

Deep Learning applications for ν_e reconstruction in the ICARUS experiment.

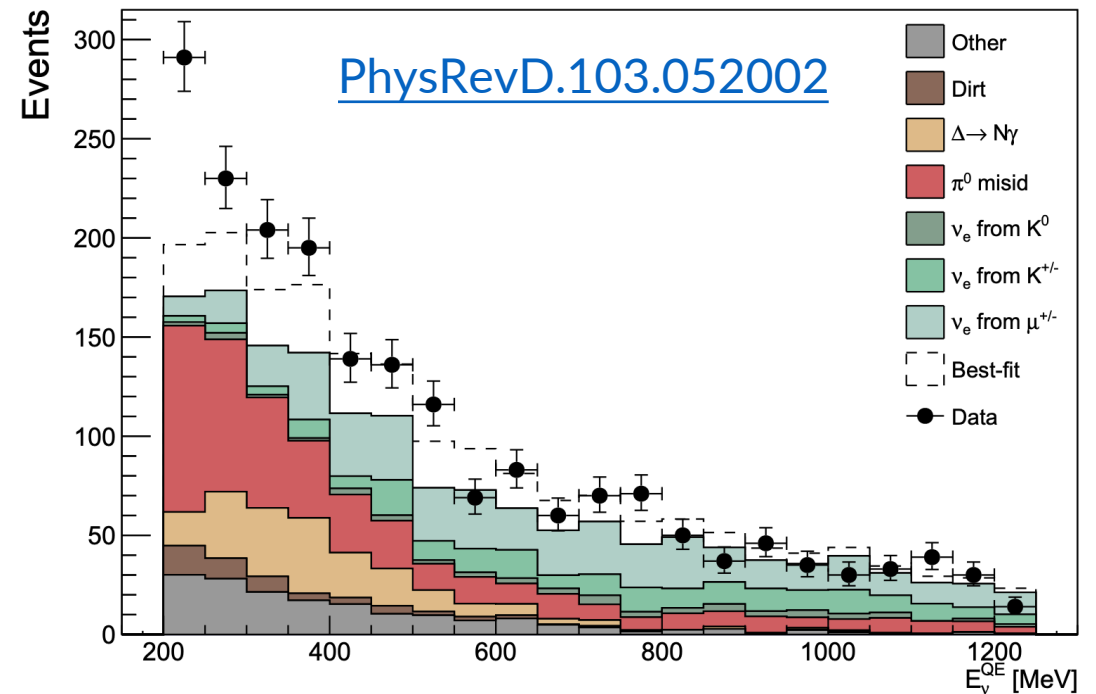
Dae Heun Koh

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Neutrino Physics and Machine Learning 2023

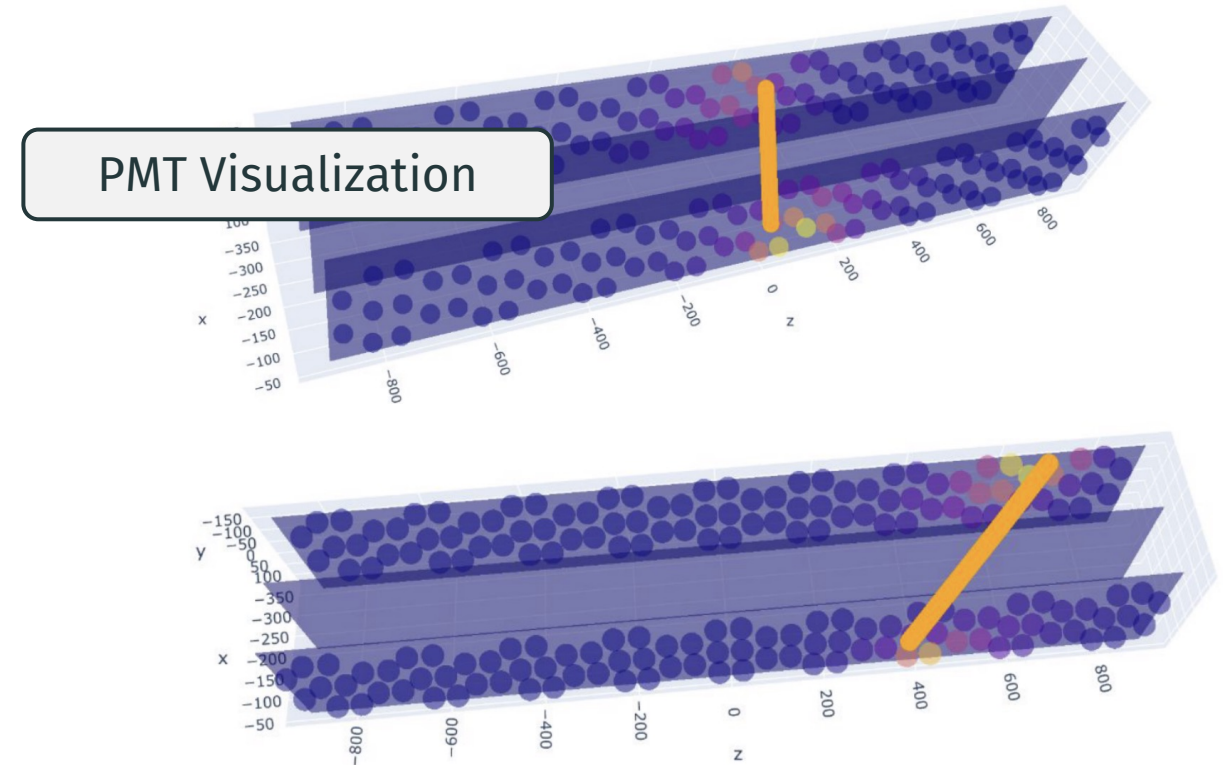
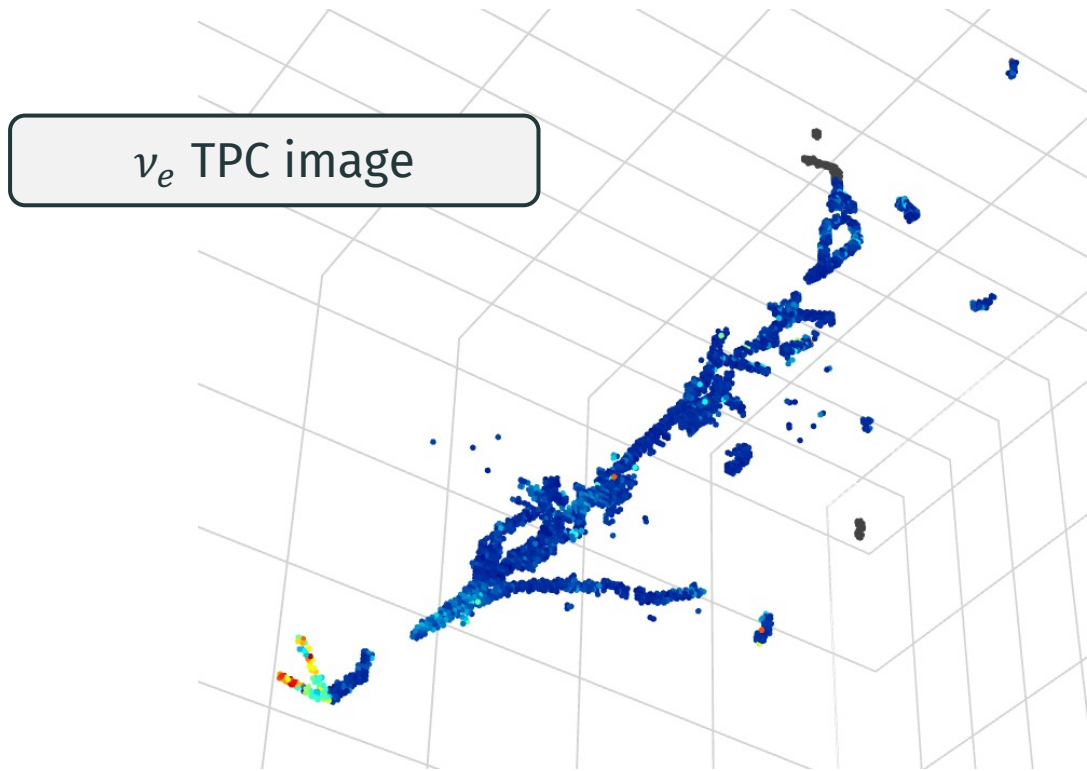
Search for sterile neutrinos with the ICARUS T600 Detector

- **Low-Energy Excess (LEE)** is an excess of electron-like neutrino events in the 200-600 MeV energy range, observed by LSND and MiniBooNE.
- Search for **sterile neutrinos** and investigating ν_e appearance/ ν_μ disappearance in these regions are one of the flagship analysis goal of the SBN program.
- The e/γ **separation capability** of the **liquid argon time projection chamber (LArTPC)** technology, and information from other subsystems will allow ICARUS to resolve the common confounding backgrounds (**CC- π^0** , **$\Delta \rightarrow N\gamma$**) in the MiniBooNE analysis.



