



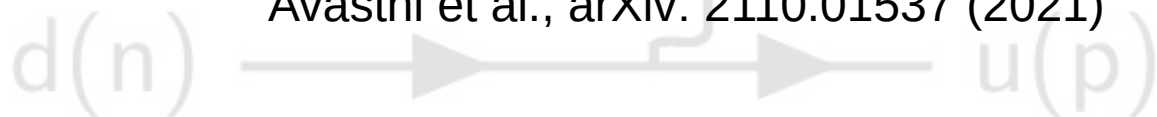
ORIGIN-X:
A $0\nu 2\beta$ experiment with $T_{1/2}$ sensitivity of 10^{30} years.

Mike Heffner

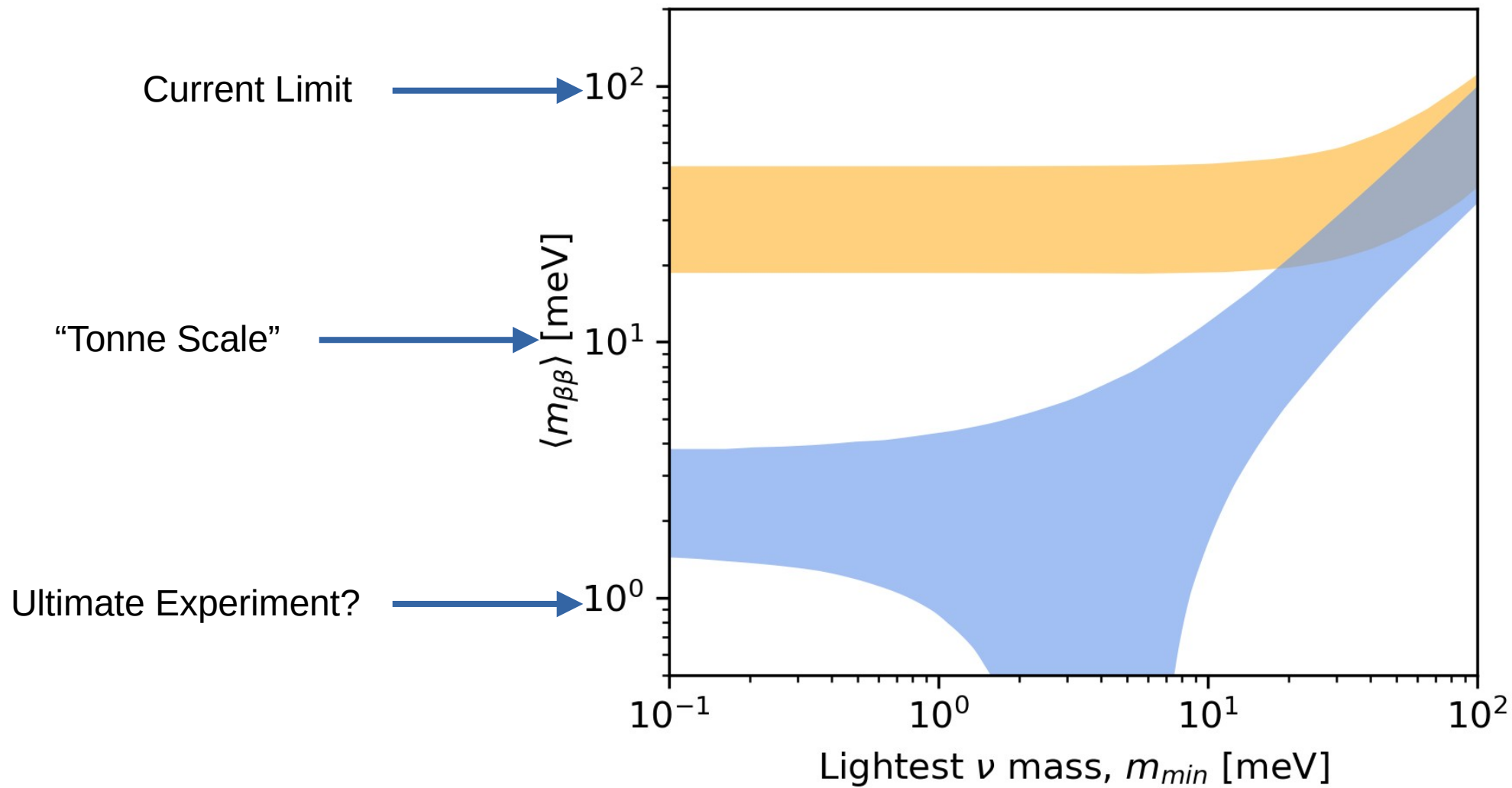
Lawrence Livermore National Lab

