

# From Anomaly to Problem to Physics: Lessons from Solar Neutrinos

- Overview
- Some Pre-History
- The Development of the ‘Problem’
- The Solution(s)
- What’s Next

# A Familiar Story

Scientists disbelieved....



# A Familiar Story

...until it was too late.



~35 Years

Why so long?

Plenty of good reasons!

But also plenty of bad ones.

Not to mention:

*Ambiguity Principle:*

~35 Years

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Plenty of good reasons!

But also plenty of bad ones.

Not to mention:

*Ambiguity Principle: For any given experimental test of a hypothesis, Nature will always strive to return the most ambiguous answer possible.*

## ~35 Years

### What it wasn't:

A chaotic period in which many results did not make sense until some unifying theory or measurement provided clarity.

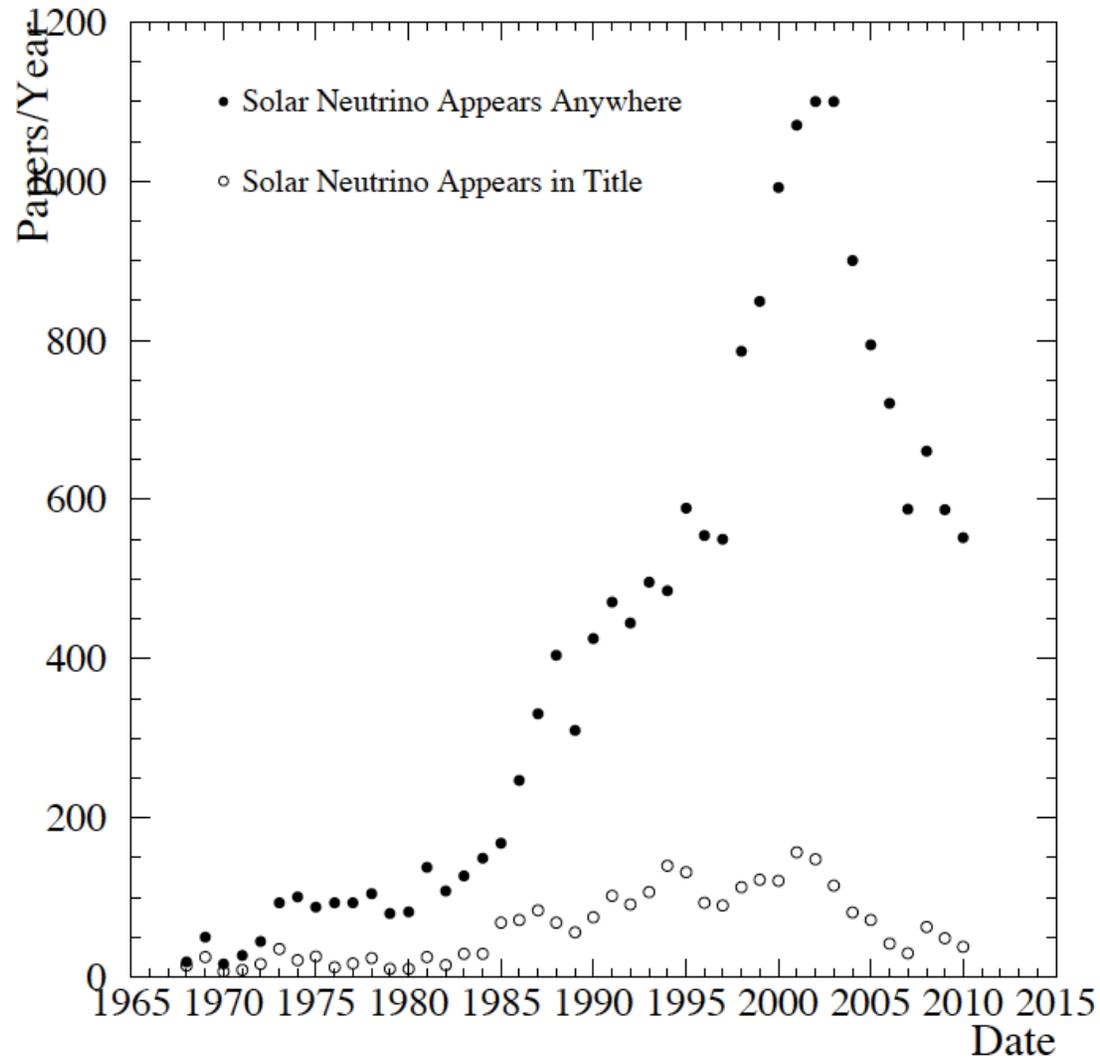
### What it was:

A long, slow, slog in which each new piece—experimental and theoretical--fell into place, with an accompanying marginal increase in 'degree-of-belief.'

Far more 'evolution' than 'revolution'

~35 Years

No really good metric....



## ~35 Years

### The cartoon version:

- **1964:** Bahcall convinces Davis to publish simultaneous papers suggesting a Chlorine-based experiment to test Bahcall's 'Standard Solar Model' predictions.
- **1968:** Davis publishes a flux of neutrinos 1/3 of Bahcall's predicted value. The Sun is broken! Neutrinos oscillate!
- **~20 years pass.** Davis continues dutifully cooking his chlorine. Bahcall becomes the party guest who stays too late.
- **1989:** Kamiokande II publish the first 'realtime' solar neutrino detection, correlated with the Sun's position. But their results disagree with Davis's.
- **1990s:** Gallium experiments publish initially confusing results, then show a deficit.
- **1998:** Super-K sees clear evidence of oscillations in atmospheric neutrinos.
- **2001-2:** SNO unambiguously sees neutral current 'appearance' of oscillated neutrinos.
- **2002:** KamLAND puts the final nail in the coffin by observing disappearance with a terrestrial source.

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# Some Pre-History

1919: Eddington suggests solar fusion:

A star is drawing on some vast reservoir of energy by means unknown to us. This reservoir can scarcely be other than the sub-atomic energy which, it is known, exists abundantly in all matter... The store is well-nigh inexhaustible, if only it could be tapped. There is sufficient [energy stored] in the Sun to maintain its output of heat for 15 billion years....

# Some Pre-History

## 1938: Bethe and stellar energy generation

MARCH 1, 1939

PHYSICAL REVIEW

VOLUME 55

### Energy Production in Stars\*

H. A. BETHE

Cornell University, Ithaca, New York

(Received September 7, 1938)

It is shown that the *most important source of energy in ordinary stars is the reactions of carbon and nitrogen with protons*. These reactions form a cycle in which the original nucleus is reproduced, *viz.*  $C^{12}+H=N^{13}$ ,  $N^{13}=C^{13}+\epsilon^+$ ,  $C^{13}+H=N^{14}$ ,  $N^{14}+H=O^{15}$ ,  $O^{15}=N^{15}+\epsilon^+$ ,  $N^{15}+H=C^{12}+He^4$ . Thus carbon and nitrogen merely serve as catalysts for the combination of four protons (and two electrons) into an  $\alpha$ -particle (§7).

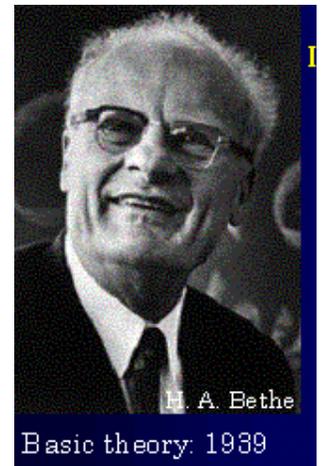
integration of the Eddington equations gives 19. For the brilliant star Y Cygni the corresponding figures are 30 and 32. This good agreement holds for all bright stars of the main sequence, but, of course, not for giants.

For fainter stars, with lower central temperatures, the reaction  $H+H=D+\epsilon^+$  and the reactions following it, are believed to be mainly responsible for the energy production. (§10)

**B<sup>8</sup>**

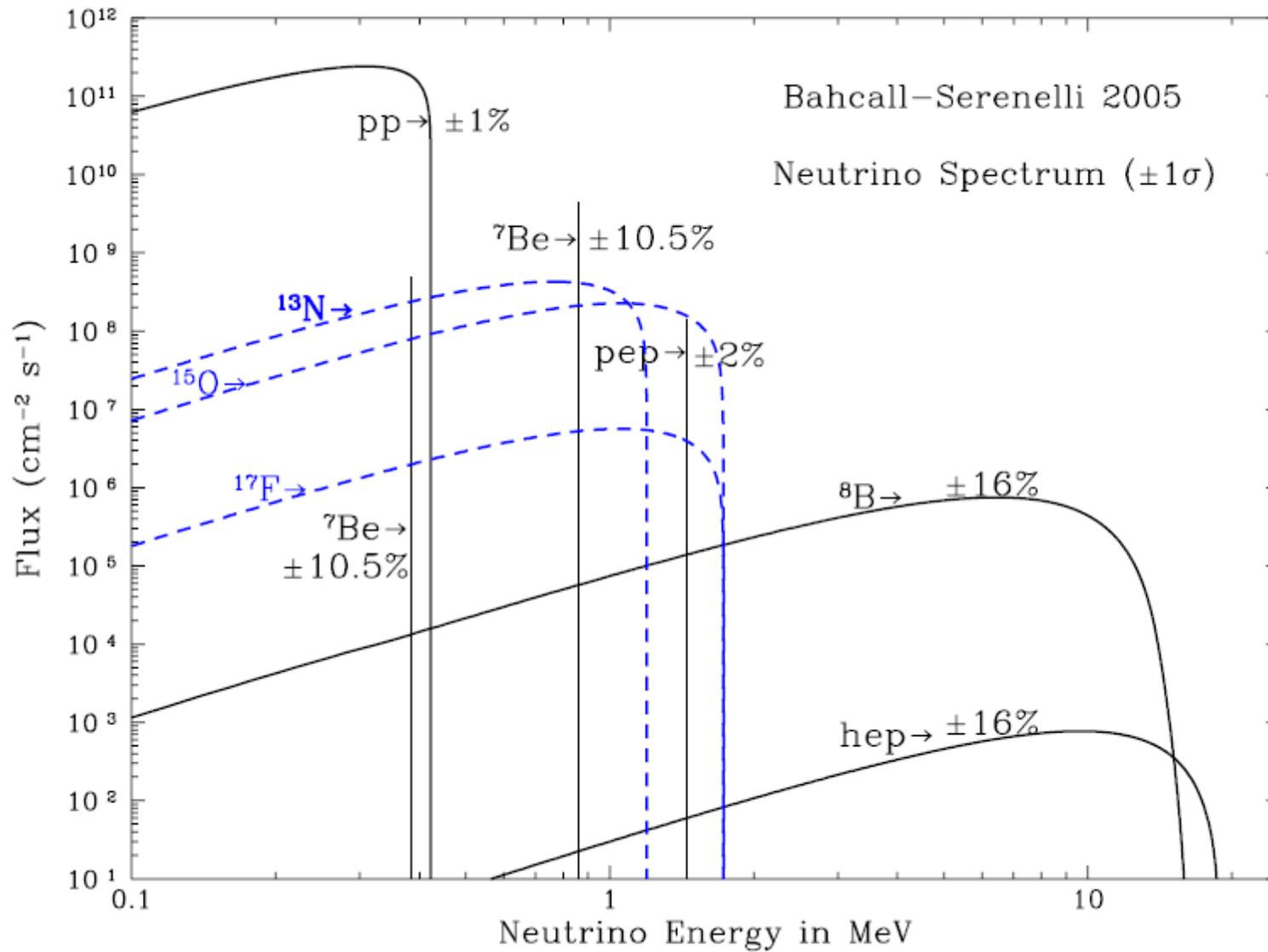
The existence is doubtful; calculation<sup>21</sup> by comparison with its isobar Li<sup>8</sup> gives a binding energy between  $-0.4$  and  $+0.4$  mMU. This nucleus is not very important for astrophysics.

Prof. Bethe will redeem himself nearly 50 years later...





# (The Modern View)



## Some Pre-History

1946: Pontecorvo suggests neutrino detection using Chlorine:



Subsequent detection of the decay of  ${}^{37}\text{Ar}$  with a  $\sim 35$  day half-life

Suggests Sun as a neutrino source, but points out that the energy from pp and CNO neutrinos is too low. Suggests reactors instead.

(His paper is classified; fear is that it could be used to detect nuclear submarines...).

1949: Alvarez discusses the same idea, with details about background rejection.

# Some Pre-History

1954-58: Davis tried Chlorine experiment with reactor and sets limits on solar flux...most sensitive neutrino experiment to date.

PHYSICAL REVIEW

VOLUME 97, NUMBER 3

FEBRUARY 1, 1955

## Attempt to Detect the Antineutrinos from a Nuclear Reactor by the $\text{Cl}^{37}(\bar{\nu}, e^-)\text{A}^{37}$ Reaction\*

RAYMOND DAVIS, JR.

*Department of Chemistry, Brookhaven National Laboratory, Upton, Long Island, New York*

(Received September 21, 1954)

Tanks containing 200 and 3900 liters of carbon tetrachloride were irradiated outside of the shield of the Brookhaven reactor in an attempt to induce the reaction  $\text{Cl}^{37}(\bar{\nu}, e^-)\text{A}^{37}$  with fission product antineutrinos. The experiments serve to place an upper limit on the antineutrino capture cross section for the reaction of  $2 \times 10^{-42}$  cm<sup>2</sup> per atom. Cosmic-ray-induced  $\text{A}^{37}$  was observed and the production rate measured at 14 100 feet altitude and sea level. Measurements with the 3900-liter container shielded from cosmic rays with 19 feet of earth permit placing an upper limit on the neutrino flux from the sun.

$$\phi_{\nu} < 1 \times 10^{14} / \text{cm}^2\text{-s} \text{ assuming CNO neutrinos}$$

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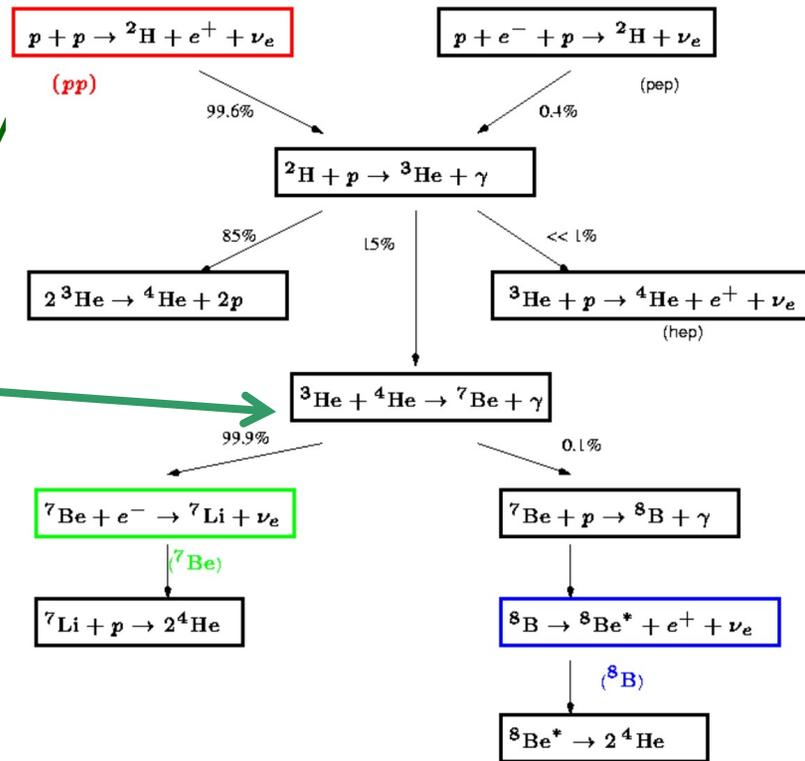
### Referee comment:

Any experiment such as this, which does not have the requisite sensitivity, really has no bearing on the question of the existence of neutrinos. To illustrate my point, one would not write a scientific paper describing an experiment in which an experimenter stood on a mountain and reached for the moon, and concluded that the moon was more than eight feet from the top of the mountain.

1958

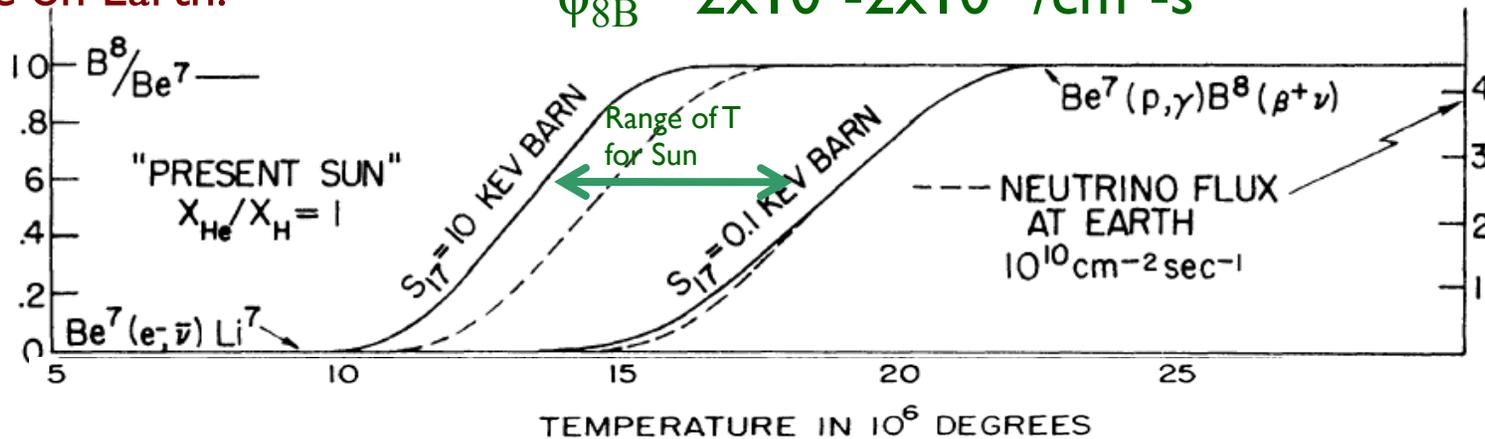
# Some Pre-History

Breakthrough: Holmgren and Johnston measure a surprisingly large value for the  ${}^3\text{He}+{}^4\text{He}$  reaction creating  ${}^7\text{Be}$ :



Fowler, Cameron independently point out that this means 'high-energy' (14 MeV endpoint)  ${}^8\text{B}$  vs are produced in the Sun and possibly detectable on Earth:

$$\phi_{8\text{B}} \sim 2 \times 10^9 - 2 \times 10^{10} / \text{cm}^2 \cdot \text{s}$$



# The Twin Papers

VOLUME 12, NUMBER 11

PHYSICAL REVIEW LETTERS

16 MARCH 1964

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## SOLAR NEUTRINOS. I. THEORETICAL\*

John N. Bahcall

California Institute of Technology, Pasadena, California  
(Received 6 January 1964)

## SOLAR NEUTRINOS. II. EXPERIMENTAL\*

Raymond Davis, Jr.

