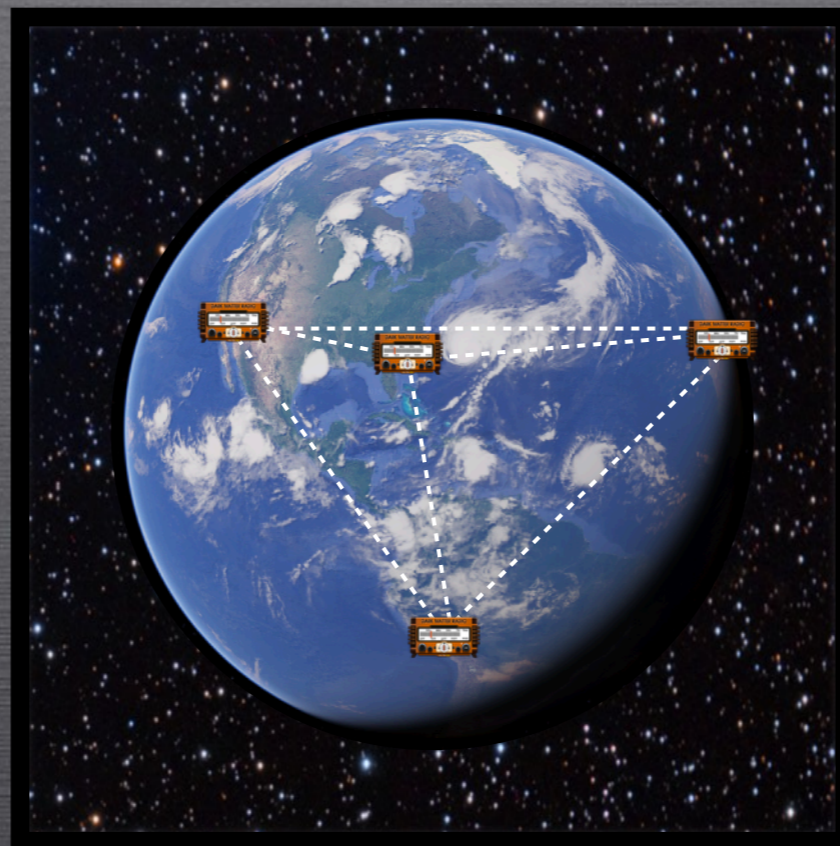




# DARK MATTER INTERFEROMETRY

## OR WHY 2 DETECTORS ARE BETTER THAN 1



**NICK RODD**

W/ JOSH FOSTER, YONI KAHN,  
RACHEL NGUYEN, AND BEN SAFDI

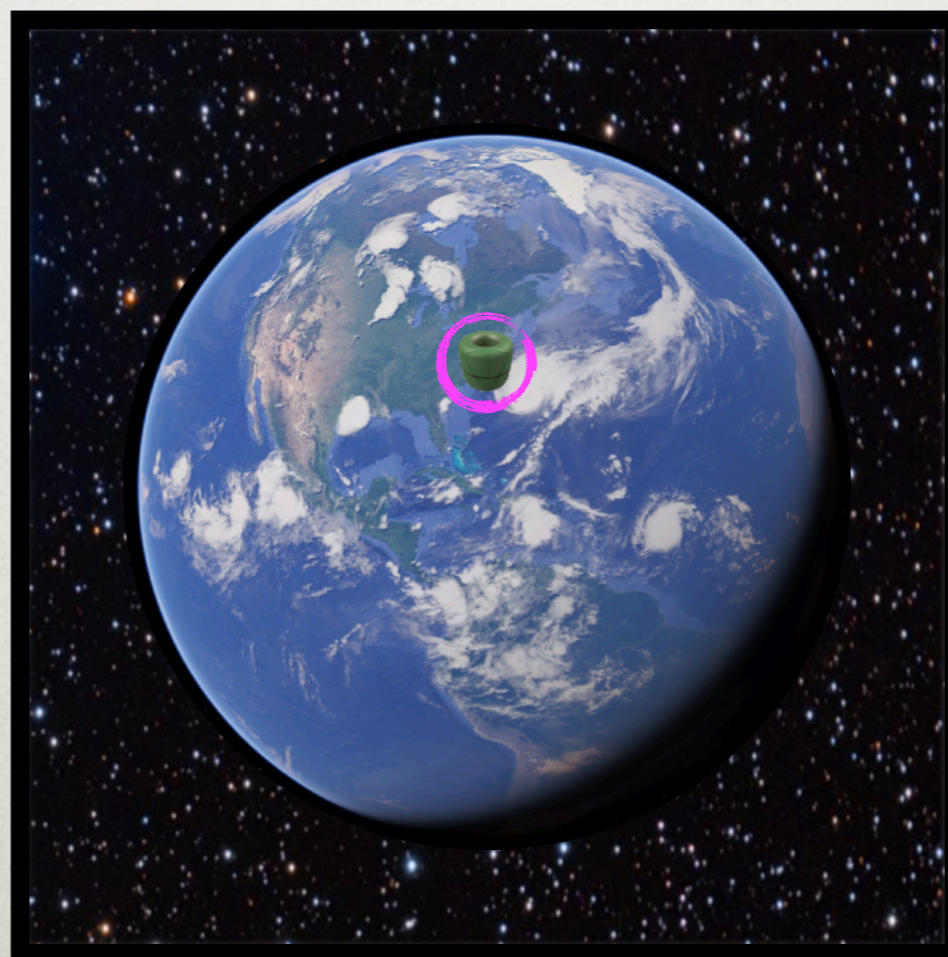
DM RADIO COLLABORATION MEETING, 14 AUGUST 2020