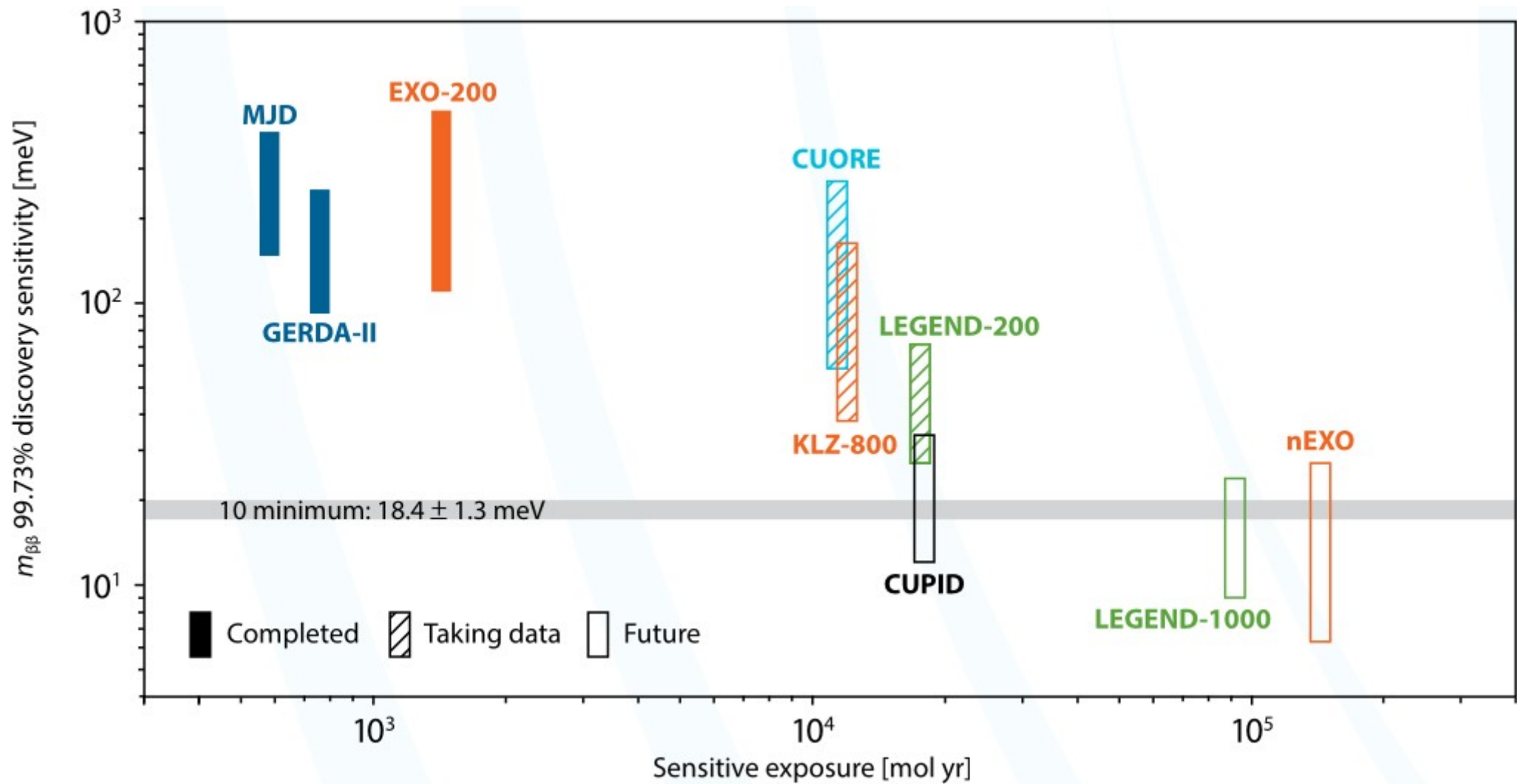
ktonne TPC Workshop SLAC 2023

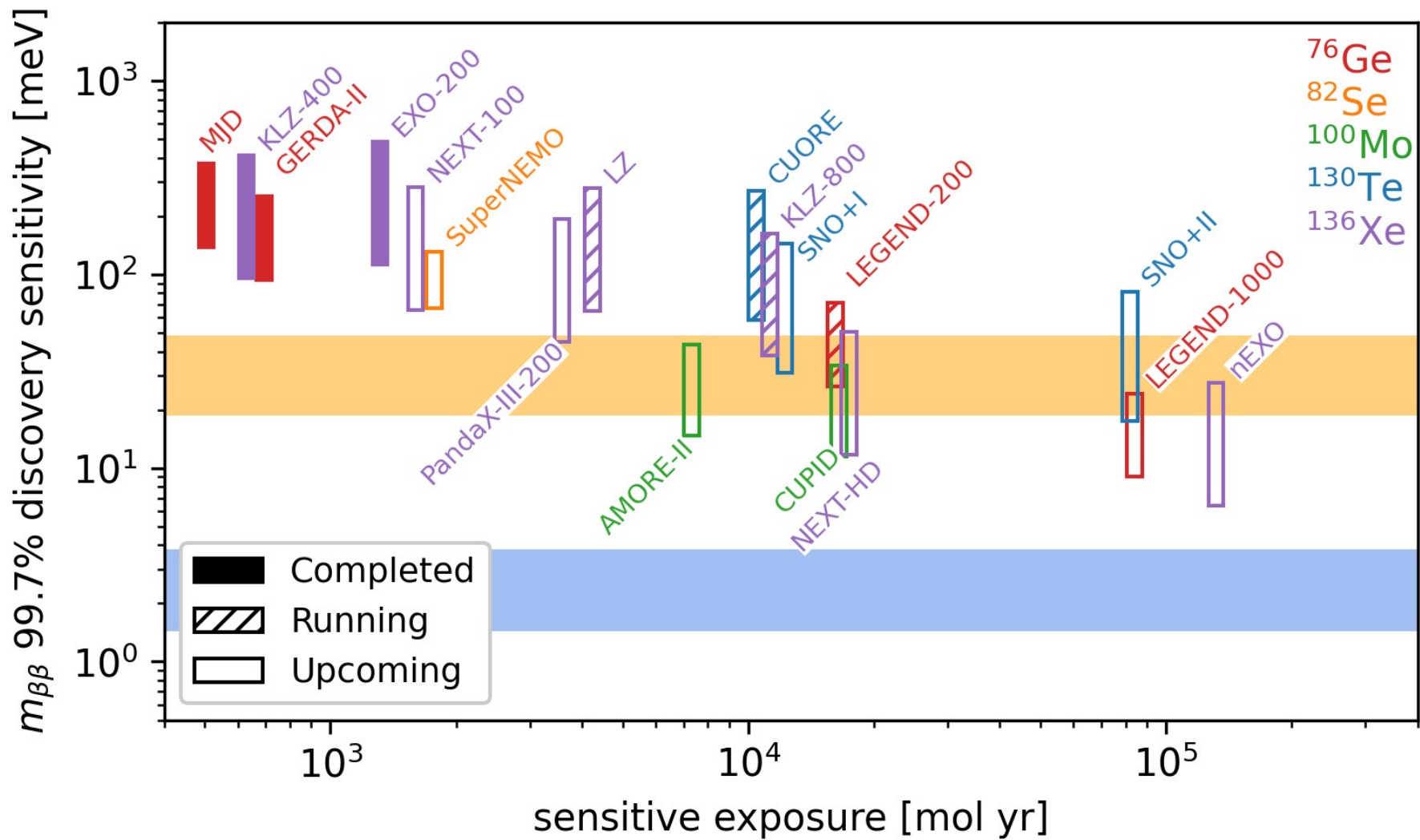
Avasthi et al., arXiv: 2110.01537 (2021)



What is the largest $0\nu 2\beta$ experiment
that should/could be built?

