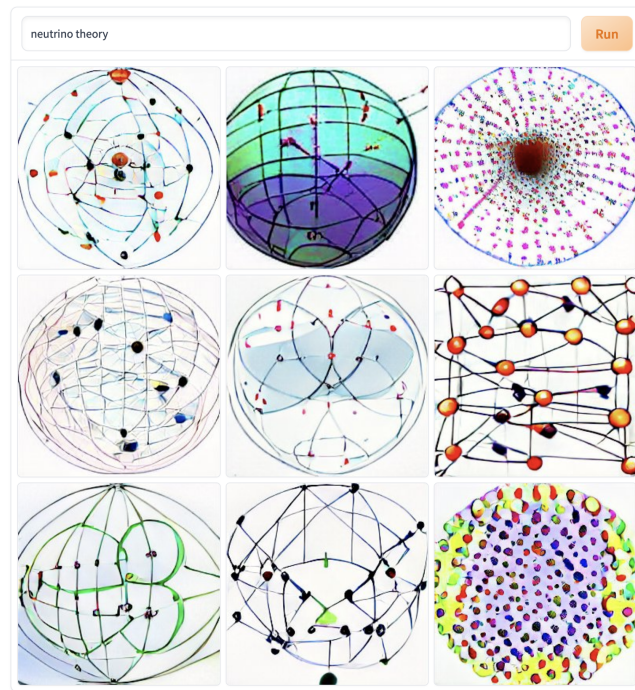


Lecture 1

Monday, 8 August 2022 17:44

Neutrino Theory

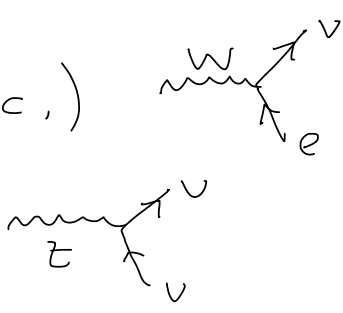
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1. Neutrino Masses1.1 Neutrinos in the SM

Quarks u c t
 d s b

Leptons e μ τ
 ν_e ν_μ ν_τ

- may elucidate origin of flavor
- produced in nuclear reactions
eg. $\pi \rightarrow \mu + \nu$
- may help explain baryon asymmetry
- may teach us about DM

$$\mathcal{L} \supset \sum_{\alpha=e,\mu,\tau} \left[\bar{\nu}_{\alpha L} i \not{\partial} \nu_{\alpha L} + \frac{g}{\sqrt{2}} (W^{\mu+} \nu_{\alpha L} \gamma_{\mu} e_{\alpha L} + h.c.) + \frac{g}{2 \cos \Theta_w} Z^{\mu} \bar{\nu}_{\alpha L} \gamma_{\mu} \nu_{\alpha L} \right] + \text{mass terms}$$


1.2 Dirac Masses

In the SM, only LH neutrinos exist

$$\nu_L = \frac{1-\gamma^5}{2} \nu = \begin{pmatrix} \chi_1 \\ \chi_2 \\ 0 \\ 0 \end{pmatrix}$$

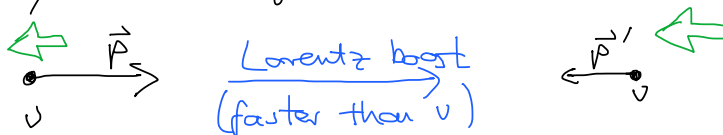
Dirac mass term makes $\nu_R = \frac{1+\gamma^5}{2} \nu = \begin{pmatrix} 0 \\ 0 \\ \chi_3 \\ \chi_4 \end{pmatrix}$

$$\mathcal{L}_m \supset - \sum_{\alpha,\beta} m_{\alpha\beta} \bar{\nu}_{\alpha L} \nu_{\beta R}$$

$\gamma_{\alpha\beta} < \nu >$

$\mathcal{L} \supset -y \bar{L} \tilde{H} \nu_R$

4 physical d.o.f.



Problem: why are $m_{\alpha\beta}$ so small

1.3 Majorana Masses

Mass terms couple LH and RH states

The antiparticle of ν_L is a RH state

Could ν_R be identical to the antiparticle of ν_L ?