Chemistry Department, Brookhaven National Laboratory, Upton, New York  
(Received 6 January 1964)

## What was new here?

1. Bahcall re-calculates cross section on chlorine---18 times larger for  $^8\text{B}$  vs than previously believed.  $\phi_{8\text{B}} \sim 2.5 \times 10^7 / \text{cm}^2\text{-s}$ , based primarily on models of other authors.
2. The goal is an explicit, unambiguous test of the solar fusion hypothesis:  

the star. Only neutrinos, with their extremely small interaction cross sections, can enable us to see into the interior of a star and thus verify directly the hypothesis of nuclear energy generation in stars.
3. Davis's earlier limits used to limit central core temp.

# The Experiment

Goldhaber convinced at Brookhaven to move forward, perhaps because result would help prove “astrophysicists do not know what they’re doing.”

Particle physicist’s view  
of astronomers, ca. 1964

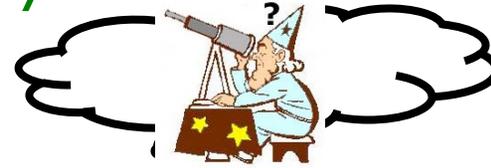


Dodson, in negotiations with the AEC for funding, writes Fowler to ask,

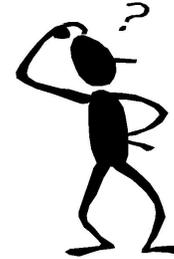
“Why spend a substantial sum trying to measure something which is calculated with great confidence by nuclear astrophysicists - and who cares about confirming the central temperature of the sun anyway?”

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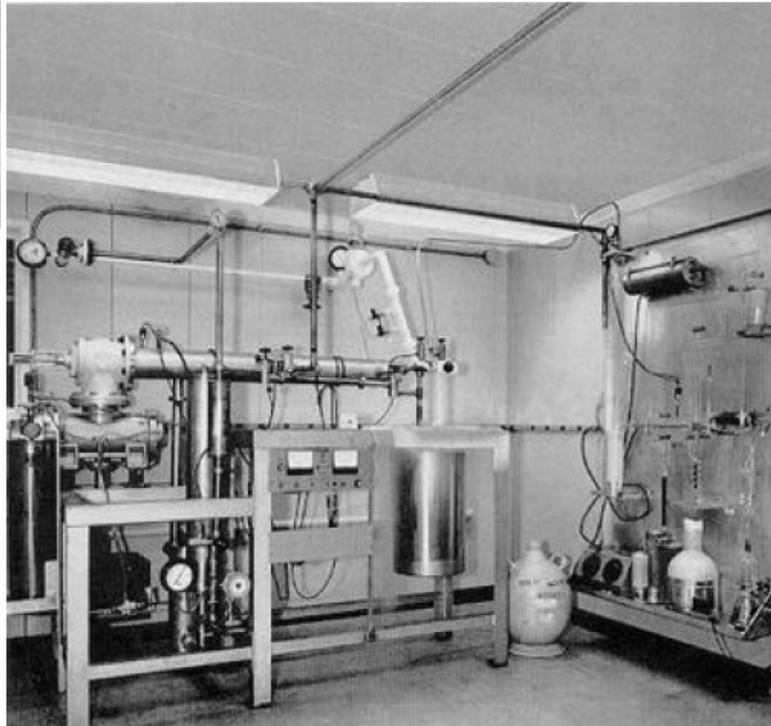
“Why spend a substantial sum trying to measure something which is calculated with great confidence by nuclear astrophysicists - and who cares about confirming the central temperature of the sun anyway?”

# The Experiment

Nevertheless...

~600 tons of  $C_2Cl_4$  at newly excavated cavity at 4850 ft level of Homestake.

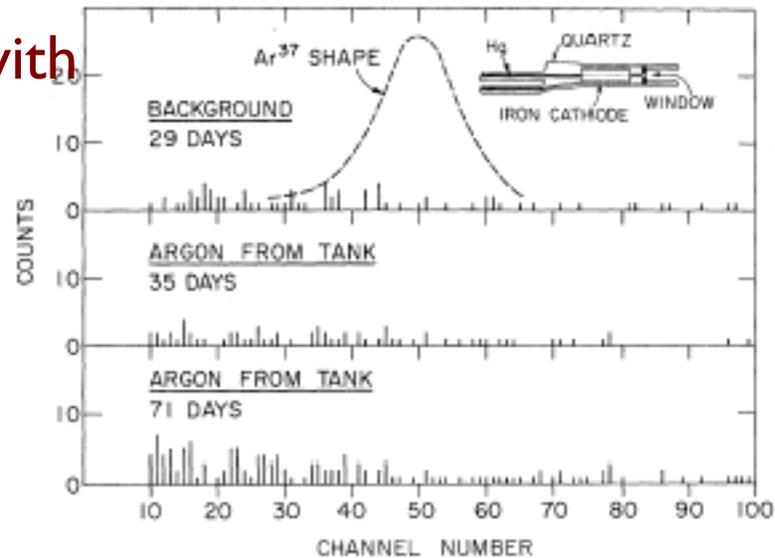
\$600k = \$5.7M in 2022, before accounting for modern safety, project management, outreach center....



Davis:  
“Just plumbing”

# The Results

All counts consistent with cosmogenic and other backgrounds.



The discovery of  $\nu$  oscillations?

FIG. 2. Pulse-height spectra.

*PRL, 20, 1205, (1968)*

Gives limit on  $^8\text{B}$  flux  $\phi_{8\text{B}} < 2 \times 10^6 / \text{cm}^2\text{-s}$

Bahcall, Bahcall, and Shaviv:

When accounting for uncertainties, "...not in obvious conflict with theory of stellar structure." Give best estimate of  $\phi_{8\text{B}} = 4.7 \times 10^6 / \text{cm}^2\text{-s}$ , with large uncertainties.

# Is there a Problem Here?

“Nice try”?

A match between prediction and experiment required:

- That solar model neutrino predictions were correct.

Relied on many keV-energy nuclear physics cross sections, assumptions about the solar core, and made no other testable prediction than the neutrino flux.

Over 30 years,  $\phi_{8B}$  went from 0 (Bethe) to  $10^{10}$  (Fowler, Cameron) to  $3 \times 10^7$  (Sears) to  $2 \times 10^7$  (Bahcall) to  $4.7 \times 10^6$  (B,B,S) and scaled as  $\sim T^{25}$ .

- That Bahcall's cross section calculation was correct.

Did make prediction of  $^{37}\text{Ca}$  state that was later discovered, but no other way of measuring this.

- That Davis and collaborators could possibly do what they claimed.

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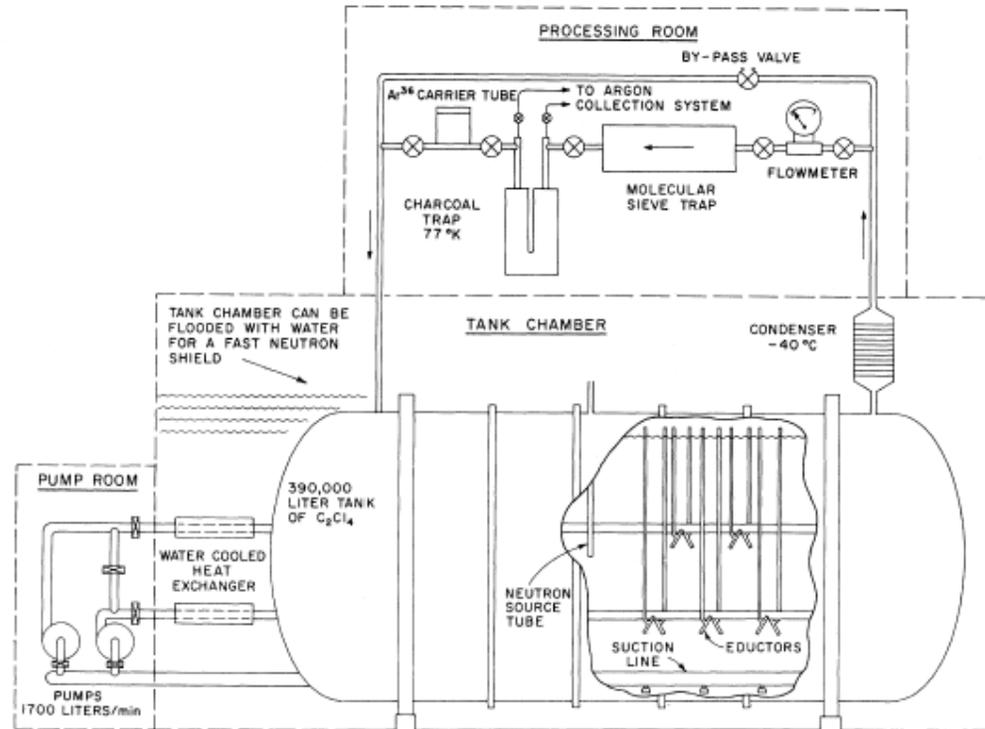
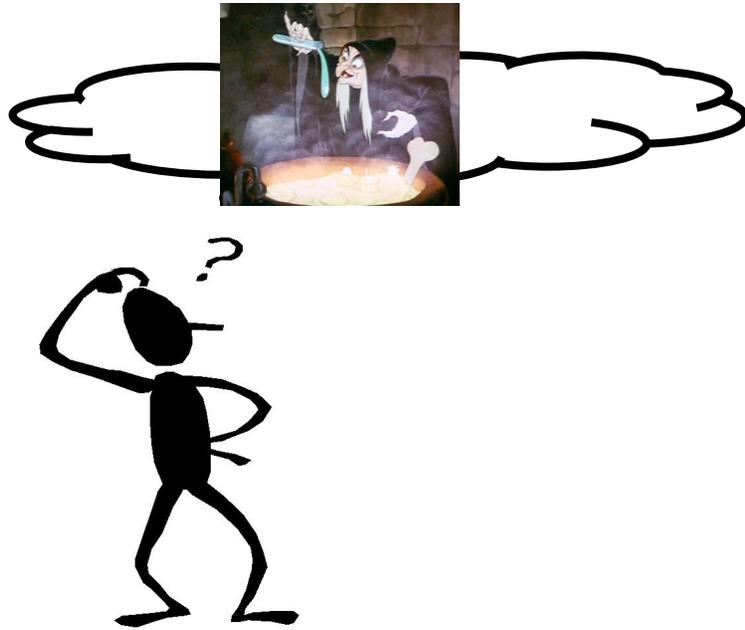
✓ Believability tip: Don't base anything on anything that depends on *anything* to the 25<sup>th</sup> power.

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# Is there a Problem Here?



To a physicist, Davis's method is incredibly complex, as he pulls the 12 atoms or so of  $^{37}\text{Ar}$  from the 600 tons of fluid in each purge, and effectively shepherds them to his counters.

No events, no cuts, no simulation, no fit to anything.

But: a serious effort to determine efficiencies is described.

# Is there a Problem Here?

Yet Davis never publishes a CI result again in PRL or PR.  
30 years pass before any results from the experiment again published in a refereed journal (ApJ). First non-zero value was in a proceedings, first significant observation in an APS Bulletin → basis for the Nobel Prize.

MU-52

Proc. 11th Int. Conf. on Cosmic Rays, Budapest 1969

## A PROGRESS REPORT ON THE BROOKHAVEN SOLAR NEUTRINO EXPERIMENT

R. DAVIS, JR.

Brookhaven National Laboratory, Upton, New York 11973, U.S.A.

The present report gives the results of an experimental search for neutrino radiation from the sun, using the neutrino capture reaction:  $\text{Cl}^{37}(\nu, e^-)\text{Ar}^{37}$ . The number of  $\text{Ar}^{37}$  decays gives a neutrino capture rate  $(2.5 \pm 1.4) \times 10^{-36} \text{ sec}^{-1}$  per terrestrial  $\text{Cl}^{37}$  atom confirming the upper limit  $3 \times 10^{-36} \text{ sec}^{-1}$  previously reported.

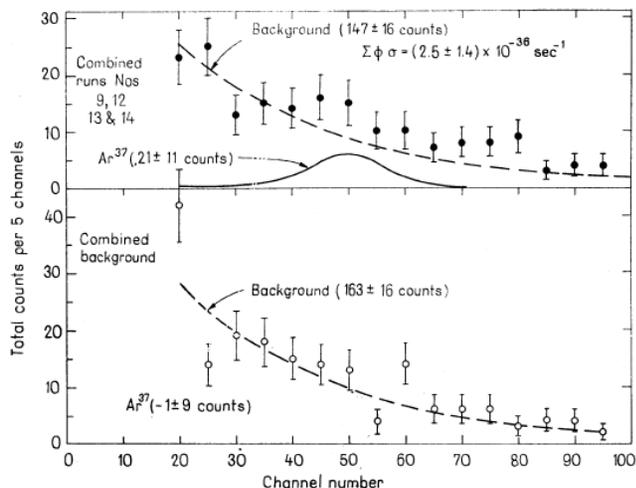


Fig. 2. The sum of counts in each five channels

## Monitor New Scientist, 1971

### The first observation of neutrinos from the Sun

After four years of effort, Raymond Davis, Jr. and his colleagues from Brookhaven National Laboratory have finally detected neutrinos which originate from the centre of the Sun. All previous measurements on the Sun give us information about its surface only. The Sun's interior is only accessible through neutrino measurements. Latest results from the Brookhaven experiment are the first direct evidence that nuclear transmutations are the internal source of energy for the Sun (Bulletin of the American Physical Society, vol 16, p 521).

measured all of the accessible reaction rates and can find no easy explanation there for the small discrepancy in the neutrino results. Yet Fowler is confident that nothing is grossly out of order.

One definite conclusion from Davis' results is the dominance of the proton-proton chain over the carbon-nitrogen cycle for solar energy. If the carbon-nitrogen cycle predominated in the Sun, the neutrino flux would have been 30 times the observed value. The neutrino experiment serves to confirm all of the terrestrial nuclear physics experiments and

✓ Believability tip: Present your biggest results in refereed journals, not proceedings, popular articles, or elevators.



# Is there a Problem Here?

A 'flurry' of papers---about seven in 4 years, three from Bahcall--- suggest possible 'solutions', among them is Gribov and Pontecorvo's suggestion of neutrino oscillations.

## NEUTRINO ASTRONOMY AND LEPTON CHARGE

V. GRIBOV\* and B. PONTECORVO  
*Joint Institute for Nuclear Research, Dubna, USSR*

Received 20 December 1968

It is shown that lepton nonconservation might lead to a decrease in the number of detectable solar neutrinos at the earth surface, because of  $\nu_e \rightleftharpoons \nu_\mu$  oscillations, similar to  $K^0 \rightleftharpoons \bar{K}^0$  oscillations. Equations are presented describing such oscillations for the case when there exist only four neutrino states.

“From the point of view of detection possibilities, an ideal object is the sun...Unfortunately, the weight of the various thermonuclear reactions in the sun, and the central temperature of the sun, are insufficiently well known in order to allow a useful comparison of expected and observed solar neutrinos, from the point of view of this article.”

They effectively assume maximal mixing. In modern terms:

$$\begin{aligned} |\nu_e\rangle &= \cos\theta|\nu_1\rangle + \sin\theta|\nu_2\rangle \\ |\nu_\mu\rangle &= -\sin\theta|\nu_1\rangle + \cos\theta|\nu_2\rangle \end{aligned} \quad \longrightarrow \quad P_{\nu_e \rightarrow \nu_e} = 1 - \sin^2 2\theta_{12} \sin^2\left(\frac{1.27\Delta m^2 L}{E}\right)$$

$$\text{For } \Delta m^2 \gg E/L, P_{\nu_e \rightarrow \nu_e} \rightarrow \frac{1}{2} \sin^2 2\theta \text{ ('washed out')}$$

$$\text{For } \Delta m^2 \ll E/L, P_{\nu_e \rightarrow \nu_e} \rightarrow 1 \text{ (no sensitivity)}$$

# Is there a Problem Here?

Also:  $\nu$  magnetic moments (Cisneros),  $\nu$  decay (Bahcall et al.), mixing of solar material (Ezer and Cameron), solar helium abundance (Iben), the solar cycle (Sheldon) and uncertainties on the now-named 'Standard Solar Model' (Bahcall et al.).

Bahcall invents the SNU(= $10^{-36}$  captures/target-s).

✓ Believability tip: Don't invent your own unit.

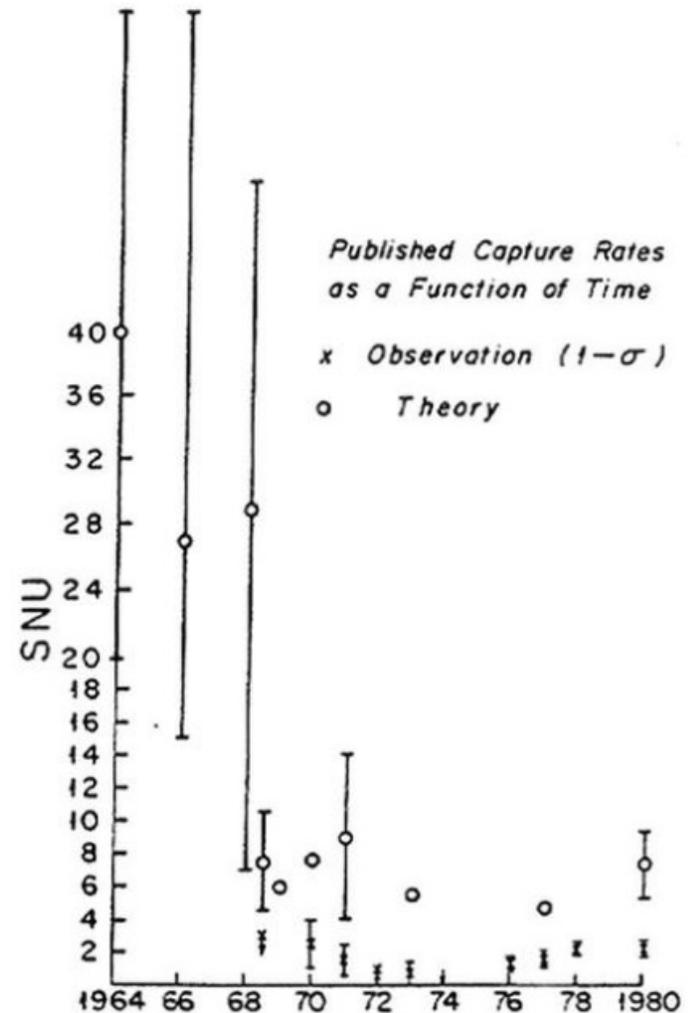
# The Wilderness

New Scientist, 1971

The number of argon-37 decays recorded by the Brookhaven team averaged about 0.5 counts per day with a statistical uncertainty of 0.2 counts. They calculate a corrected solar neutrino capture rate per chlorine-37 atom of  $(1.5 \pm 1.0) \times 10^{-36}$  per second. This is the first positive signal to emerge above the background. It can be compared with the current theoretical calculations of  $10^{-35}$  per second. Although slightly low, the experimental result probably will not cause any serious problems.

measured all of the accessible reaction rates and can find no easy explanation there for the small discrepancy in the neutrino results. Yet Fowler is confident that nothing is grossly out of order.

More data from Cl, but no one plans to redo the experiment.



New experiments are suggested;  
Davis's is the last solar  $\nu$  experiment to  
ever run in the US.

