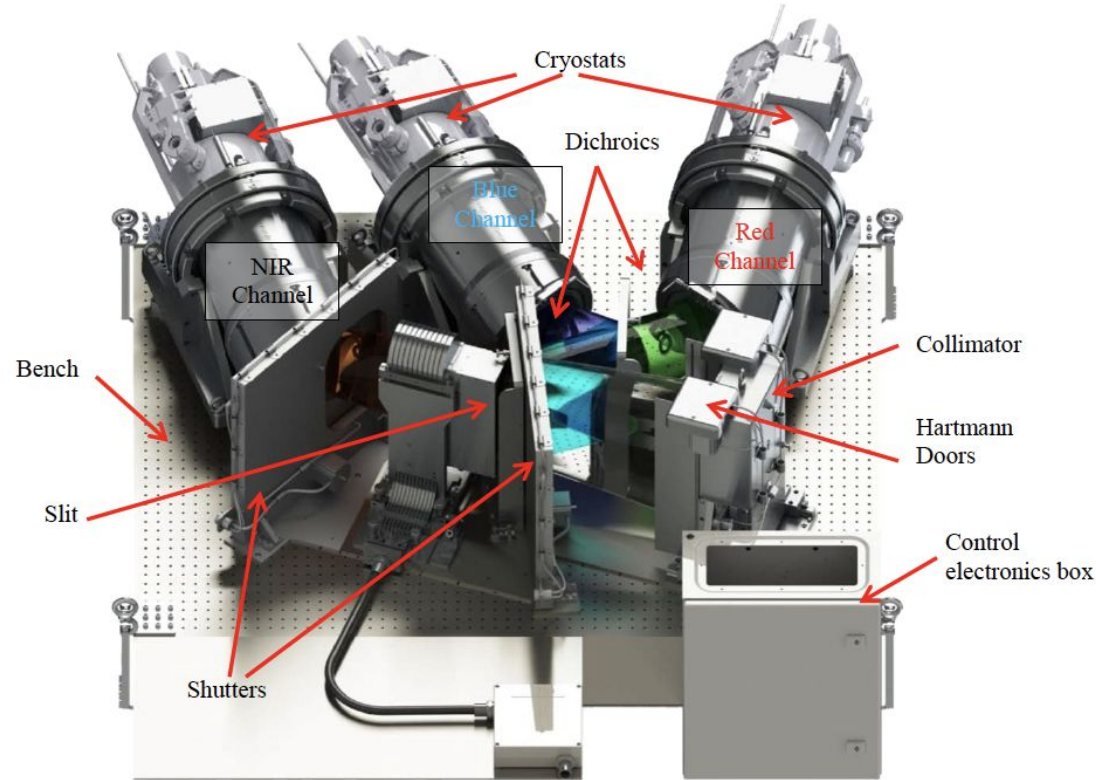
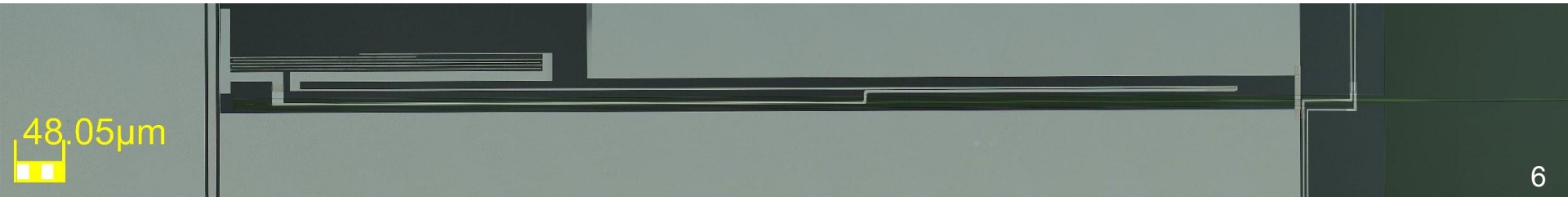
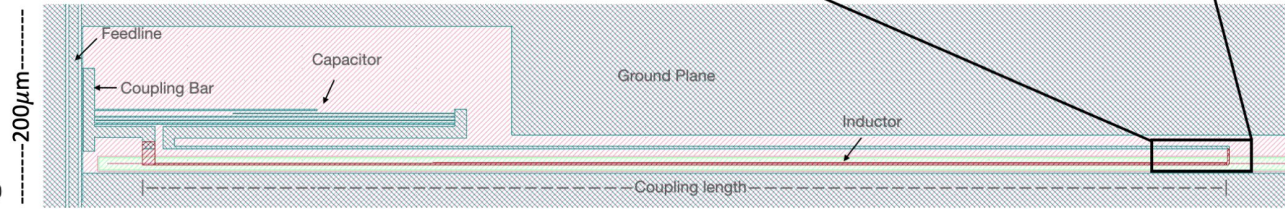
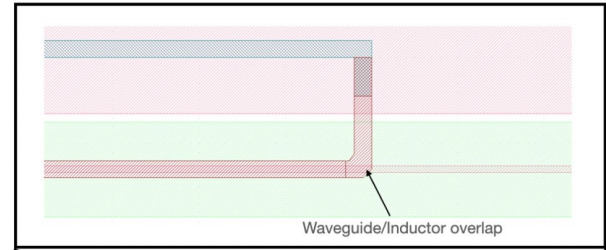
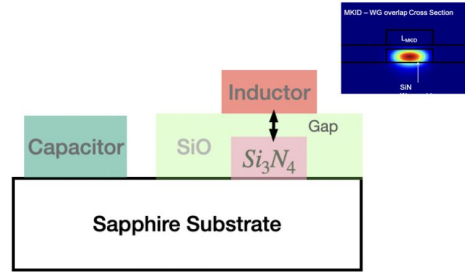
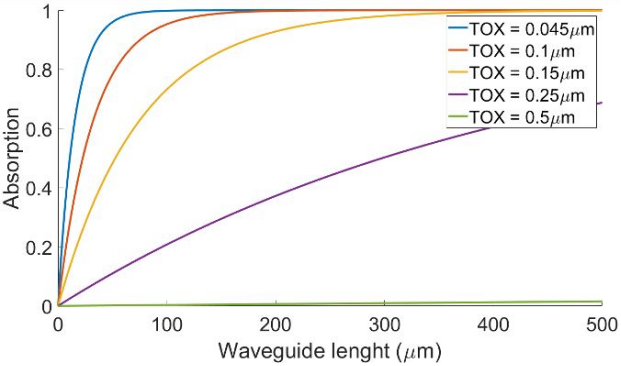