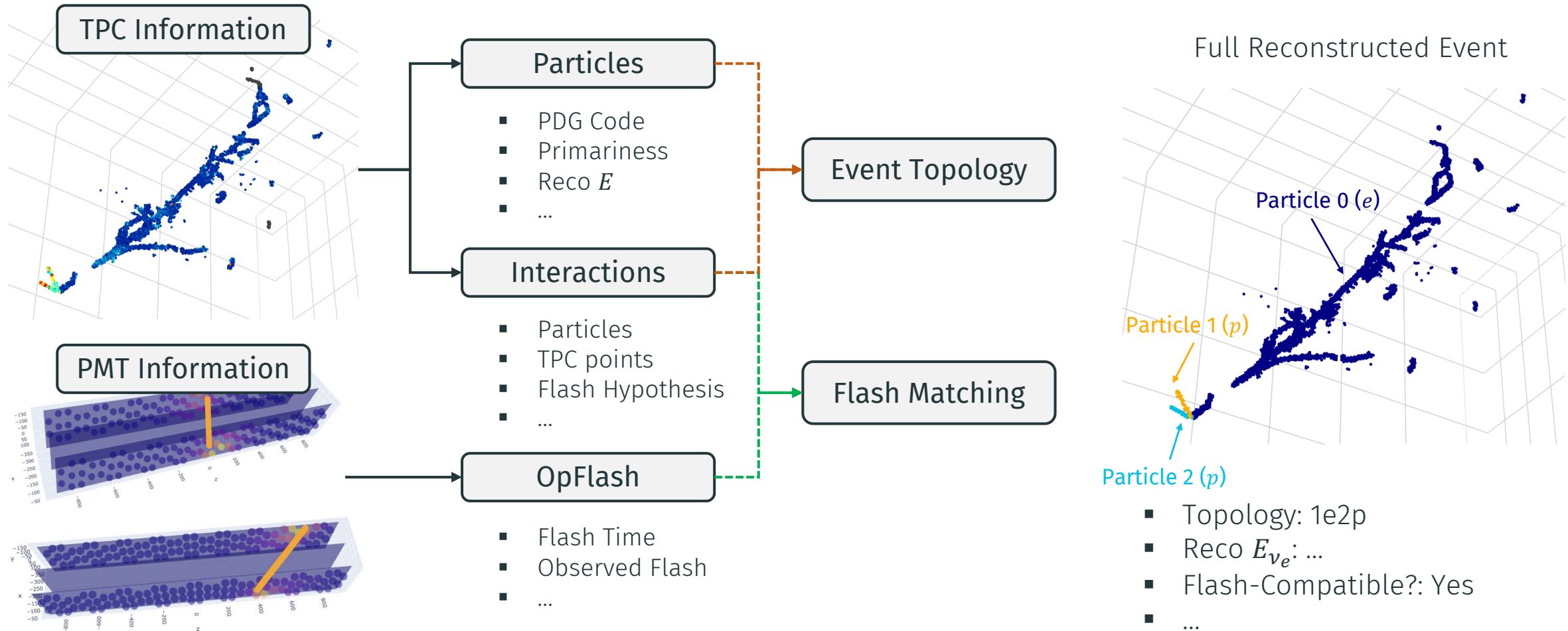
The ICARUS T600 Detector

- The ICARUS T600 detector is composed of three major subsystems:
 - **Time Projection Chambers (TPCs):** allow high resolution imaging of particle trajectories.
 - **Photomultiplier Tubes (PMTs):** scintillation light from charged particles used for interaction timing information
 - **Cosmic Ray Tagger (CRTs):** tagging system for crossing and exiting particles



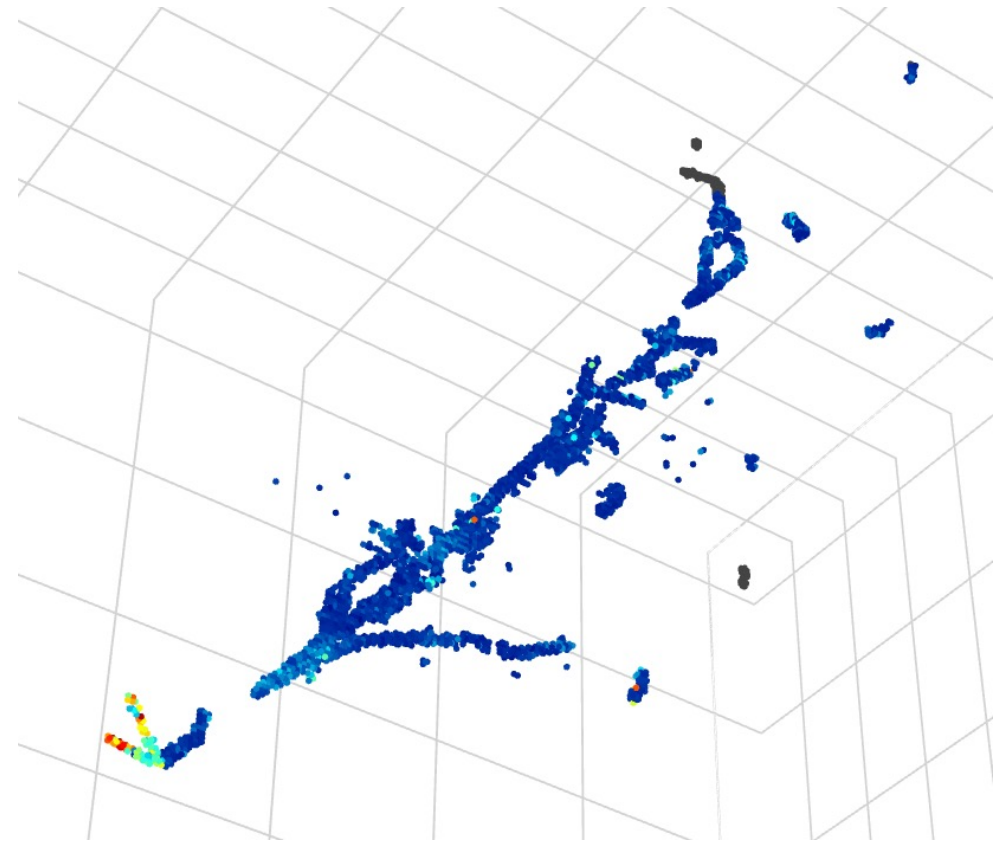
Electron Neutrino Reconstruction

- In this talk, we use TPC and PMT information to identify candidate electron neutrino events.



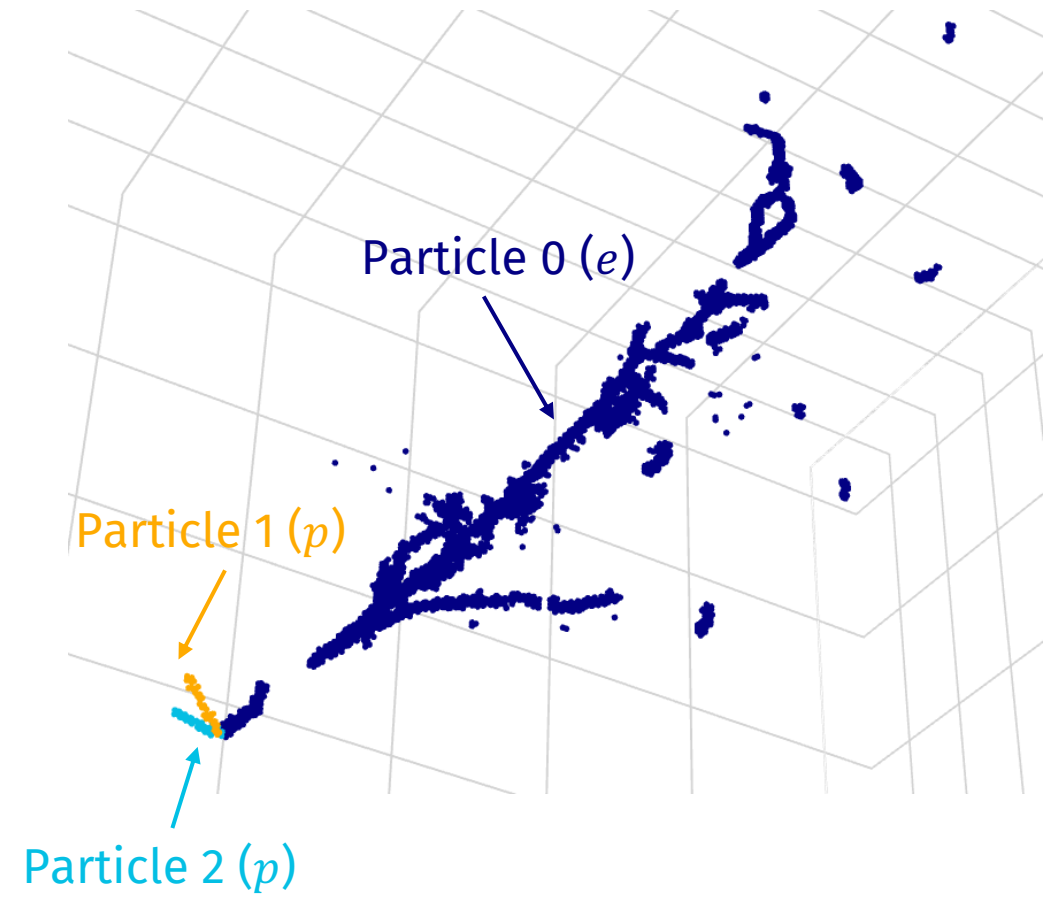
Particle and Interaction Reconstruction using ML