# The Wilderness

Occasionally, the honest feelings slip out:

## NEUTRINO MOMENTS, MASSES AND CUSTODIAL SU(2) SYMMETRY\*

Howard GEORGI and Michael LUKE

*Lyman Laboratory of Physics, Harvard University, Cambridge, MA 02138, USA*

### 1. The problem

Most likely, the solar neutrino problem [1] has nothing whatever to do with particle physics. It is a great triumph that astrophysicists are able to predict the number of  $B^8$  neutrinos coming from the sun as well as they do, to within a factor of 2 or 3 [2]. However, one aspect of the solar neutrino data, the apparent

# The Wilderness

Overheard in the Ludwig A. Wittgenstein Memorial Library:

# The Wilderness

Overheard in the Ludwig A. Wittgenstein Memorial Library:

“Yeah, as if they know what is going on inside the Sun”  
“Sounds like an `observational` problem to me”  
“Like we’re exactly the right distance from the Sun?”  
“How do they even know they’re from the Sun?”  
“Even so, the mixing angles have to be small, not large”  
“Never believe a disappearance experiment”  
“It’s—it’s—it’s chemistry!”  
“What do you expect? There’s no calibration.”

The less informed one was, the more likely to assume the entire thing was nonsense.

# Theorists Help Out

## *Oscillations in Matter (MSW Effect)*

Wolfenstein (1978) points out birefringence of neutrinos within the Sun, from coherent forward scattering in  $\nu_e$  CC reactions:

$$\sigma(\nu_{\mu,\tau}) = 0.15\sigma(\nu_e)$$

**All neutrino flavors**
**Only electron neutrinos**

$$\langle \nu_e | H_W | \nu_e \rangle = \sqrt{2} G_F N_e$$

$$\tilde{H} = \tilde{H}_f + \tilde{H}_W$$

**Bulk matter just treated as a potential term!**

# Theorists Help Out

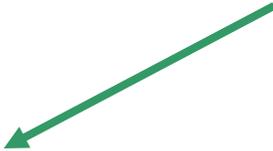
## *Oscillations in Matter (MSW Effect)*

If we calculate mixing matrix elements in flavor basis:

$$|\nu_1\rangle = \sin\theta|\nu_e\rangle + \cos\theta|\nu_\mu\rangle$$

$$|\nu_2\rangle = \cos\theta|\nu_e\rangle - \sin\theta|\nu_\mu\rangle$$

$$\begin{aligned}\langle\nu_1|H|\nu_1\rangle &= \langle\nu_1|H_f|\nu_1\rangle + \langle\nu_1|H_W|\nu_e\cos\theta - \nu_\mu\sin\theta\rangle \\ &= E_1 + \sqrt{2}G_f N_e \cos\theta\end{aligned}$$


$$i\frac{\partial}{\partial t}\begin{pmatrix}\nu_e \\ \nu_\mu\end{pmatrix} = \begin{pmatrix}\sqrt{2}G_f N_e - \frac{\Delta m^2}{2p}\cos^2\theta & \frac{\Delta m^2}{4p}\sin 2\theta \\ \frac{\Delta m^2}{4p}\sin 2\theta & -\frac{\Delta m^2}{2p}\sin^2\theta\end{pmatrix}\begin{pmatrix}\nu_1 \\ \nu_2\end{pmatrix}$$

# Theorists Help Out

## *The Standard Scenario in Matter (MSW)*

Hamiltonian matrix now has new 'matter' eigenvalues and –vectors:

$$i \frac{\partial}{\partial t} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \sqrt{2} G_f N_e - \frac{\Delta m^2}{2p} \cos^2 \theta & \frac{\Delta m^2}{4p} \sin 2\theta \\ \frac{\Delta m^2}{4p} \sin 2\theta & -\frac{\Delta m^2}{2p} \sin^2 \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$|\nu_{1m}\rangle = \cos \theta_m |\nu_e\rangle - \sin \theta_m |\nu_\mu\rangle$$

$$|\nu_{2m}\rangle = \sin \theta_m |\nu_e\rangle + \cos \theta_m |\nu_\mu\rangle$$

Which evolve again as  $P(E_{\nu_e}, x, \theta, \Delta m^2) = 1 - \sin^2 \theta_m \sin^2 \frac{\pi x}{L_m}$

But with

$$\tan 2\theta_m = \frac{\frac{\Delta m^2}{2E} \sin 2\theta}{\frac{\Delta m^2}{2E} \cos 2\theta - \sqrt{2} G_F N_e}$$

Depends on sign of  $\Delta m^2$



# Theorists Help Out

Mixing angle and oscillation length in matter are different from vacuum:

$$\tan 2\theta_m = \frac{\frac{\Delta m^2}{2E} \sin 2\theta}{\frac{\Delta m^2}{2E} \cos 2\theta - \sqrt{2} G_F N_e} \quad l_m = \frac{2\pi}{\sqrt{\left(\frac{\Delta m^2}{2E} \cos 2\theta - \sqrt{2} G_F N_e\right)^2 + \left(\frac{\Delta m^2}{2E}\right)^2 \sin^2 2\theta}}$$

But Wolfenstein concludes:

In general, if one is considering the possibility of large vacuum oscillation lengths, as in the discussion of solar neutrinos, the oscillations should be calculated for the actual vacuum path<sup>16</sup> ignoring the passage through matter. Thus, in the detailed solar neutrino calculations<sup>10</sup> the effective distance over which neutrino oscillations take place is from the solar surface to the earth's surface; there are no significant oscillations inside the sun or in traversals through the earth.

If you read just the abstract (about flavor-changing neutral currents) and the conclusion, you wouldn't have thought much more about the paper.

# Theorists Help Out

Eight years later...Mikheyev and Smirnov point out that

$$\tan 2\theta_m = \frac{\frac{\Delta m^2}{2E} \sin 2\theta}{\frac{\Delta m^2}{2E} \cos 2\theta - \sqrt{2} G_F N_e} \quad \xrightarrow{\text{Resonance when}} \quad \sqrt{2} G_F N_e - \frac{\Delta m^2}{2E} \cos 2\theta = 0$$

(Eight years? Really??)

MS point out that the resonance condition can be satisfied either by the broad energy spectrum, or by the varying matter density in the Sun.

Bethe publishes MS ideas in a PRL:

VOLUME 56, NUMBER 12

PHYSICAL REVIEW LETTERS

24 MARCH 1986

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## Possible Explanation of the Solar-Neutrino Puzzle

H. A. Bethe

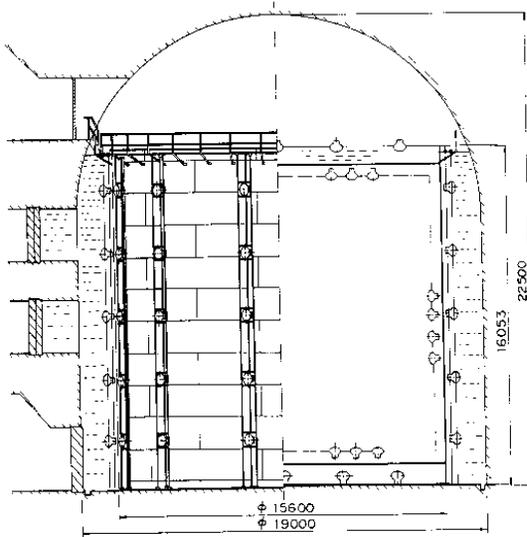
Biggest point perhaps that Langacker found a sign error carried over from one of VV's papers that would have required  $m_2 < m_1$ . But maybe this was really just good P-R?

# Theorists Help Out

**Irony 1:** The excitement over MSW was in large part because it provided a way to get large  $\nu_e$  suppression out of small mixing angles.

**Irony 2:** The sign error pointed out by Langacker reassured people that the MSW effect was a real possibility, because of course the hierarchy had to 'look like' the charged leptons (i.e., 'normal', not 'inverted').

# From Out of the East

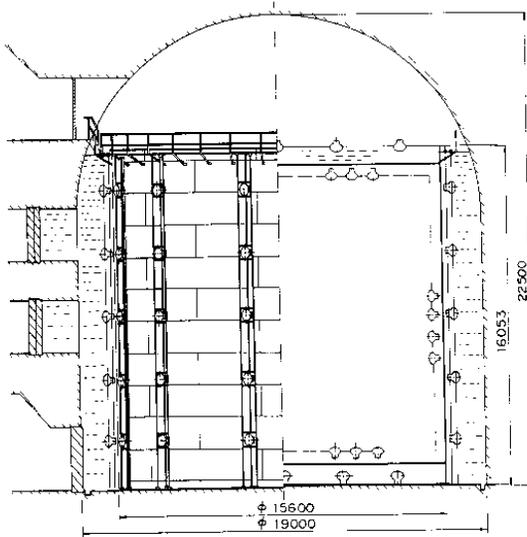


Kamiokande I designed with threshold of  $\sim 30$  MeV, for a search for proton decay.

Penn group joined, upgraded electronics, for threshold  $< 10$  MeV, making detection of  ${}^8\text{B}$  vs via  $\nu + e \rightarrow \nu + e$  with Cherenkov light possible (Kamiokande II).

Beier:

# From Out of the East



Kamiokande I designed with threshold of  $\sim 30$  MeV, for a search for proton decay.

Penn group joined, upgraded electronics, for threshold  $< 10$  MeV, making detection of  ${}^8\text{B}$  vs via  $\nu + e \rightarrow \nu + e$  with Cherenkov light possible (Kamiokande II).

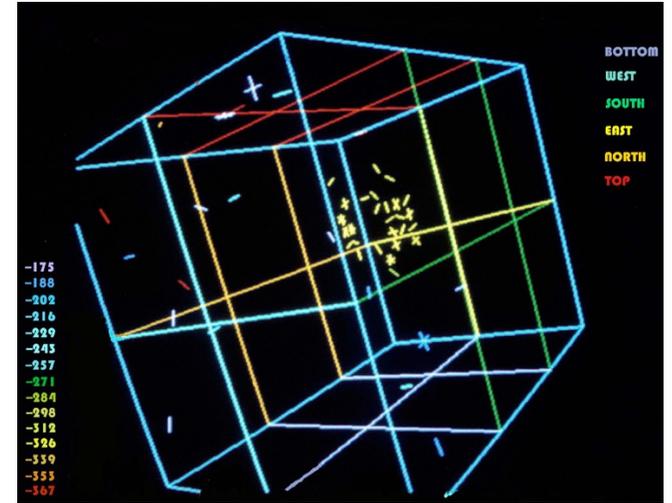
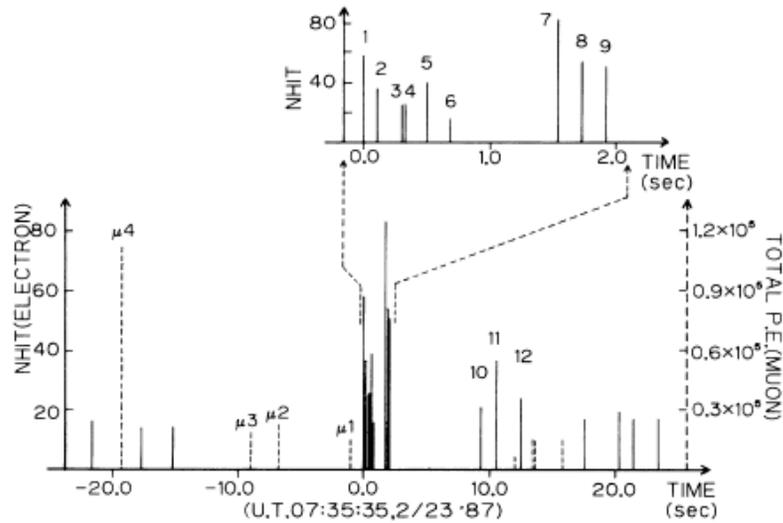
Beier: “Seemed interesting.”

But would anyone believe another crazy cosmic ray experiment, particularly one with a (mostly) unknown Japanese collaboration?

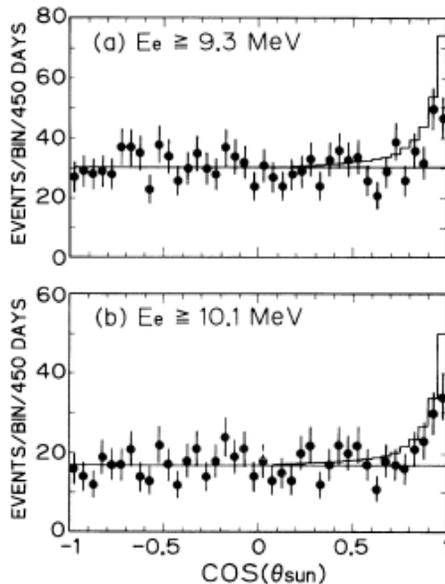
# From Out of the East

Fortunately...

**SNI 1987A**  
establishes  
KII 'creds'



Two years later first solar neutrino observation:



First demonstration vs are from the Sun;

$$\frac{\text{Kam-II data}}{\text{SSM}} = 0.46 \pm 0.13(\text{stat.}) \pm 0.08(\text{syst.}),$$

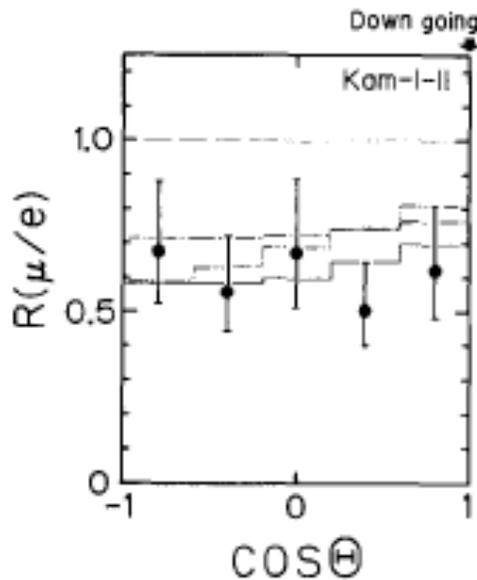
This is larger than Davis's value, but consistent if looking at just  $^8\text{B}$

# Meanwhile...

## Atmospheric neutrino 'anomaly'

'Ratio of ratios' supposedly robust to flux uncertainties:

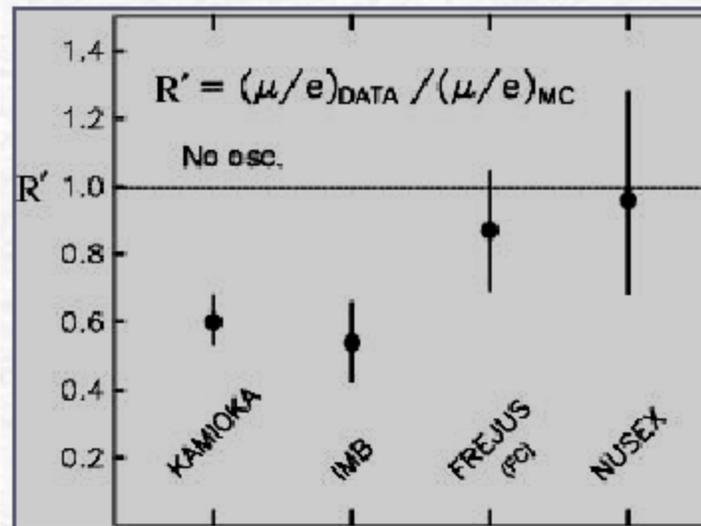
$$R(\mu/e) = (\mu/e)_{\text{data}} / (\mu/e)_{\text{MC}}$$



We are unable to explain the data as the result of systematic detector effects or uncertainties in the atmospheric neutrino fluxes. Some as-yet-unaccounted-for physics such as neutrino oscillations might explain the data.

*Hirata et al, PLB, 205, 416, (1988) and Hirata et al, PLB, 280, 146, (1992)*

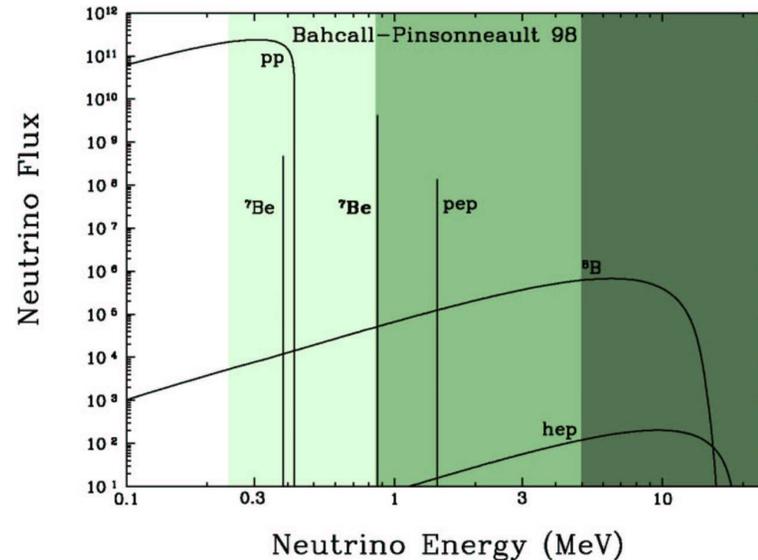
But...



*“No one ever lost money betting a neutrino experiment was wrong.”  
(Don Perkins)*

# Finally Gallium

1965 Kuzmin points out reaction  $\nu_e + {}^{71}\text{Ga} \rightarrow e^- + {}^{71}\text{Ge}$  has threshold of 233 keV, low enough to detect pp neutrinos (and everything else):



pp neutrinos are primary fusion products---

- fixed lower limit based on solar power output,  $\sim 79$  SNU.
- Standard Solar Model prediction is  $\sim 130$  SNU.

So: suppressed rate like Cl or K-II (40-60 SNU), unambiguous discovery of non-astrophysical solution.

But if rate  $\sim 130$  SNU, then everyone else is wrong.

# Finally Gallium

Two independent experiments

SAGE



First results (1991):

$^{71}\text{Ga}$  capture rate =  $20 \pm_{20}^{15}(\text{stat}) \pm 32(\text{syst})$  SNU

Big suppression?

