DUNE OPPORTUNTIES

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Thanks to Maria Elena et al. for organizing this It's a great way to see what people are thinking as we go into the Community Summer Meeting in July and bring people together to talk about science

In the following, I will try to:

- Overview of LBNF/DUNE
- Why is the experiment the way it is? Why is it needed?
- Near Detector

LBNF/DUNE OVERVIEW

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Near Site:

1.2 MW neutrino beam

Near Detector Hall

Far Site:

2 detector halls 1 mile underground Houses 4 x 17 kton LArTPCs

Phase 1 Detectors Near Detector (ND)

- LArTPC + muon spectrometer
- Moveable up to 30 m off-axis
- On-axis beam monitor

Far Detectors (FD)

- 1 horizontal drift 17 kton LArTPC, wire readout
- 1 vertical drift 17 kton LArTPC, PCB readout

PHYSICS

$$\begin{pmatrix} \nu_{e} \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_{1} \\ \nu_{2} \\ \nu_{3} \end{pmatrix} \quad \text{What} \quad$$

• What is the magnitude and sign of Δm^2_{32} ?

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- What is the value of δ_{CP} ?
- What are the values of θ_{13} and θ_{23} ?
- m^2 Is this three-flavor " ν -SM" picture correct?

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{22} & s_{22} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & e^{-i\delta_{\rm CP}}s_{13} \\ 0 & 1 & 0 \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \end{pmatrix} 4$$

REQUIREMENTS

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- Highly effective neutrino flavor tagging and energy resolution
- Precise control of systematics \rightarrow near detector

In the first few years of the experiment, goals include:

- Definitive determination of the neutrino mass ordering
- 3 σ sensitivity to CPV in maximal case ($\delta_{CP} = \pm \pi/2$)
- World leading measurement of atmospheric mass splitting
 With the Phase 1 DUNE detector
- 2 x 17 LArTPC Far Detectors
- LArTPC + muon spectrometer

MASS ORDERING





- In ~4 years with conservative beam ramp up scenario, DUNE will definitively determine the mass ordering
- In all but the most "favorable" cases, mass ordering is an essential "gateway" measurement towards CP violation and testing the three flavor ν SM

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NEAR DETECTOR





To measure neutrino oscillation parameters, LB experiments compare:

- Observed energy spectrum of flavor-tagged neutrinos at the far detector
- Prediction as a function of neutrino oscillation parameters (both "signal" and background).

This requires

- "following" neutrinos: production, oscillation, interaction, detection in far detector (FD)
- the measurement is only as good as the prediction

• Systematic errors in the prediction result in degradation in precision/sensitivity Each element is critical in producing the prediction but has large (~10%) a priori uncertainty

SLA

Interactions in LArTPC



ND IN PHASE 1

Phase I ND DUNE MO Sensitivity All Systematics Normal Ordering $\Delta \chi^2$ Years

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- ND includes: LArTPC (ND-LAr) with a muon spectrometer (TMS)
 - Key requirement:
 - Observe neutrino interactions in a LArTPC with comparable performance to FD LArTPC in a high rate environment (~50 interactions in 10 μ s spill)
 - Design features:
 - Sufficiently large to contain hadronic recoil
 - Muons will punch through \rightarrow TMS
 - Pixel readout for native 3D charge representation
 - Segmented readout to facilitate reconstruction in high pileup environment (charge/ light signals)
- ~by design, DUNE will accumulate statistics quickly in FDs and systematics become important even in ~first year

No physics is possible without systematic constraints from ND

SUMMARY:

 LBNF/DUNE has unique capabilities that will make critical measurements of neutrino oscillation parameters within its first few years

- It will also set the foundation for a continuing program of neutrino measurements with unprecedented precision and scope that will test ν SM
- It will also support a broader program that I didn't talk about.
- As the program continues, additional detectors will be needed that offer exciting opportunities for SLAC and opportunities to expand DUNE's physics program
 - Addition Far Detector modules
 - Upgrades to the Near Detector (Please see ND-GAr whitepaper)

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