- Analysis using ML tools is a two-stage process:
 1. **Organize neural network predictions** to human-readable objects:



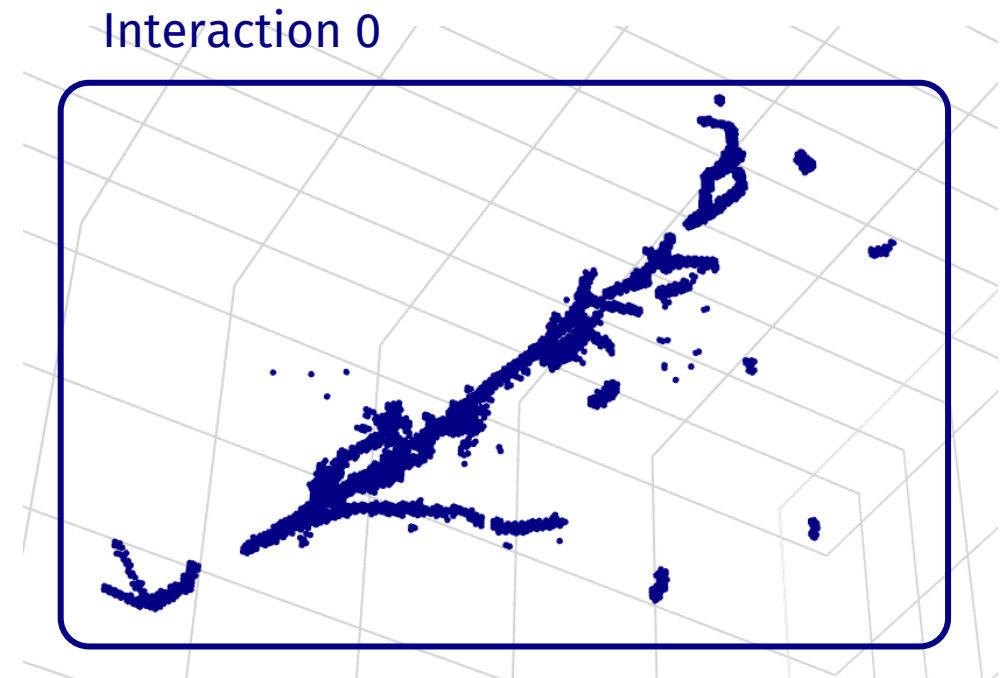
Particle and Interaction Reconstruction using ML

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 - **Particle:** depositions that are predicted to belong to the same particle
 - ex. PDG code, primary indicator, interaction clustering labels



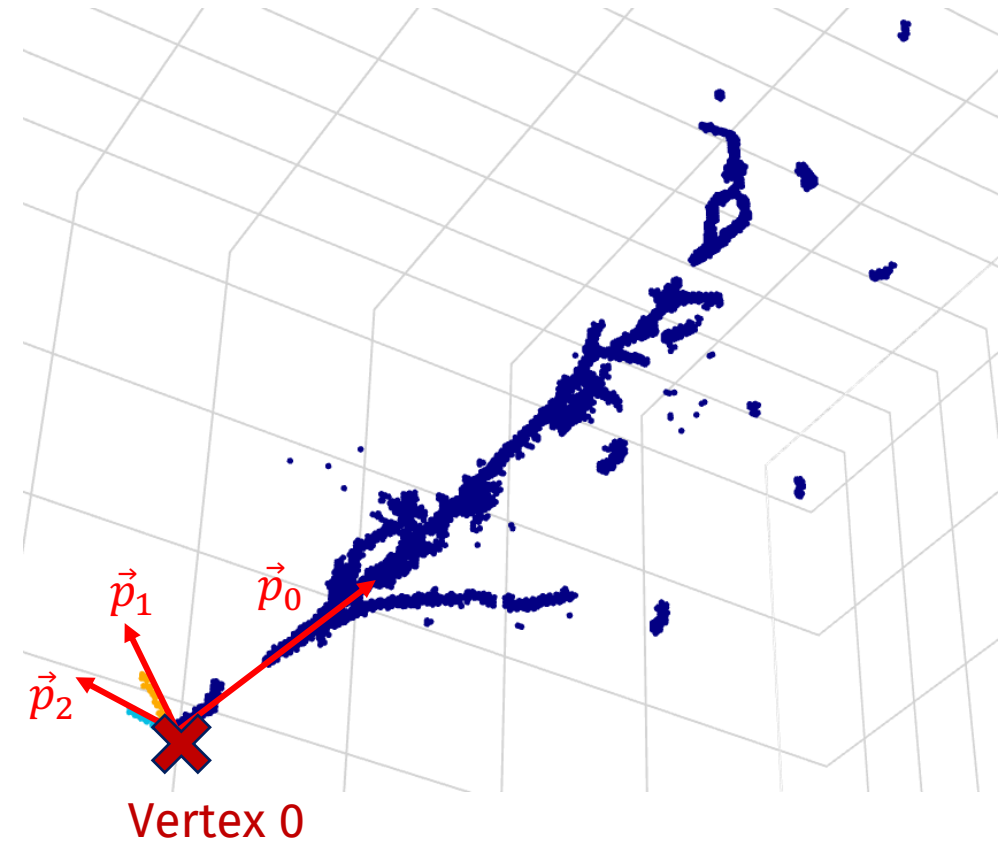
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 - ex. PDG code, primary indicator, interaction clustering labels
 - **Interaction:** collection of particles that originate from the same vertex
 - ex. constituent particles, PMT flash timing,



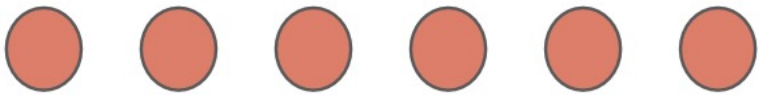
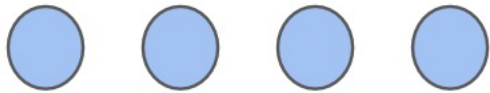
Particle and Interaction Reconstruction using ML

- Analysis using ML tools is a two-stage process:
 1. **Organize neural network predictions** to human-readable objects:
 - **Particle:** depositions that are predicted to belong to the same particle
 - **Interaction:** collection of particles that originate from the same vertex
 2. **Run post-processing algorithms** to further compute useful quantities for reconstruction and append information to **Particle** and **Interaction** instances.
 - ex. Range-based track energy estimation, vertex reconstruction, particle direction estimation



How OpT0Finder works

TPC Interactions
(spatially clustered particle trajectories)

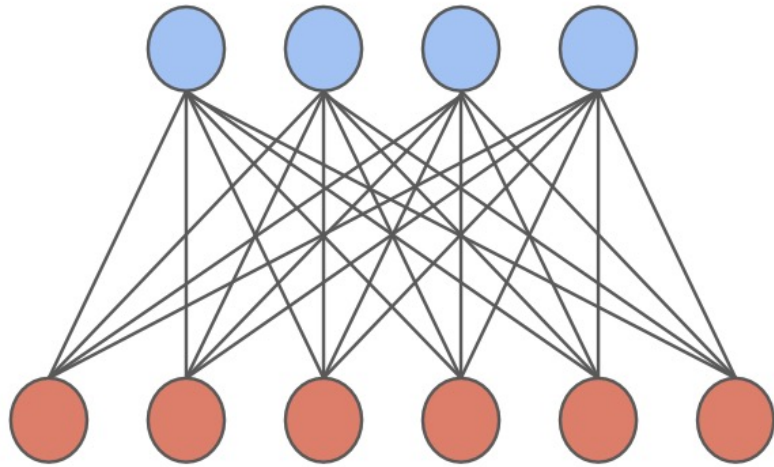


PMT Interactions
("OpFlash" = time-coincident optical signal)

- **TPC Interactions:** group of particle depositions that have the same originating parent (**Qcluster_t**)
- **PMT Interactions:** collection of time-coincident optical signals across PMTs. (**Flash_t**)
- **Goal:** which TPC and PMT interaction share the same underlying interaction / root particle?