Gallex



First results (1992):

$83 \pm 19$  (stat.)  $\pm 8$  (syst.) SNU

Consistent with SSM

# Finally Gallium

Two independent experiments

SAGE



First results (1991):

$${}^{71}\text{Ga capture rate} = 20 \pm_{20}^{15}(\text{stat}) \pm 32(\text{syst}) \text{ SNU}$$

Eventually (1999):

$$67.2 \begin{matrix} +7.2+3.5 \\ -7.0-3.0 \end{matrix} \text{ SNU}$$

Gallex



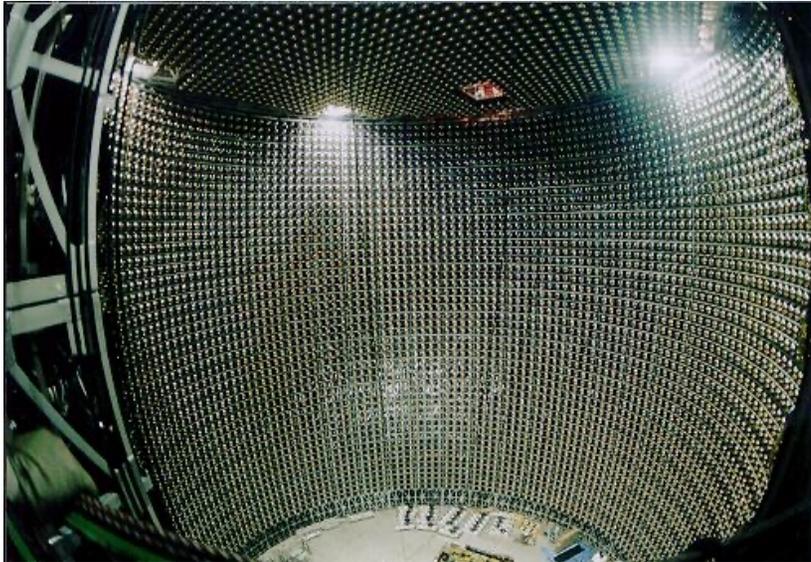
First results (1992):

$$83 \pm 19 (\text{stat.}) \pm 8 (\text{syst.}) \text{ SNU}$$

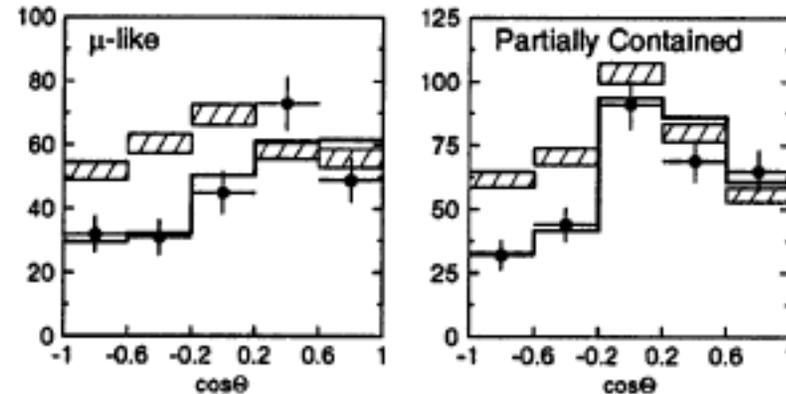
$$76.4 \pm 6.3 \begin{matrix} +1.5 \\ -4.9 \end{matrix} \text{ SNU}$$

→ Including calibrations with  ${}^{51}\text{Cr}$

# Super-Kamiokande

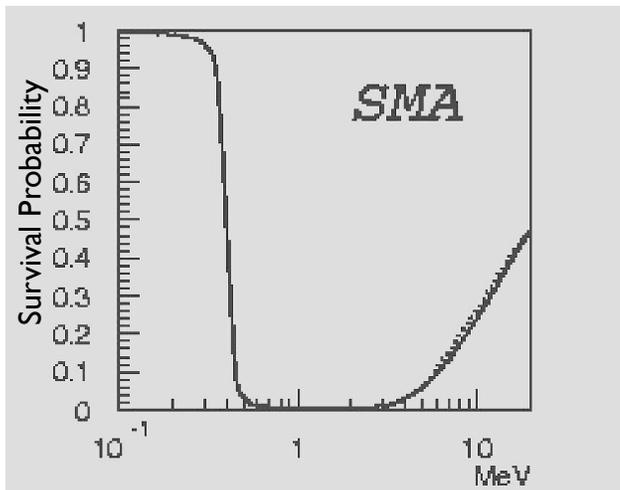


The big news (1998):



Atmospheric  $\nu$  disappearance fit extremely well by oscillations

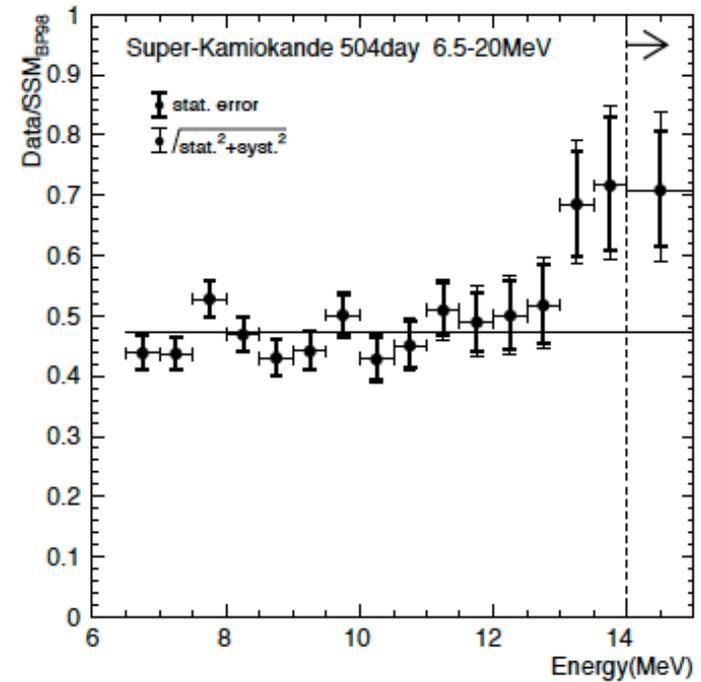
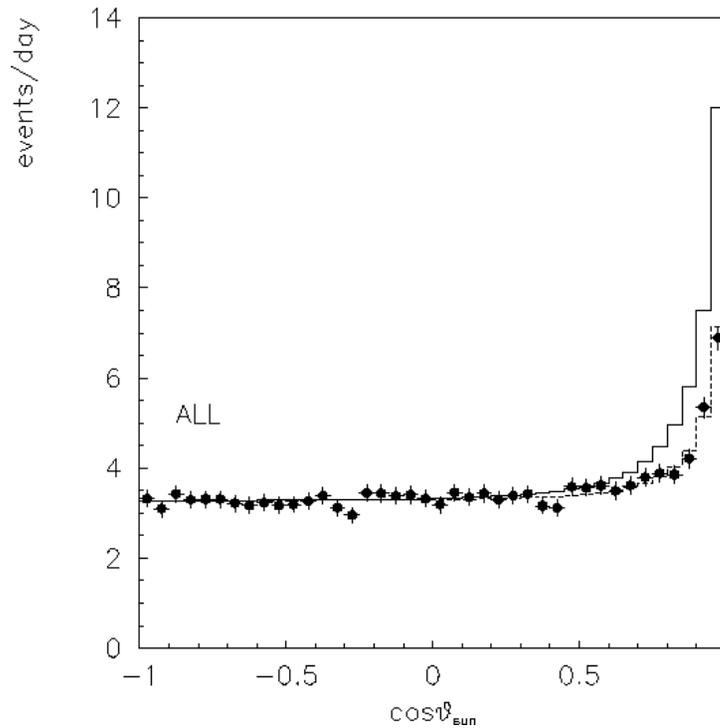
For solar  $\nu$ 's, MSW+small angle mixing predicts:



Big distortion in  $\nu$  energy spectrum!  
Or could see Day/Night asymmetry via MSW in Earth!

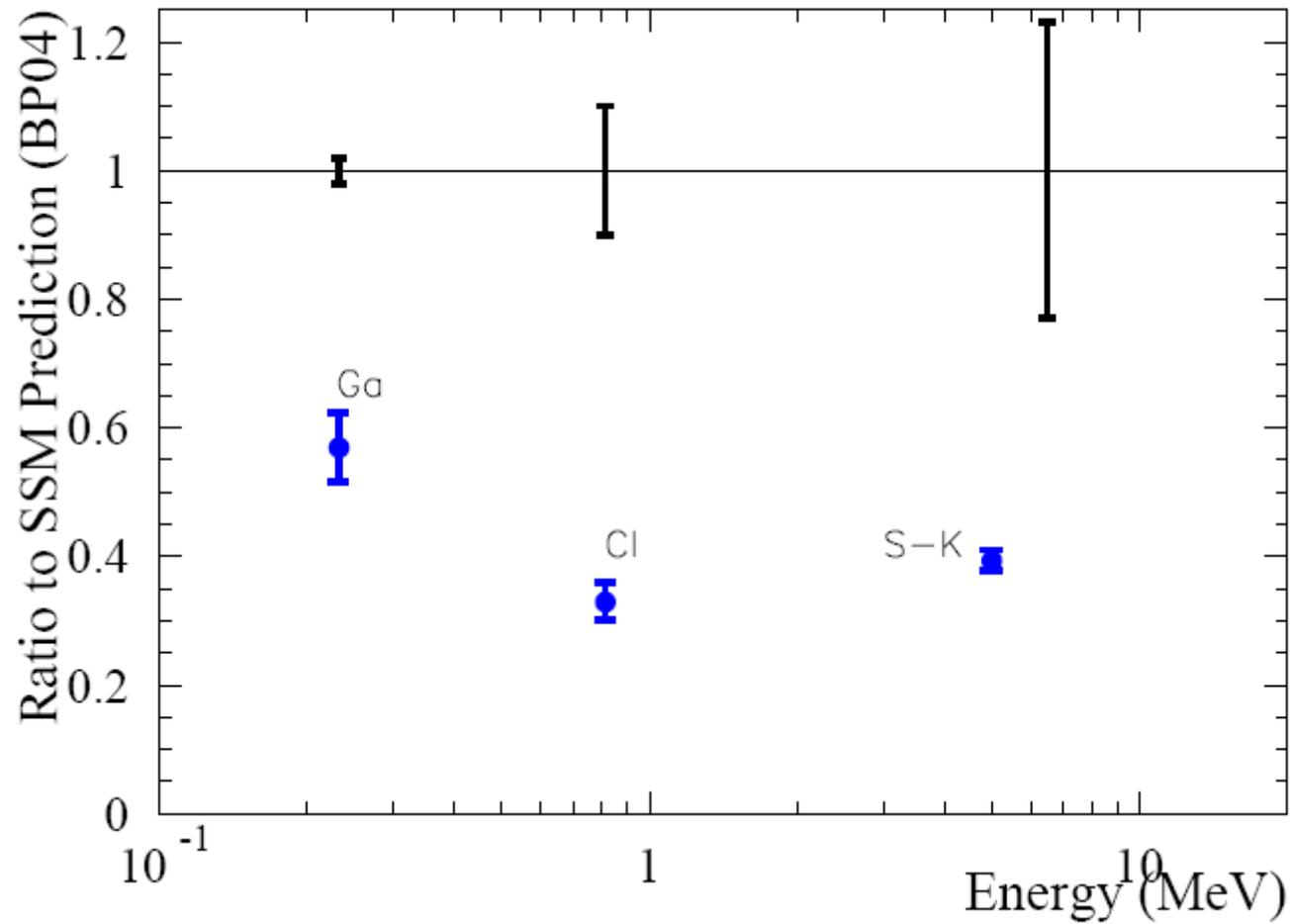
Either would be unambiguous solution to solar  $\nu$  problem...

# Super-Kamiokande



Nope. Clear confirmation of K II results with high precision, but no smoking gun.

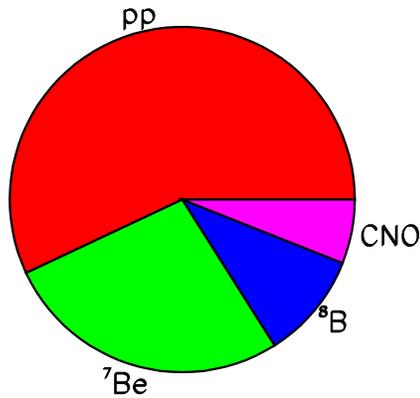
# The Story So far:



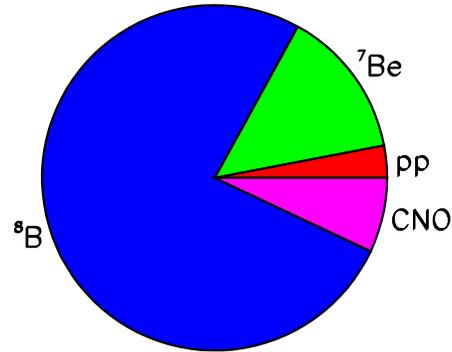
# After Six Solar $\nu$ Experiments

## Expectations

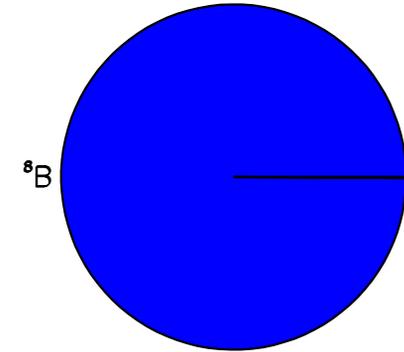
Gallium



Chlorine

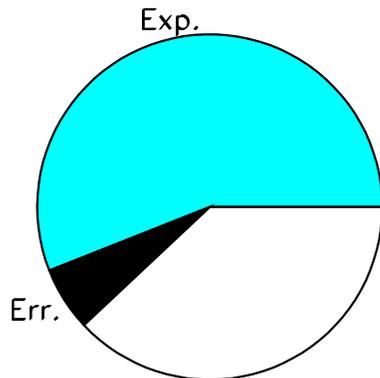


SuperKamiokande

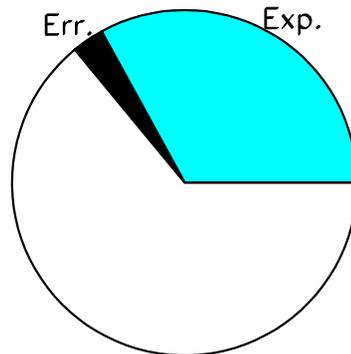


## Measurements

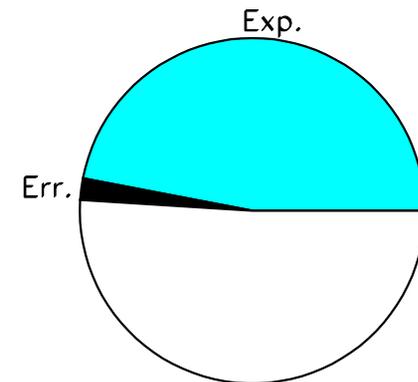
Gallium



Chlorine



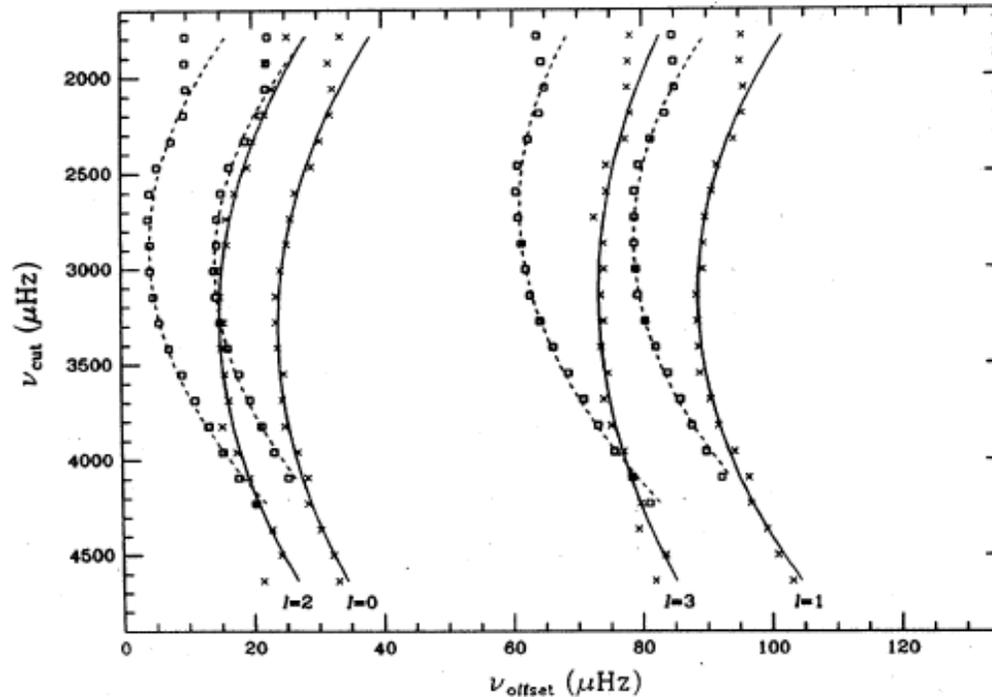
SuperKamiokande



➤ Looks like complete absence of <sup>7</sup>Be neutrinos.

# A Second Prediction of the SSM

Acoustical modes of Sun can be measured very precisely



And agreement with SSM predictions good to 0.5%!

(Note: depends critically on CNO content!)

# Desperation?

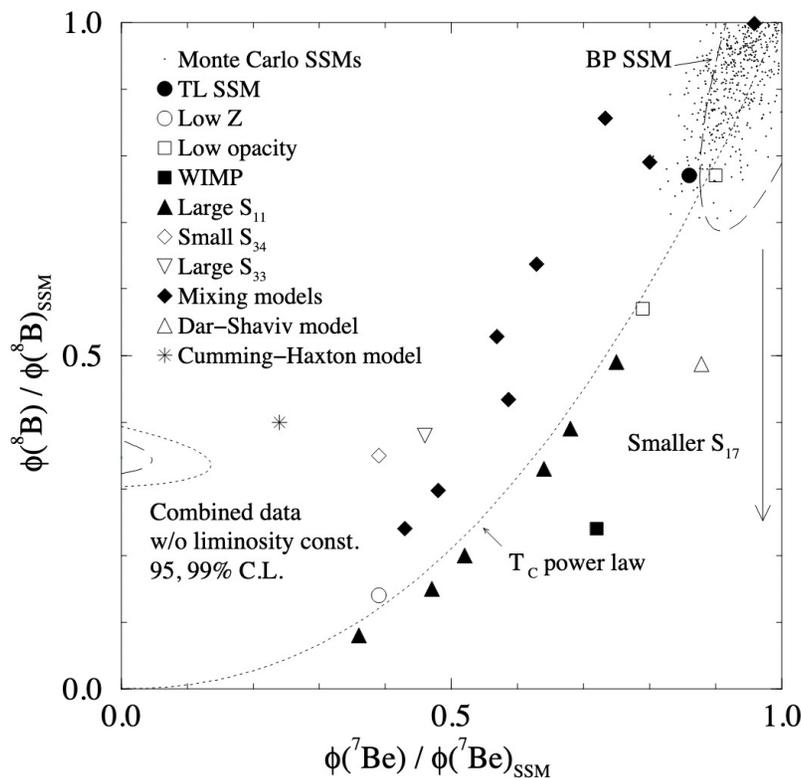
Or Anticipation?

People begin running Monte Carlo simulations of 1000s of solar models, to prove astrophysics is not to blame....

Bahcall begins quoting  $3\sigma$  errors on everything.

Papers often contain phrases like, “Even without the CI results..” and “Allowing the CI rate to float...”

With the Super-K atmospheric results, these arguments are compelling...but not quite the smoking gun people need.