More formally: charge conjugation

$$\hat{C} : \psi \rightarrow \psi^c \equiv -i\gamma^2 \gamma^0 \bar{\psi}^T$$

Effect on chirality:

$$\begin{aligned} \partial^- \Psi^- &= \gamma^0 (-i \gamma^1 \gamma^2) \Psi \\ &= +i \gamma^2 \gamma^5 \Psi^* \\ &= -(\gamma^5 \Psi)^c \end{aligned}$$

↳ \hat{C} flips chirality

Identify $\nu_R \equiv (\nu_L)^c$

In 4-component notation: $\nu_L = \begin{pmatrix} \chi \\ \chi^* \\ 0 \\ 0 \end{pmatrix} \equiv \begin{pmatrix} \chi \\ 0 \end{pmatrix}$

$$(\nu_L)^c = -i \begin{pmatrix} 0 & \sigma^2 \\ -\sigma^2 & 0 \end{pmatrix} \begin{pmatrix} \chi^* \\ 0 \end{pmatrix} = \begin{pmatrix} 0 \\ i\sigma^2 \chi^* \end{pmatrix}$$

$\Rightarrow \nu = \begin{pmatrix} \chi \\ i\sigma^2 \chi^* \end{pmatrix}$ 4-component Majorana spinor but only 2 d.o.f.

A new type of mass term:

$$\mathcal{L} \supset -\frac{1}{2} \sum_{\alpha, \beta} m_{\alpha\beta} \overline{(\nu_L)_\alpha} \nu_{L\beta} + h.c.$$

Complex Symmetric 3x3 matrix

Problem:

- why is $m_{\alpha\beta}$ so small
- How can this arise from an $SU(2)$ -invariant UV completion

1.4 The Seesaw Mechanism

Augment SM with 3 RH neutrinos N_R , singlet under the SM gauge group

For simplicity, consider first 1 LH ν_L , 1 RH N_R :

$$\mathcal{L} \supset -\underbrace{m_D}_{\text{from Higgs mechanism}} \overline{\nu_L} N_R - \frac{1}{2} \underbrace{m_M}_{\text{no natural scale could be very large}} \overline{(N_R)^c} N_R$$

natural scale is $m_D \sim m_H$

Write $\psi = \begin{pmatrix} \nu_L \\ N_R^c \end{pmatrix}$

$$\mathcal{L} \supset -\frac{1}{2} \bar{\nu}^c \underbrace{M}_{\substack{m_D \\ m_M}} \nu + h.c.$$

$$\begin{pmatrix} 0 & m_D \\ m_D & m_M \end{pmatrix}$$

$$= -\frac{1}{2} \begin{pmatrix} \bar{\nu}_L^c & \bar{N}_R \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D & m_M \end{pmatrix} \begin{pmatrix} \nu_L \\ N_R^c \end{pmatrix}$$

Diagonalize M to obtain physical states (with definite mass)

eigenvalues $\sim m_M \left(1 \pm \frac{m_D^2}{m_M^2} \right)$ (for $m_M \gg m_D$)
 absorbed by rephasing $\nu_L' \rightarrow i \nu_L'$

eigenvectors $N_R' = N_R + O\left(\frac{m_D}{m_M}\right)$

$\nu_L' = \nu_L + O\left(\frac{m_D}{m_M}\right)$

Effective low-E Lagrangian

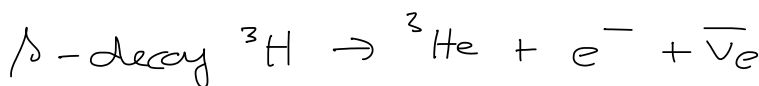
$$\mathcal{L} \supset -\frac{1}{2} \frac{m_D^2}{m_M} \overline{(\nu_L')^c} \nu_L'$$

With $m_D \sim 100 \text{ GeV}$ and $m_M \sim 10^{14} \text{ GeV}$

$\hookrightarrow m_\nu = \frac{m_D^2}{m_M} \sim 0.1 \text{ eV}$

1.5 Measuring ν masses

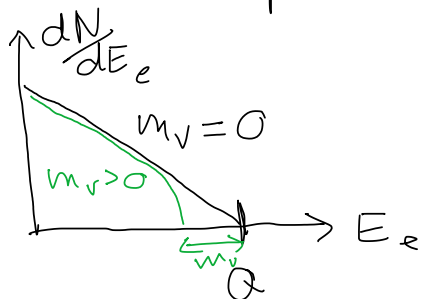
1.5.1 Kinematics



$$E_{e, \text{max}} = Q - m_\nu$$

$\underbrace{Q}_{m_H - m_{\text{He}}}$

Measure e^- spectrum, look at endpoint



1.5.2 Neutrinoless Double Beta Decay

\hookrightarrow see experimental lecture

1.5.3 Cosmology

Sensitive through large-scale structure

ν carry energy from overdense to underdense regions of the Universe

the larger m_ν , the larger this effect

↳ look for suppression on small scales

Current limits

$$\sum m_\nu < 0.12 \text{ eV}$$