How OpT0Finder works

TPC Interactions
(spatially clustered particle trajectories)

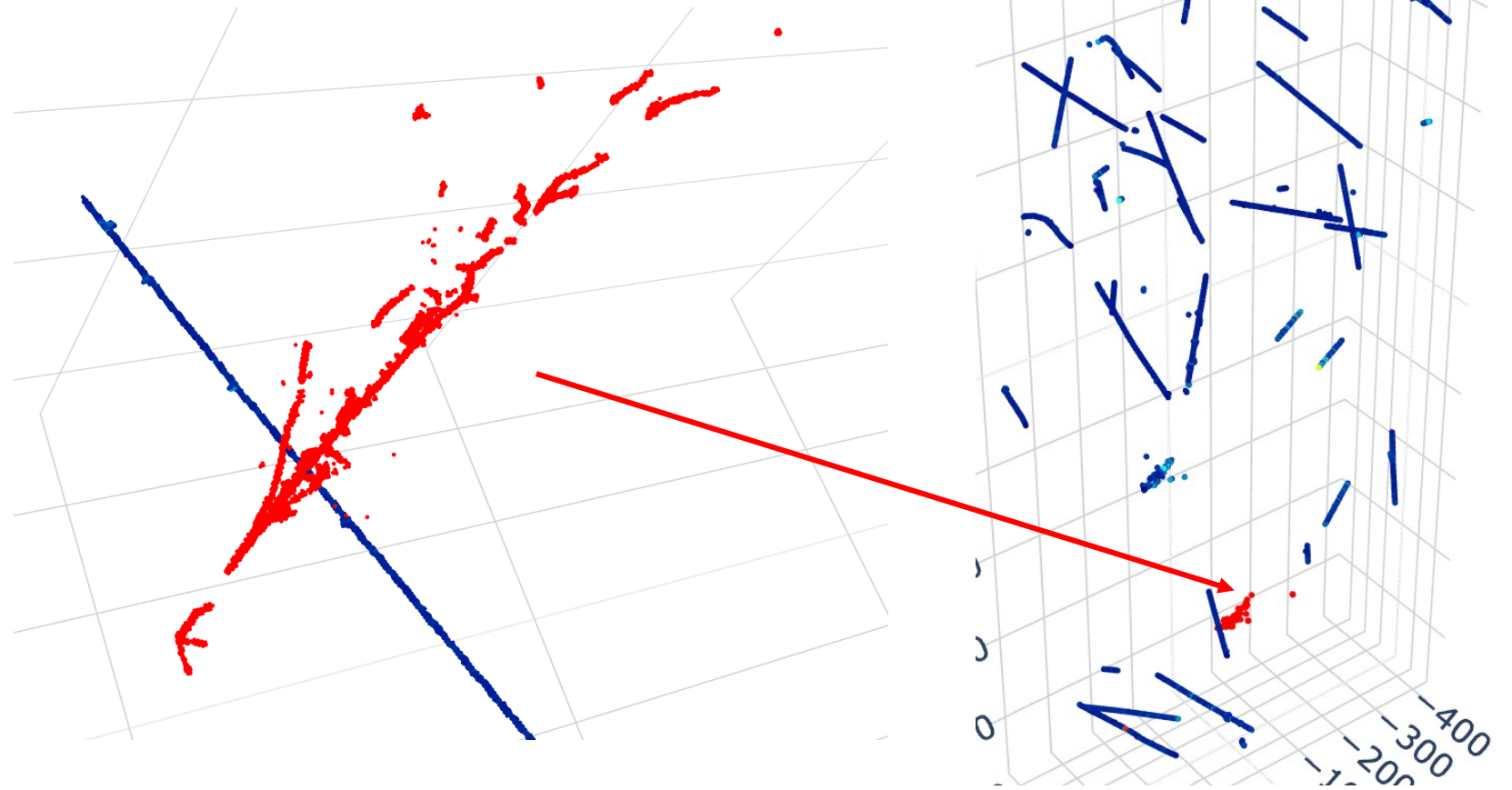


PMT Interactions
("OpFlash" = time-coincident optical signal)

- A form of **weighted-edge bipartite matching problem**.
- For each pair (**QCluster_t**, **Flash_t**):
 - Exclude space/time-wise impossible matches
 - Compute the charge-based log-likelihood (LL) score
 - Find the combination of pairs that maximize the combined LL from all chosen pairs (Greedy, or Munkres/Hungarian Matching algorithm)

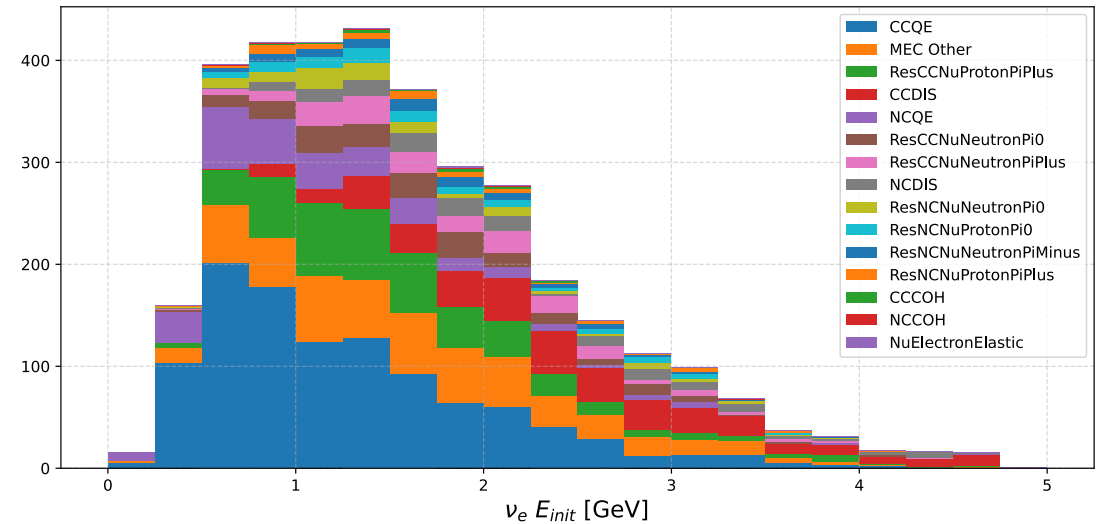
Electron Neutrino Selection: Dataset

- Dataset: BNB ν_e + CORSIKA (3.6k)
 - One $\nu_e + Ar$ interaction, ≈ 36 out-of-time cosmic interactions per image.

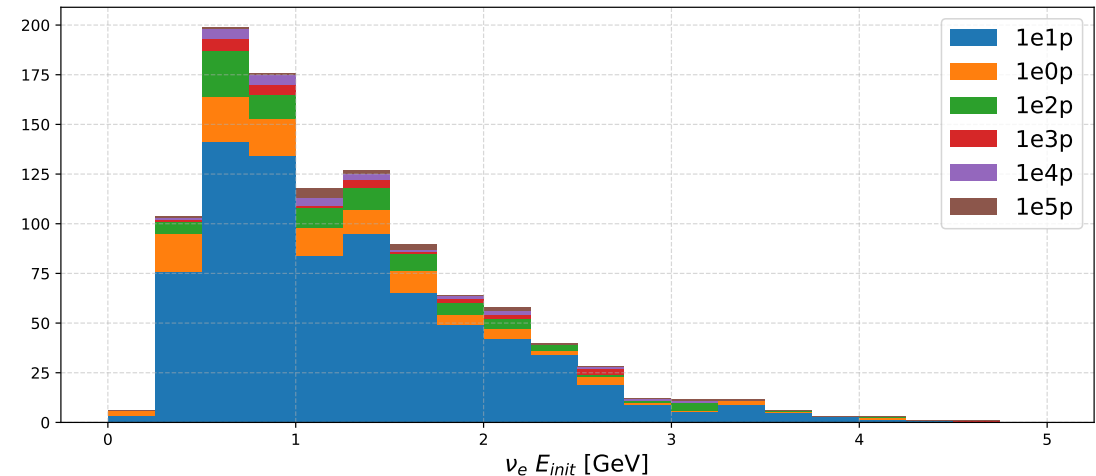


Electron Neutrino Selection: Dataset

- Dataset: BNB ν_e (G4, simulated) + CORSIKA (3.6k)
 - One $\nu_e + Ar$ interaction and ≈ 36 out-of-time cosmic interactions per image.
 - In 200-1000 MeV region, CCQE interaction mode with 1e1p topologies are dominant.