# BASIC IDEA

*Wave-like DM*

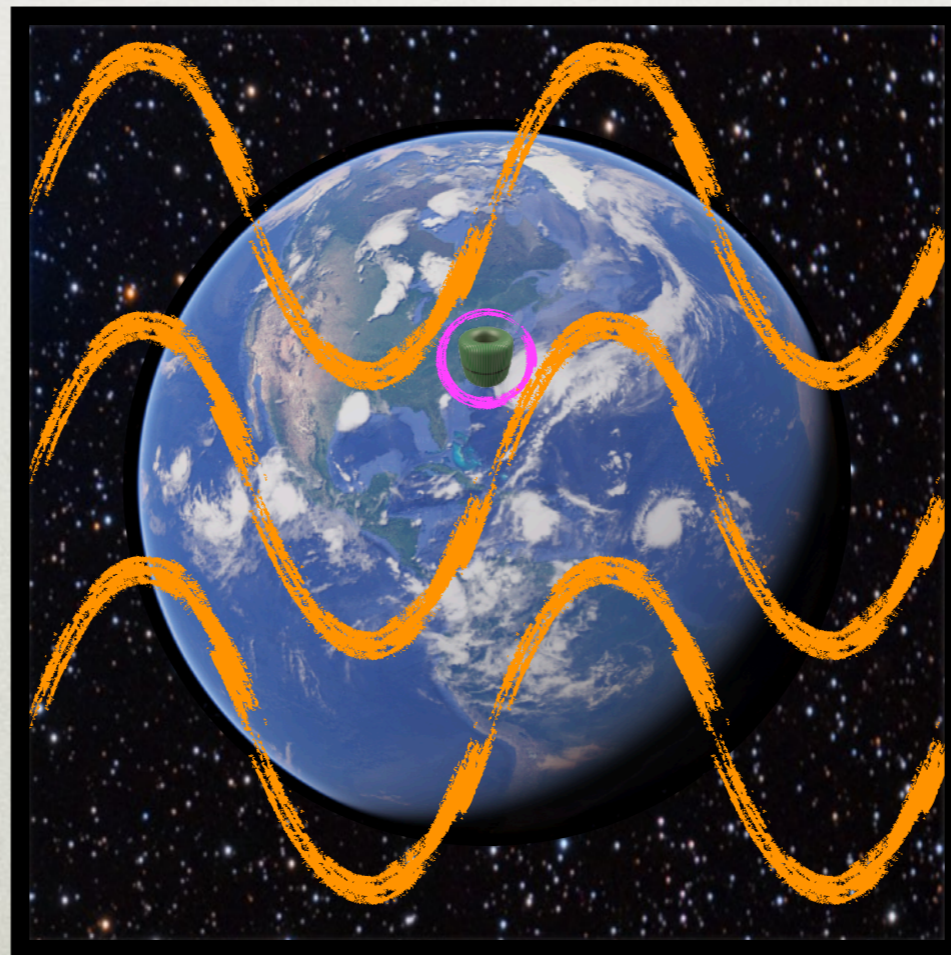




# BASIC IDEA

*Wave-like DM*

$$a \sim \cos(m_a t)$$

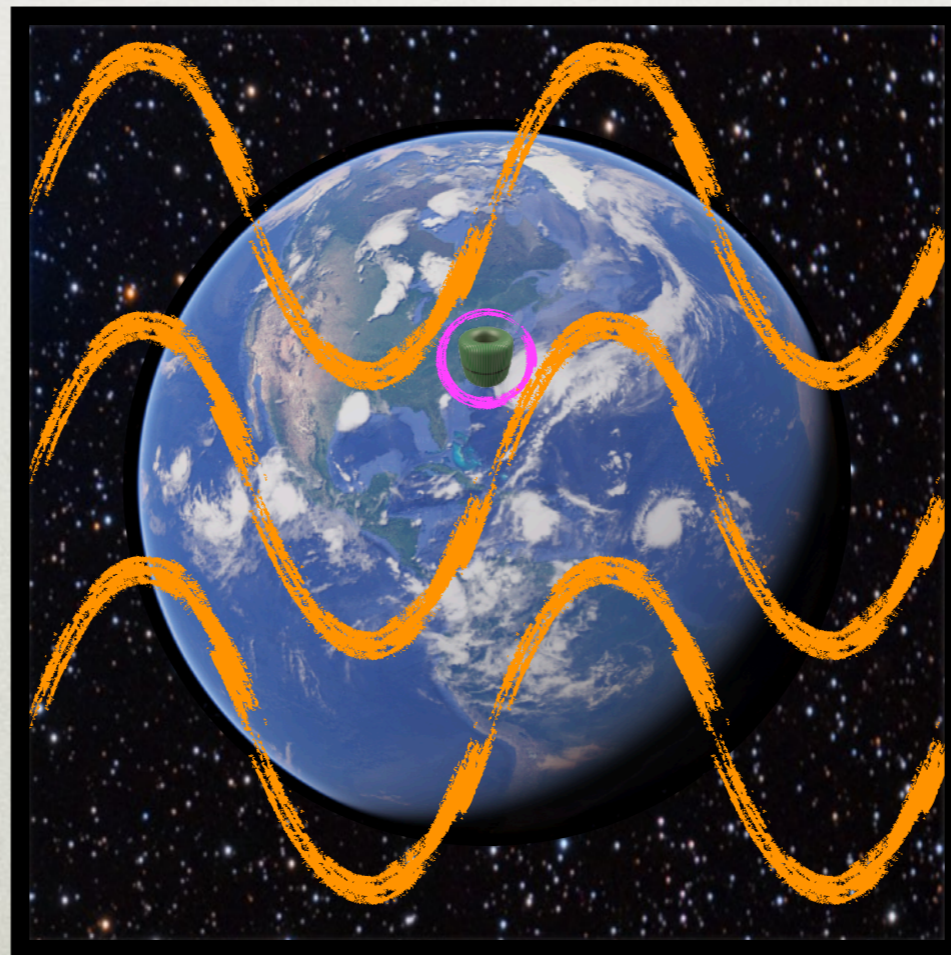




# BASIC IDEA

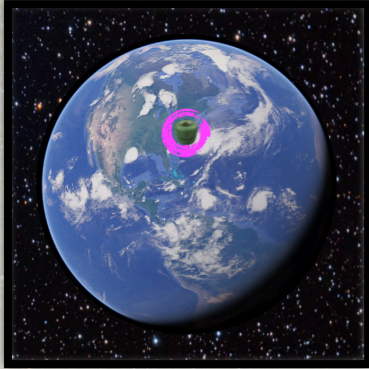
*Wave-like DM*

$$a \sim \cos(\omega t - \mathbf{k} \cdot \mathbf{x})$$





# BASIC IDEA

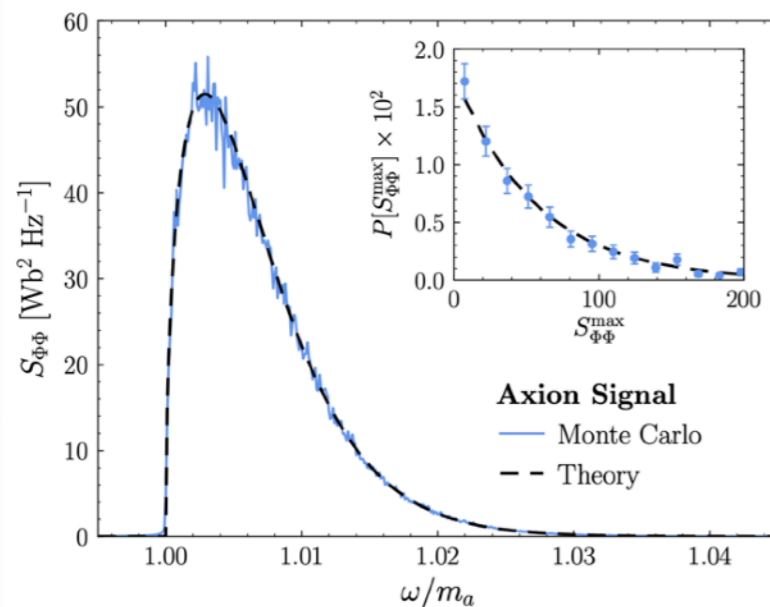


*Wave-like DM*

$$a \sim \cos(\omega t - \mathbf{k} \cdot \mathbf{x})$$

*Speed*  $\omega \approx m_a(1 + |\mathbf{v}|^2/2)$

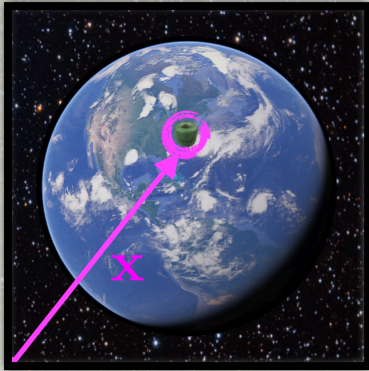
Determined by  $f(v)$



[Foster, NLR, Safdi 17]



# BASIC IDEA

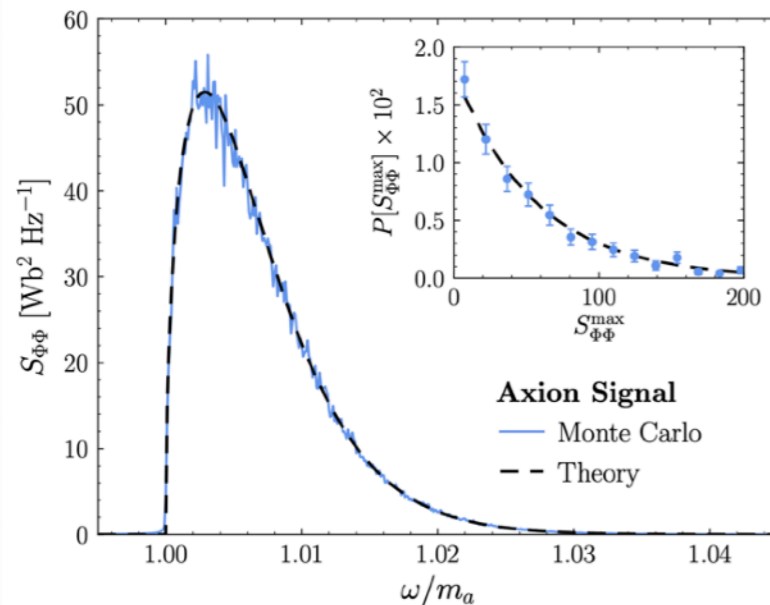


*Wave-like DM*

$$a \sim \cos(\omega t - \mathbf{k} \cdot \mathbf{x})$$

*Speed*  $\omega \approx m_a(1 + |\mathbf{v}|^2/2)$

Determined by  $f(v)$



[Foster, NLR, Safdi 17]

*Velocity*  $\mathbf{k} = m_a \mathbf{v}$

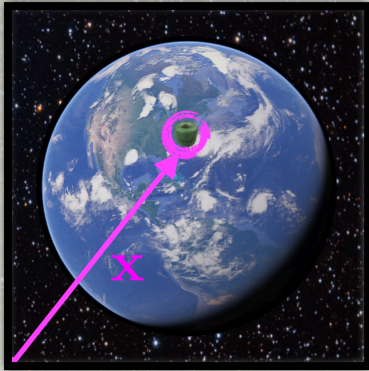
Determined by  $f(\mathbf{v})$

*But usually invisible to detectors*

1.  $\nabla a \sim \mathbf{k} \ll \omega \sim \partial_t a$
2. For 1 experiment, can choose  $\mathbf{x} = 0$



# BASIC IDEA

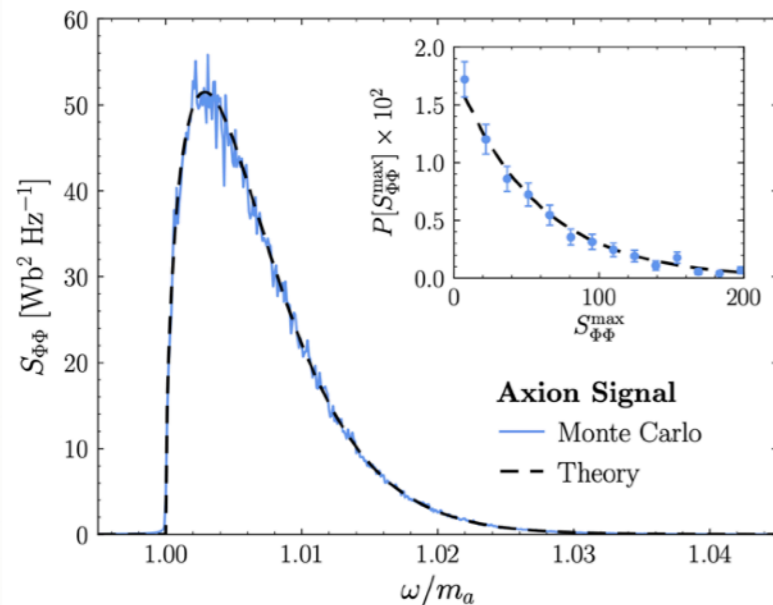


Wave-like DM

$$a \sim \cos(\omega t - \mathbf{k} \cdot \mathbf{x})$$

**Speed**  $\omega \approx m_a(1 + |\mathbf{v}|^2/2)$

Determined by  $f(v)$



[Foster, NLR, Safdi 17]

**Velocity**

$$\mathbf{k} = m_a \mathbf{v}$$

Determined by  $f(\mathbf{v})$

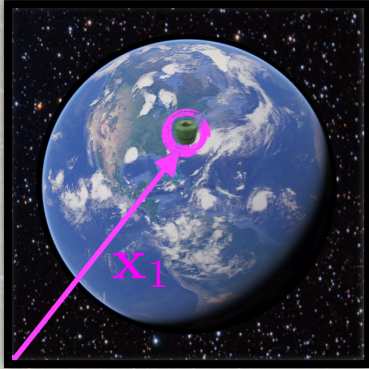
*But usually invisible to detectors*

1.  $\nabla a \sim \mathbf{k} \ll \partial_t a$
2. For 1 experiment, can choose  $\mathbf{x} = 0$

**OUTLINE**



# BASIC IDEA



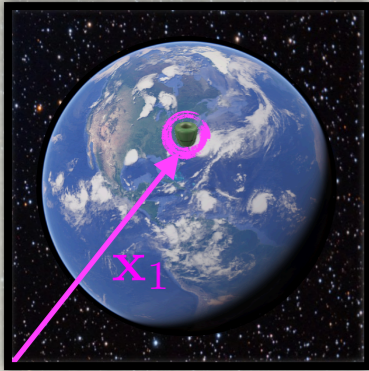
*Two Detectors*

Measurement at MIT

$$\Phi_1 \sim a(\mathbf{x}_1, t) \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t - \mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_1 + \phi_{\mathbf{v}}]$$



# BASIC IDEA



*Two Detectors*

Measurement at MIT

$$\Phi_1 \sim a(\mathbf{x}_1, t) \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t - \mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_1 + \phi_{\mathbf{v}}]$$

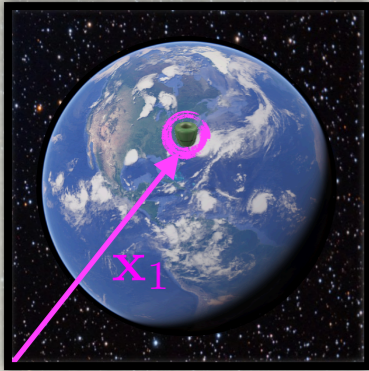
Superposition of  
velocities

$$a_0 \propto \sqrt{\rho_{\text{DM}}}$$

Random phase



# BASIC IDEA



*Two Detectors*

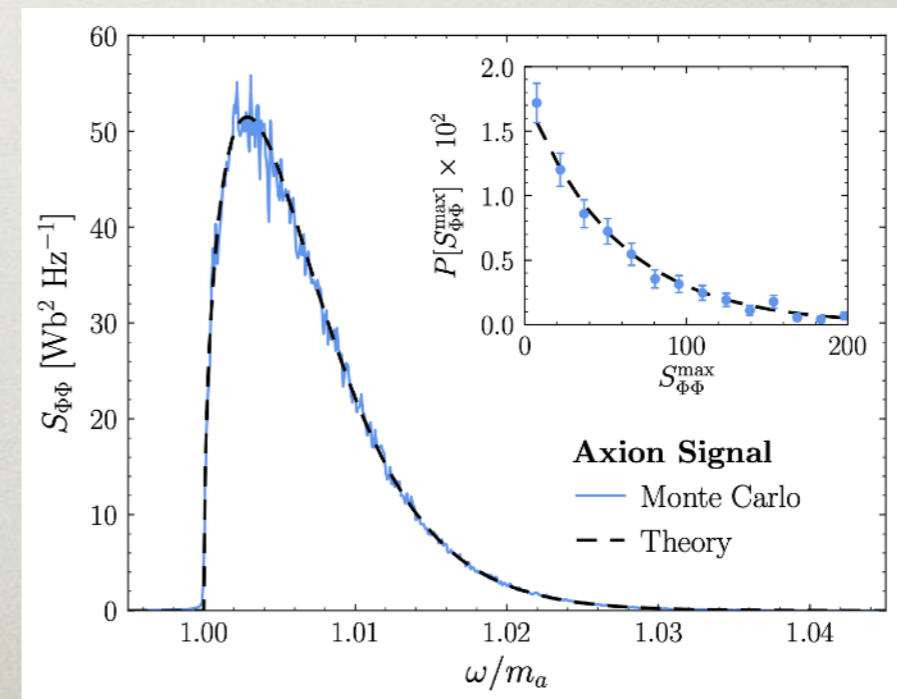
Measurement at MIT

$$\Phi_1 \sim a(\mathbf{x}_1, t) \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t - \mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_1 + \phi_{\mathbf{v}}]$$

$$\Rightarrow S_{\Phi\Phi}(\omega) \sim \frac{f(v_\omega)}{v_\omega}$$



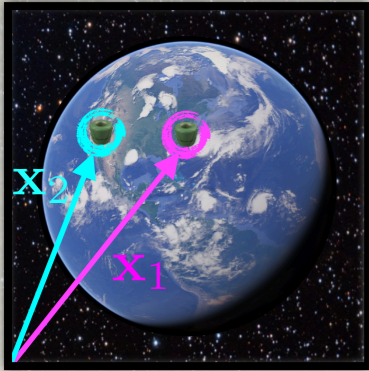
$$v_\omega = \sqrt{2\omega/m_a - 2}$$



[Foster, NLR, Safdi 17]