Figure 1. Top: Optical layout of the spectrographs. Bottom: mechanical implementation. The spectrograph is 1.8 m wide \times 1.4 m deep \times 0.6 m high. <https://hal.science/hal-02003974>

Photon Detection Through Evanescent Coupling

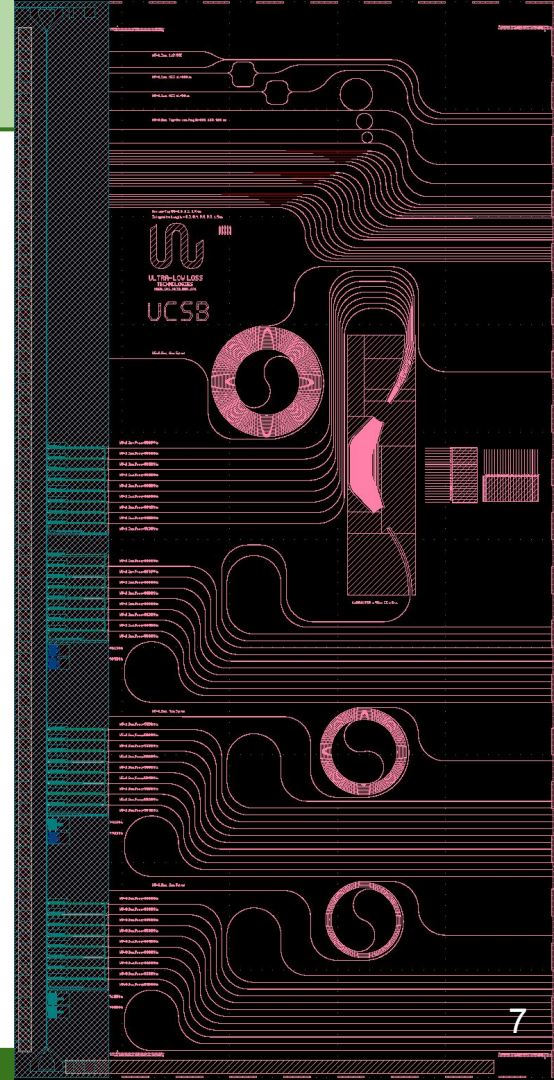
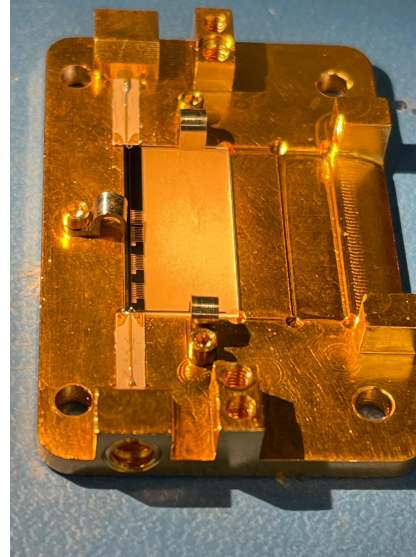
- Photonic waveguides run over MKID inductor

Coupling Efficiency



Test Chips

- 32 α - and/or β - tantalum MKIDs
- Arrayed Waveguide Gratings (AWGs) connected to MKIDs
- Spirals, Ring resonators, Mach-Zehnder Interferometers (MZIs)
- Sapphire substrate / 'Platform'
- Single mode fiber couple-able



Fiber to Chip Coupling

Pros

Cons

Fiber Array

Standard Available

< 65% coupling
> \$2.5K to make + > \$1.5K to attach
Rigid connection of dissimilar materials