Hata and Langacker, Phys. Rev. D56, 6107 (1997)

# Heavy Water and Neutrinos

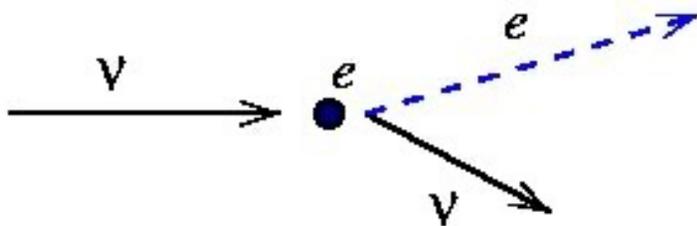
VOLUME 55, NUMBER 14

PHYSICAL REVIEW LETTERS

30 SEPTEMBER 1985

## Direct Approach to Resolve the Solar-Neutrino Problem

Neutrino-electron scattering

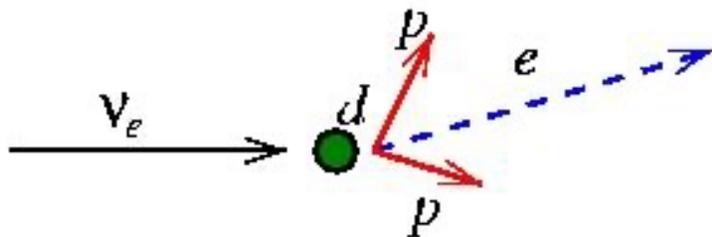


**Elastic Scattering (ES)**

Chen suggests an (inclusive) appearance measurement using  $D_2O$

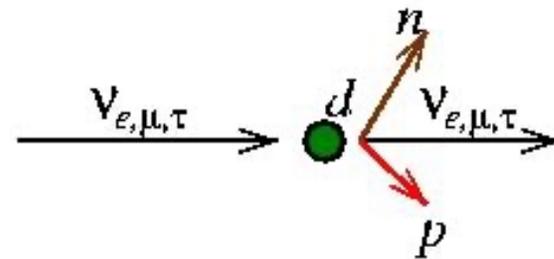


Neutrino absorption by deuteron



**Charged Current (CC)**

Neutrino breakup of deuteron

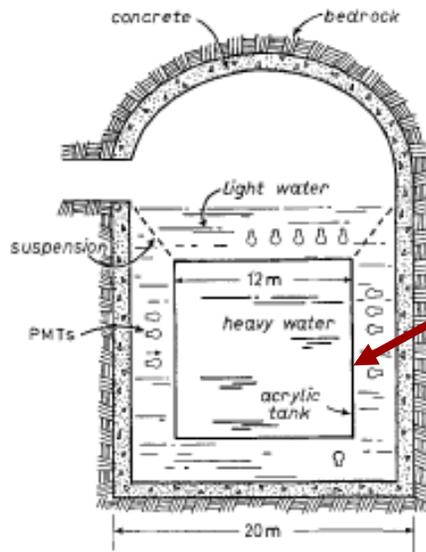


**Neutral Current (NC)**

# Sudbury Neutrino Observatory

All that's needed is:

- About 1 kton of heavy water (~\$250M)
- 2 km underground in a mine
- With enough PMTs to see  $\gamma$ -rays from neutron capture

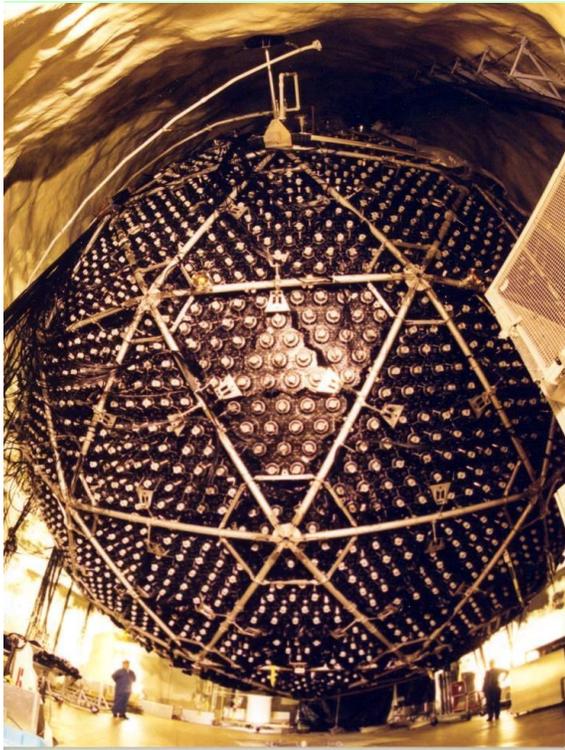


'Sea World tank' to use H<sub>2</sub>O as shielding

Fig. 5. -- Conceptual design of the light/heavy-water detector. The heavy water is contained in an acrylic tank and is shielded from activity in the rock by low activity concrete and light water. The Čerenkov light produced in the water is detected by an array of 2400 50 cm diameter phototubes.

Sure. Review committee says “Physics goals...are of outstanding value.”  
Funding agency award (1985): \$0.

# Sudbury Neutrino Observatory



1000 tonnes D<sub>2</sub>O

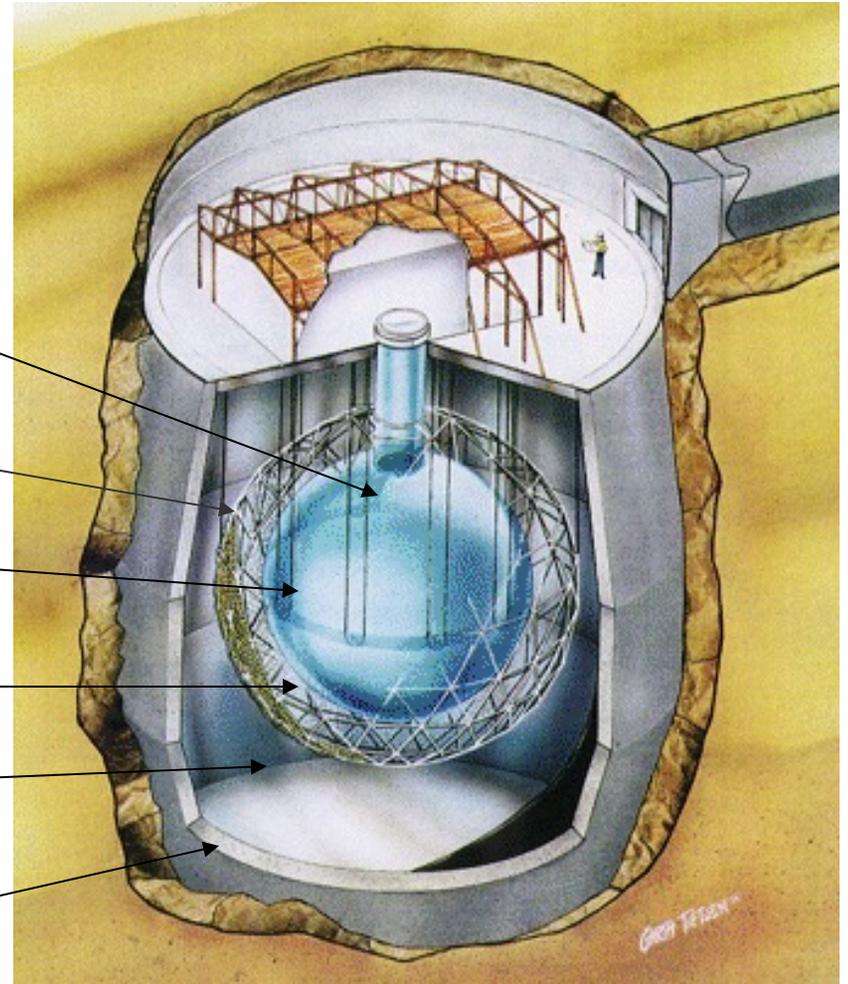
Support Structure  
for 9500 PMTs,  
60% coverage

12 m Diameter  
Acrylic Vessel

1700 tons Inner  
Shielding H<sub>2</sub>O

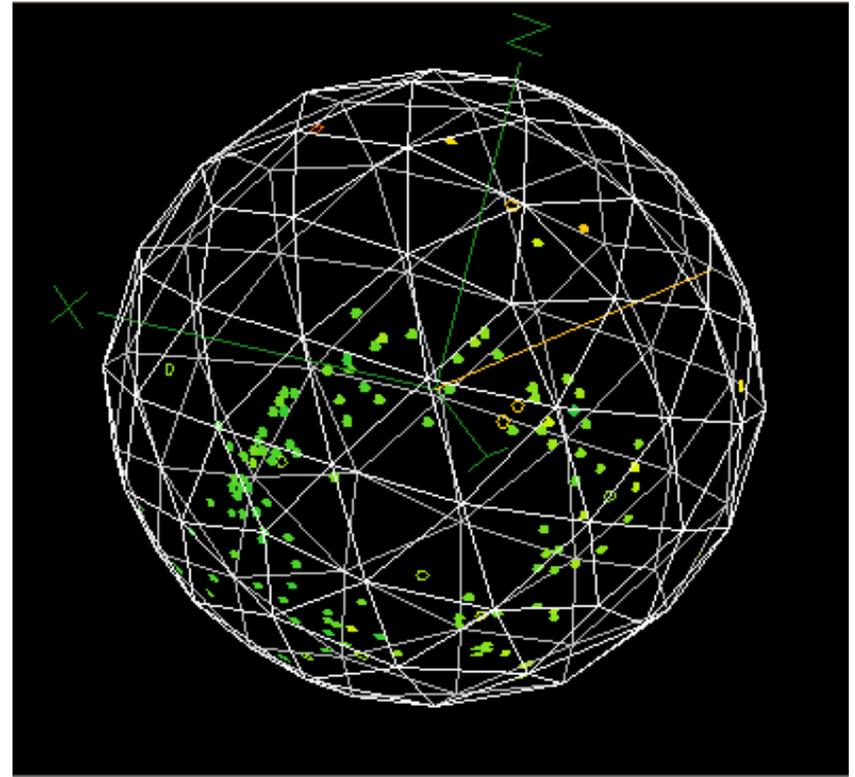
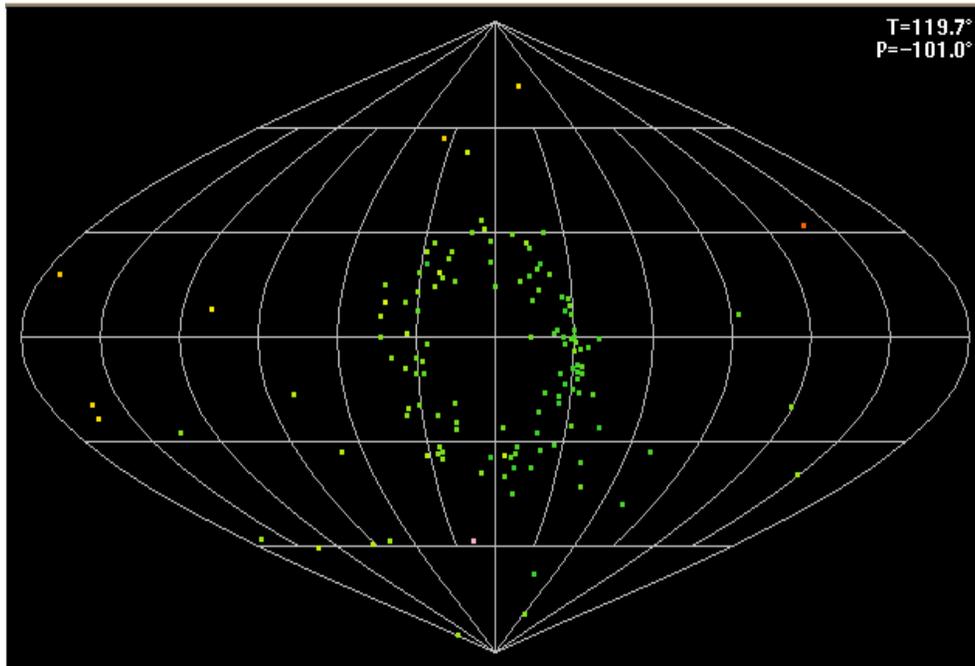
5300 tons Outer  
Shield H<sub>2</sub>O

Urylon Liner and  
Radon Seal



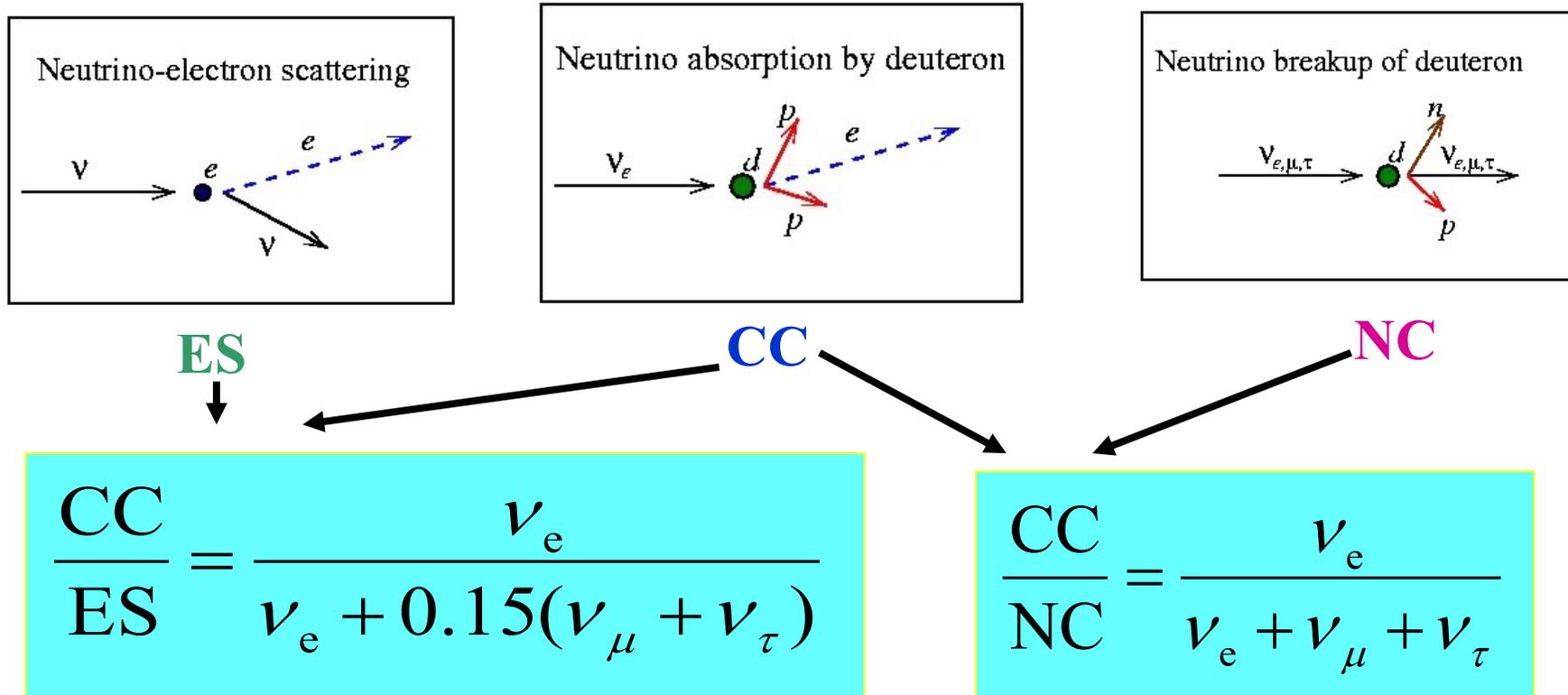
# Sudbury Neutrino Observatory

## $^8\text{B}$ Solar $\nu$ Event



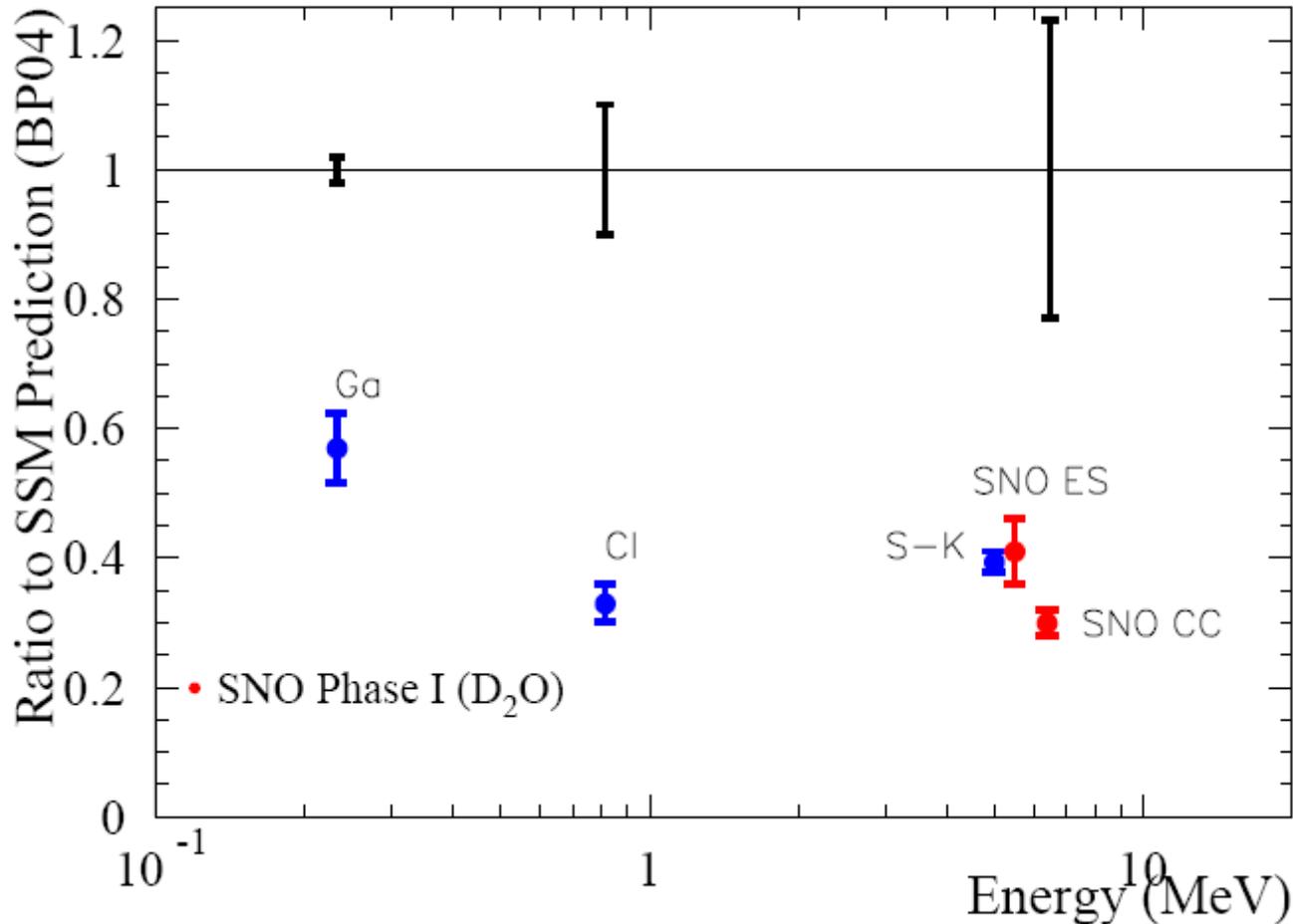
# Sudbury Neutrino Observatory

Two ways of doing 'inclusive appearance' measurement:



# Sudbury Neutrino Observatory

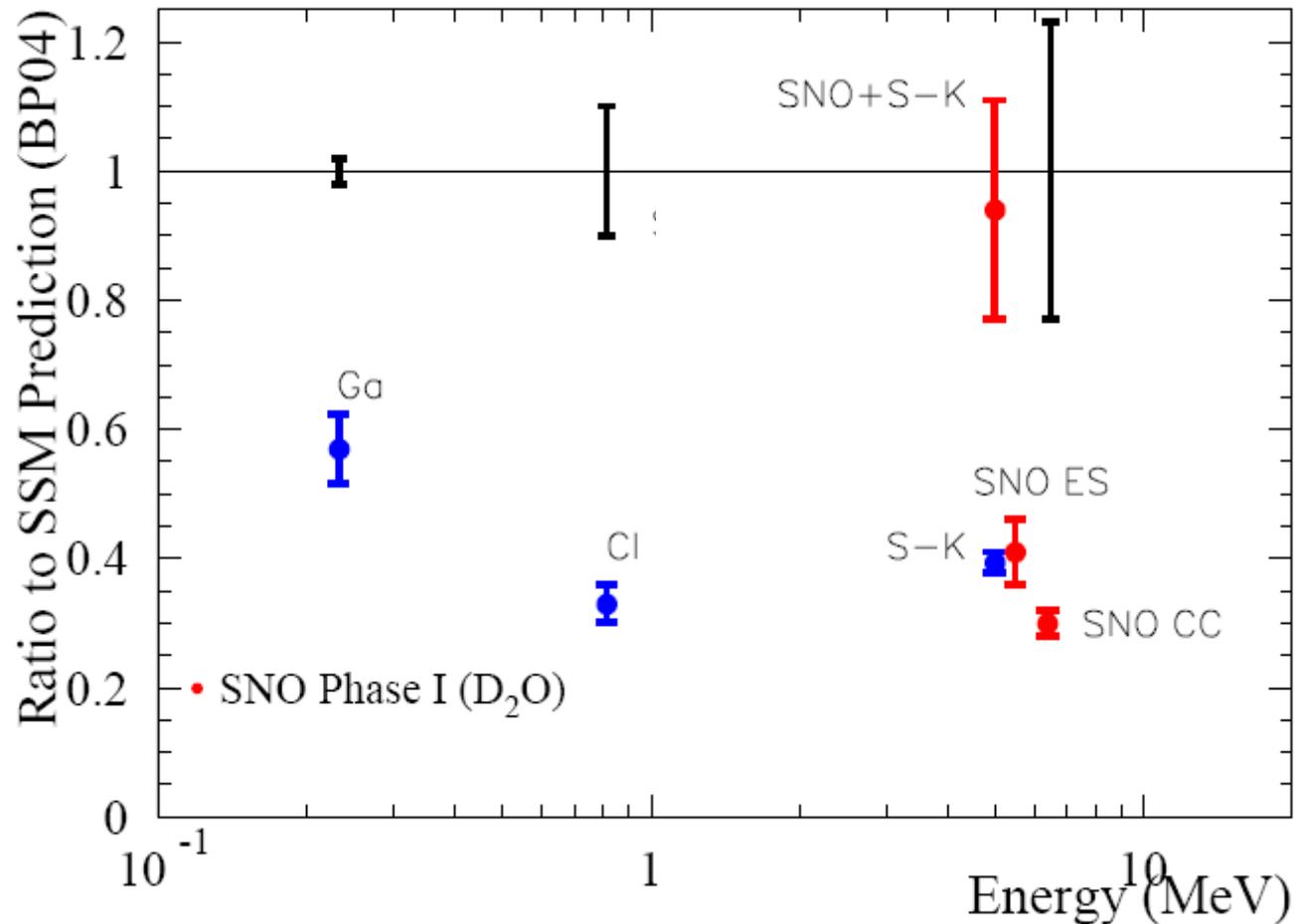
First Results (2001)



$$\Phi_{ES}^{SK} - \Phi_{CC}^{SNO} = 0.57 \pm 0.17 \Rightarrow 3.3\sigma > 0$$

# Sudbury Neutrino Observatory

First Results (2001)



Total flux agrees with Standard Solar Model!

# A Bayesian World

VOLUME 87, NUMBER 7

PHYSICAL REVIEW LETTERS

13 AUGUST 2001

Measurement of the Rate of  $\nu_e + d \rightarrow p + p + e^-$  Interactions Produced  
by  $^8\text{B}$  Solar Neutrinos at the Sudbury Neutrino Observatory

Reaction to SNO results more positive than perhaps collaboration  
expected for just  $3.3\sigma$ , e.g.:

“...the standard model of the Sun appears to be in better shape  
than the standard model of electroweak interactions.”

Fogli, Lisi, Montanino, Palazzo, Phys. Rev. D64, 2001



“I feel very much like the way I expect that these prisoners that are  
sentenced for life do when a D.N.A. test proves they're not guilty," Dr.  
Bahcall said. "For 33 years, people have called into question my  
calculations on the Sun.”

New York Times, June 19, 2001

Nevertheless, some skepticism for the next 10 months...

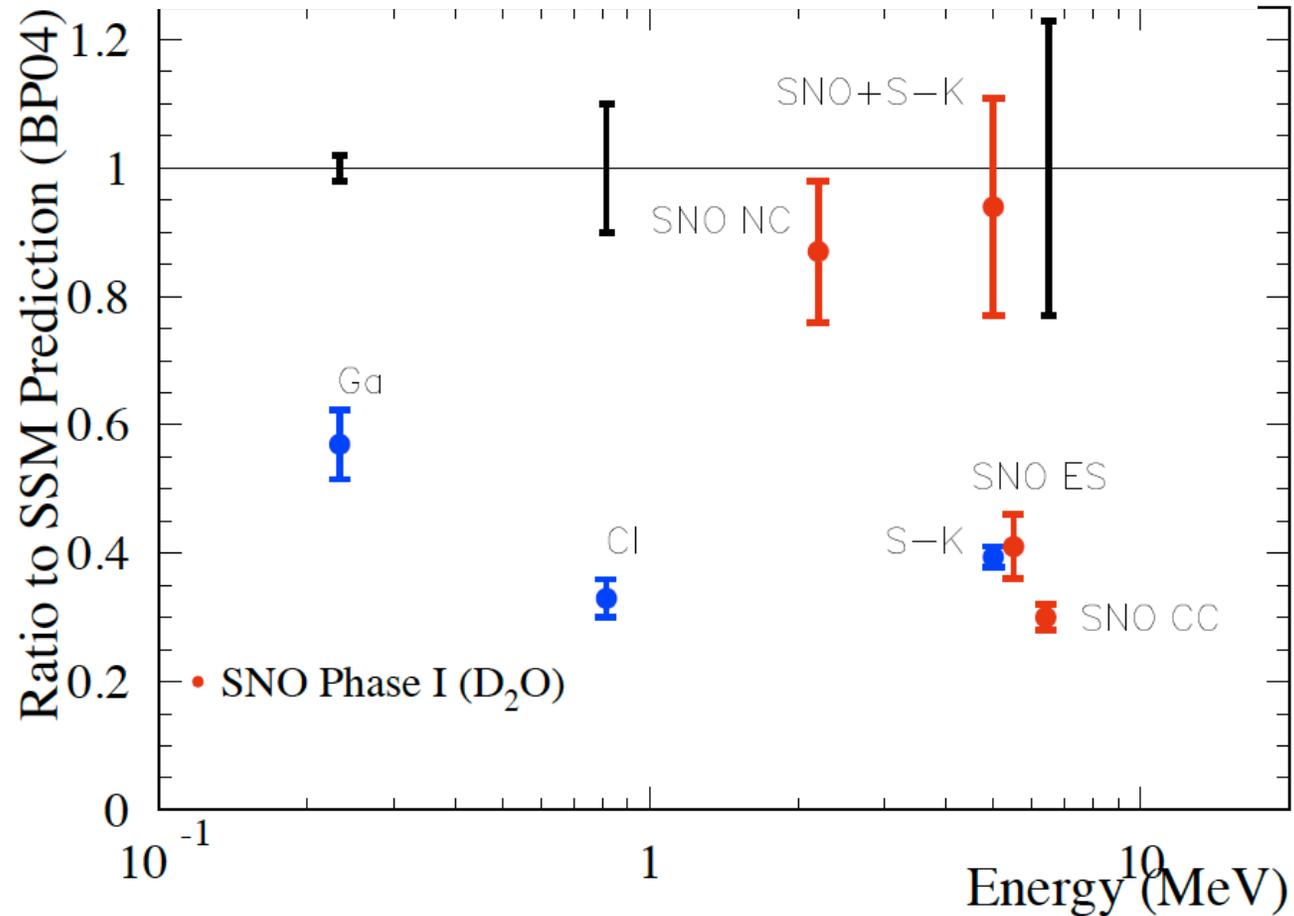
# Sudbury Neutrino Observatory

VOLUME 89, NUMBER 1

PHYSICAL REVIEW LETTERS

1 JULY 2002

Direct Evidence for Neutrino Flavor Transformation from Neutral-Current Interactions in the Sudbury Neutrino Observatory



5.3 $\sigma$  difference between NC and CC

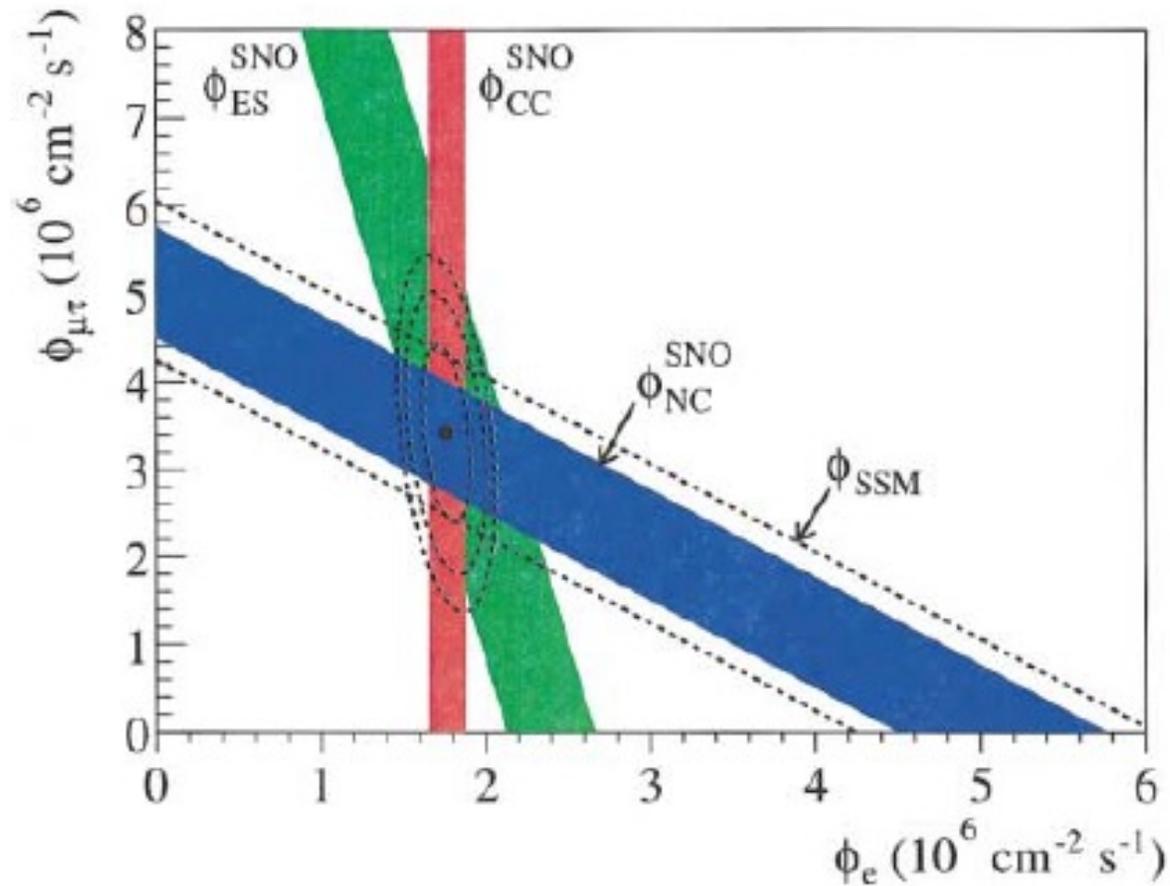
# Sudbury Neutrino Observatory

VOLUME 89, NUMBER 1

PHYSICAL REVIEW LETTERS

1 JULY 2002

## Direct Evidence for Neutrino Flavor Transformation from Neutral-Current Interactions in the Sudbury Neutrino Observatory



5.3 $\sigma$  difference between NC and CC

# Back to 1964

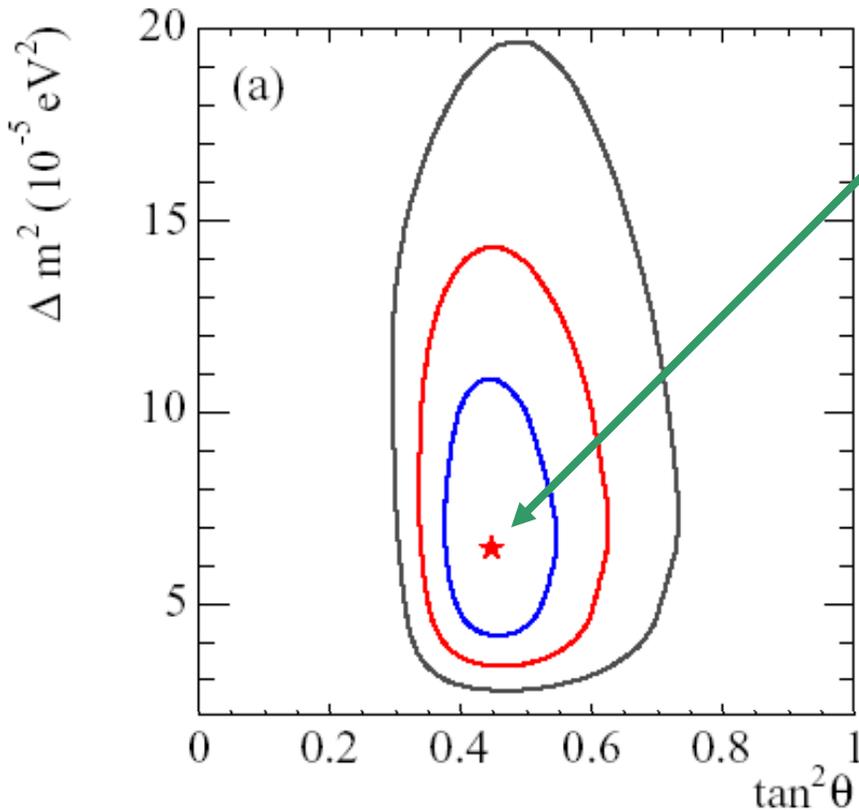
Standard Solar Model Predicts

$$\phi_{\nu}^{8B} \propto T^{25}$$

▶  $T_{\text{Sun}}^{\text{core}} \cong 15.6 \times 10^6 \text{ K}$

# Closing the Loop

Solar data:



This is the best fit value of the **vacuum** mass difference and mixing angle under the **hypothesis** that **observations** are due to **MSW** (matter) oscillations.



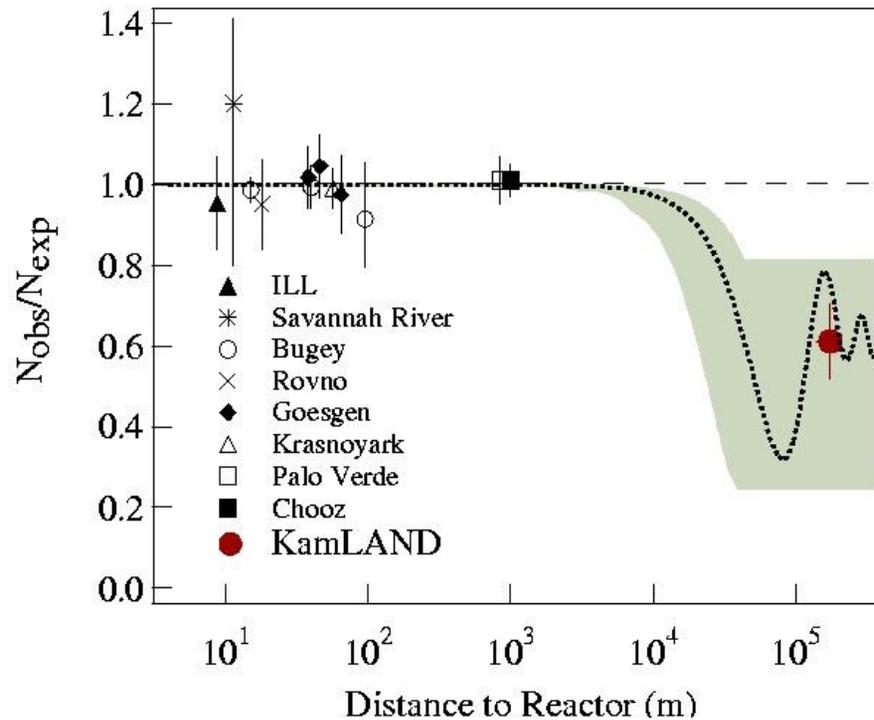
But if we plug these values into the vacuum survival probability:

$$P_{\nu_e \rightarrow \nu_e} = 1 - \sin^2 2\theta_{12} \sin^2 \left( \frac{1.27 \Delta m_{12}^2 L}{E} \right)$$

We find that oscillation length  $L \sim 100$  km for  $E = 4$  MeV.

# Solar $\nu$ Oscillations with No Sun

➤ KamLAND—Reactor antineutrinos



(This could have been done in 1968!)