# BASIC IDEA



*Two Detectors*

Measurement at MIT

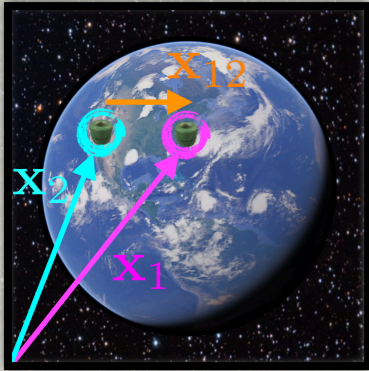
$$\Phi_1 \sim a(\mathbf{x}_1, t) \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t - \mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_1 + \phi_{\mathbf{v}}]$$

Measurement at SLAC

$$\Phi_2 \sim a(\mathbf{x}_2, t) \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t - \mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_2 + \phi_{\mathbf{v}}]$$



# BASIC IDEA



*Two Detectors*

Fruits of a collaboration\*

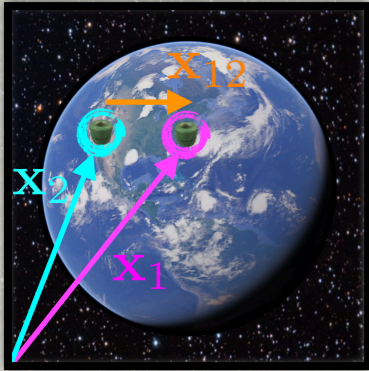
$$\Phi_1 + \Phi_2 \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t + \phi_{\mathbf{v}}] \cos[\mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_{12}/2]$$

$$\mathbf{x}_{12} = \mathbf{x}_1 - \mathbf{x}_2$$

\*Requires  $\mu\text{-ns}$  level synchronization  
(exact value depends on  $m$ )



# BASIC IDEA



*Two Detectors*

Fruits of a collaboration

$$\Phi_1 + \Phi_2 \sim \sum_{\mathbf{v}} \sqrt{f(\mathbf{v})} \cos[\omega_{\mathbf{v}} t + \phi_{\mathbf{v}}] \cos[\mathbf{k}_{\mathbf{v}} \cdot \mathbf{x}_{12}/2]$$

$$\Rightarrow S_{\Phi\Phi}(\omega) \sim \frac{4}{v_{\omega}} \int d^3\mathbf{v} f(\mathbf{v}) \cos^2[m_a \mathbf{v} \cdot \mathbf{x}_{12}/2] \delta[|\mathbf{v}| - v_{\omega}]$$

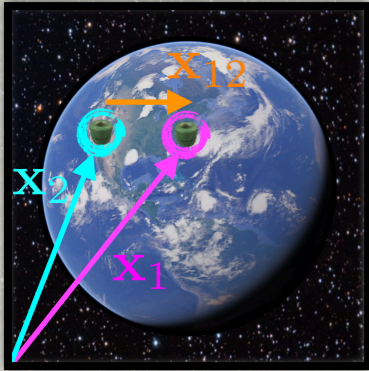
PSD remains exponentially distributed

$\Rightarrow$  full likelihood formalism

c.f. [Foster, NLR, Safdi 17]



# BASIC IDEA



## *Two Detectors*

c.f. 1 detector PSD:

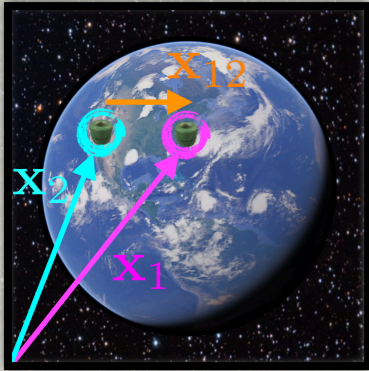
$$f(v) \rightarrow \mathcal{F}(v) = 4 \int d^3 \mathbf{v} f(\mathbf{v}) \cos^2 [m_a \mathbf{v} \cdot \mathbf{x}_{12} / 2] \delta[|\mathbf{v}| - v_\omega]$$

Interference leads to a modulated speed distribution

- *Contains information invisible to one detector*
- Result varies with  $x_{12}$  - daily modulation
- Simple generalization to  $\mathcal{N}$  detectors



# BASIC IDEA



*Two Detectors*

## OUTLINE

(FOR THE LAST FEW MINUTES)

1. Toy Example: scalings & intuition
2. Realistic Case: estimating  $\mathbf{v}_\odot$

c.f. 1 detector

$$f(v) \rightarrow \mathcal{F}(v)$$

$$\delta[|\mathbf{v}| - v_\omega]$$

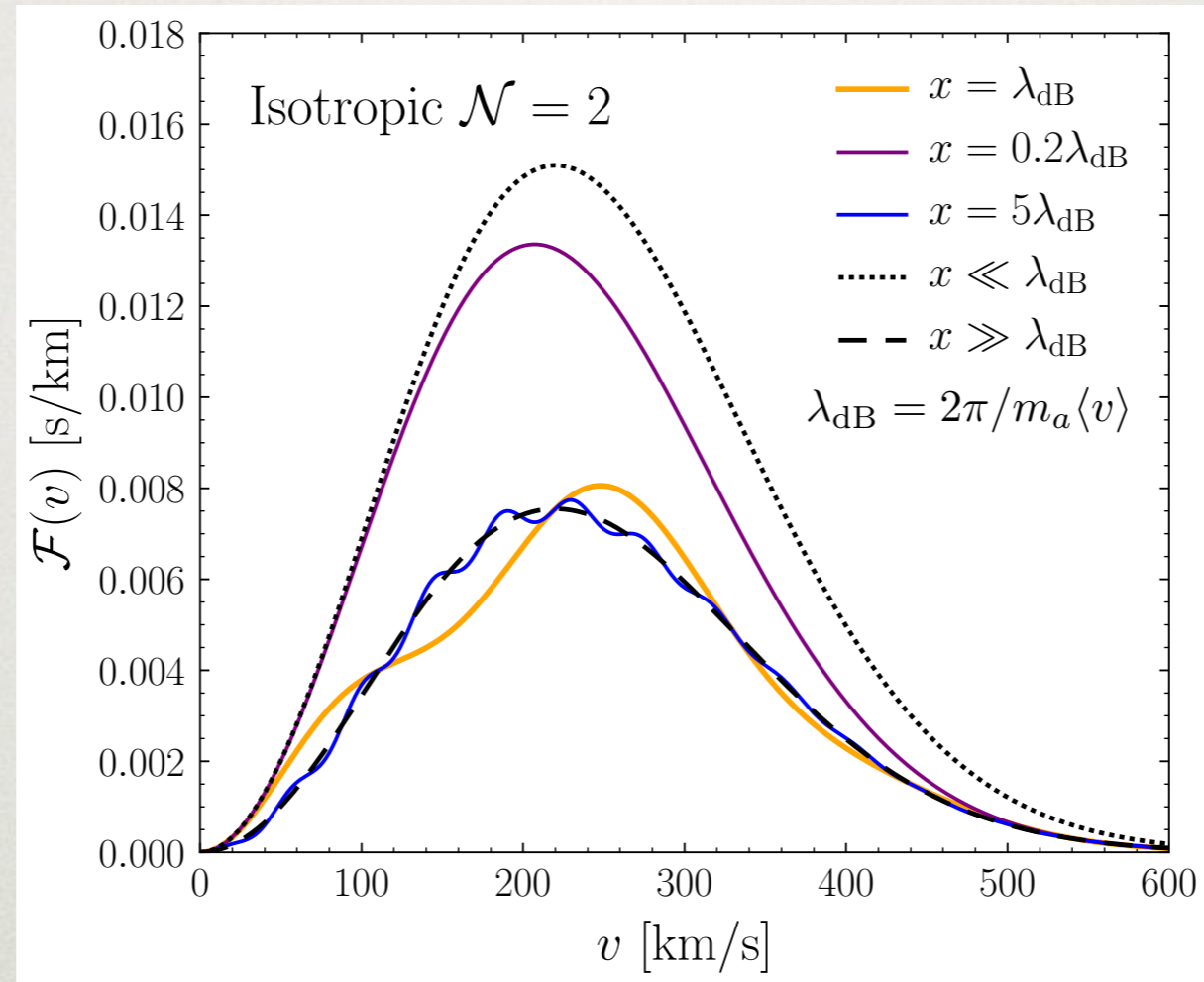
Interference leads to a modulated speed distribution

- *Contains information invisible to one detector*
- Result varies with  $x_{12}$  - daily modulation
- Simple generalization to  $\mathcal{N}$  detectors





# TOY EXAMPLE

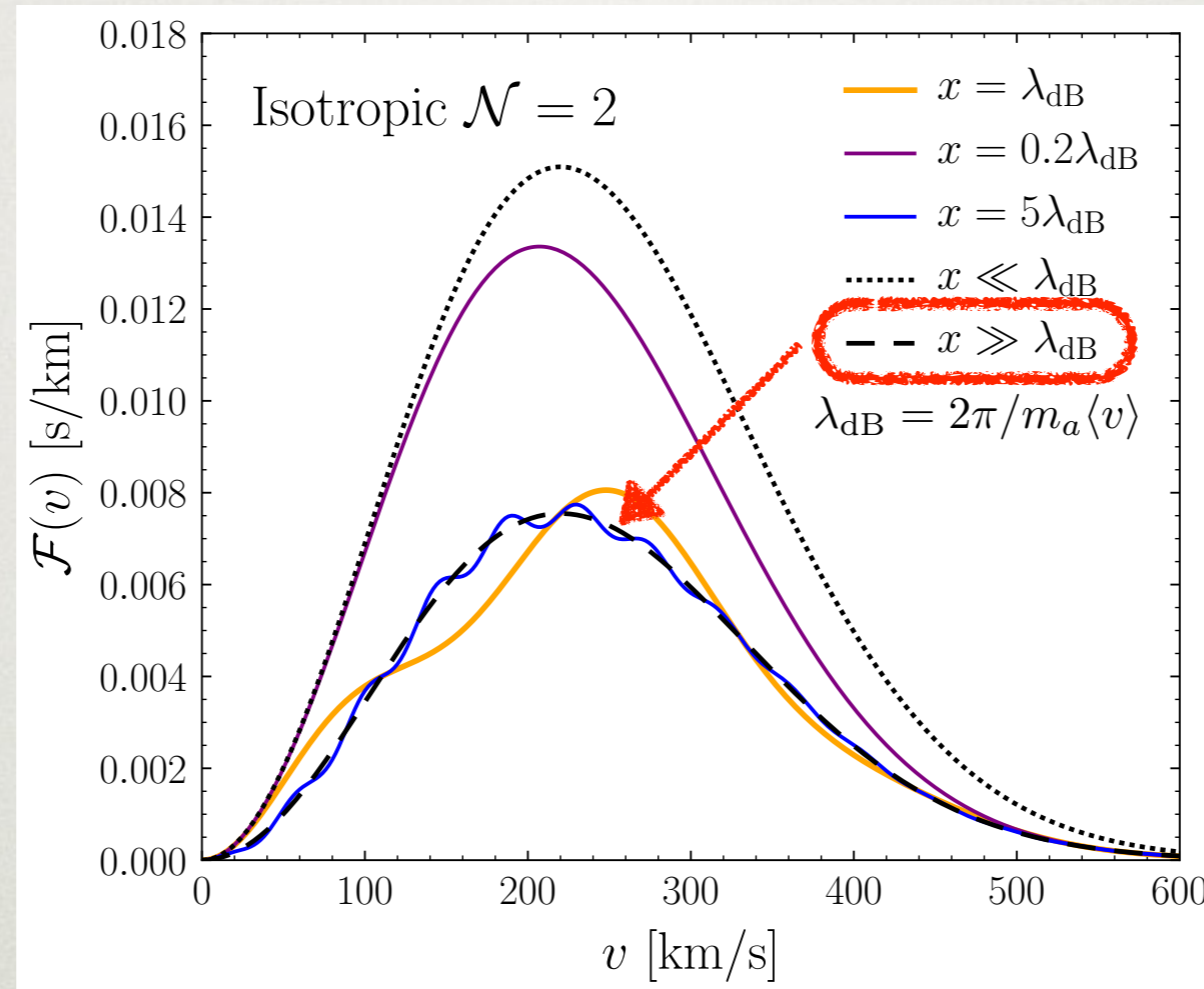


$$\mathcal{F}(v) = 4 \int d^3\mathbf{v} f(\mathbf{v}) \cos^2[m_a \mathbf{v} \cdot \mathbf{x} / 2] \delta[|\mathbf{v}| - v]$$