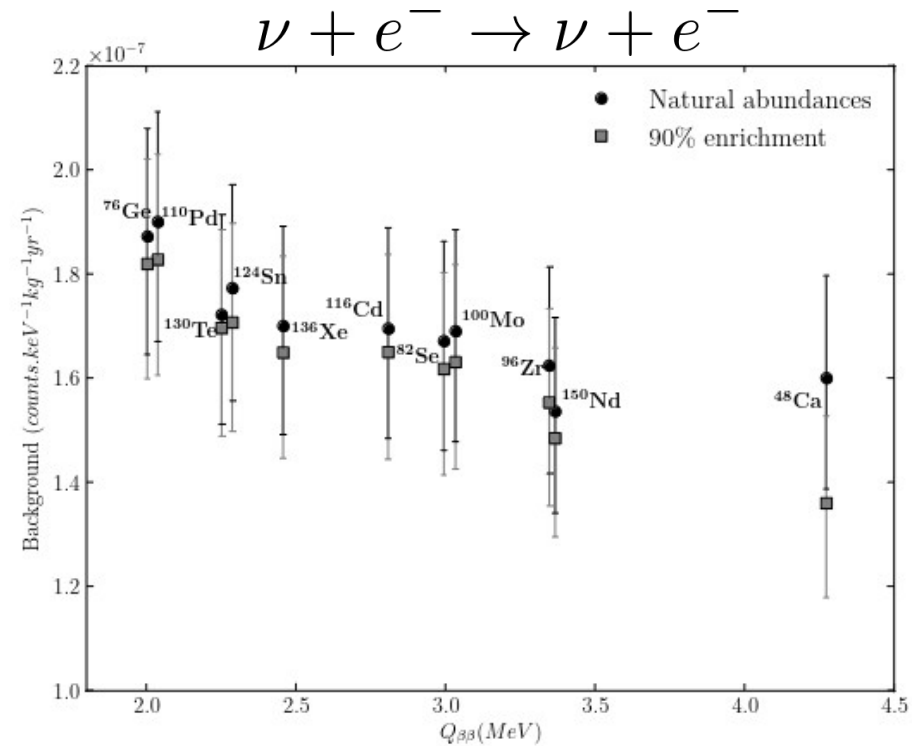
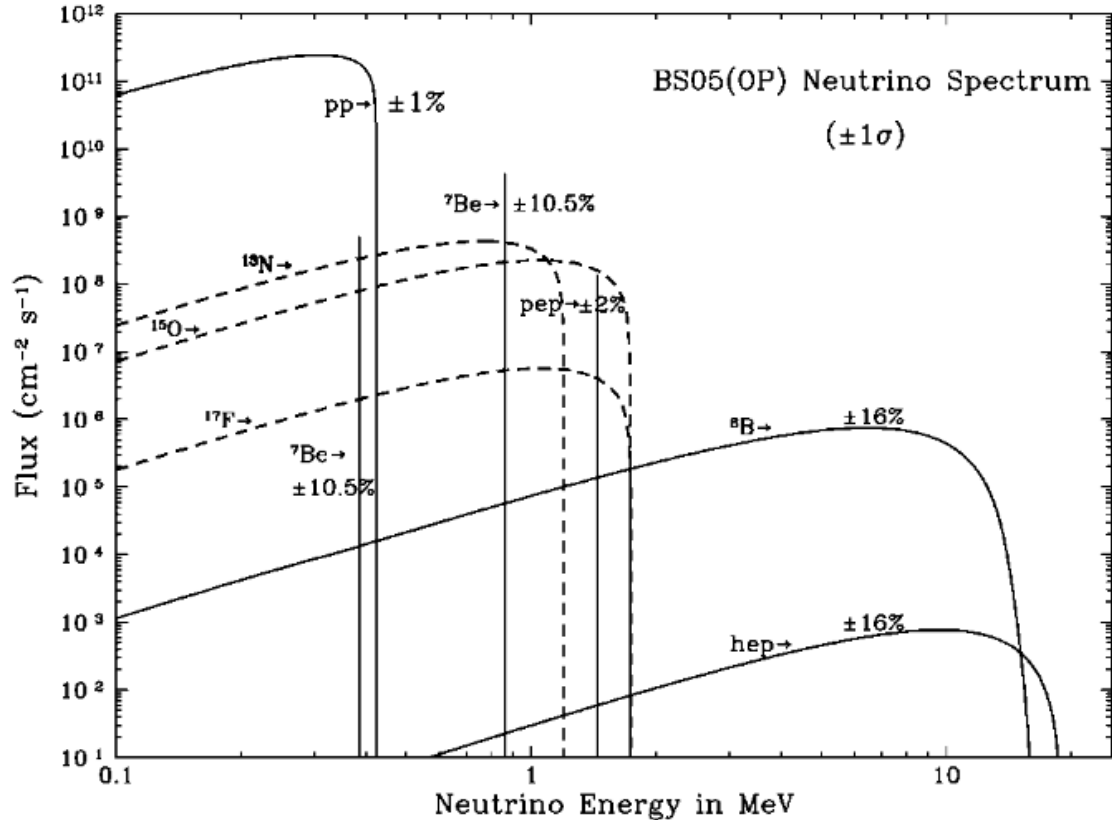