Event numbers for 766 ton-years of data:

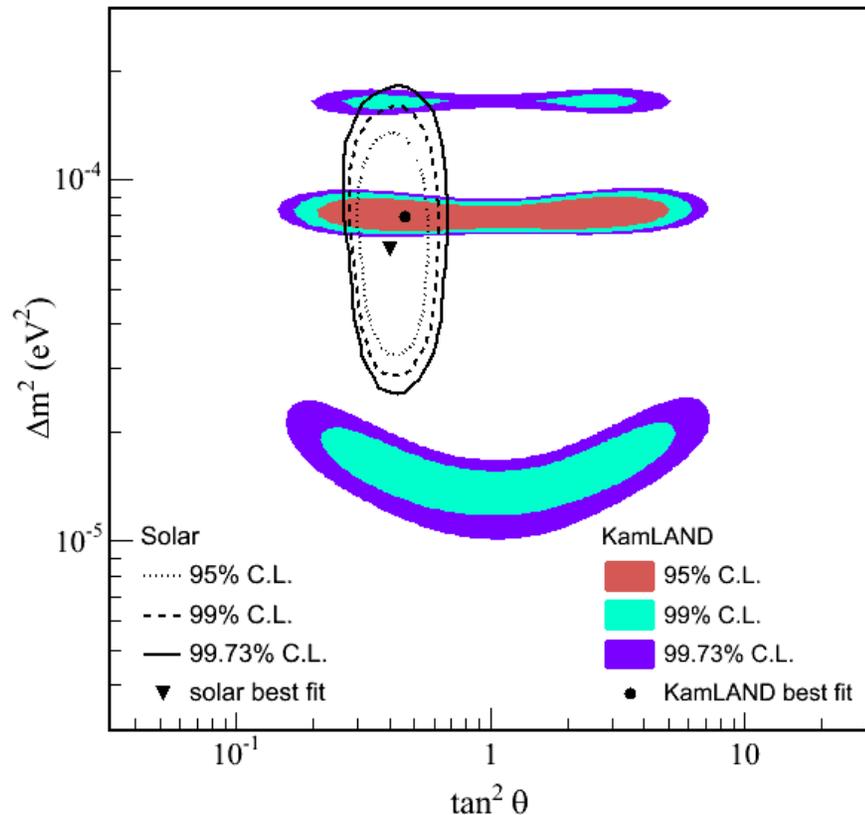
Expected (no oscillation)=365.2+17.8 background

Observed =258

Disappearance with significance of 99.998%

# Solar $\nu$ Oscillations with No Sun

➤ 'Precision' Comparison Across Regimes



	Reactor	Solar
<b>E</b>	2-10 MeV	0.1-15 MeV
<b>L</b>	150 km	$1.5 \times 10^8$ km
<b>MSW</b>	No	Yes
<b><math>\nu</math></b>	Anti- $\nu_e$	$\nu_e$

Large mixing means all transformation happens inside Sun

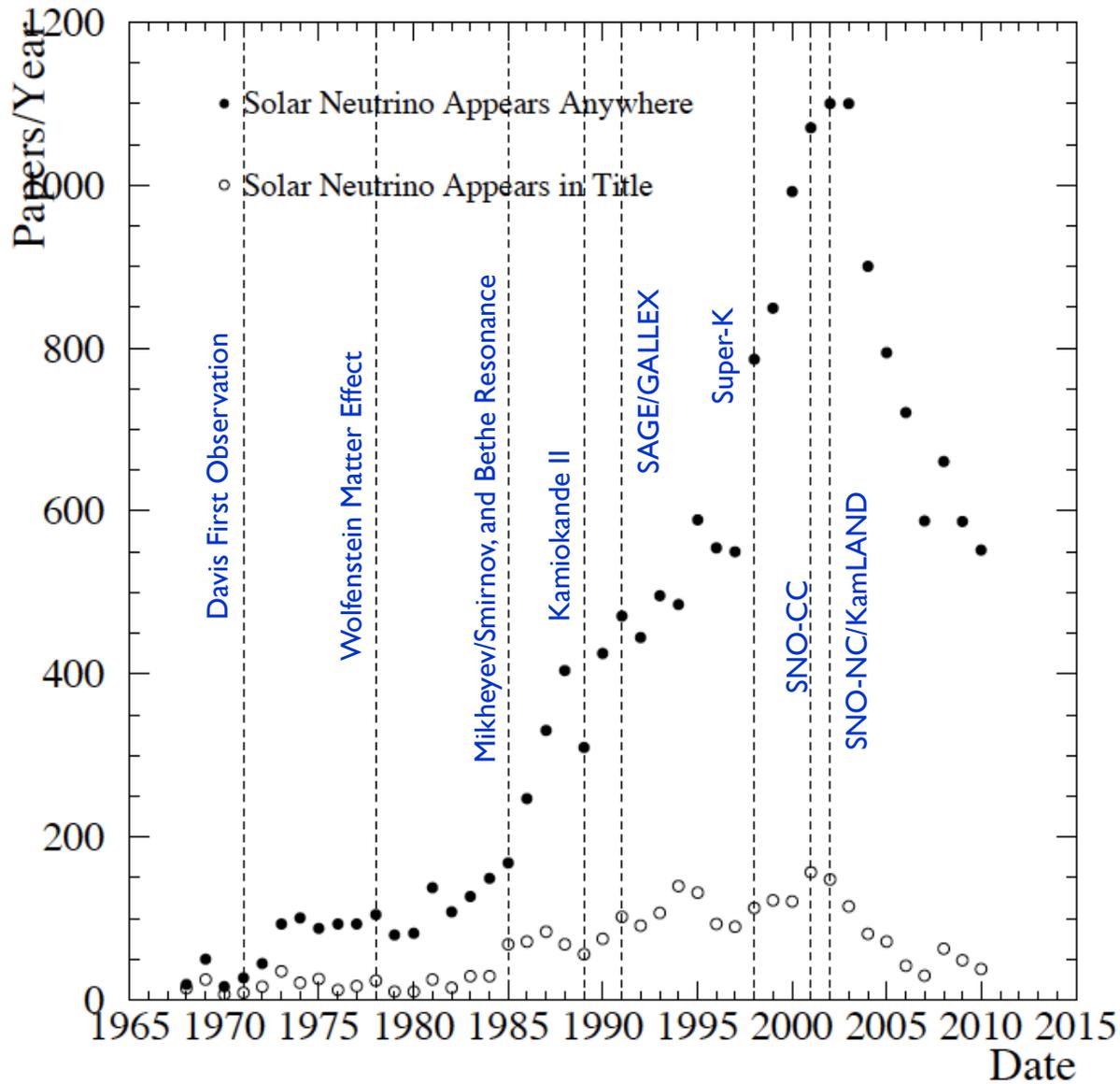
# A Long Story

## Punished by Success

*“For 35 years people said to me: ‘John, we just don’t understand the Sun well enough to be making claims about the fundamental nature of neutrinos, so we shouldn’t waste time with all these solar neutrino experiments.’ Then the SNO results came out. And the next day people said to me, ‘Well, John, we obviously understand the Sun perfectly well! No need for any more of these solar neutrino experiments.’”*

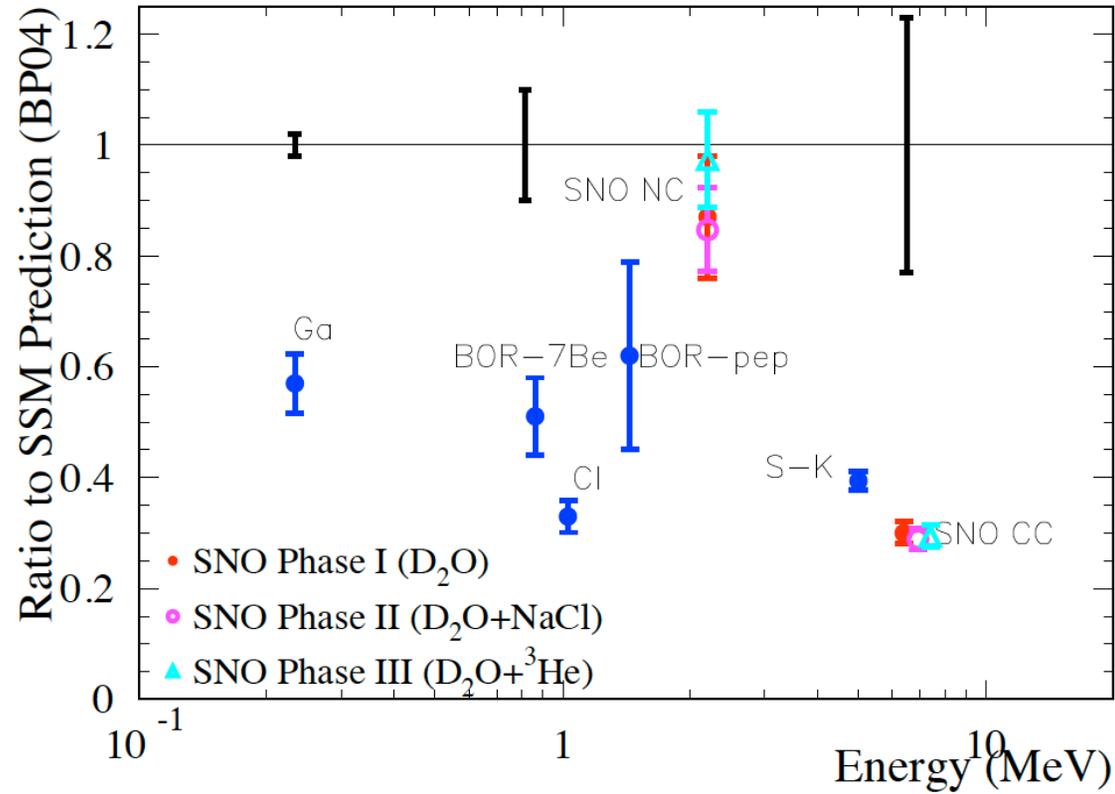
---John Bahcall, 2003

~35 Years



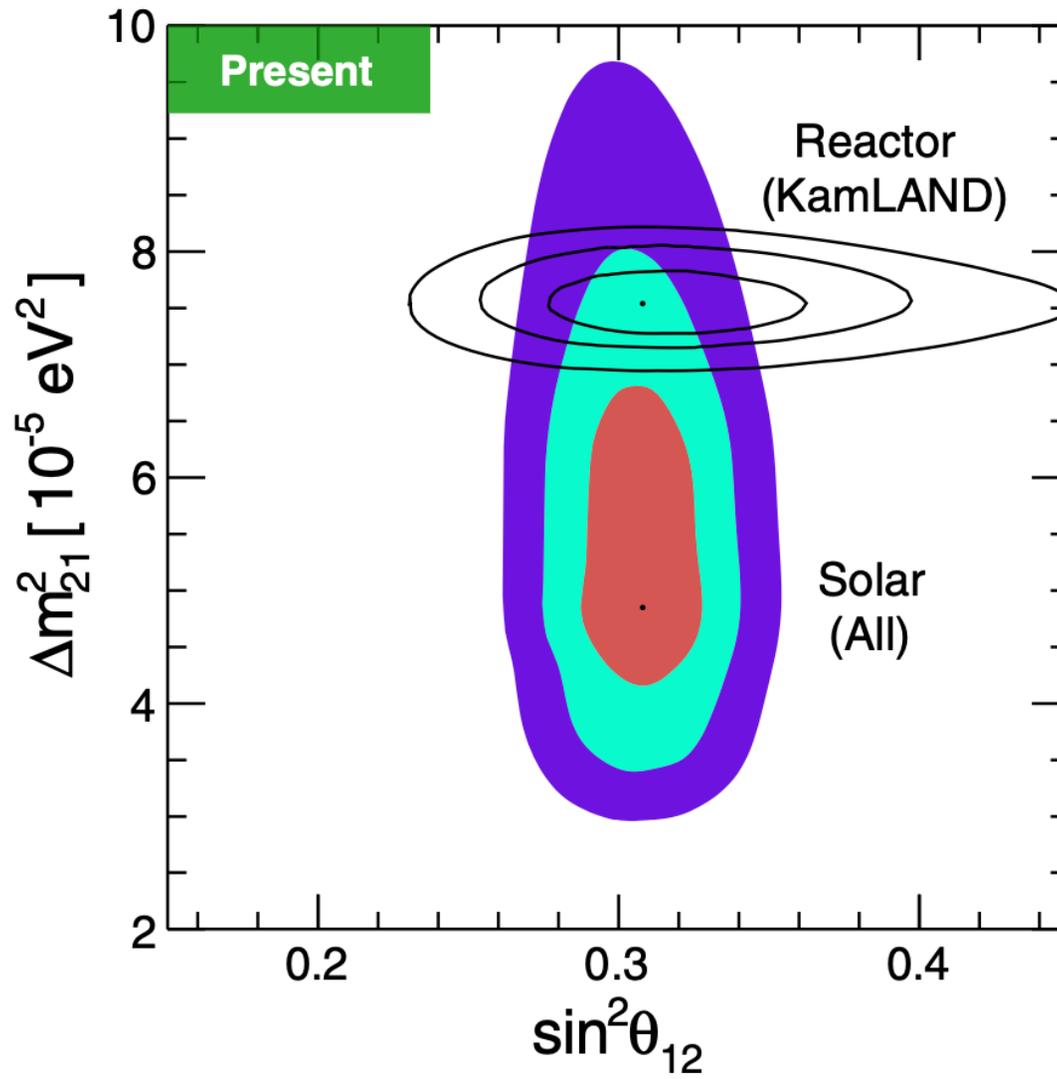
# Current Status

## Residual Paranoia



Three ways of detecting NC in SNO, to be 'really sure.'

# Current Status

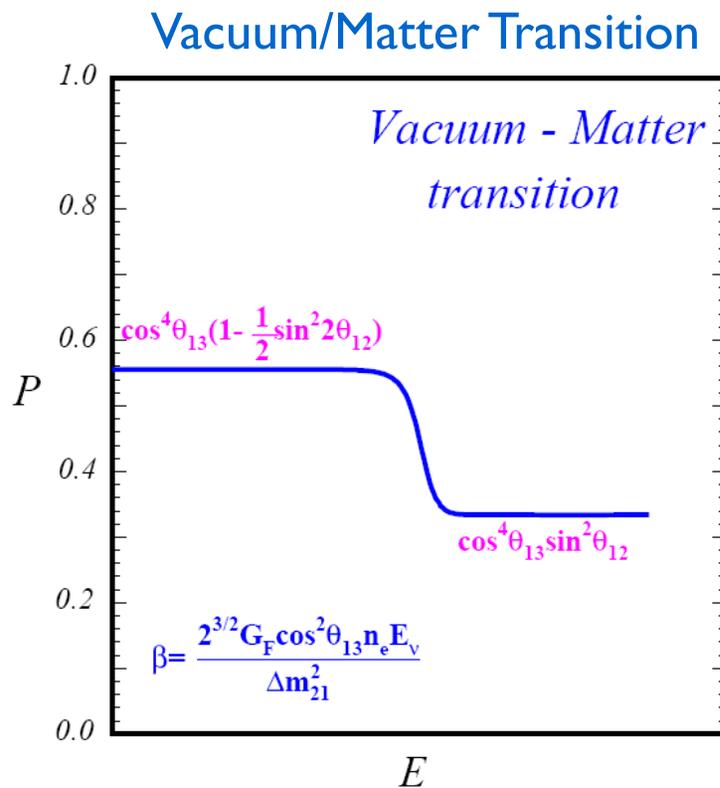


Terrestrial vs. solar  
measurements of  $\Delta m^2$   
Agree at  $\sim 2\sigma$  level

Should get new  
reactor  
measurements from  
SNO+ and JUNO  
soon.

# Anything Left?

- Test the model of massive neutrino mixing



Day/Night  $\nu_e$  Asymmetry

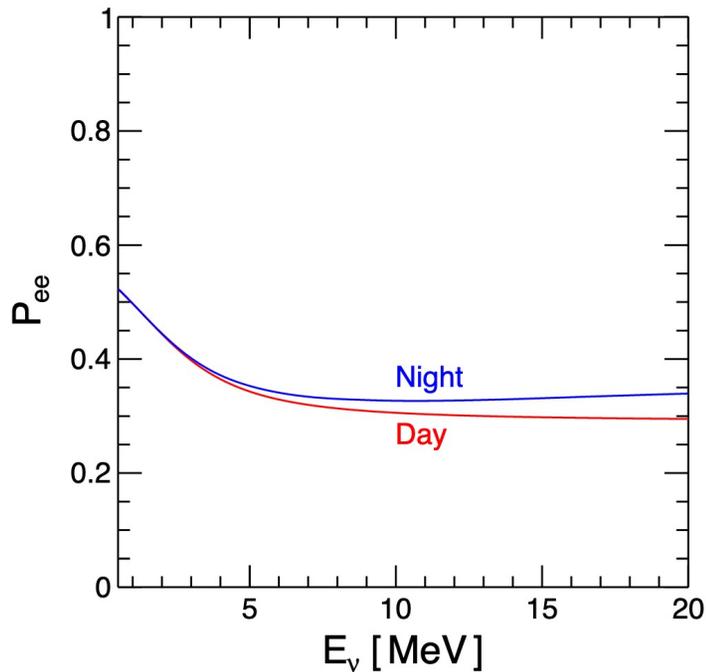


MSW effect predicts transition from matter to vacuum-dominated mixing, and Day/Night asymmetry

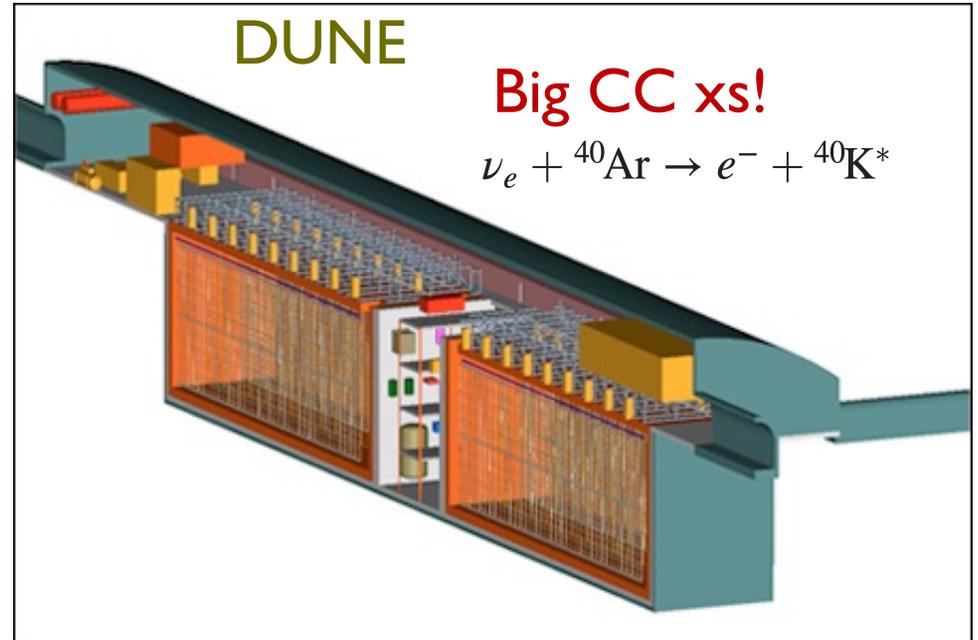
# Anything Left?

- Test the model of massive neutrino mixing

Day/Night  $\nu_e$  Asymmetry



Might be first solar neutrinos detected in US since Davis!