# TOY EXAMPLE



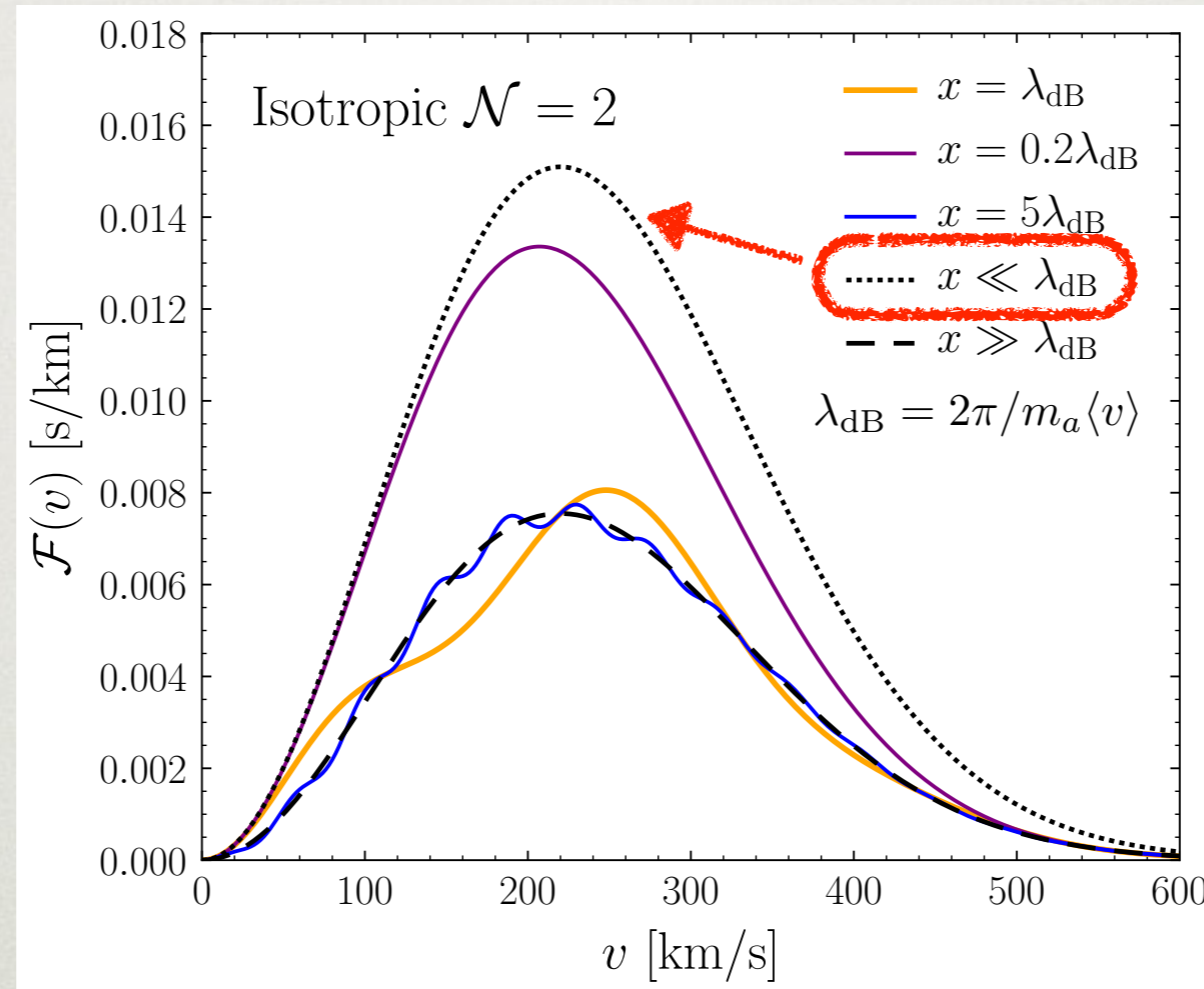
$$\mathcal{F}(v) = 4 \int d^3\mathbf{v} f(\mathbf{v}) \underbrace{\cos^2[m_a \mathbf{v} \cdot \mathbf{x}/2]}_{\rightarrow 1/2} \delta[|\mathbf{v}| - v] = 2f(v)$$

**INCOHERENT**





# TOY EXAMPLE



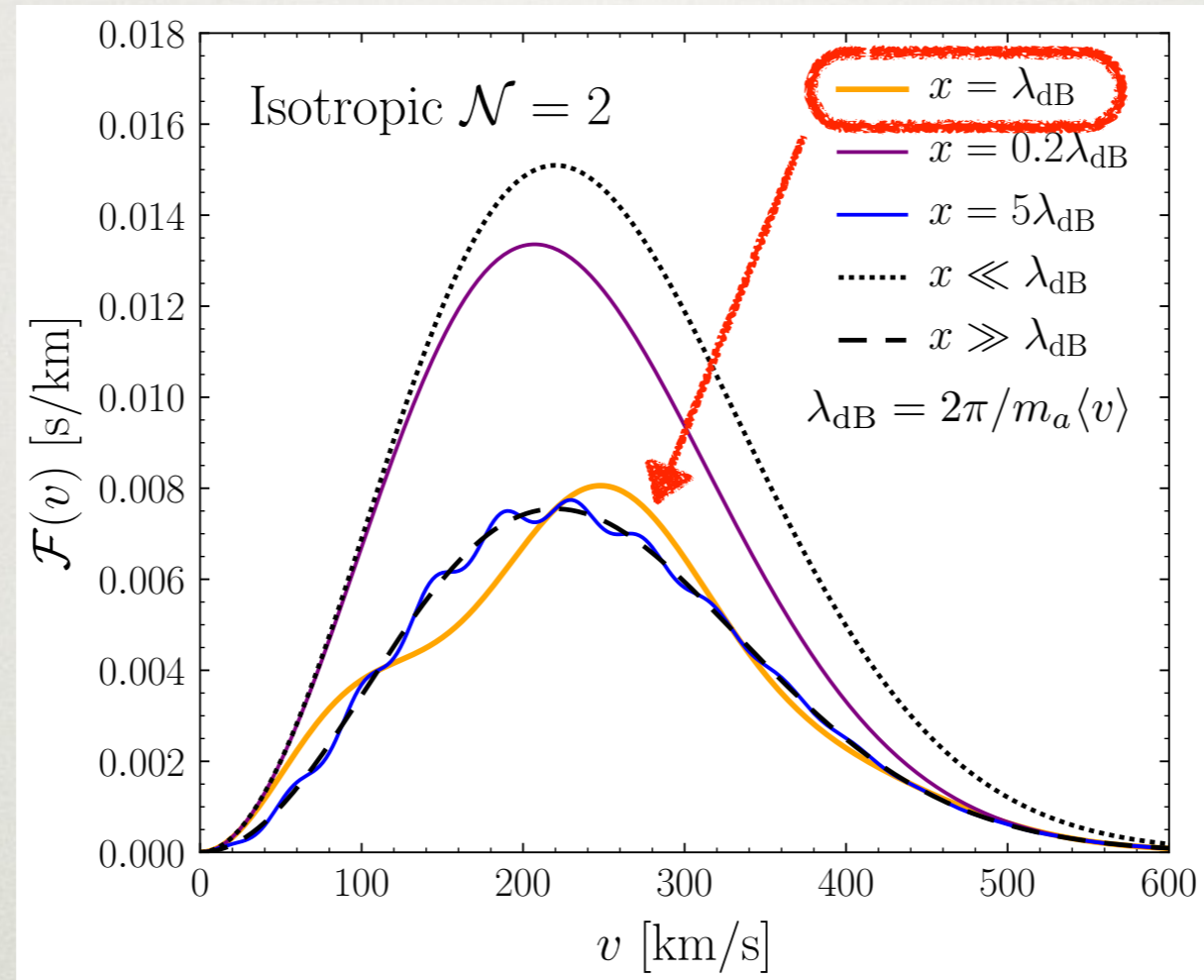
$$\mathcal{F}(v) = 4 \int d^3\mathbf{v} f(\mathbf{v}) \underbrace{\cos^2[m_a \mathbf{v} \cdot \mathbf{x}/2]}_{\rightarrow 1} \delta[|\mathbf{v}| - v] = 4f(v)$$

**COHERENT**





# TOY EXAMPLE

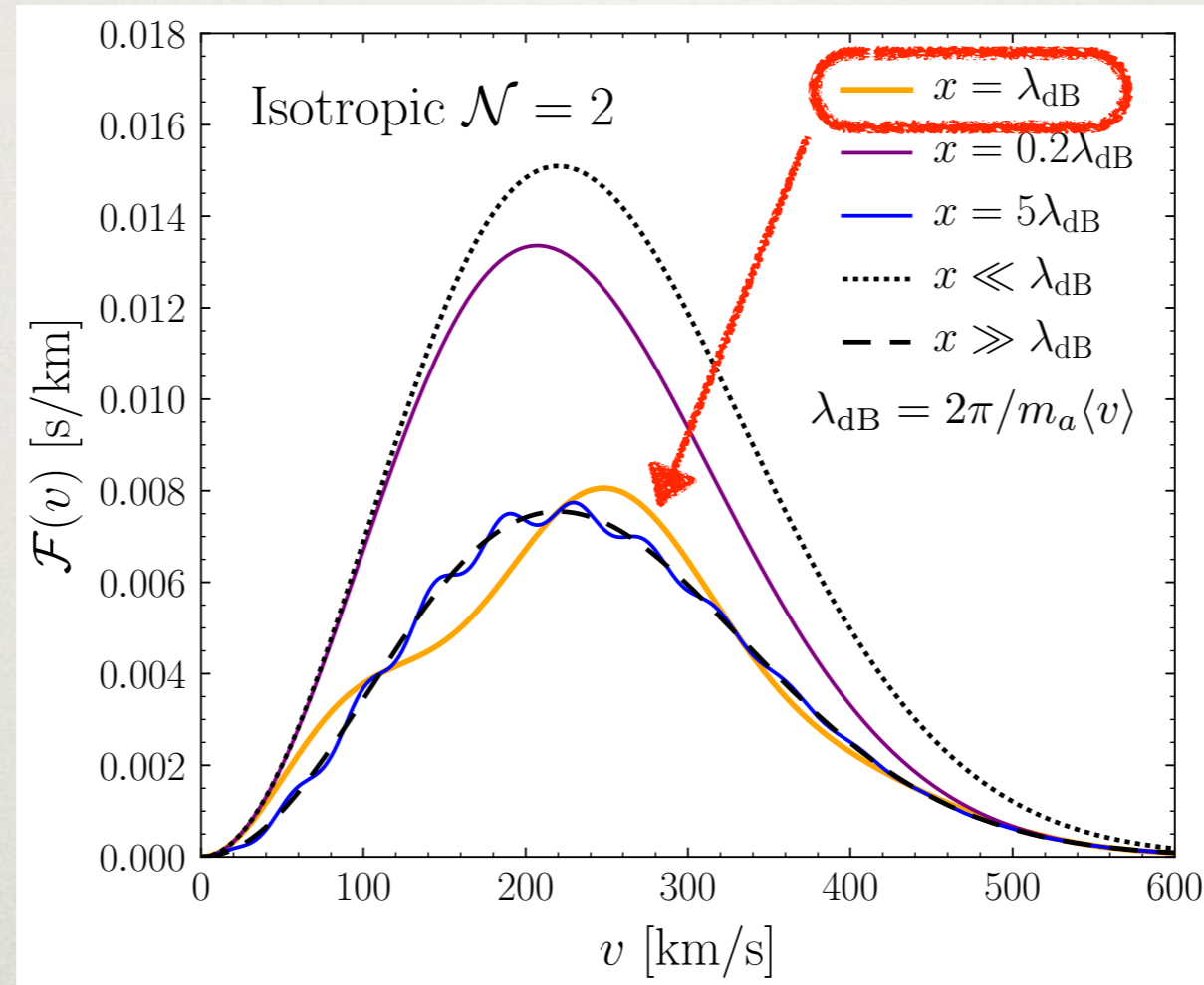


$$\mathcal{F}(v) = 4 \int d^3\mathbf{v} f(\mathbf{v}) \cos^2[m_a \mathbf{v} \cdot \mathbf{x}/2] \delta[|\mathbf{v}| - v] \propto f(v)$$