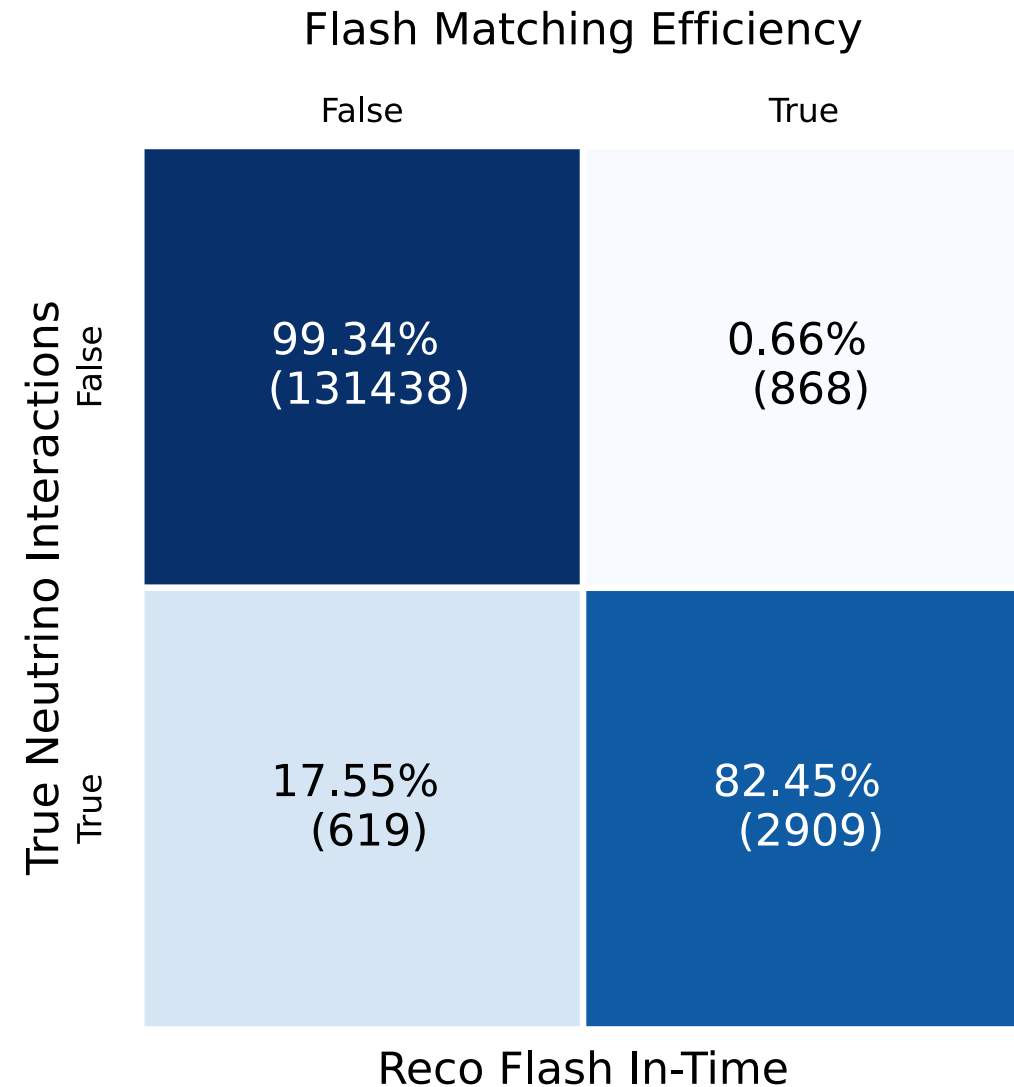


CCQE, Inclusive



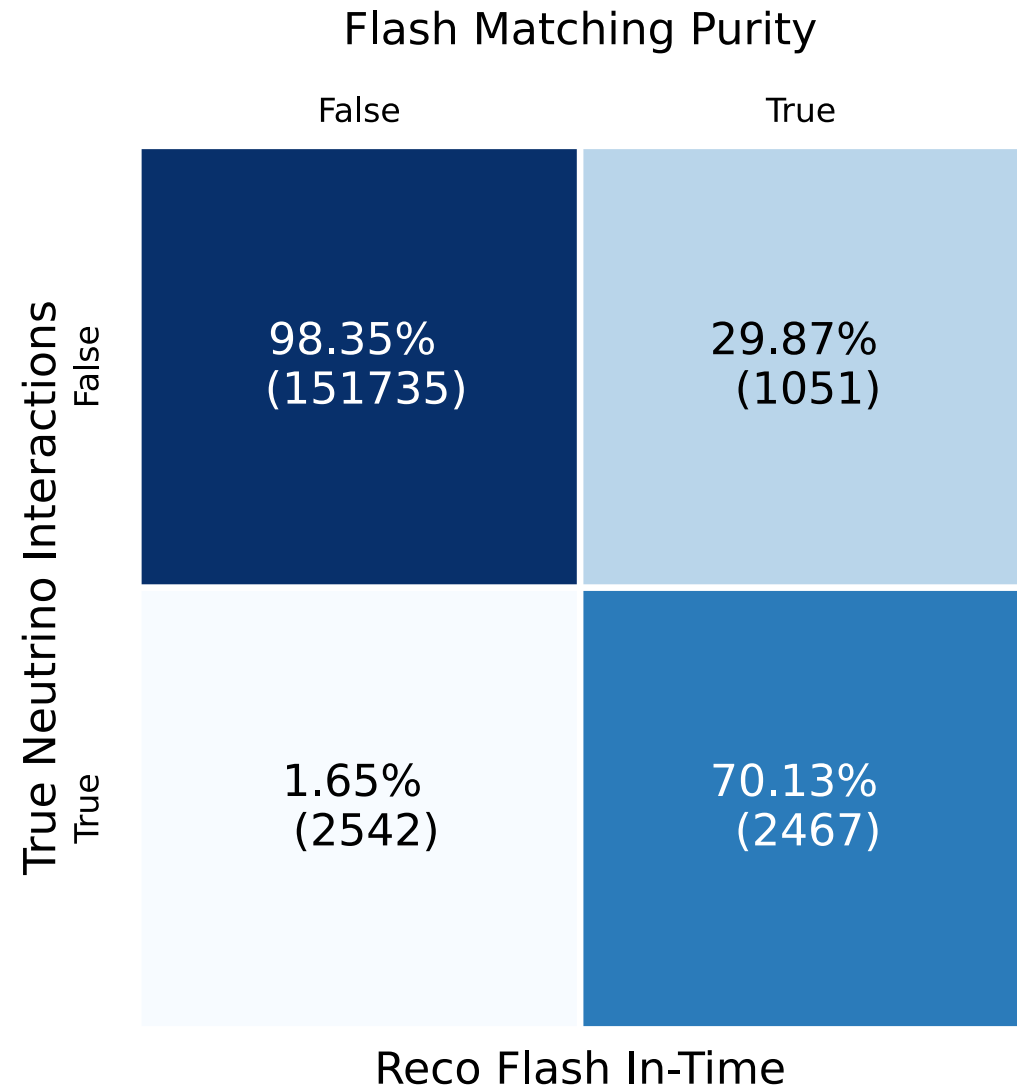
Electron Neutrino Selection: Flash Matching

- Using PMT information alone, we can reject a significant number of cosmic interactions:
 - 82.45%** of all True ν_e 's are matched to a PMT flash within the beam window (0, 1.6 μ s).



Electron Neutrino Selection: Flash Matching

- Using PMT information alone, we can reject a significant number of cosmic interactions:
 - 82.45%** of all True ν_e 's are matched to a PMT flash within the beam window (0, 1.6 μ s).
 - 70.13%** of all reconstructed PMT flashes that are within the beam window are true ν_e 's.