Grating / Free Space Coupler

Can possibly give >95% coupling

Narrow band, $\Delta\lambda/\lambda < 5\%$
Double sided

Photonic Wire Bond / 3D Lithography

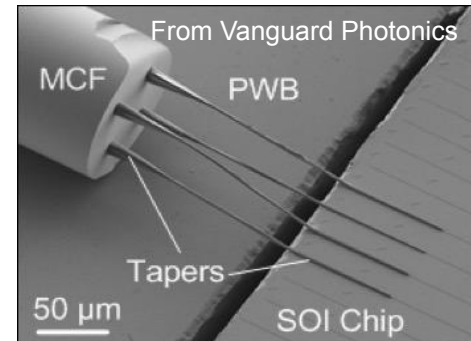
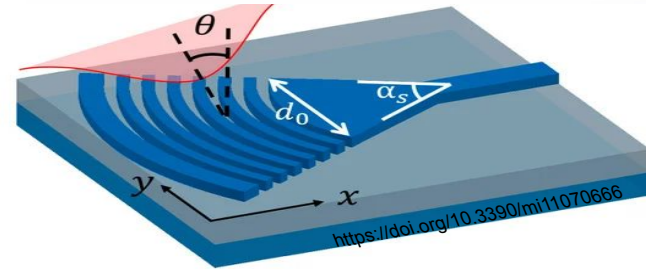
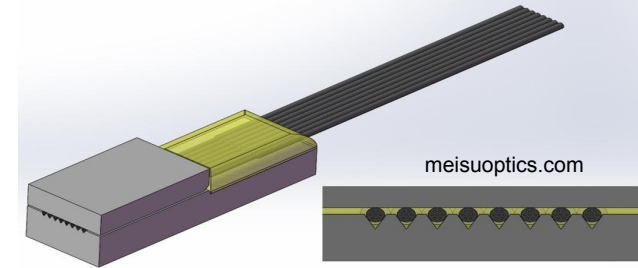
>80% coupling and highly optimizable
Probably good at lot temp, Scalable

Expensive new unavailable
~\$50k NRE + ~\$1k/job

Edge Fire / Butt Coupling

Fast, Free and Nondestructive

Expensive, Large positioning stage in the cryostat



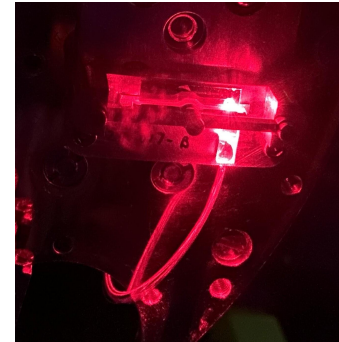
Key Challenges

- Monolithic WG-MKID integration
 - WGs need glassy dielectric processed at high temperature
 - Degrades MKIDs performance
 - Fab solutions accessible

- Light Coupling
 - Fiber array thermal mismatch → progressive decline
 - Insertion loss
 - Stray light
 - Fridge heating (e.g. 0.5mW input @ 686 nm → 10mK to 125mK)

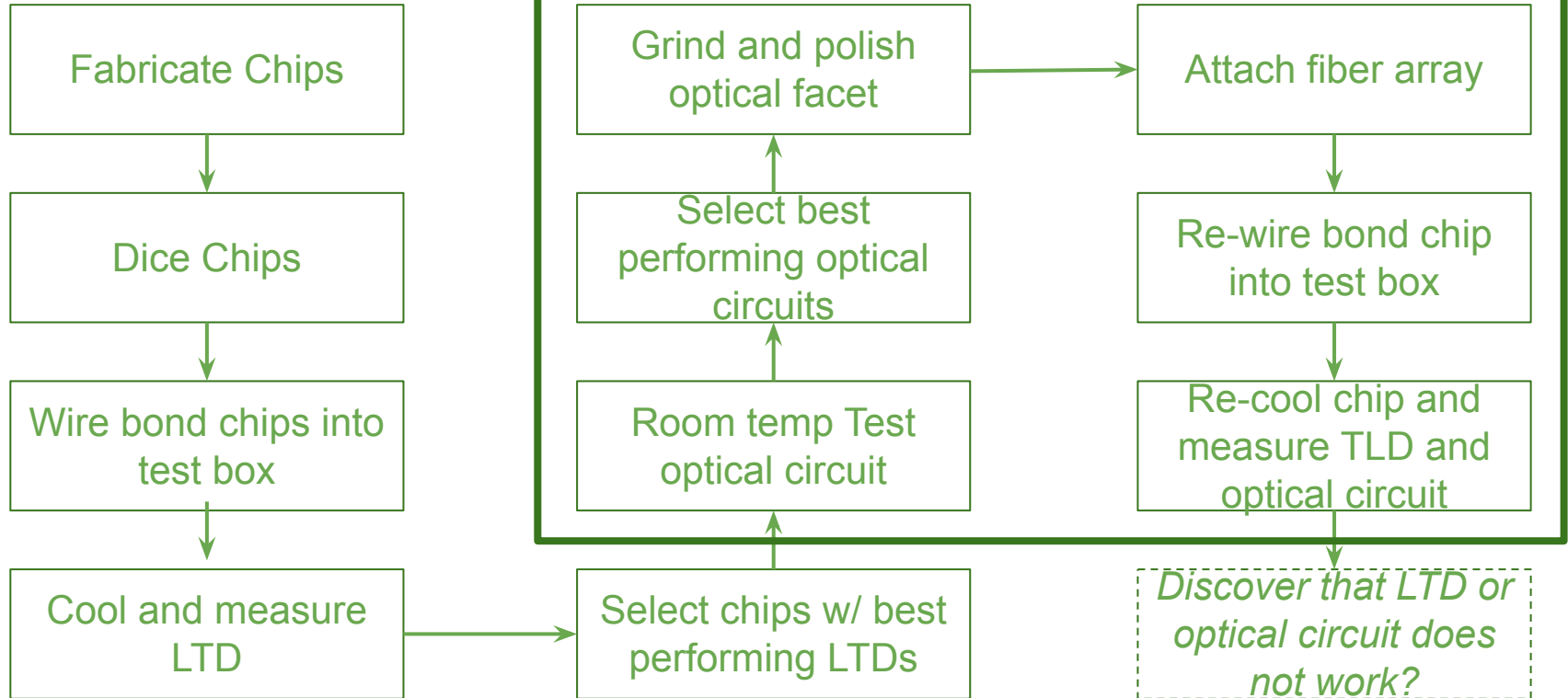
Our devices on sapphire platform @ 686 nm