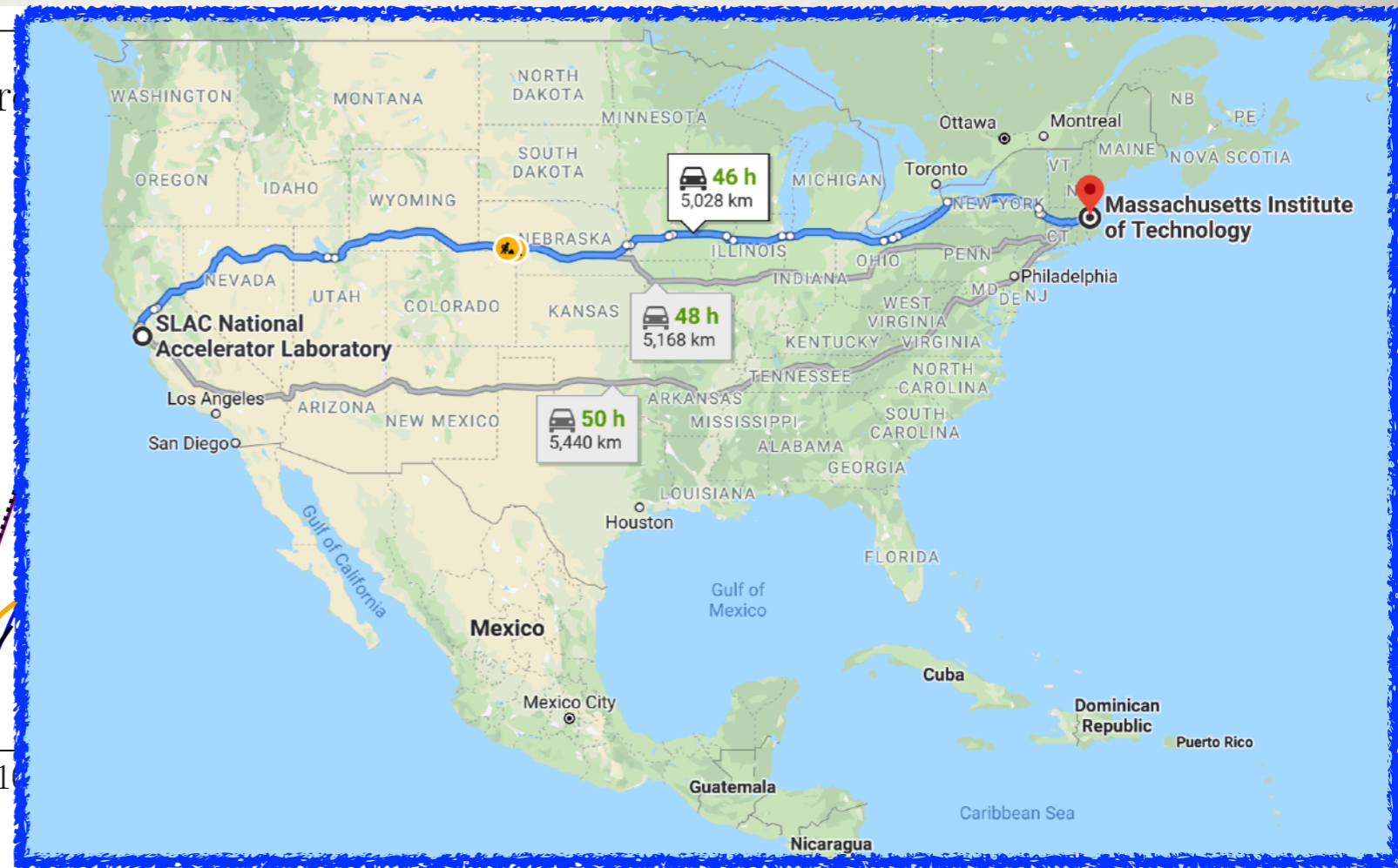
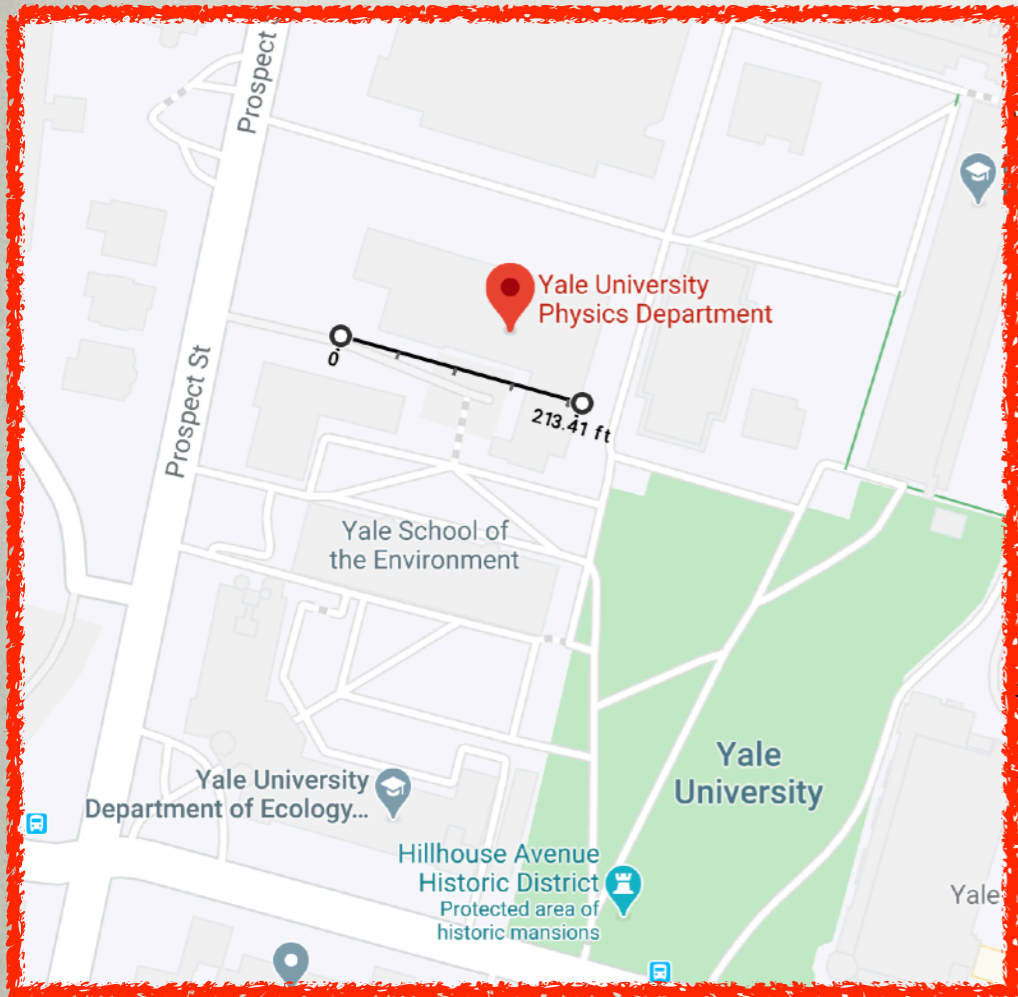
# TOY EXAMPLE



1.  $m_a = 25 \mu\text{eV} \Rightarrow \lambda_{\text{dB}} \approx 50 \text{ m}$
2.  $m_a = 1 \text{ neV} \Rightarrow \lambda_{\text{dB}} \approx 2,000 \text{ km}$



# TOY EXAMPLES



1.  $m_a = 25 \mu\text{eV} \Rightarrow \lambda_{\text{dB}} \approx 50 \text{ m}$

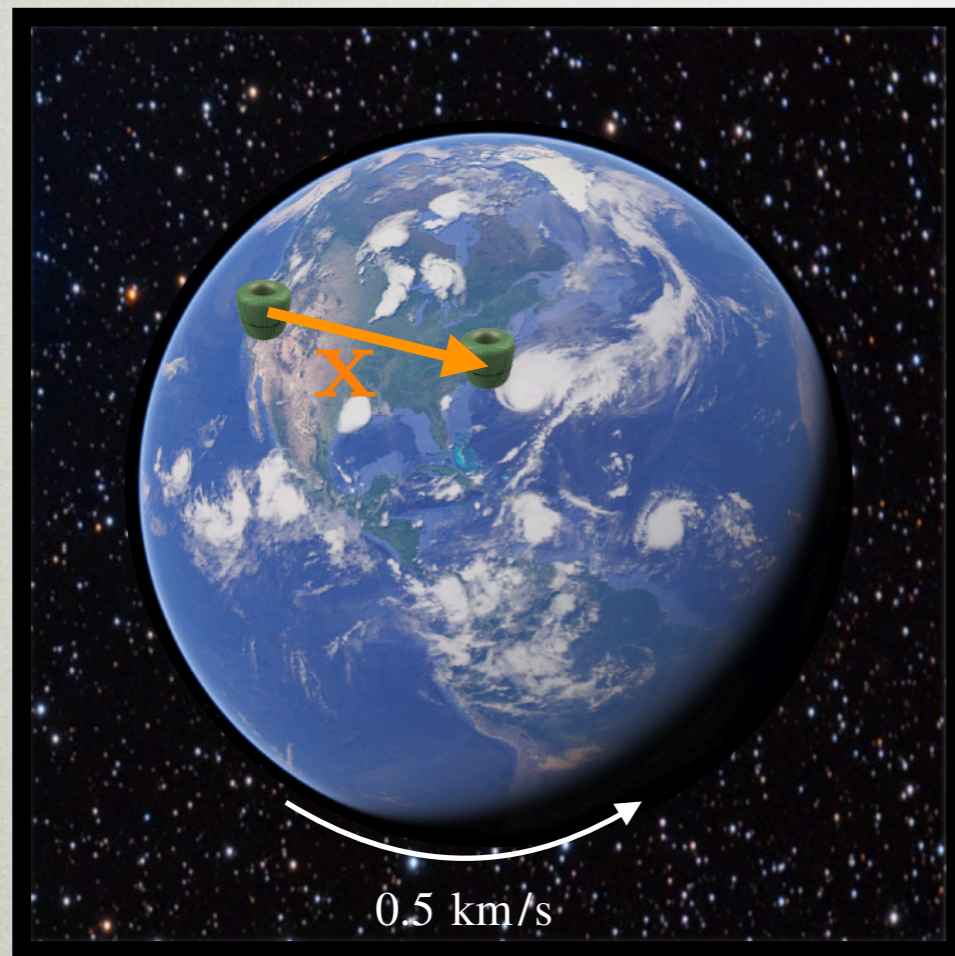
2.  $m_a = 1 \text{ neV} \Rightarrow \lambda_{\text{dB}} \approx 2,000 \text{ km}$