Electron Neutrino Selection: Particle Identification

- Particle Identification with GNNs
- Left: **BNB ν_e Primaries Only**
- e vs. γ separation** is comparable to generic dataset, in good shape.
- Significant $\pi \rightarrow \mu$ and $p \rightarrow \pi$ confusion.
 - Issue in particle energy distribution within interaction at fixed total energy (WIP)

	γ	e	μ	π	p
γ	92.55% (1392)	6.85% (103)	0.00% (0)	0.53% (8)	0.07% (1)
e	12.60% (335)	86.42% (2297)	0.00% (0)	0.98% (26)	0.00% (0)
μ	0.00% (0)	0.00% (0)	0.00% (0)	0.00% (0)	0.00% (0)
π	0.08% (1)	0.00% (0)	24.19% (293)	66.89% (810)	8.84% (107)
p	0.00% (0)	0.00% (0)	1.22% (62)	11.25% (574)	87.53% (4466)

True Particle Type (rows), Reco Particle Type (columns)

Electron Neutrino Selection: $1eNp0\pi^\pm$

- **Signal Definition:** $1eNp0\pi^\pm$ Predicted Interactions with PMT flash time within the beam window ($1eNp0\pi^\pm + FM$)
- **Efficiency:** measure of how well the reconstruction method captures true $\nu_e 1eNp0\pi^\pm$ interaction as signal.
 - *Efficiency = # of true positives / # of True $1eNp0\pi^\pm$*

Row-Normalized
Truth -> Reco

		1eNp0 π^\pm +FM	Background
		True Neutrino Interactions	
Cosmic	0.01% (17)	99.99% (132122)	
True $\nu_e 1eNp0\pi^\pm$	55.34% (948)	44.66% (765)	
True ν_e Other	5.34% (97)	94.66% (1718)	
		Reconstructed	

Electron Neutrino Selection: $1eNp0\pi^\pm$

- **Signal Definition:** $1eNp0\pi^\pm$ Predicted Interactions with PMT flash time within the beam window ($1eNp0\pi^\pm + FM$)
- **Efficiency:** measure of how well the reconstruction method captures true $\nu_e 1eNp0\pi^\pm$ interaction as signal.
 - *Efficiency = # of signal / # of True $1eNp0\pi^\pm$*
- **Purity:** measure of how many predicted signal interactions actually correspond to a true signal.
 - *Purity = # of true positives / # of all signal interactions*

Column-Normalized
Reco -> Truth

		1eNp0 π^\pm + FM	Background
		True Neutrino Interactions Cosmic	0.36% (4)
True Neutrino Interactions True $\nu_e 1eNp0\pi^\pm$	True $\nu_e 1eNp0\pi^\pm$	95.13% (1054)	0.81% (1557)
	True ν_e Other	4.51% (50)	1.90% (3669)
		Reconstructed	

Electron Neutrino Selection: $1eNp0\pi^\pm$

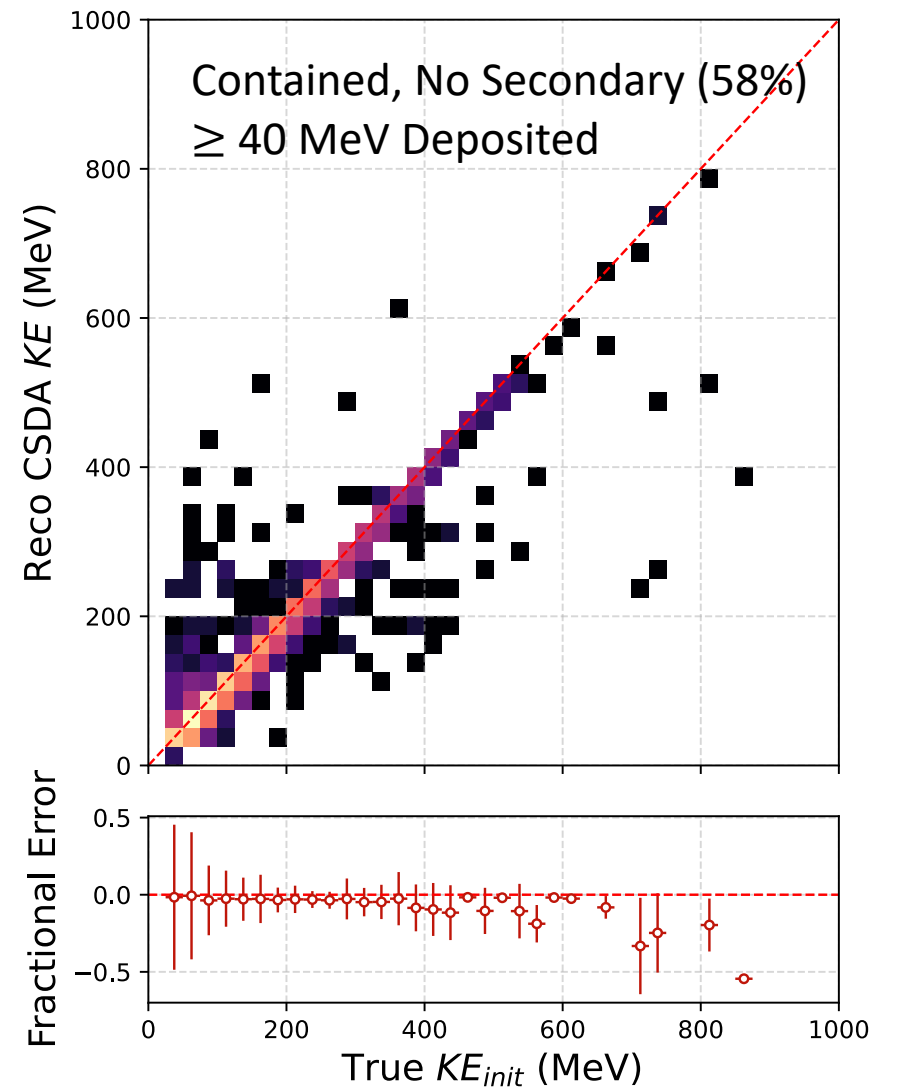
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 - *Efficiency = # of signal / # of True $1eNp0\pi^\pm$*
- **Purity:** measure of how many predicted signal interactions actually correspond to a true signal.
 - *Purity = # of true positives / # of all signal interactions*
- **Summary: 55% Efficiency, 95% Purity for $1eNp0\pi^\pm$**

Column-Normalized
Reco -> Truth

		$1eNp0\pi^\pm + FM$	Background
True Neutrino Interactions	Cosmic	0.36% (4)	97.30% (188035)
	True $\nu_e 1eNp0\pi^\pm$	95.13% (1054)	0.81% (1557)
	True ν_e Other	4.51% (50)	1.90% (3669)
		Reconstructed	

Electron Neutrino Selection: Proton Energy Estimation

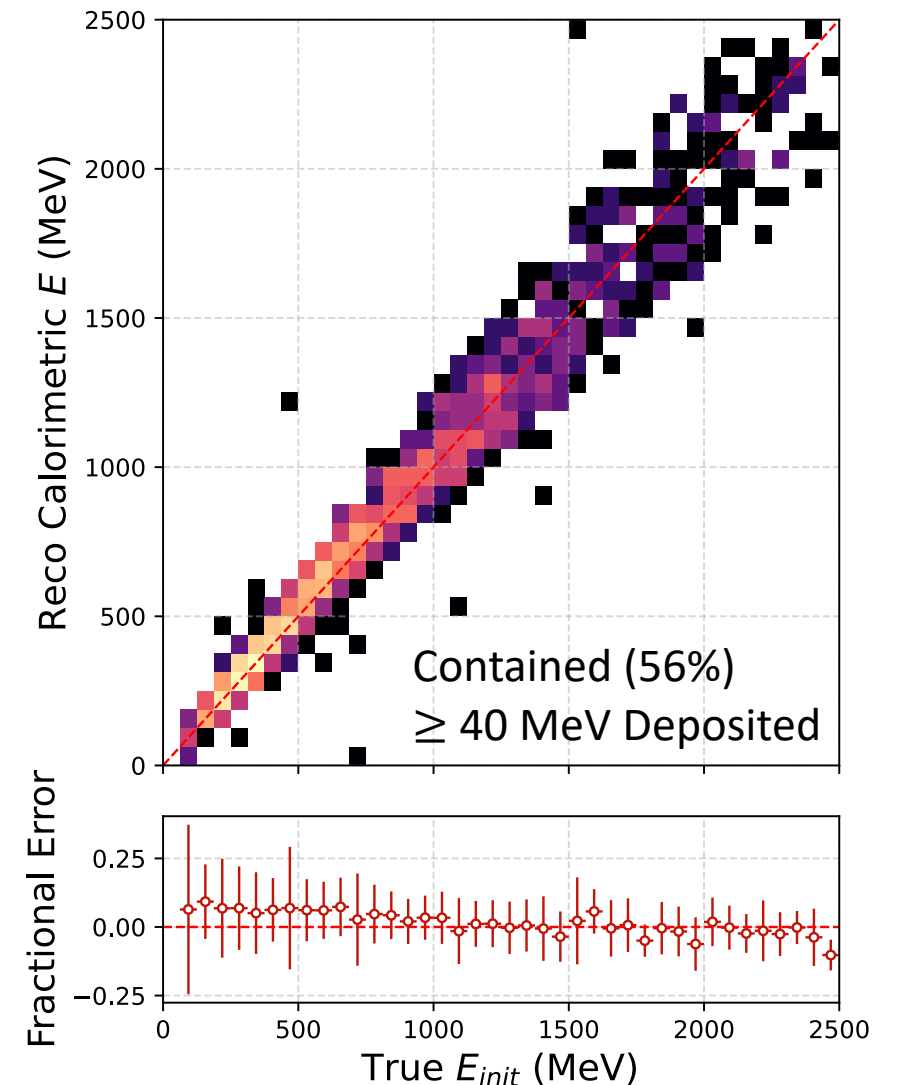
- For **protons**, we use the integrated Bethe-Bloch relation to compute the initial kinetic energy KE in terms of the particle travel range R .
- Fractional Error:** $\epsilon = (\widehat{KE} - KE_{true})/KE_{true}$