MKID Qi	MKID Yield	WG Loss	Coupling Loss
3,000 - 200,000	25% - 80%	1.2 to 24 dB/cm	> 4dB/facet (60% photon loss)



Lengthened, Costly Test & Iterate Cycle

Due to Expensive + Time Consuming Optical Attachment



Is efficient light coupling a common problem?

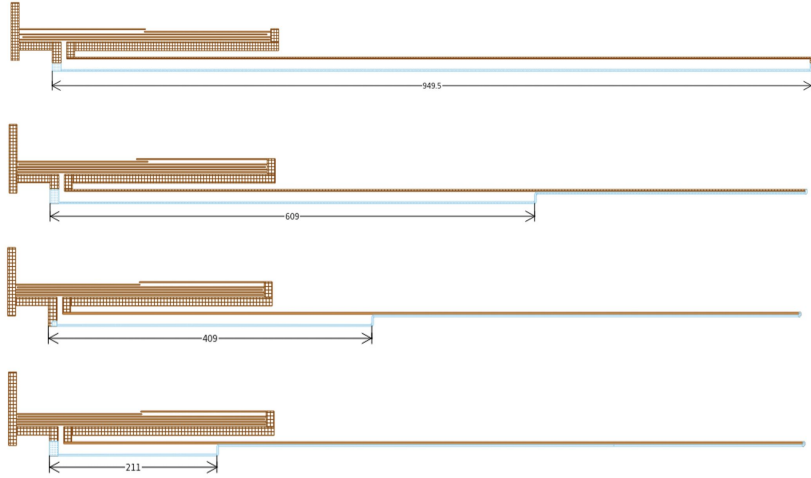
- Problem shared by folks moving light around on chip \Rightarrow single mode operation.
 - Relatively new capability, but increasingly needed \Rightarrow problem will be increasingly encountered by the community
- Astrophotonics
 - Single mode fiber systems (unlike DESI) allowing for on chip light processing
- Quantum information
 - Photonic quantum computing, communication, metrology, cryptography
 - Systems favor compact size, scalability and low tolerance for loss of light
- Other Applications: Strain measurement..., ?

Suggestion: RDC Work Package \supset system integration challenges \ni light coupling

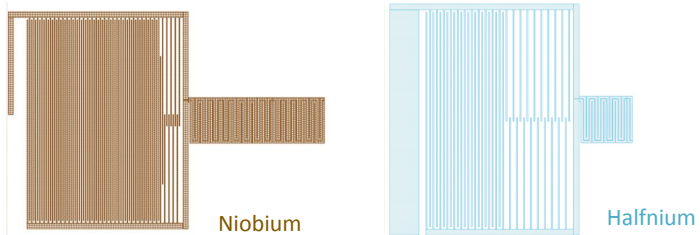
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Resonator Design