# DAILY MODULATION

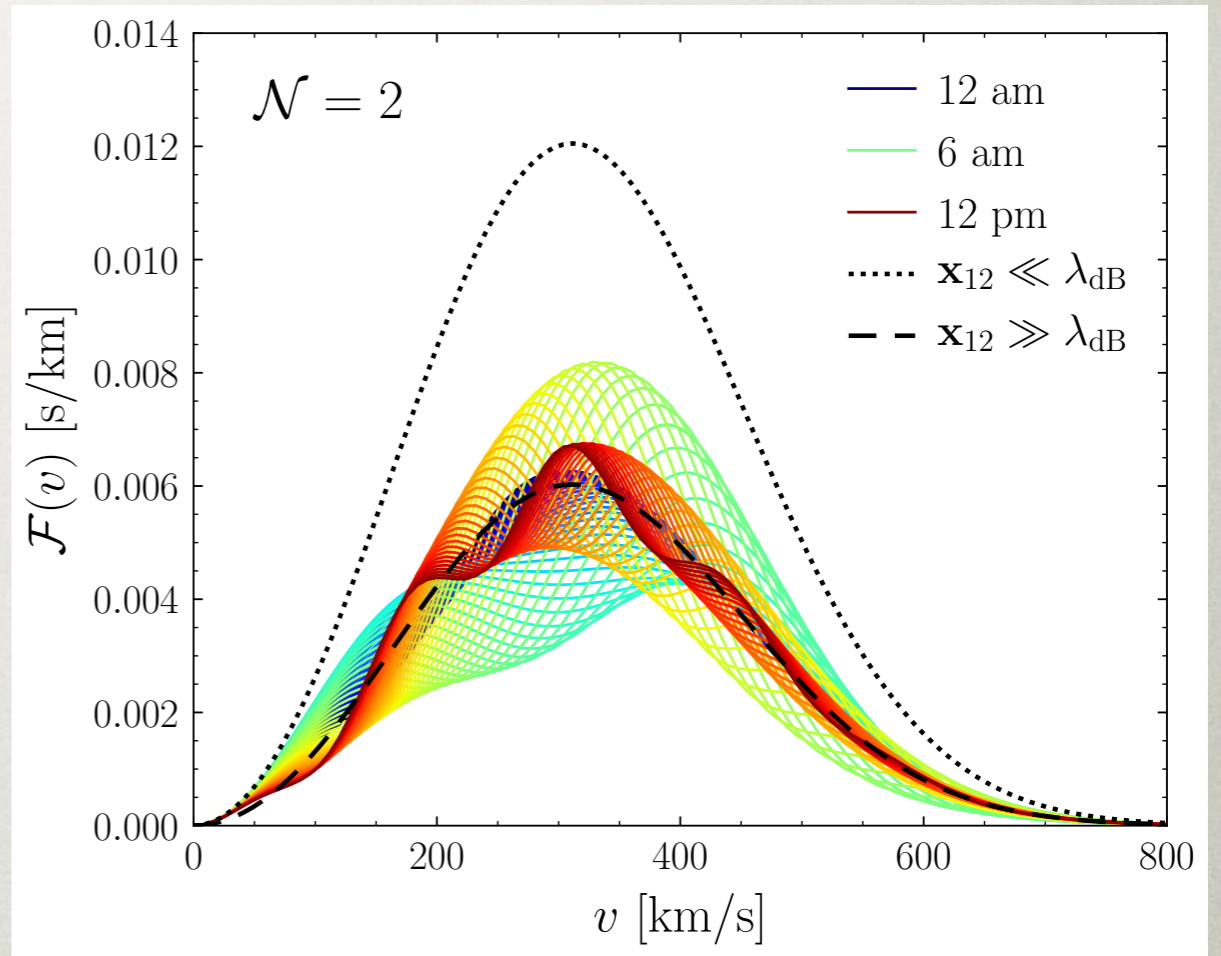
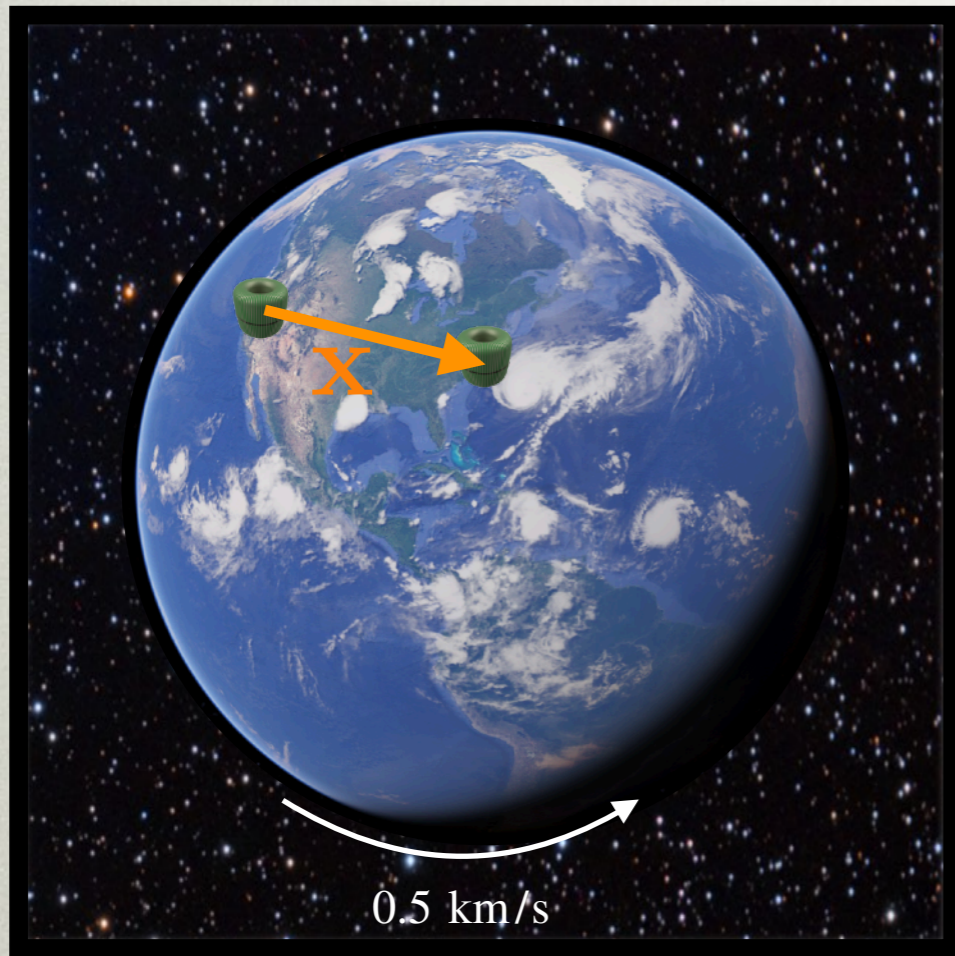
*A Unique Wave Dark Matter Signal*





# DAILY MODULATION

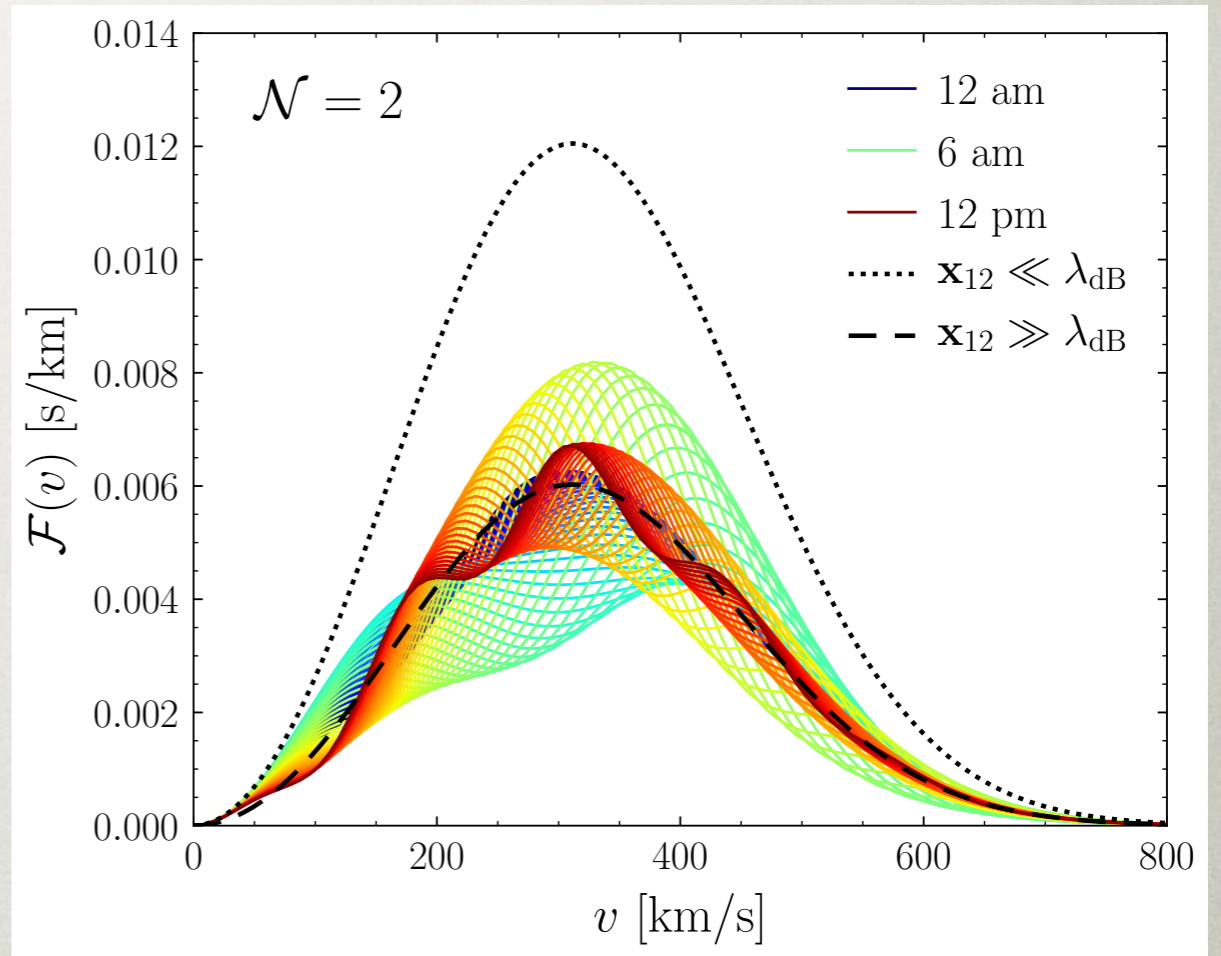
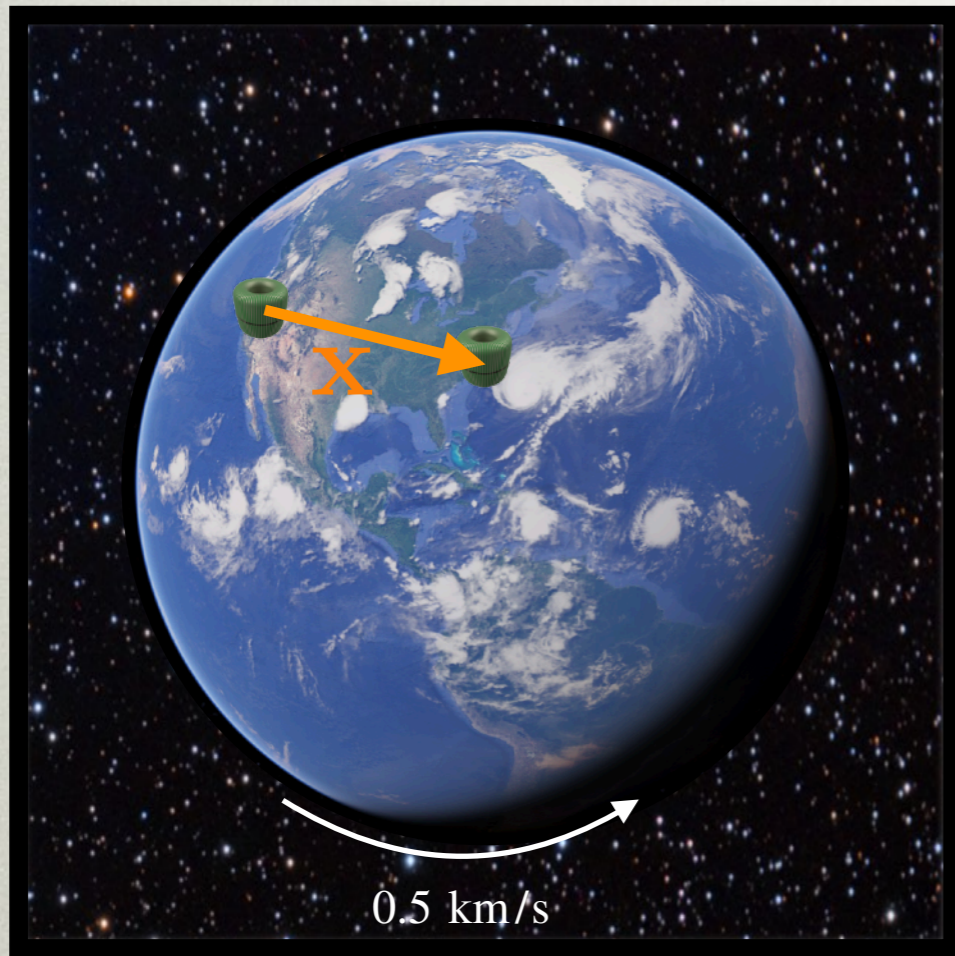
*A Unique Wave Dark Matter Signal*