For DUNE, expect

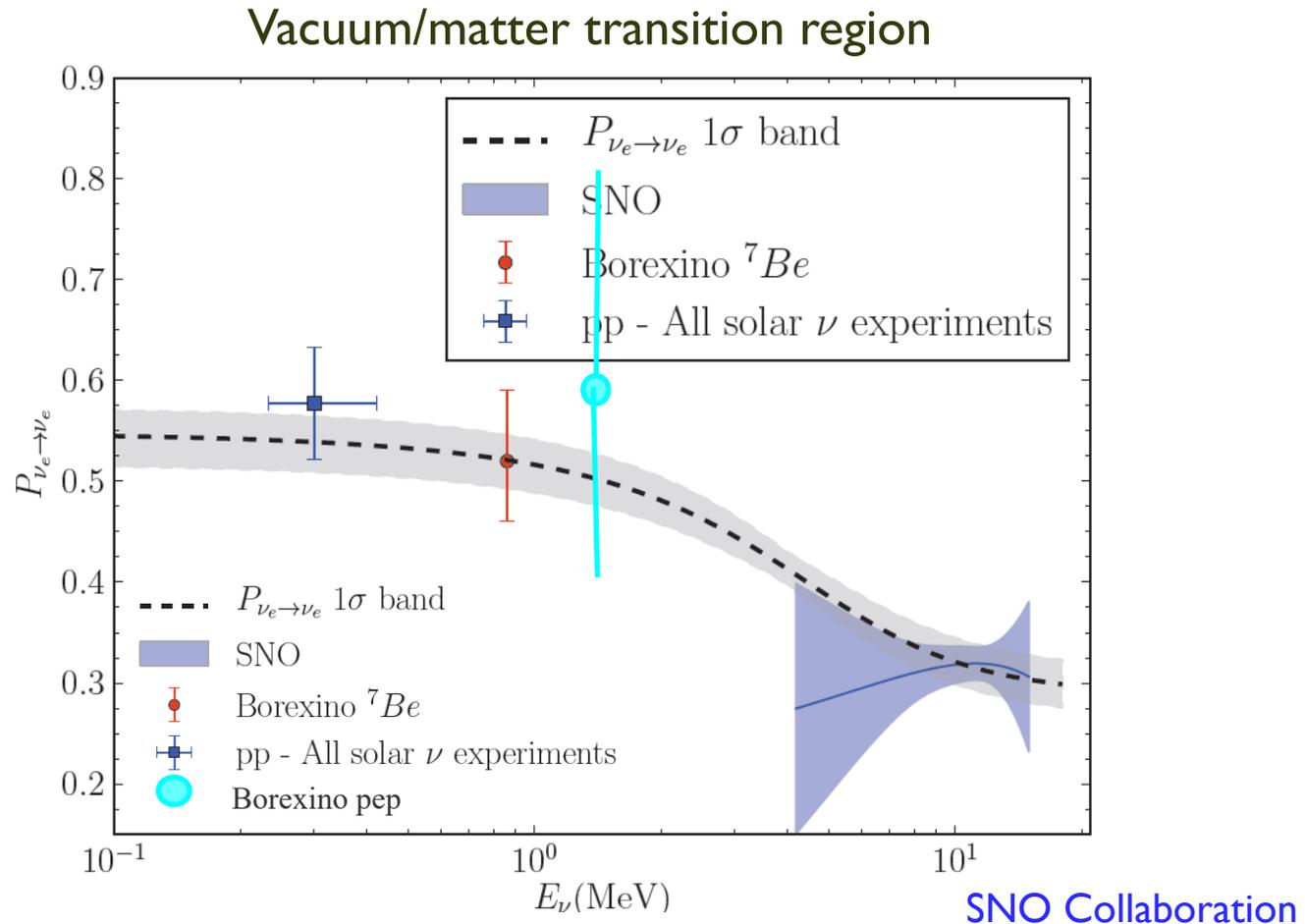
$$A_{D/N} = -(7.67 \pm 0.74)\%$$

With  $\sim 40,000$  events/day, can measure this to  $\sim 1\%$  statistically

PHYSICAL REVIEW LETTERS **123**, 131803 (2019)

Francesco Capozzi<sup>1,2,3,\*</sup> Shirley Weishi Li<sup>1,2,4,†</sup> Guanying Zhu<sup>1,2,‡</sup> and John F. Beacom<sup>1,2,5,§</sup>

# Observing MSW Phenomenology

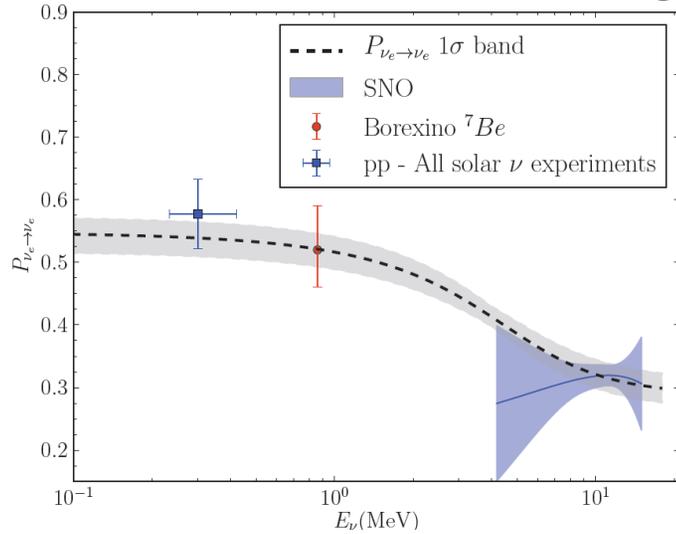


This is a little frustrating.

Solar vs give us our only huge, observed matter effect.  
The situation is ripe for a precision measurement program.

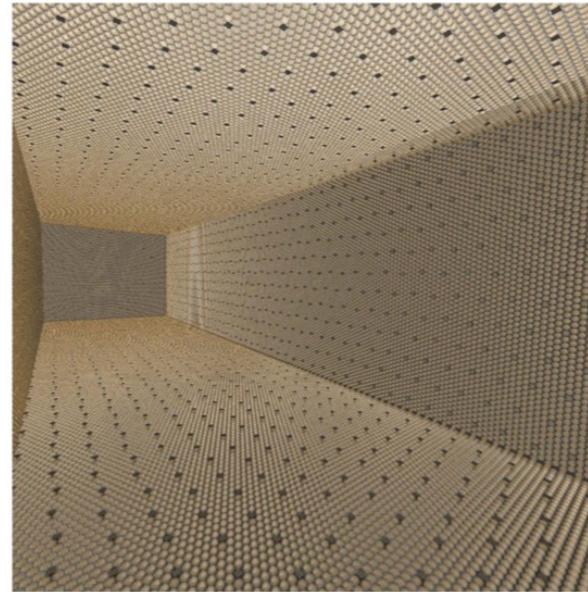
# Observing MSW Phenomenology

## Vacuum/matter transition region

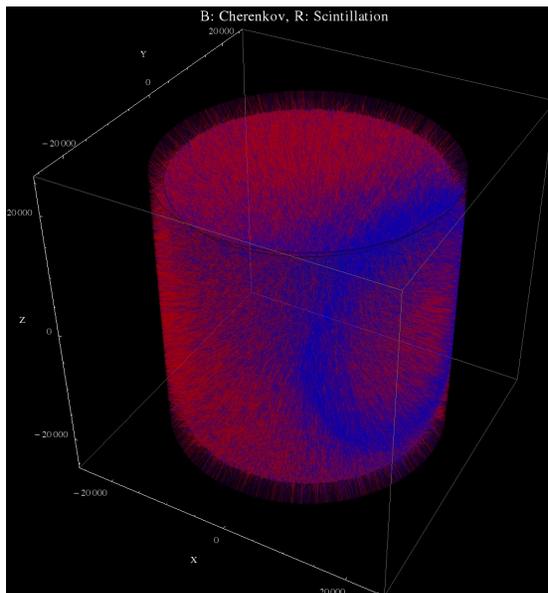


“Hybrid” Cherenkov/scintillation detector could allow threshold in 1 MeV region with statistics beyond Super-K

Theia25

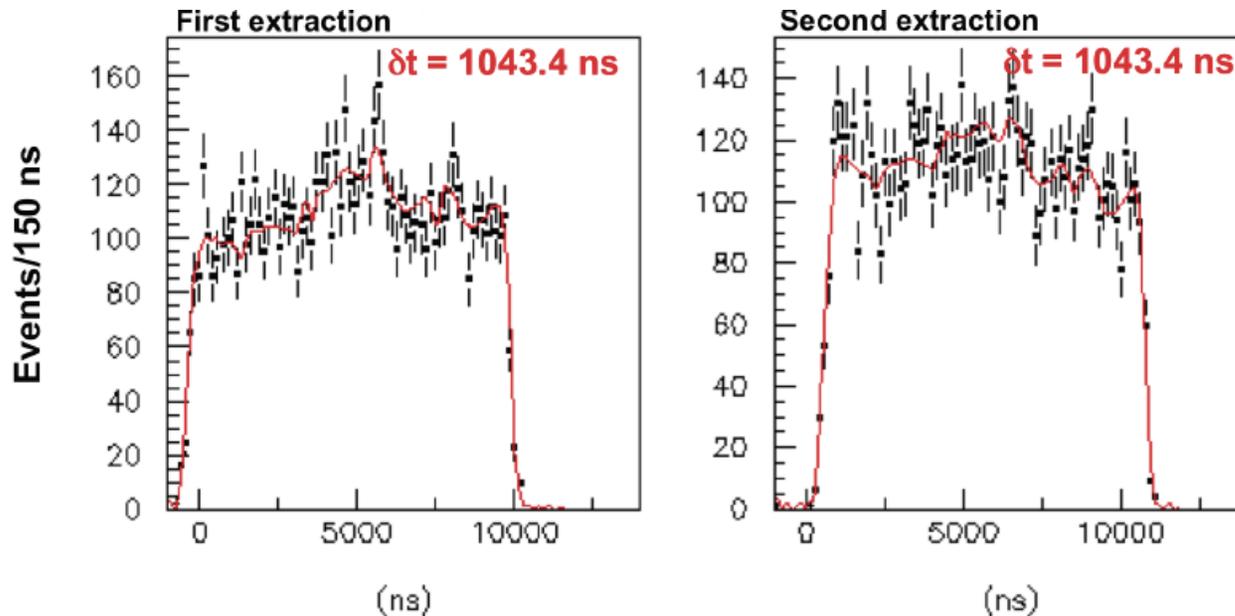


Theia100



# Have We Learned the Lessons Too Well?

Measurement of the neutrino velocity with the OPERA detector in the CNGS beam<sup>\*</sup>

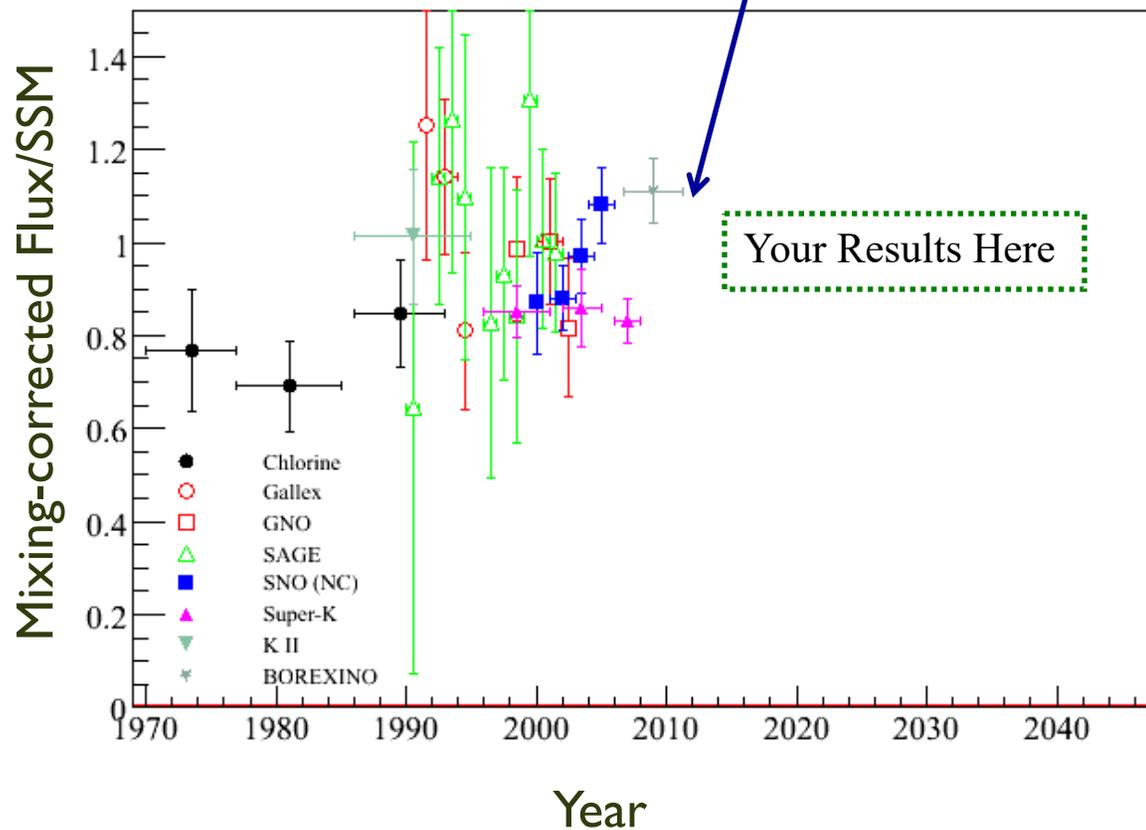


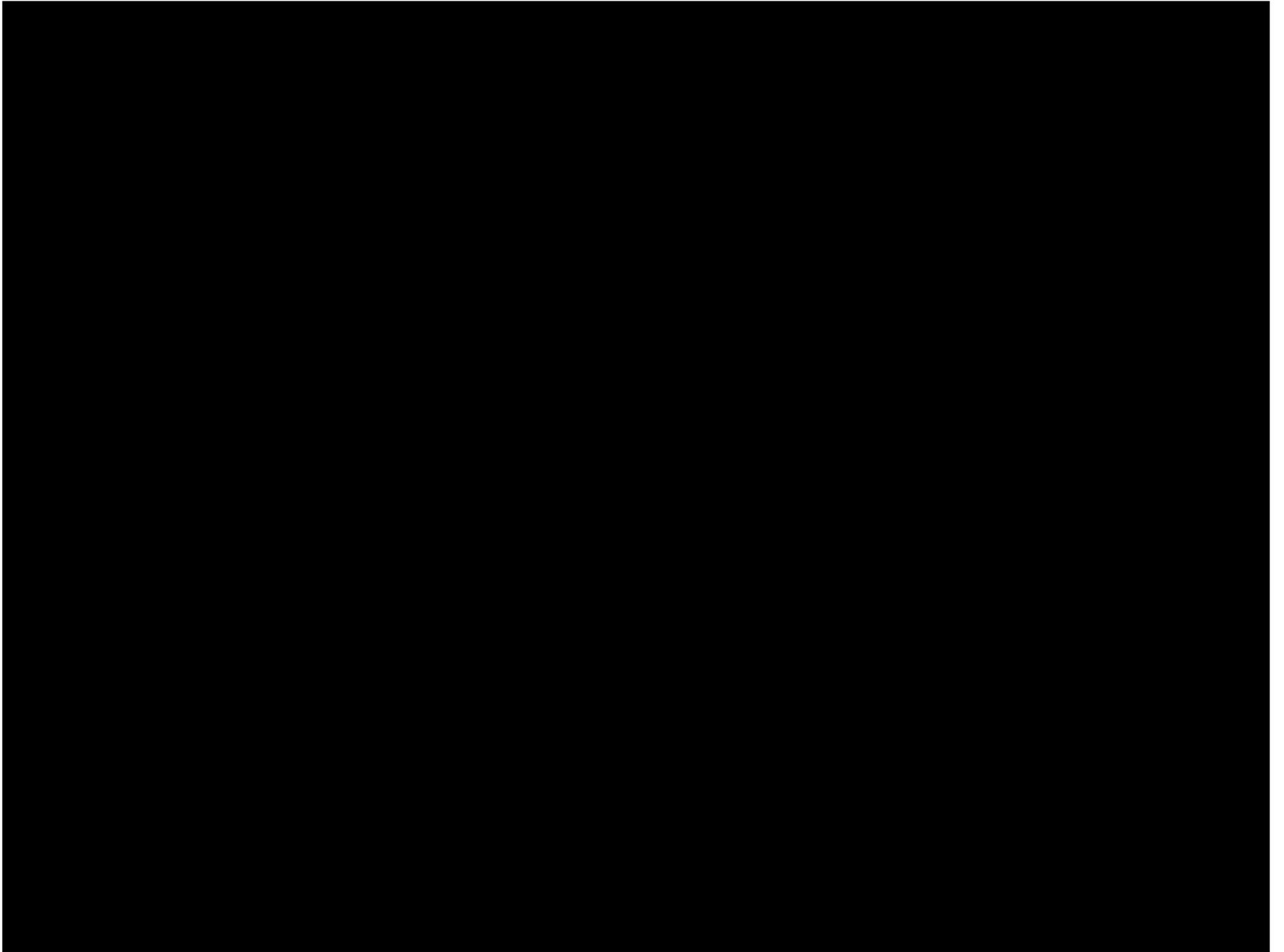
$$\delta t = \text{TOF}_c - \text{TOF}_\nu = 1043.4 \text{ ns} - 985.6 \text{ ns} = (57.8 \pm 7.8 \text{ (stat.)}) \text{ ns.}$$

~250 citations...

# Summary

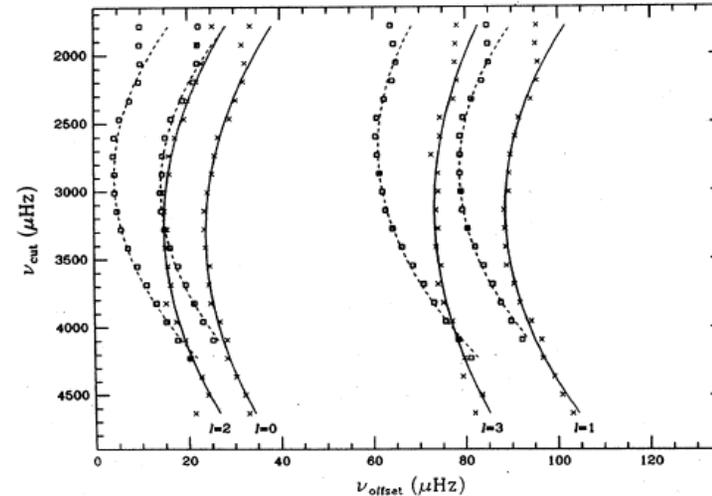
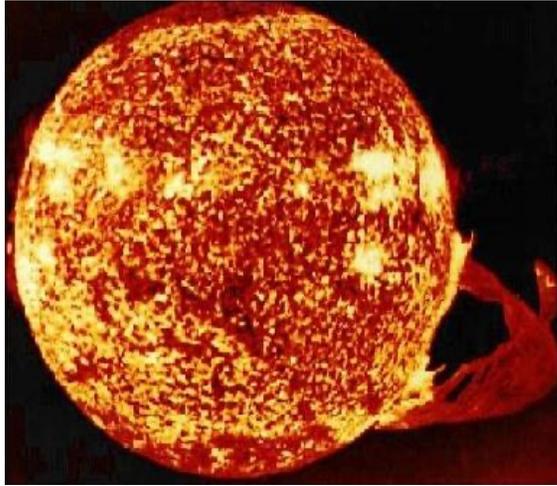
- A long road from `surprising' result to mystery to physics
- Most important element: tenacity
- Still things left to do...
- And we should at least keep watching...





# Anything Left?

## The solar 'metallicity problem'



- Helioseismology convinced 'everyone' that SSM was correct
- Modern measurements of surface metallicity are lower than before
- Which makes SSM helioseismologic predictions wrong...

But! CNO neutrinos tell us metallicity of solar core

→ Flux may differ by factor of 2 between old/new metallicity

(Maybe Jupiter and Saturn 'stole' metals from solar photosphere?)

---Haxton and Serenelli, *Astrophys.J.* 687 (2008)

# Gallium “Anomaly”