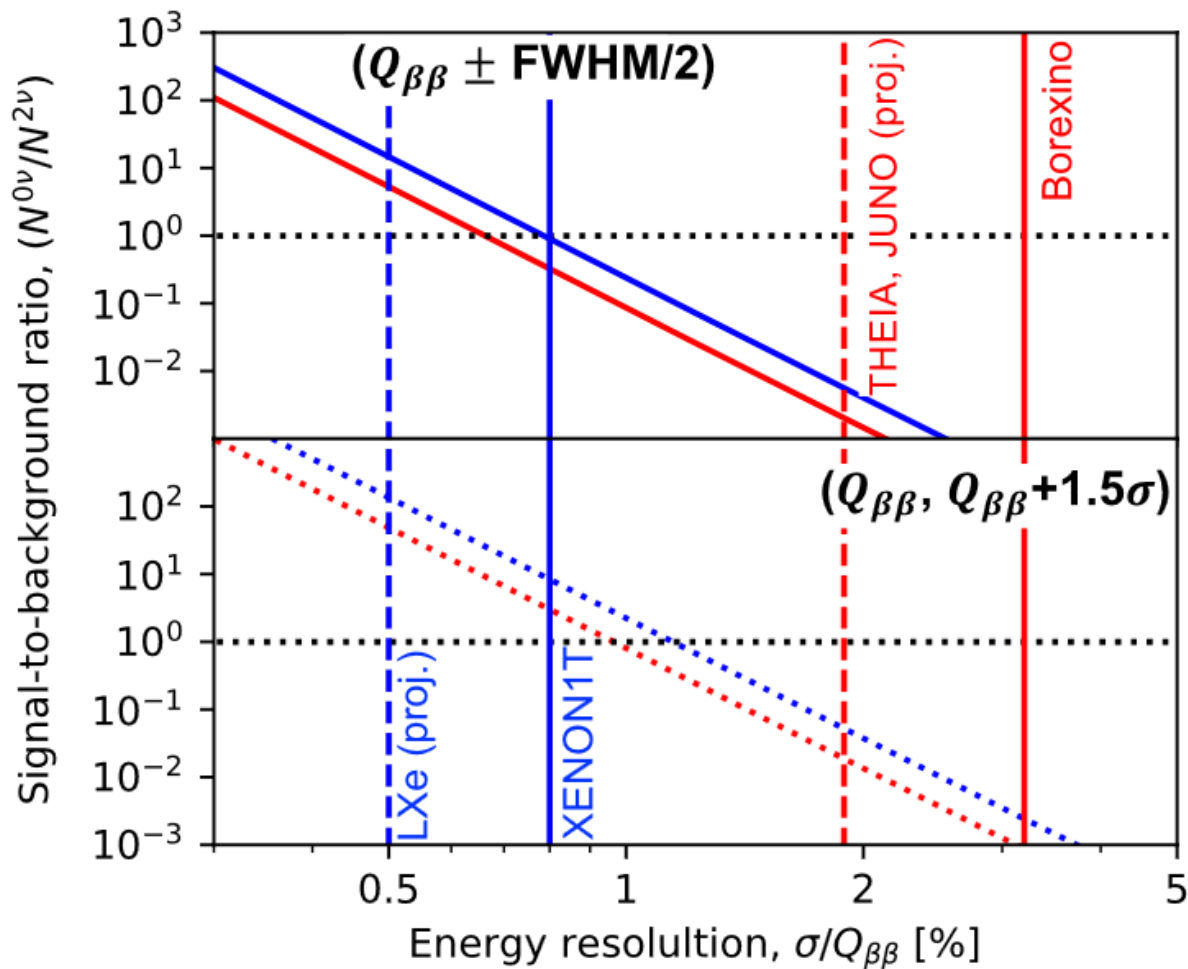
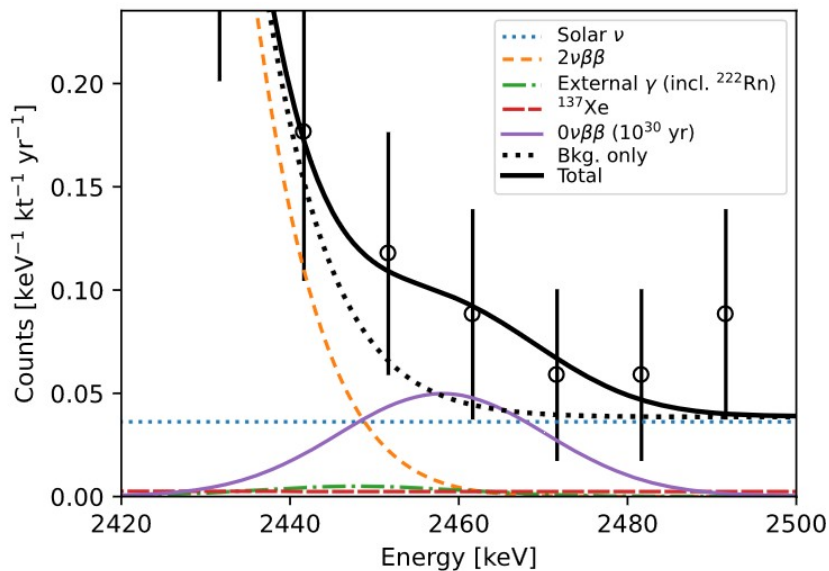


The limits: solar neutrinos

5 (2) evts/(FWHM kTonne yr) @ 0.5% (0.2%)



The limits: 2ν2B



Assume we can perfectly reject background, how much mass do we need?

Radioactive Decay

$$\frac{dN}{dt} = -\frac{\ln(2)}{T_{1/2}} N$$

<u>Half life(yr)</u>	<u>Events/tonne/yr*</u>
10^{26}	30
10^{27}	3
10^{28}	0.3
10^{29}	0.03
10^{30}	0.003

*for xenon

Amount of Xenon that Could be Collected

~2 Gtonne of Xenon in Atmosphere