# DAILY MODULATION

*A Unique Wave Dark Matter Signal*

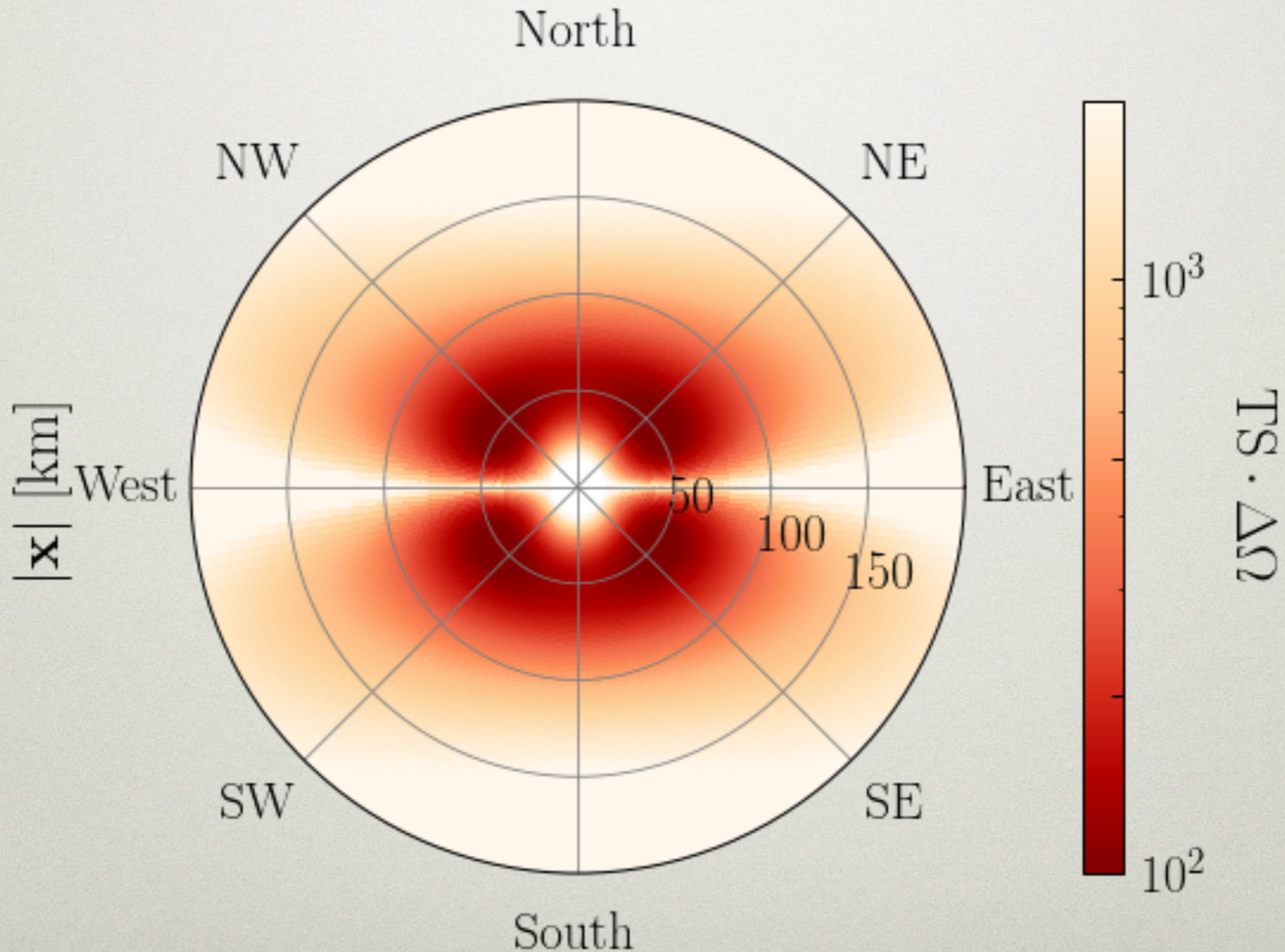


How well can we estimate the direction of  $\mathbf{v}_{\odot}$ ?