Electron Neutrino Selection: Electron Shower Energy Estimation

- For **protons**, we use the integrated Bethe-Bloch relation to compute the initial kinetic energy KE in terms of the particle travel range R .
- Fractional Error:** $\epsilon = (\widehat{KE} - KE_{true})/KE_{true}$
- For **electron showers**, the ADC total sum is converted to MeV via a calibration factor:

$$\widehat{KE}_e = \rho * (\text{total sum of ADC})$$

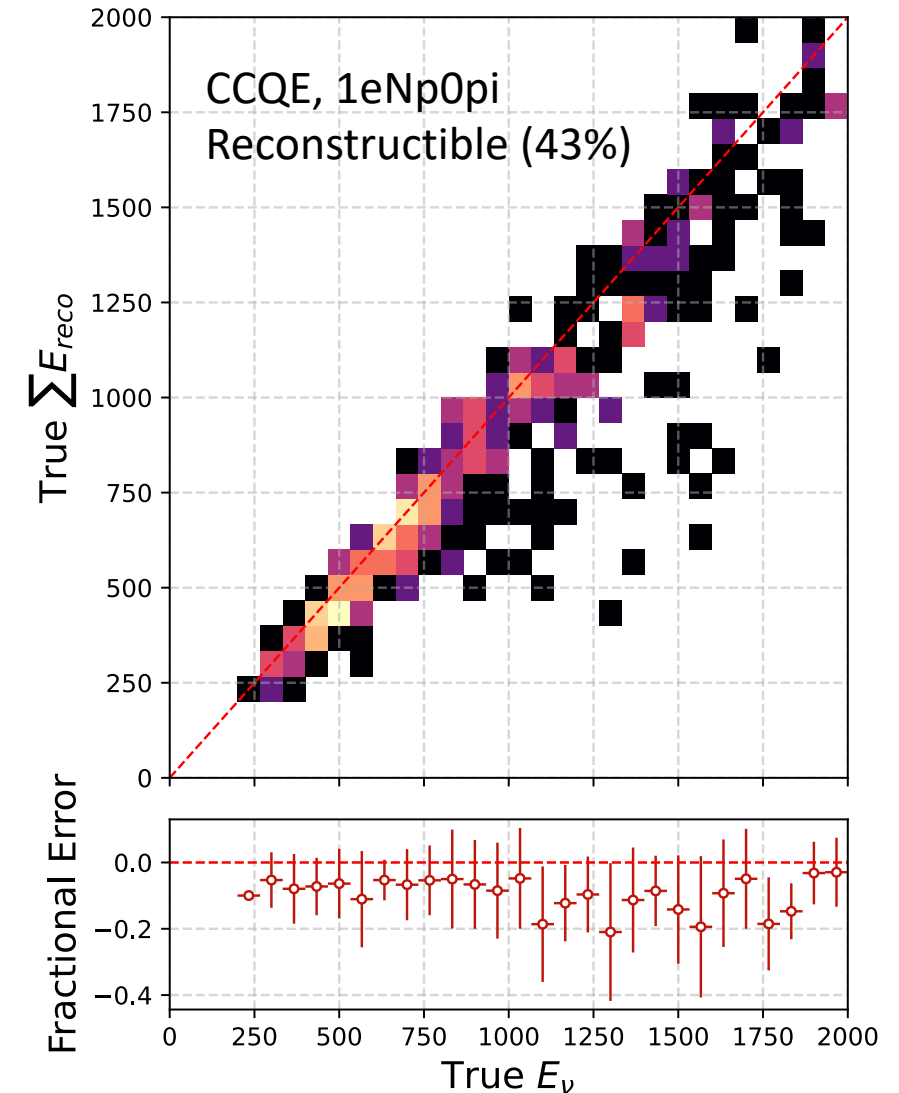


Electron Neutrino Selection: Neutrino Energy Reconstruction

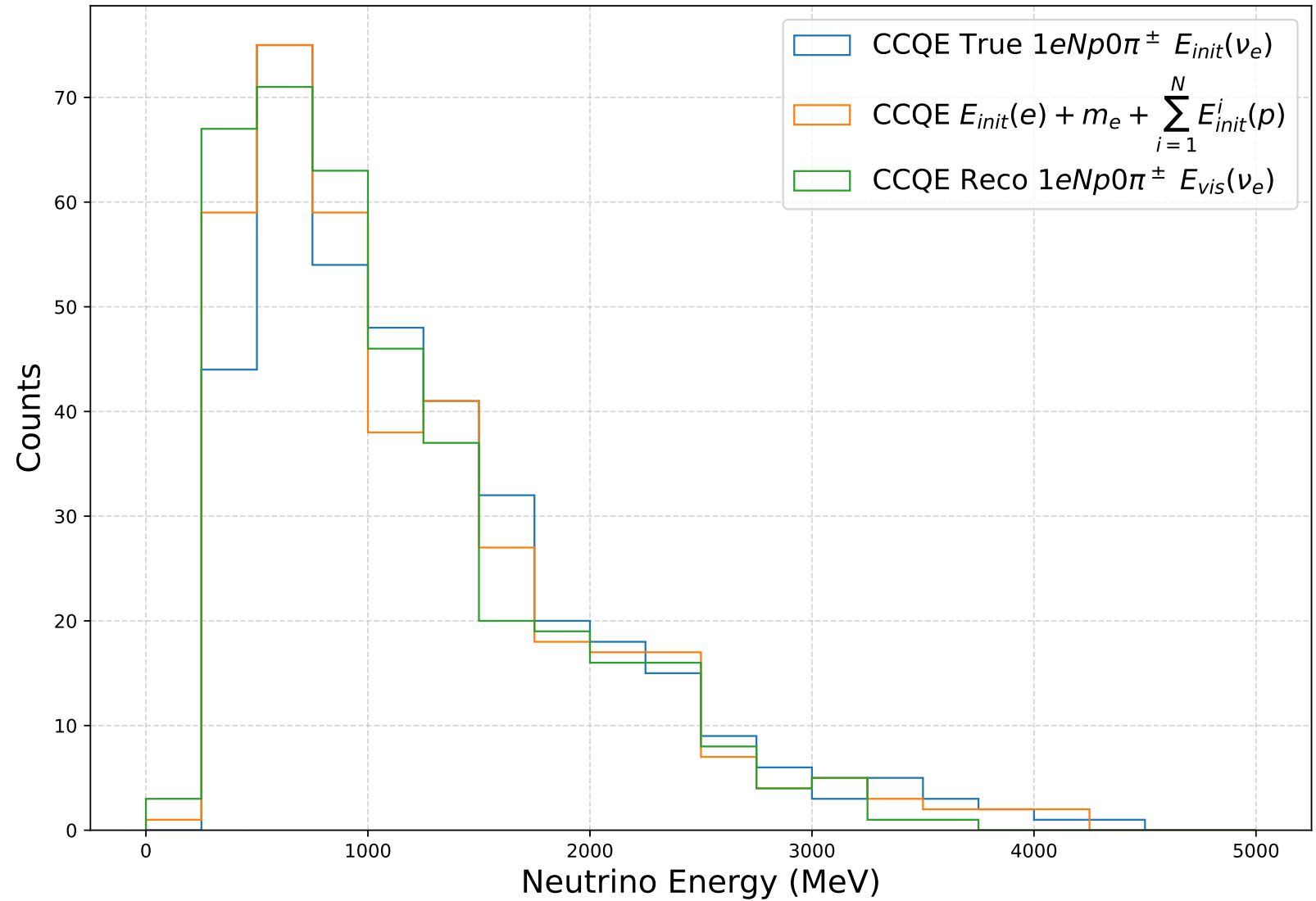
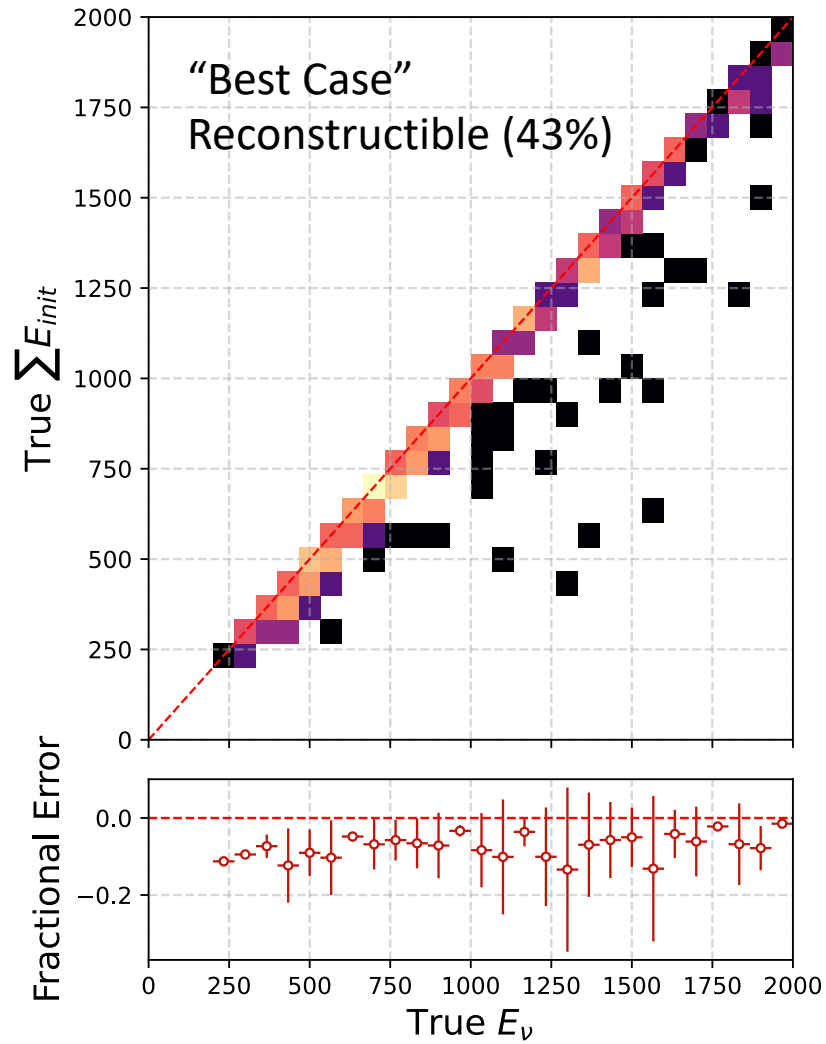
- Since the initial KE of protons and electrons are known, the total visible energy is given as:

$$E_{vis} = KE_e + m_e + \sum_{i=1}^N KE_p^i$$

- Any energy carried by outgoing neutrons or nucleon binding energy is not accounted for.
- “Reconstructible”: all constituent electrons and protons pass the cuts.

