

WEBVTT

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00:00:00.000 --> 00:00:00.480

charlie young: Everyone

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00:00:04.290 --> 00:00:05.339

charlie young: Welcome everyone.

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00:00:06.839 --> 00:00:16.800

charlie young: We have the will start today with the third lecture from Tim Tate on dark matter and dark sector theory. So, Tim.

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00:00:18.240 --> 00:00:18.690

Tim Tait: All right.

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00:00:18.779 --> 00:00:20.160

Tim Tait: Thank you very much. It's

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00:00:20.520 --> 00:00:26.100

Tim Tait: Nice to be here again I really wish I was actually seeing you all in person. But this is the era that we live in.

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00:00:27.120 --> 00:00:28.110

Tim Tait: So,

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00:00:30.960 --> 00:00:36.450

Tim Tait: Alright, so this is the third and final lecture about the theory of dark matter and dark sectors.

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00:00:37.050 --> 00:00:45.690

Tim Tait: As a recap of what we've done so far on the first day we talked about some preliminaries and how to build a theory of dark matter that looks like it has a chance of describing dark matter.

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00:00:46.440 --> 00:00:55.470

Tim Tait: We talked about how dark matters produced in the early universe. And then, yesterday we finished up some of those topics and moved on to talk about specific theories of dark matter.

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00:00:56.340 --> 00:01:03.030

Tim Tait: In particular we spent a lot of time talking about supersymmetry, which is of course the most fleshed out and best understood theory of dark matter that we have

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00:01:03.870 --> 00:01:12.990

Tim Tait: And I think if you were to take a poll. It may still be the most popular though I think other ideas are also very exciting and get a lot of attention as well.

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00:01:13.560 --> 00:01:20.280

Tim Tait: Today we're going to finish up talking about specific theories of dark matter, talk about theories of dark sectors and that'll be the end of these lectures.

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00:01:22.440 --> 00:01:34.650

Tim Tait: So just to recap what we did yesterday. They're actually very many specific detailed models of dark matter. They range from theories designed to solve other mysteries two theories for which dark matter is the principal reason for it to exist.

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00:01:35.160 --> 00:01:40.080

Tim Tait: supersymmetry. Well, often maligned these days. And we even got a question that kind of revealed that yesterday.

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00:01:40.590 --> 00:01:46.140

Tim Tait: Is an interesting theory which easily accommodate a wimp or freeze out relic and other words, which is typically a neutralino

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00:01:46.620 --> 00:01:59.550

Tim Tait: Its specific properties are highly dependent on a complicated parameter space and we saw a little bit about how people deal with that issue. And there's a rich program of particle physics experiments that can access a wide swath of supersymmetric parameters space.

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00:02:01.020 --> 00:02:01.830

Tim Tait: So,

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00:02:03.450 --> 00:02:05.940

Tim Tait: We're going to talk about a couple other theories and these are not

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00:02:07.290 --> 00:02:11.880

Tim Tait: These are chosen to be a little different than supersymmetry some particular let's start talking about the accion

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00:02:12.480 --> 00:02:19.380

Tim Tait: Now the accion is another particle whose existence was proposed really independently of need for dark matter. In fact, it was posed.

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00:02:20.070 --> 00:02:27.240

Tim Tait: Back before we really understood that we needed a large amount of dark matter in the universe. It's instead motivated by the strong CP problem.

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00:02:27.750 --> 00:02:35.460

Tim Tait: The strong CP problem, which unfortunately don't have time to go through in great detail, but I guess you have lectures with accion, the more specifically that may talk about

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00:02:36.360 --> 00:02:46.650

Tim Tait: This a little bit more is basically the problem that if I write down the QC do a garage in the garage and describes a strong nuclear force there are terms in it which violate

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00:02:47.700 --> 00:02:56.580

Tim Tait: CP charge conjugation and parity and they would predict that the neutron should get an electric dipole moment. So over here, the plot that you see.

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00:02:57.210 --> 00:03:05.940

Tim Tait: shows you a history of measurements of the neutron EDM and also plot. It is the pink band is sort of the range that you'd expect if you just wrote down

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00:03:06.660 --> 00:03:17.310

Tim Tait: Using just dimensional analysis order one parameters and the QC Lagrangian. So you can see that from the very beginning, it was clear that the measurements were far below what they were expected, but even a factor of 1000

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00:03:17.850 --> 00:03:23.130

Tim Tait: And nowadays. These experiments have gotten down to the point where they're like 10 orders of magnitude smaller than they should be.

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00:03:24.540 --> 00:03:34.650

Tim Tait: So the parameter that describes the CP violation in the QC Lagrangian in some normalization should be something like 10 to the minus 10 and that seems like it's too small to be just an accident.