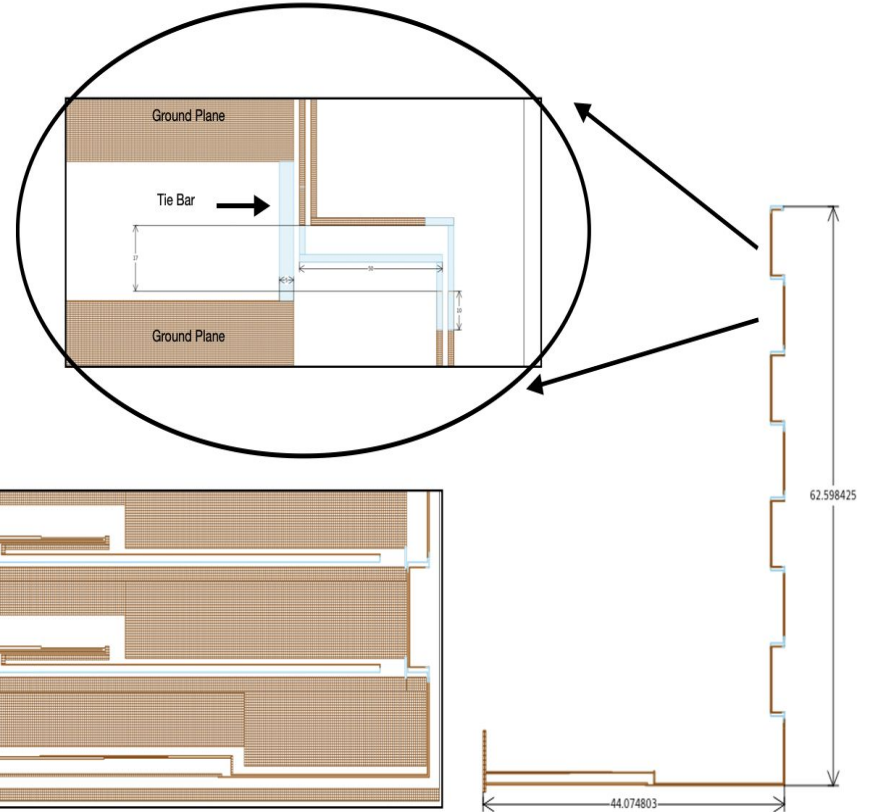
Hybrid Resonator



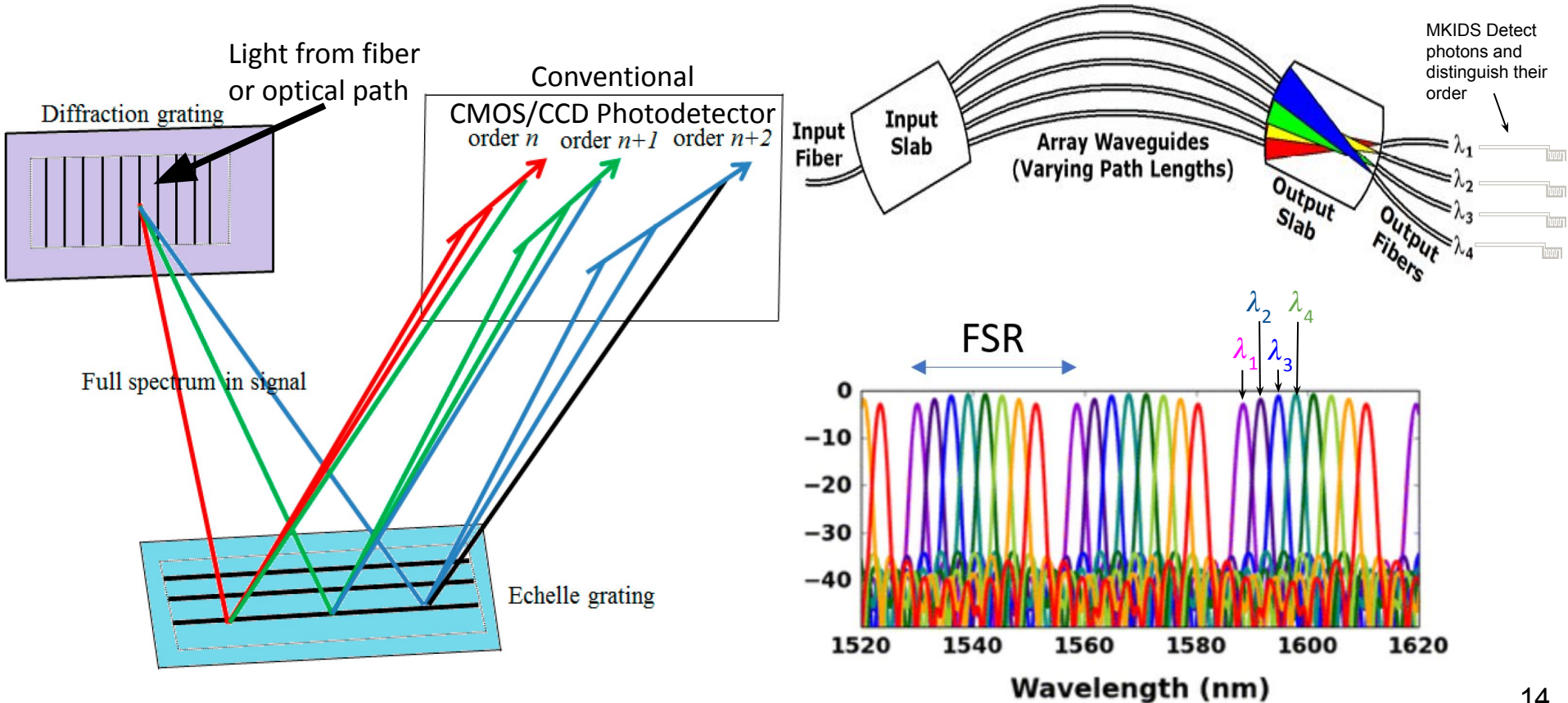
Reference Resonator

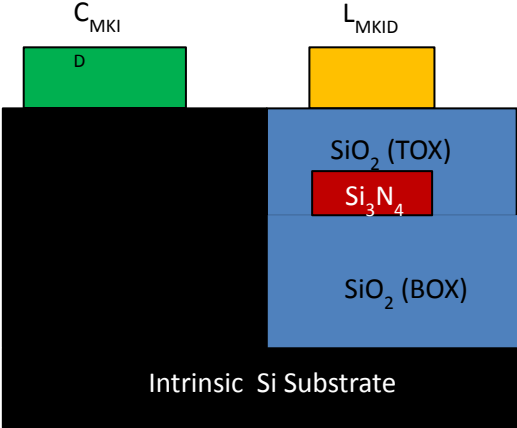
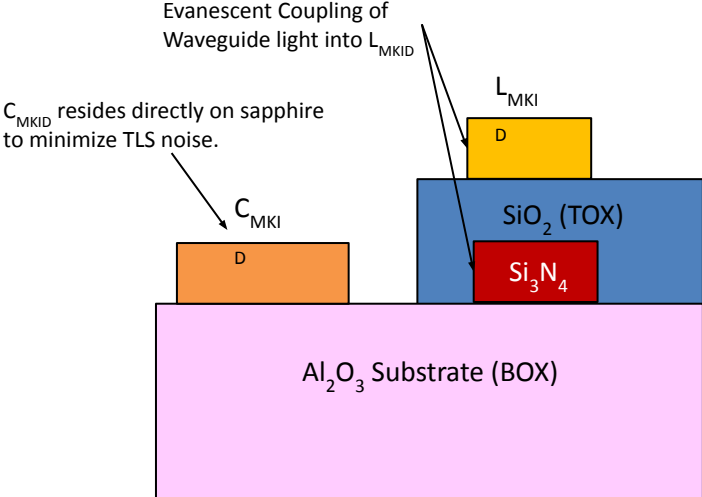


Sense Resonator



Order Sorting





Integration of quantum and superconducting sensors with photonic circuits enables applications and solutions that are out of reach when the two technologies are instrumented separately. For example, it is not difficult to imagine that the size, weight and cost of an integrated photonic and sensing instrument would be less than a non-integrated instrument. Among the applications benefiting from such integration are high resolution spectrometers, photon number counters, quantum computers. The challenges of integration include ensuring the efficient coupling of light into the instrument. The sensors require low temperature operation so the integrated photonic circuit must be placed at the cold stage of a cryostat, where loss of light could create a background for the sensor, warm the cryostat and/or result in