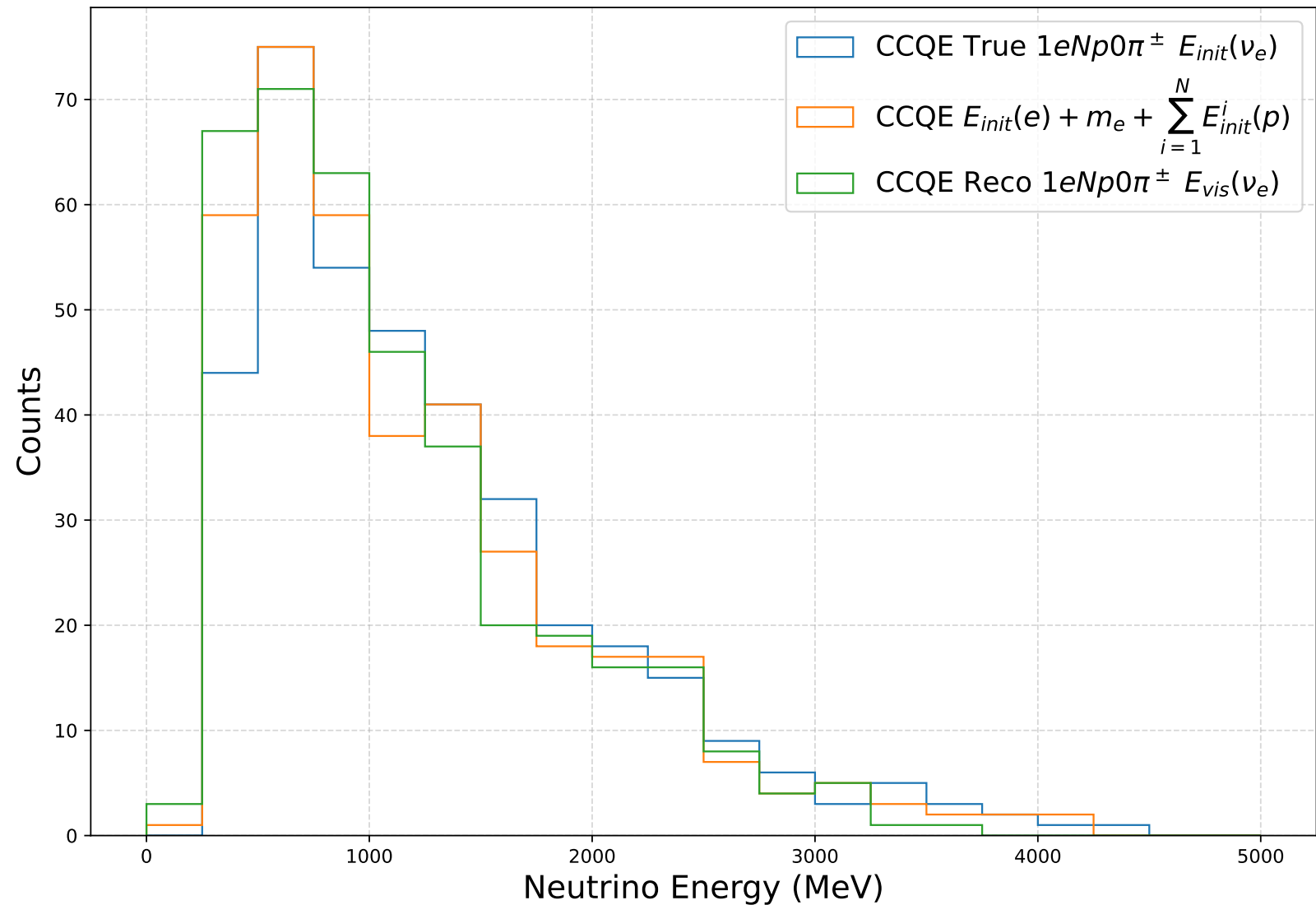


Electron Neutrino Selection: Neutrino Energy Reconstruction



Electron Neutrino Selection: Neutrino Energy Reconstruction

- Currently, E_{vis} falls short of estimating the true initial energy of the neutrino.
- Work in progress on more robust calorimetry methods for proton length estimation.
- Shower Energy Reconstruction (L. Kashur) validation using CC- π^0



Conclusion

- Outline for CCQE electron neutrino selection and energy reconstruction using ML tools and traditional calorimetry algorithms.
- 55% Efficiency and 95% Purity in $1eNp0\pi$ Exclusive Selection on BNB ν_e + OOT(out of time) cosmic background simulation dataset.
- Higher statistics study with both out-of-time and in-time cosmic background
- Simulation vs. Data Study with run9435 data and hand-scanned events.
- Further fine-tuning of calorimetric post-processing algorithms.