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00:03:35.820 --> 00:03:40.920

Tim Tait: So the accion is introduced as part of the pitch a Quinn mechanism, which was invented there at slack.

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00:03:42.120 --> 00:03:51.480

Tim Tait: To solve the strong CP problem by introducing a new field, which actually compensates for that term in the Lagrangian and makes it effectively zero just through dynamical arguments.

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00:03:52.470 --> 00:04:00.450

Tim Tait: For our purposes. The important thing is that leaves behind the quantum field that's right quantum particles from every quantum field has expectations which are particle

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00:04:01.080 --> 00:04:13.560

Tim Tait: And that's the particle that we've all the accion its mass and coupling are determined as part of the way that the theory is constructed. It's a pseudo Goldstone, both on if you know what that means. If you don't, it's not really that important for this discussion.

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00:04:14.940 --> 00:04:19.020

Tim Tait: It just means that it's interactions scale like energy and momentum and

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00:04:19.560 --> 00:04:24.450

Tim Tait: They're all higher dimensional interactions and they're characterized by some energy scale, which is very large.

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00:04:24.840 --> 00:04:32.370

Tim Tait: So the accion decay constant F survey is got to be greater than 10 to the nine GV or I run into problems from astrophysical bounce.

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00:04:32.790 --> 00:04:41.520

Tim Tait: Basically stuff in the universe would lose too much energy by cooling by admitting axioms. The accion mass is predicted by virtue of its connection to Q CD.

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00:04:42.030 --> 00:04:49.200

Tim Tait: It's actually very roughly given by something like the ratio of the pile on decay constant which you might know is something about point one GB

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00:04:49.740 --> 00:04:58.950

Tim Tait: Divided by the accion decay constant, we just mentioned, and then times the pie on mass. So if you put all that together, you get masses that are really, really tiny on the scale of particle physics.

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00:04:59.670 --> 00:05:06.720

Tim Tait: It's unstable. It has an interaction with two photons, but because it has a very tiny mass and because the coupling, which is characterized by FCA

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00:05:07.590 --> 00:05:16.710

Tim Tait: Is so weak basically its lifetime is so long that it lives much, much longer than the age of the universe. So as of now, we should have expected to see no accion sticking

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00:05:18.030 --> 00:05:24.360

Tim Tait: More generally, actually string theories and other beyond the standard model theories often contain accion like particles which are long lived

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00:05:24.810 --> 00:05:33.810

Tim Tait: And could play the role of dark matter, but of course they have less tight correlation between masses and couplings because they're not really introduced to solve a specific problem, like the QC accion is

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00:05:34.530 --> 00:05:41.370

Tim Tait: And some of the original references are listed here on the on the slide. They're actually very well written. Some of them are very good places to work.

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00:05:43.590 --> 00:05:44.280

Tim Tait: So,

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00:05:45.630 --> 00:05:56.880

Tim Tait: accion has a model dependent coupling to the electromagnetic field. It's a little bit smaller than one over that the scale F 's today, but I've written down what what Rajan term here at the top of the slide.

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00:05:57.360 --> 00:06:01.860

Tim Tait: So I have the coupling constant one over f . And then the accion field multiplied by

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00:06:02.340 --> 00:06:13.290

Tim Tait: The electromagnetic field strength contracted with it's doable. And if you turn this back into the ordinary language of Jackson e&m you get something that looks like an accion coupling to the electric field and the magnetic field.

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00:06:14.040 --> 00:06:22.680

Tim Tait: So in order to do things with accion you want to invoke both and electric and magnetic field. The one of them could be the initial state and one of them could be in the final state.

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00:06:23.310 --> 00:06:32.430

Tim Tait: So there's actually a very rich and very program of accion searches based on this coupling. So one particular search looks for the ambient accion that is the axioms that make up the dark matter.

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00:06:33.030 --> 00:06:40.050

Tim Tait: converting into electromagnetic signals in the presence of a strong background magnetic field, right. So the idea is I use this interaction.

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00:06:40.590 --> 00:06:48.990

Tim Tait: By setting up a strong magnetic field. The accion comes in, it turns it self into an electric field. And then I looked for the effect of the electric field that produced

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00:06:49.830 --> 00:07:00.420

Tim Tait: You can see some of the experiments that have been running plotted here in this plot as a function of the accion mass, which will notice is naturally expressed in micro Ed.

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00:07:01.110 --> 00:07:10.680

Tim Tait: And it's coupling the photons is on the Y axis, you can see that there are bounced from white dwarfs and supernovas that control.