loss of quantum information. The technique used for light coupling must be stable against vibrations, thermal expansion, and compatible with sensor and waveguide fabrication so as not to degrade the performance of either. Multiple techniques are needed to accommodate the diversity of applications of these integrated instruments, for example applications permitting cryostat windows and applications with fibers fed into the cryostat. This talk describes work we would like to do in the context of a CPAD RnD Collaboration to find fabrication techniques and manufacturers of use to achieving efficient light coupling into instruments with integrated photonic and sensing circuits.

– END BACKUP SLIDES –



ULTRA-LOW LOSS
TECHNOLOGIES

Project: SiN-MKID

Prepared by: Gopi

Institution: ULLT

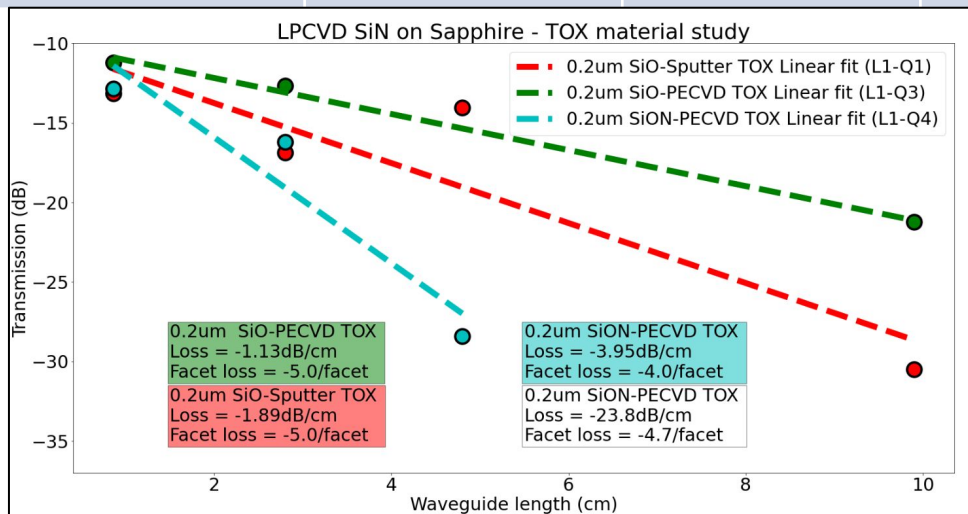
Date: 2023-11-08

*SiN on Sapphire measurements
for Miguel's presentation*

Loss measurement plots for LPCVD
Nitride

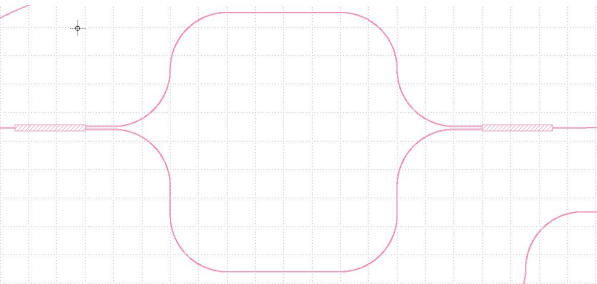
Loss measurements for 200nm LPCVD SiN on Sapphire: Wavelength = 686nm