5.2×10^{21} g Handbook of Chemistry and Physics



NASA Photo of Earth Atmosphere

1kTonne is only $\sim 10^{-6}$ of the total abundance

87 nL/L Xenon concentration

Assume ideal gas

29 g/mol for air

136 g/mol for xenon

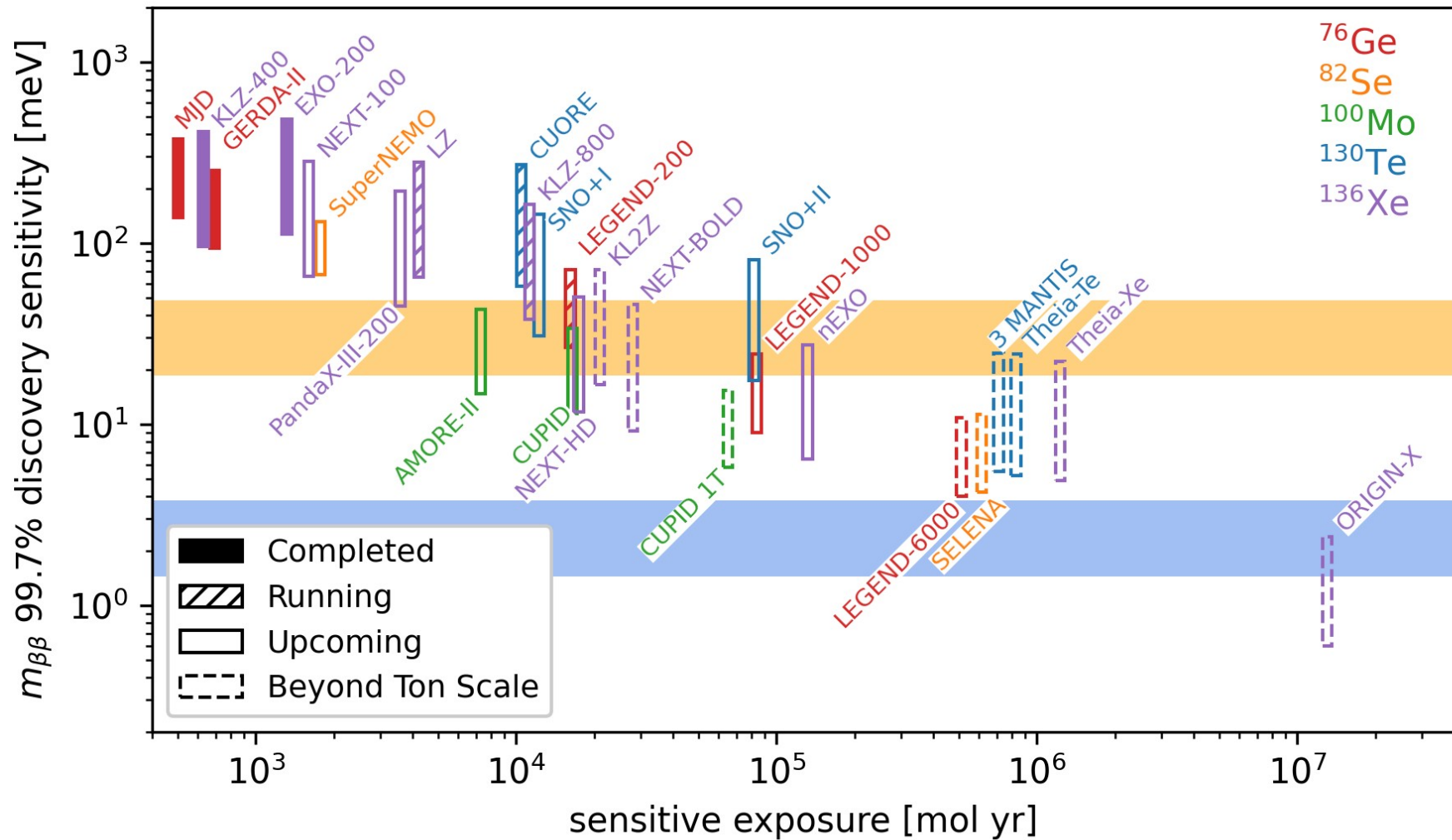
$$5.2 \times 10^{21} \text{ g} \times \frac{\text{mol}}{29 \text{ g}} \times \frac{22.4 \text{ L}}{\text{mol}} \times 87 \times 10^{-9} \frac{\text{L}}{\text{L}} \times \frac{\text{mol}}{22.4 \text{ L}} \times \frac{136 \text{ g}}{\text{mol}} = 2122 \times 10^{12} \text{ g}$$