# DAILY MODULATION

*A Unique Wave Dark Matter Signal*



PRELIMINARY

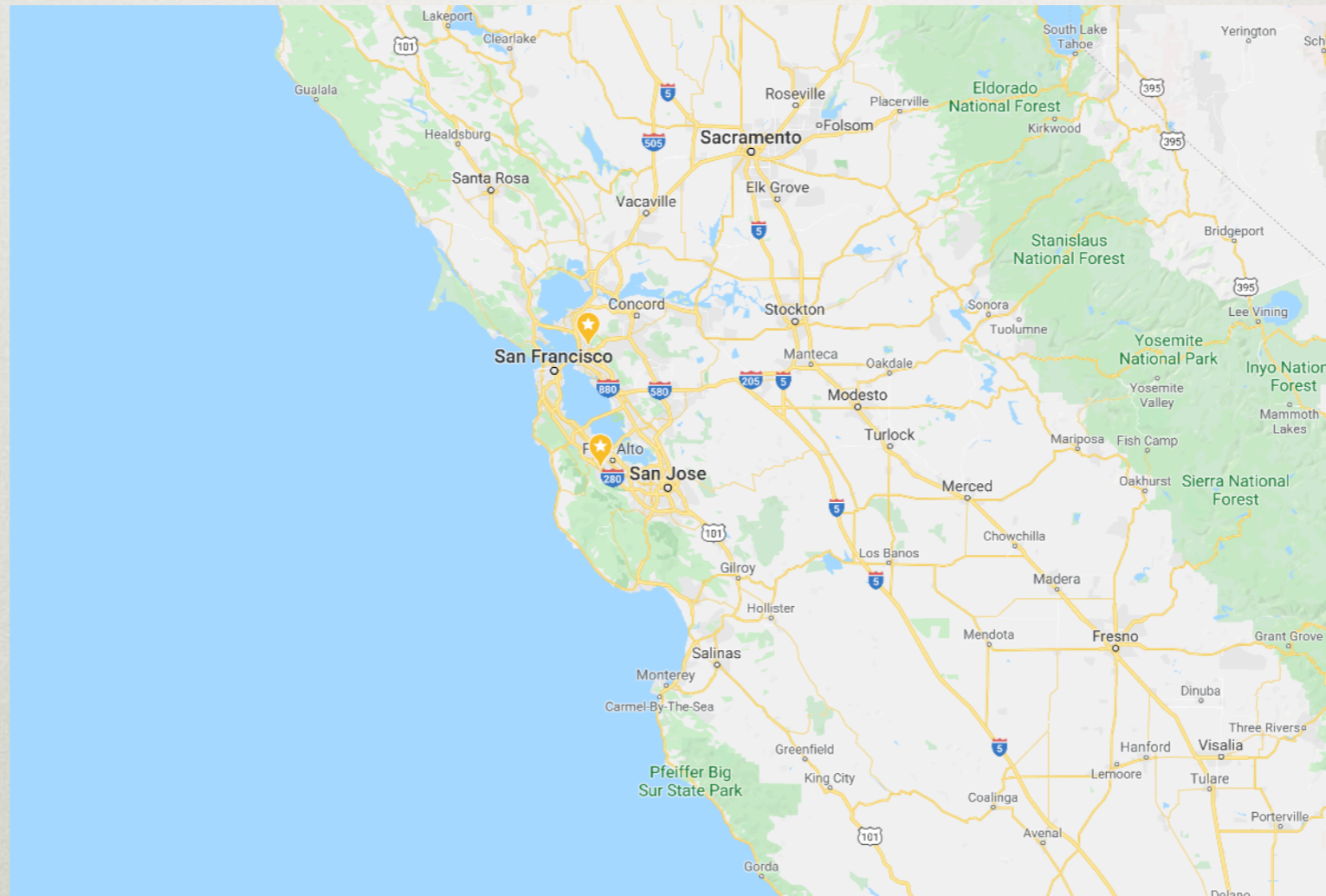
PRELIMINARY

$$m_a = 10 \text{ neV}$$



# DAILY MODULATION

*A Unique Wave Dark Matter Signal*



PRELIMINARY

PRELIMINARY

$$m_a = 10 \text{ neV}$$

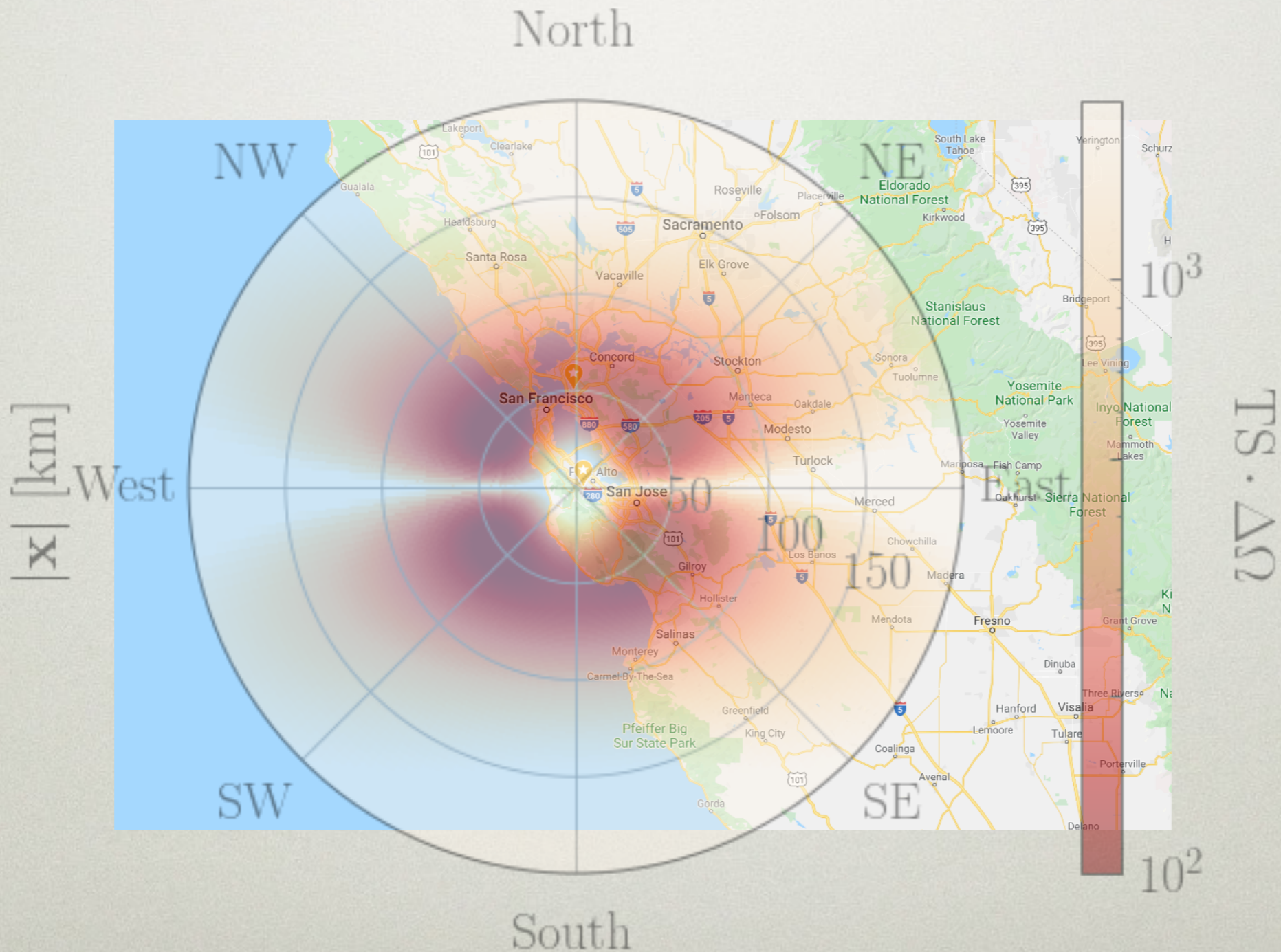


# DAILY MODULATION

*A Unique Wave Dark Matter Signal*

PRELIMINARY

PRELIMINARY



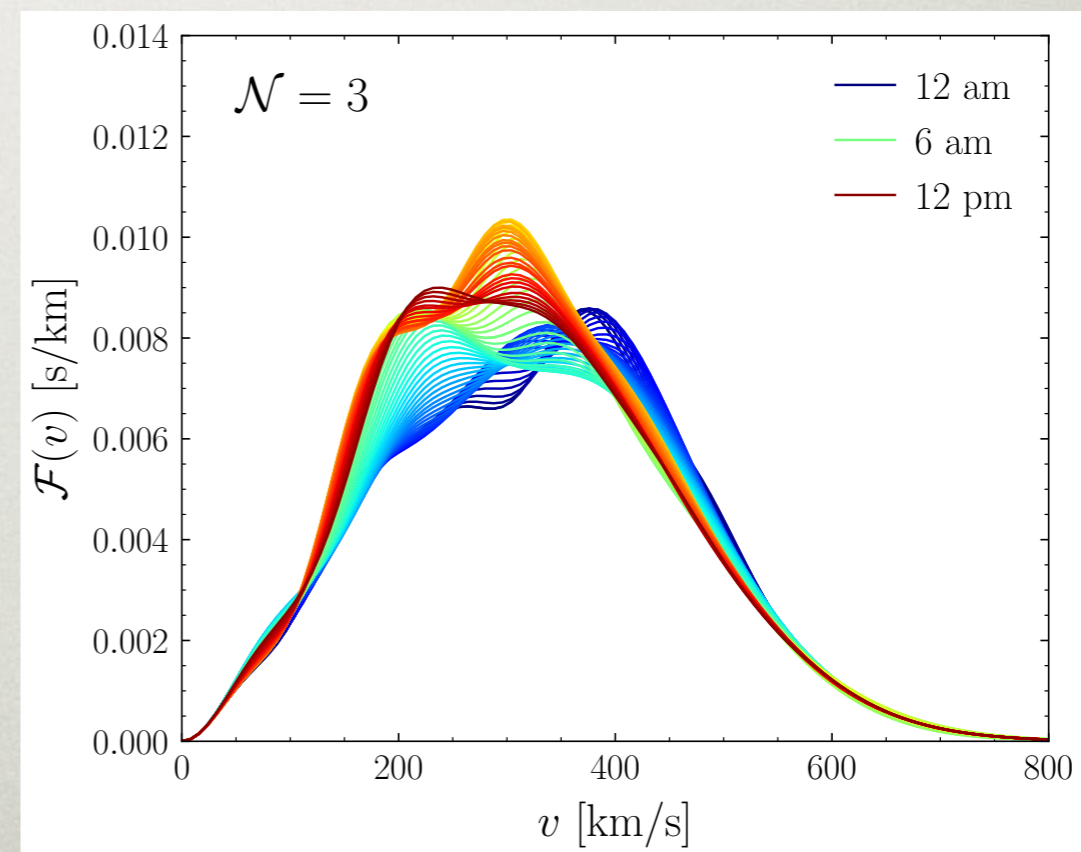
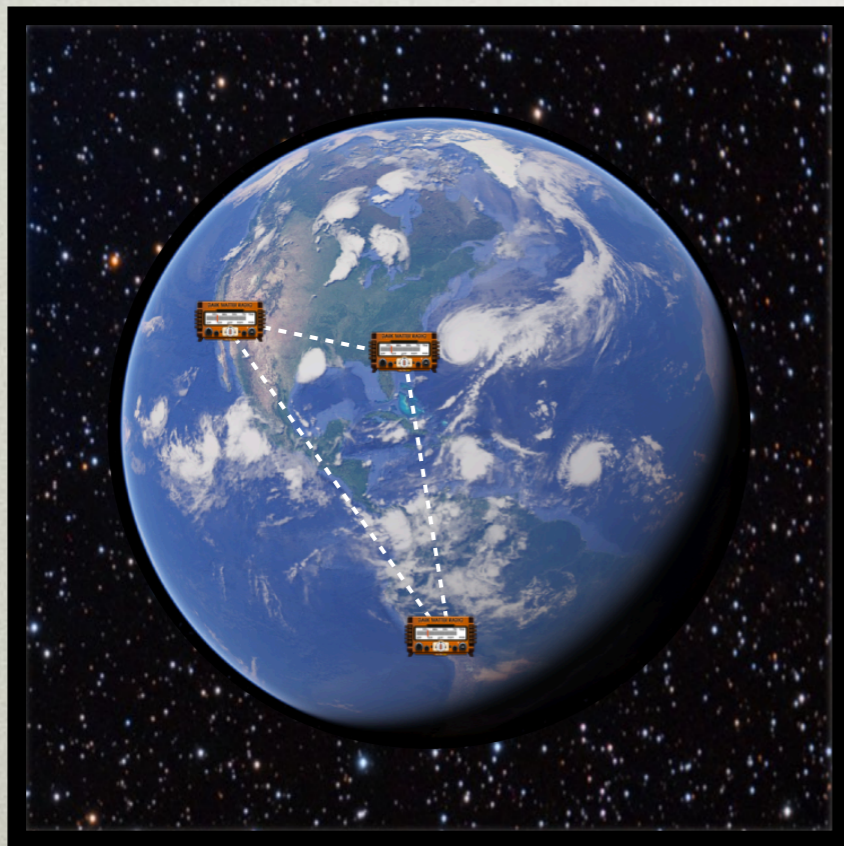
$$m_a = 10 \text{ neV}$$



# CONCLUSION

## *A Unique Wave Dark Matter Signal*

- Encodes information invisible to 1 detector, e.g.  $v_{\odot}$
- Carries the unique fingerprint of dark matter







# BACKUP SLIDES



# RESULT FOR A GENERAL $\mathcal{N}$



$$\Phi(\mathbf{x}, t) = \sqrt{A} a(\mathbf{x}, t)$$

$$\langle S_{\Phi\Phi}(\omega) \rangle = A \frac{\pi \mathcal{F}(v_\omega)}{m_a v_\omega}$$

Exponentially Distributed

$$\mathcal{F}(v) = \mathcal{N} \int d^3\mathbf{v} f(\mathbf{v}) \mathcal{X}(\mathbf{v}) \delta[|\mathbf{v}| - v]$$

$$\mathcal{X}(\mathbf{v}) = 1 + \frac{2}{\mathcal{N}} \sum_{(ij) \in \mathcal{C}} \cos[m_a \mathbf{v} \cdot \mathbf{x}_{ij}]$$

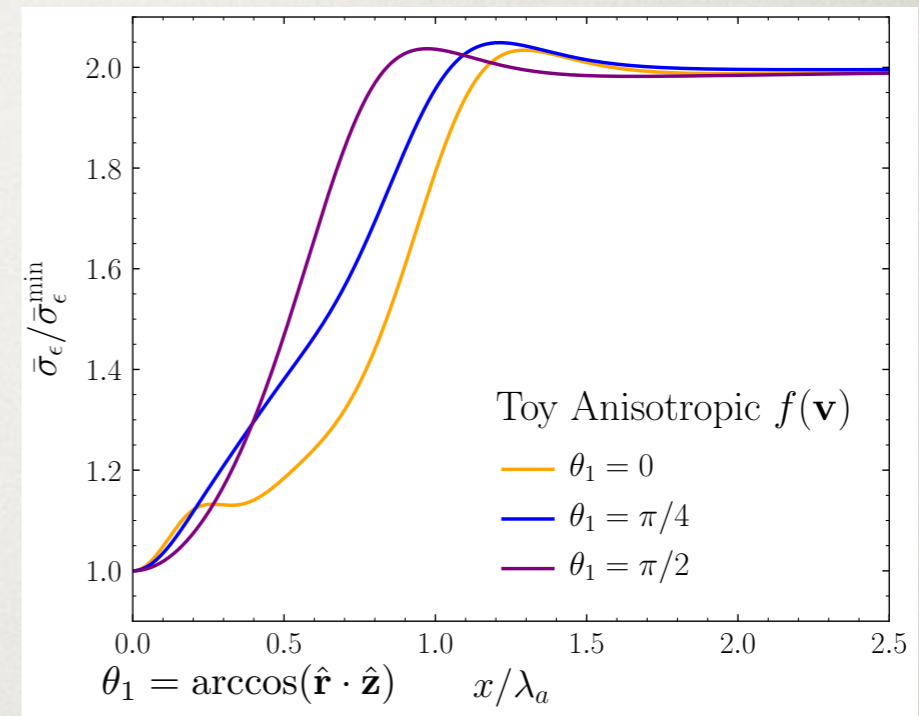
Unique pairs



# TOY EXAMPLES

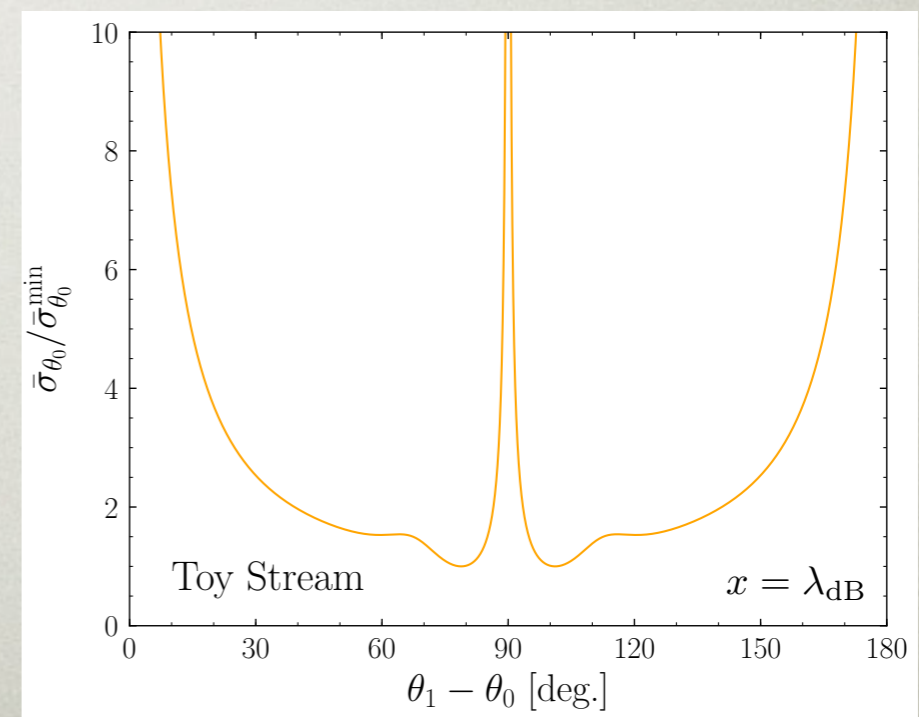
Anisotropy? Interferometry ✗

$$f(\mathbf{v}) = \frac{\sqrt{1 + \epsilon}}{\pi^{3/2} v_0^3} e^{-(v_x^2 + v_y^2 + (1 + \epsilon)v_z^2)/v_0^2}$$



Stream? Interferometry ✓

$$f(\mathbf{v}) = \frac{4}{\sqrt{\pi} v_0^3} \frac{e^{-v^2/v_0^2}}{\sin \theta_0} \delta(\theta - \theta_0) \delta(\phi)$$





# PARAMETERS FOR REALISTIC CASE



- Data over 24 hours
- Read out every 30 minutes, 75 averages / read out
- 0.04 Hz resolution