Wafer	BOX	Core	TOX	Loss	Coupling loss	Comments
L1-Q1	Sapphire	SiN LPCVD 200nm Thick, 1.8um wide	SiO Sputter 200nm Thick	-1.89dB/cm	-4.99dB/Facet	
L1-Q2	Sapphire	SiN LPCVD 200nm Thick, 1.8um wide	SiON Sputter 200nm Thick	-23.8dB/cm	Could not measure	
L1-Q3	Sapphire	SiN LPCVD 200nm Thick, 1.8um wide	SiO PECVD 200nm Thick	-1.13dB/cm	-5.26dB/Facet	
L1-Q4	Sapphire	SiN LPCVD 200nm Thick, 1.8um wide	SiON PECVD 200nm Thick	~3.95dB/cm	-4.00dB/Facet	
L1-Q4	Sapphire	SiN LPCVD 200nm Thick, 1.8um wide	SiON PECVD 1000nm Thick	~1.61dB/cm	-4.50dB/Facet	

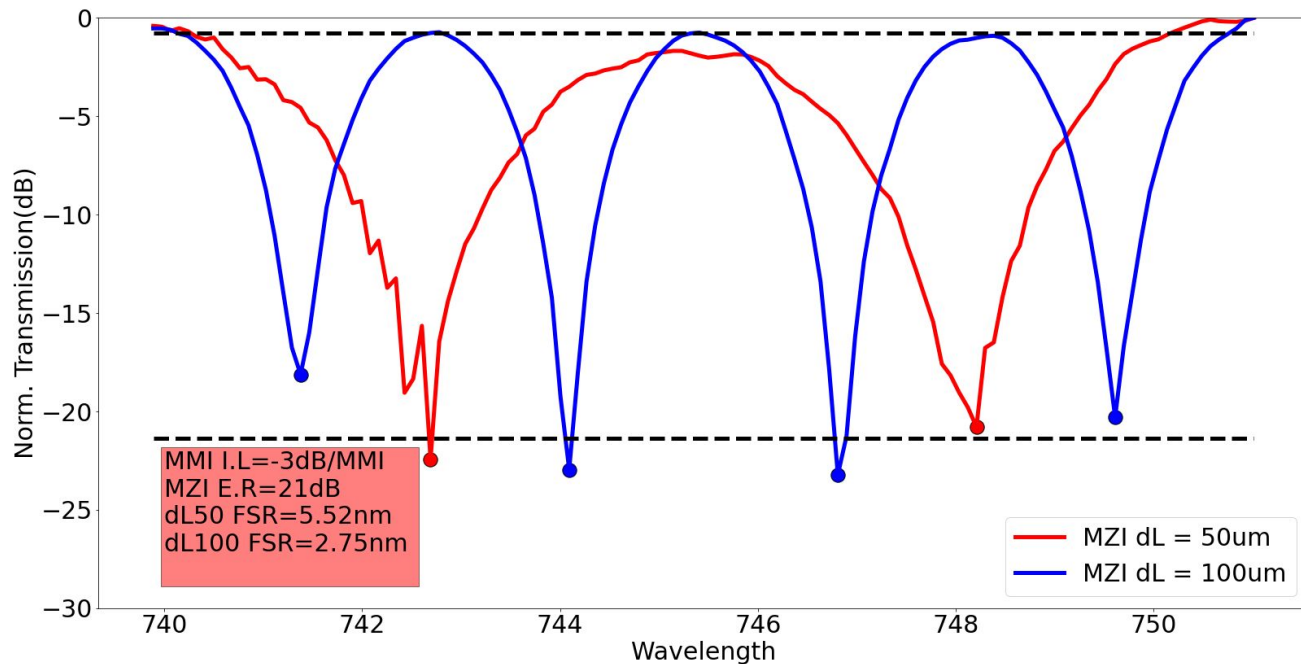


LPCVD SiN on sapphire
Photonic components

MZI Measurements

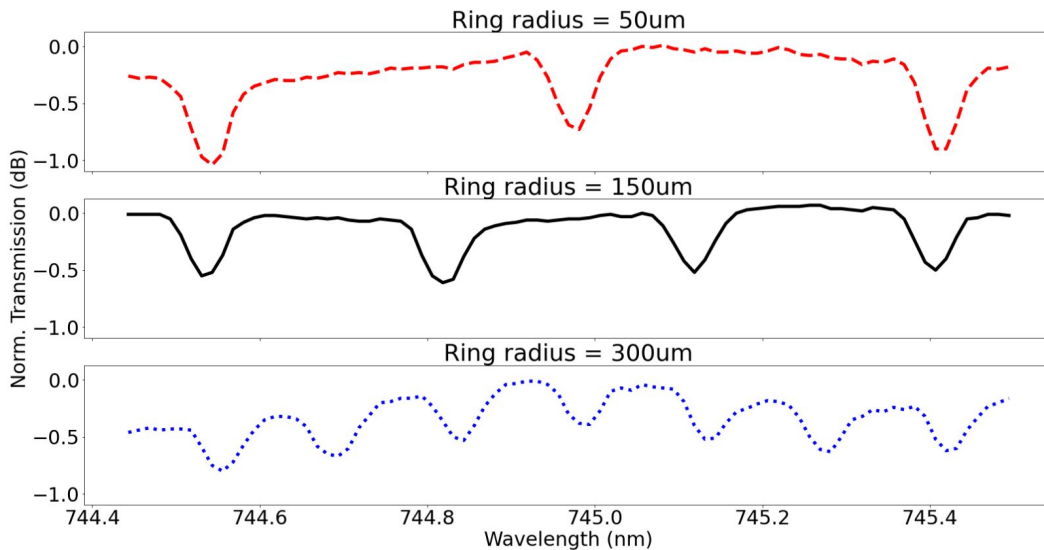
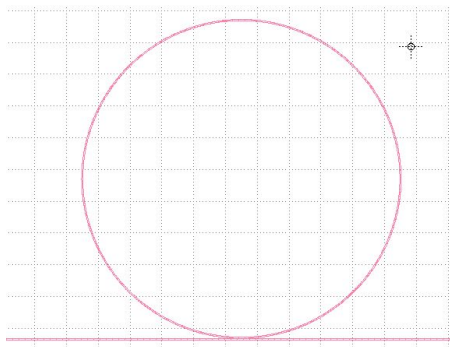