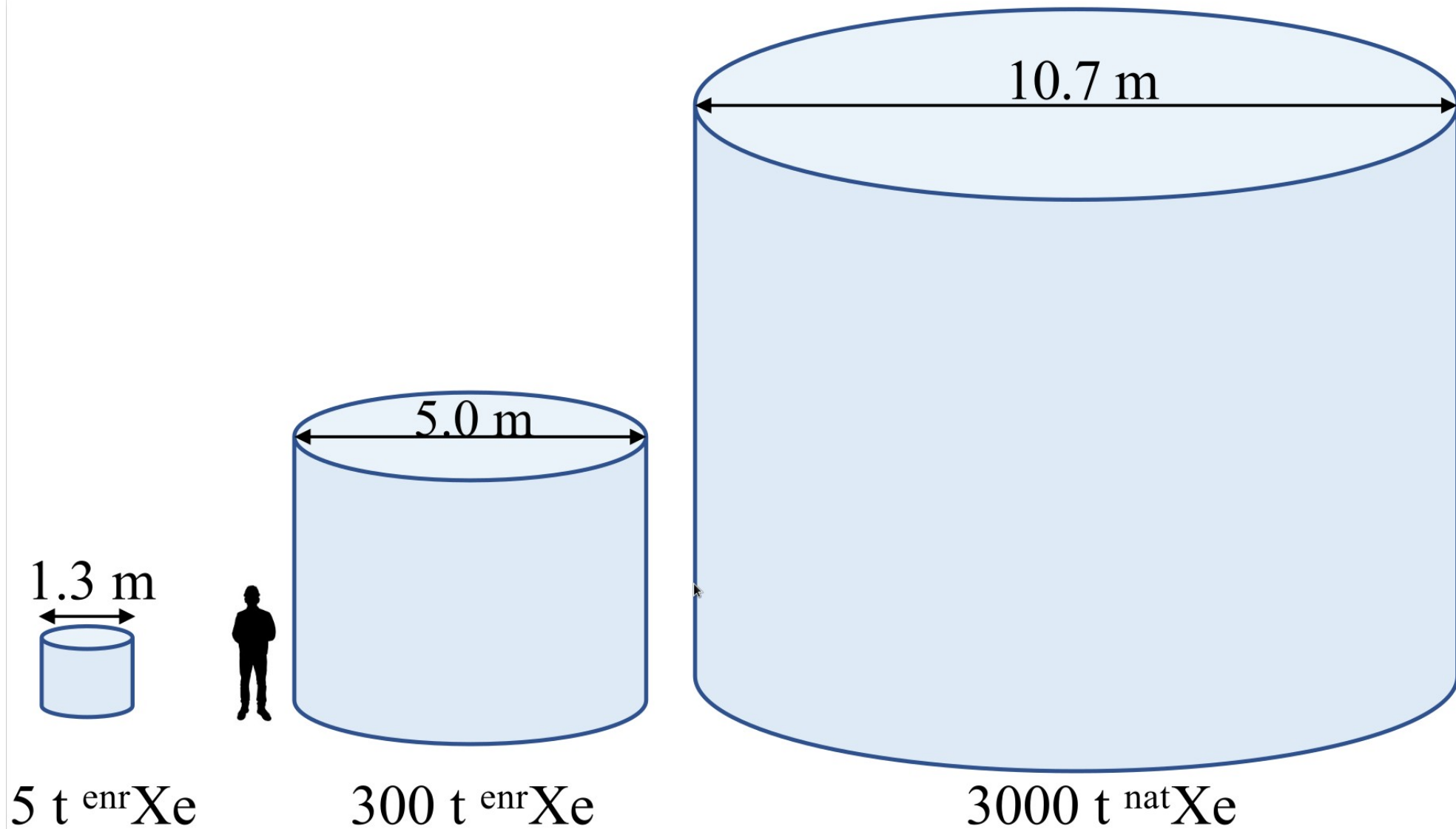
ORIGIN-X

Observing Rare Interactions with a GLaNt Xenon experiment

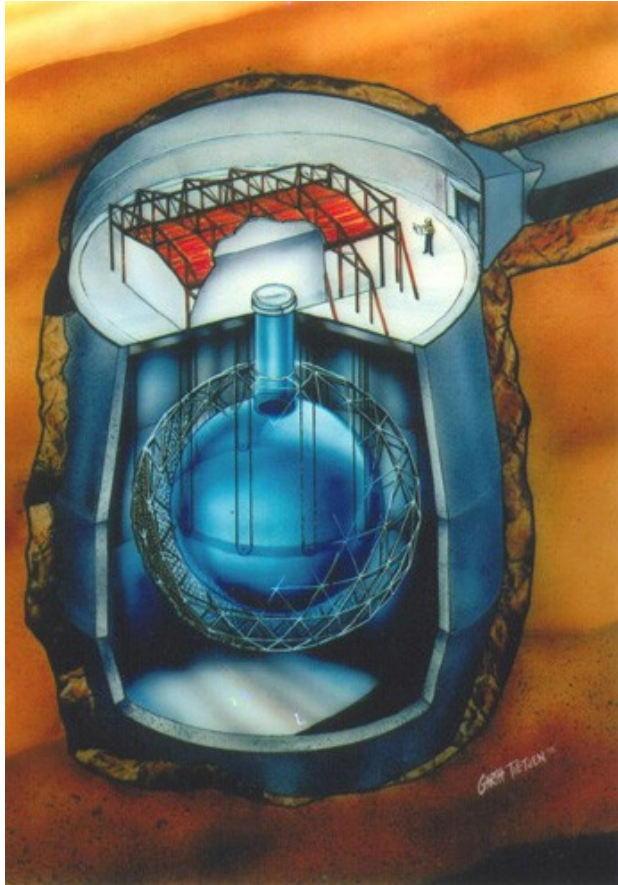


<https://origin-x.lnl.gov/>





Cavern size has been demonstrated



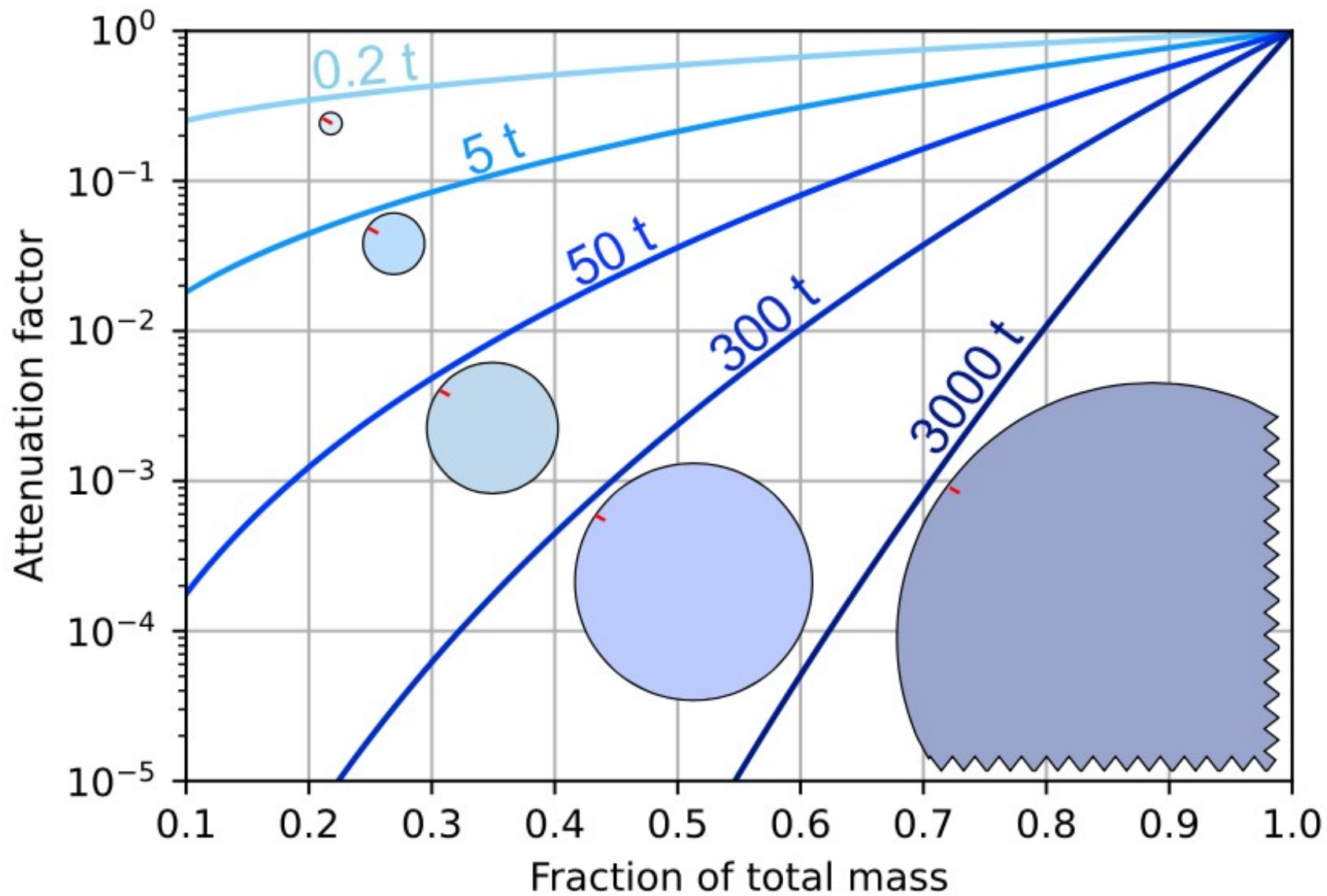
"The detector cavity, 22 metres in diameter and 30 metres high, is the largest cavity at that depth anywhere in the world." – R. Brewer

15m dia. 20m tall

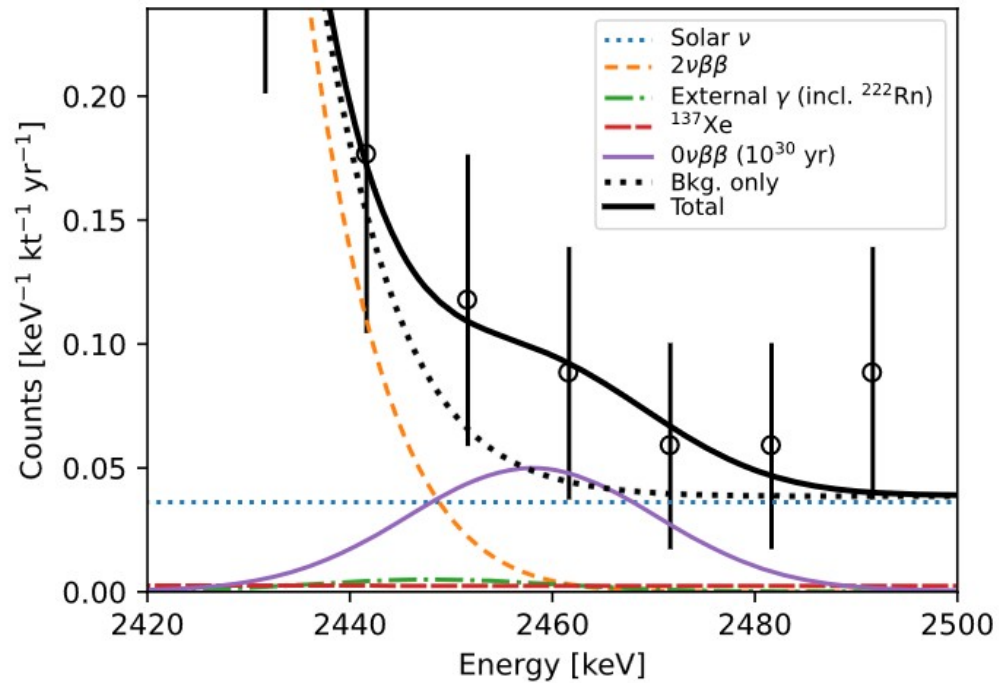
SNO Cavern



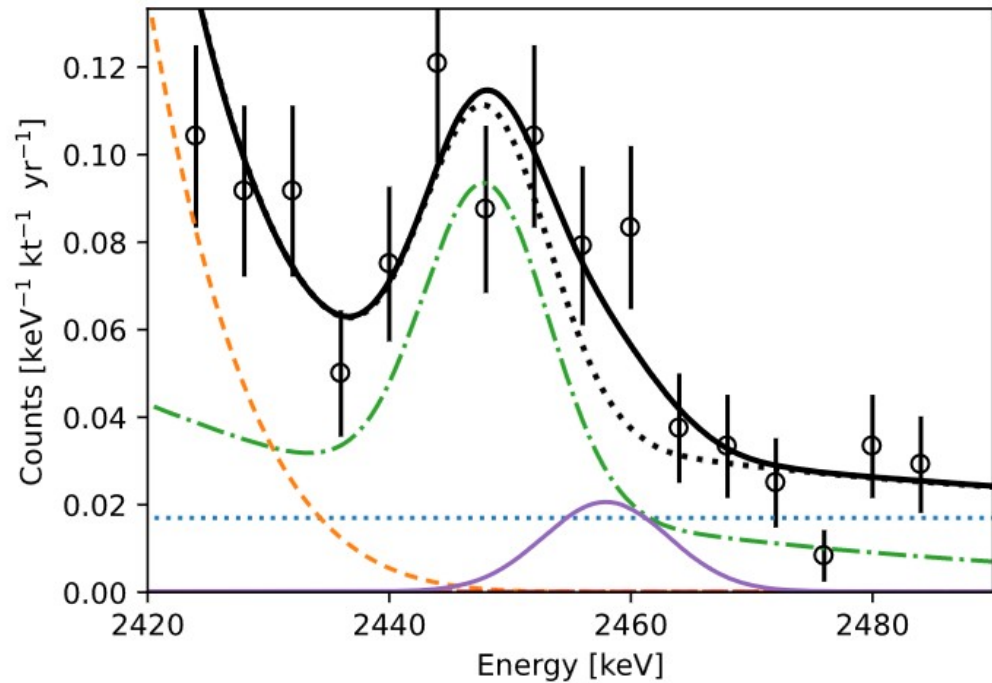
SNOLAB Cryopit

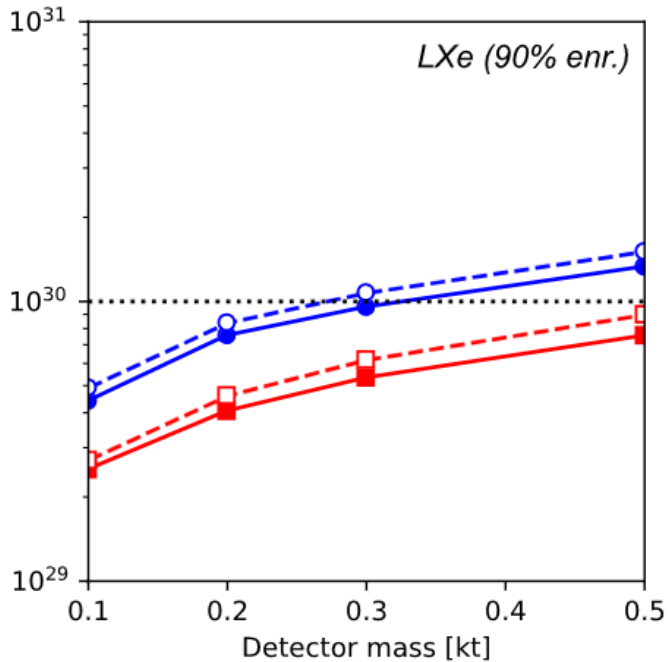
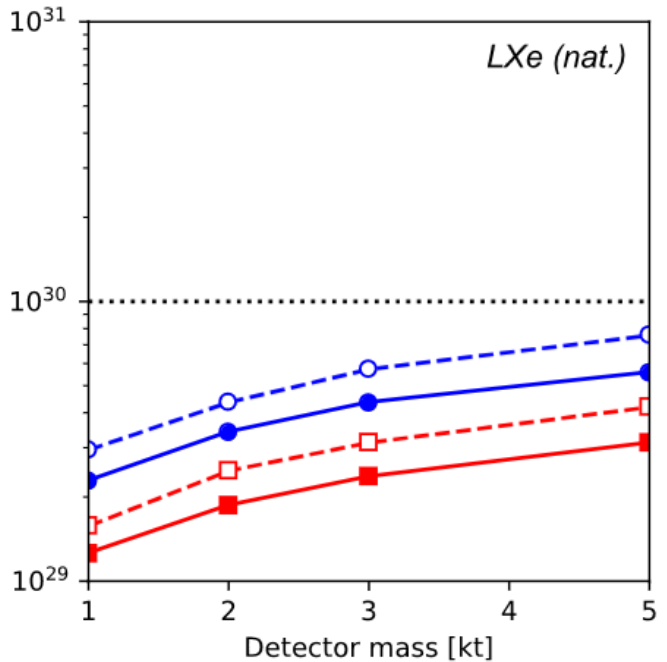
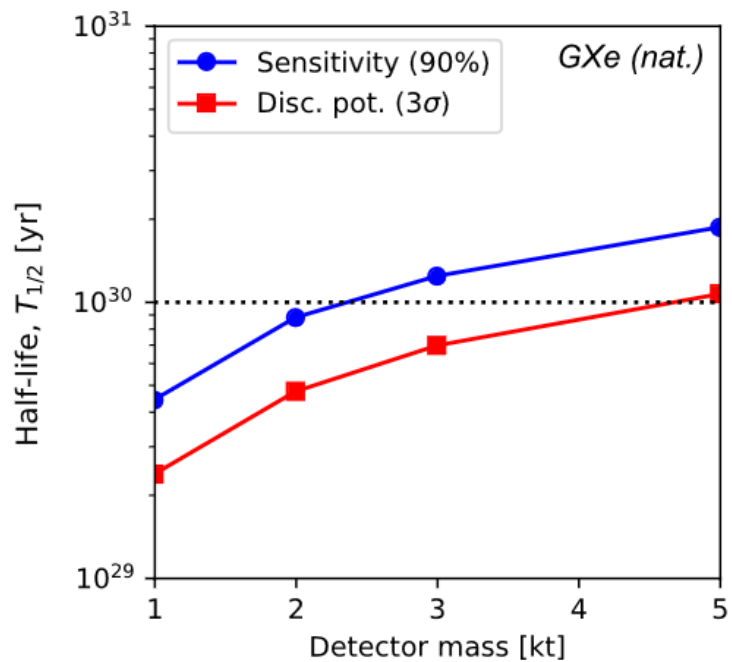