MZI Insertion loss =
-3dB
MZI Extinction ratio =
21dB



MZI dL (um)	Measured FSR (nm)	Expected FSR (nm)
50um	5.52nm	6.21nm
100um	2.75nm	3.12nm

Ring Measurements

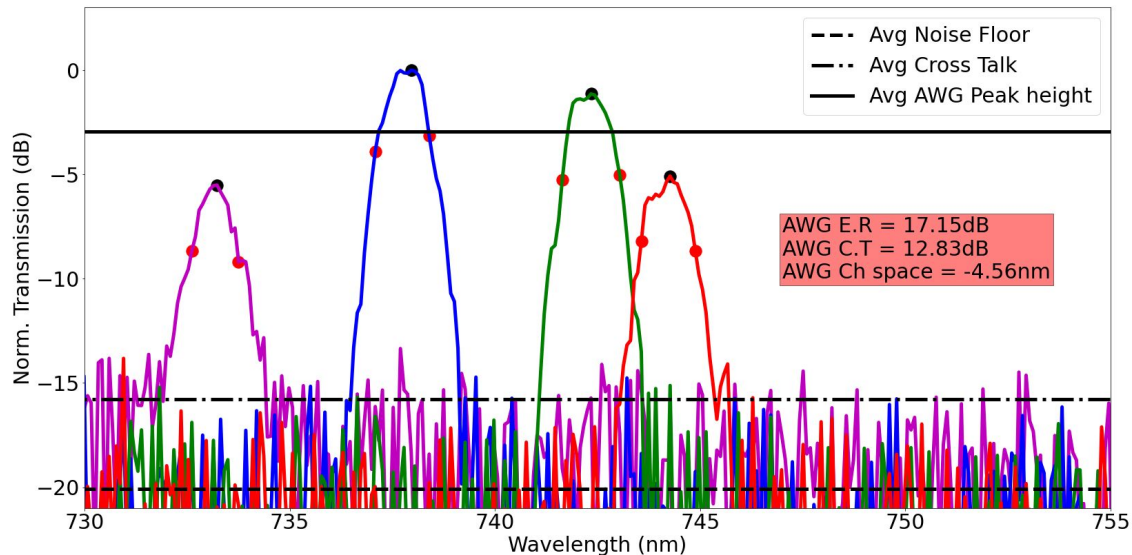
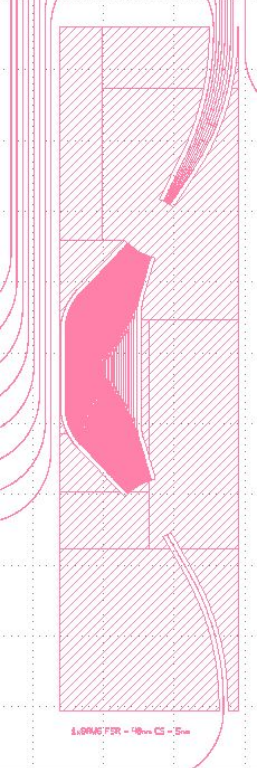


Could not do proper curve fitting

Ring radius (um)	Measured FSR (nm)	Expected FSR (nm)
50um	0.44nm	0.98nm
150um	0.28nm	0.33nm
300um	0.14nm	0.16nm

Ring radius (um)	Kappa(n m)	Loss (dB/cm)
50um	??	??
150um	??	??
300um	??	??

1x8 AWG (1.2um WVG)
FSR=40nm, CS=5nm



- Could get transmission only from Ch6, Ch7, Ch8, Ch1(second order)
- Light from broad band source not guiding in Ch2, Ch3, Ch4
- The AWG center wavelength seemed to have shifted
 - Average AWG Extinction ratio = 17.15 dB
 - Average AWG Cross Talk ratio = 12.83 dB
 - Measured AWG Channel space = 4.56 nm

Cost of Photonic wire bonds

- Not sure how much is the cost. But it is very easily scalable to large arrays in a single chip because it is done using two photon lithography
- It would be a completely new research project on its own

Fiber array - PM\$4000 to make (16channel), \$1,750 to install; non-pm \$2600, attach \$1700 + polish

Freedom Photonics - not standard - \$10K based on what chips are bonded, mode vmatching involved, mostly set up



ULTRA-LOW LOSS
TECHNOLOGIES