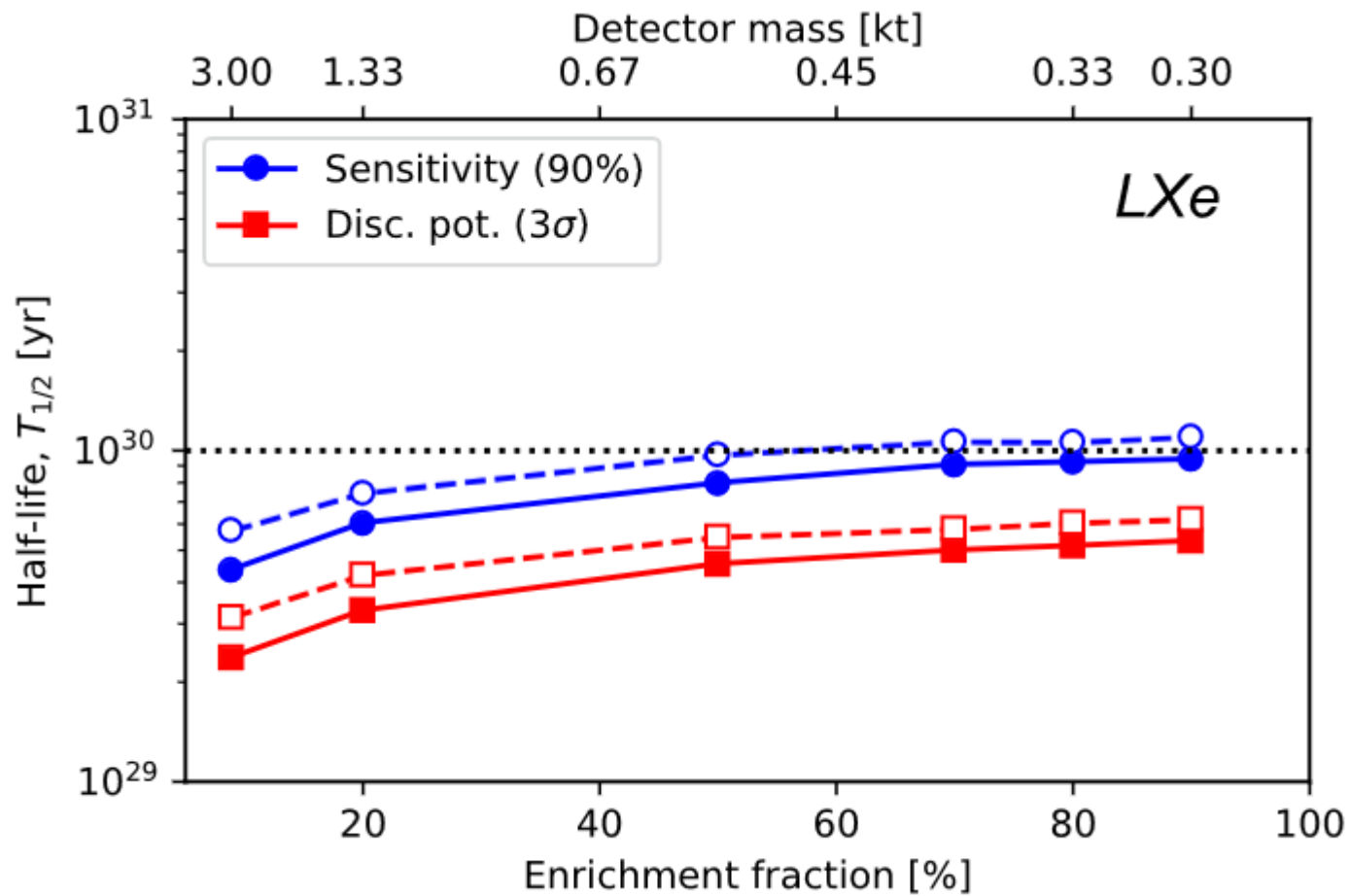
300T enriched Liquid Xenon



3kT natural Gas Xenon







Detector technology		Isotope acquisition	External backgrounds	Internal backgrounds	Energy resolution ($2\nu\beta\beta$)	Isotope mass fraction (solar ν)	Detector technology maturity (kton scale)
Segmented detectors							
	HPGe	?	×	?	✓	✓/!	×
	Bolometers	✓/!	×	?	✓	✓/!	×
	Tracking/CCDs	Se based	✓	?/×	?	✓	×
Monolithic detectors							
	Liquid scintillator	Te doped	✓	✓	?	×	✓
		Xe doped	×	✓	✓	×	✓
	TPCs	Gas Xe	×	✓	✓	✓	✓/!
		Liquid Xe	×	✓	✓	✓/!	✓
		Xe doped Ar	×	✓	×/!	×	✓
		SeF ₆ (ion drift)	✓	✓	?	✓/!	×

Continuing the story in the next talks

Material science R&D
for xenon capture



Extracting xenon from the atmosphere
at low cost and high volume.



Summary

- It is plausible to build a ~ktonne 0vBB detector
- Xenon acquisition is the key challenge
- Adsorption development is likely a breakthrough R&D path (TSA+MOF+structured bed)
- More details in arXiv paper:
<http://arxiv.org/abs/2110.01537>

END