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00:07:11.160 --> 00:07:18.870

Tim Tait: That restrict the parameter space. If the coupling is too large. And then there are experiments like the DMX detector.

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00:07:19.380 --> 00:07:26.250

Tim Tait: Which have got limits where it searches for axons and so far hasn't actually discovered in the axons converting themselves in the magnetic field.

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00:07:26.850 --> 00:07:36.240

Tim Tait: And the range, you expect from a QC accion is shown here as this band in purple lines. And you can see that there are plans to go to even

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00:07:36.960 --> 00:07:44.400

Tim Tait: Cover even more of the parameter space. There's actually a lot of other ideas searching for axioms is a very active area of theoretical research.

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00:07:45.240 --> 00:07:48.780

Tim Tait: I think you're going to have a whole lecture about it. So I'm not going to spend much time about it right now.

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00:07:49.500 --> 00:08:00.120

Tim Tait: But there are very interesting new ideas, looking for, like, things like the time variation of the neutron electric dipole moment or an induced current and an LLC circuit produced by axioms.

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00:08:03.810 --> 00:08:09.270

Tim Tait: Another interesting alternative to dark matter that doesn't look like a wimp would be a sterile neutrino.

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00:08:09.780 --> 00:08:15.150

Tim Tait: So I suspect that this was discussed somewhat in both Andre and Scott's lectures that you've already had.

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00:08:15.780 --> 00:08:24.150

Tim Tait: Dark matter might be connected to the other mystery of the Standard Model neutrino masses, we require physics beyond the standard model to explain neutrino masses.

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00:08:24.720 --> 00:08:31.530

Tim Tait: And I would say the leading theory of how that happens is the seesaw mechanism and you can understand that by looking

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00:08:31.920 --> 00:08:39.060

Tim Tait: At this diagram that I've shown here at the top of the slide where you see two neutrinos to left on double. It's really coming in.

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00:08:39.420 --> 00:08:46.650

Tim Tait: The couple to do Higgs doublets. But there needs to be something that can bridge between the two sides of the diagram, that would be a firm eon

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00:08:47.340 --> 00:08:55.410

Tim Tait: Right, because it's connecting another for me on and it goes on. It has no standard model charges and so therefore it has the charges or right handed neutrino.

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00:08:56.100 --> 00:09:02.460

Tim Tait: So this particle, you know, of course, there are other ways of realizing the Juno mass. So this is not necessarily the way nature works.

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00:09:03.390 --> 00:09:18.360

Tim Tait: But this is a very attractive way for it to work. And so we would have to add new fields to our theory, the right hand of your Trina fields. It's actually not clear how many of them, you need at least one but very likely there would be more just like there are more neutrinos themselves.

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00:09:20.040 --> 00:09:30.240

Tim Tait: If these additional states are light and not very strongly mixed with active neutrinos. And that's actually a requirement because we have precision data about neutrinos from things like

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00:09:30.780 --> 00:09:39.090

Tim Tait: SLC and left. And they basically tell us that we can't have any light active neutrinos that are strongly. Next, we would have seen them in our experiments.

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00:09:40.080 --> 00:09:46.110

Tim Tait: Then they would be stable on the scale of the age of the universe and other words they will decay, but they want to keep fast enough, just like the accion

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00:09:46.560 --> 00:09:56.970

Tim Tait: And they could play the role of dark matter. So arriving at the right amount of dark matter via neutrino oscillations and the early universe typically requires delicately choosing the massive and mixing angle.

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00:09:57.360 --> 00:10:04.350

Tim Tait: Or invoking some other new physics. And you can see a plot of the parameter space here and I'm not going to go through in great detail. But you see the massive the neutrino.

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00:10:05.220 --> 00:10:13.410

Tim Tait: Is on the x axis looks like it actually got cut off. But what's plotted on the y axis is the mixing angled between the active in the sterile neutrino.

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00:10:14.280 --> 00:10:25.710

Tim Tait: And you can see different predictions for for where you would get the right relic density as these colored lines that make different assumptions about other properties of the universe and also shown our

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00:10:26.790 --> 00:10:33.840

Tim Tait: Constraints that have been flooded here at about one ke VI and below where the neutrinos are to light and they wouldn't form galaxies, the way we expect

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00:10:34.230 --> 00:10:42.870

Tim Tait: There are some constraints, looking from X rays, which I'm going to talk about on the next slide is showing us the red region that's constrained. And so you have to live in this sort of white a sliver

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00:10:44.100 --> 00:10:52.620

Tim Tait: So in other words, masses between about one KGB, and maybe up to a few 10s of kV and mixing angles that are really very tiny somewhere around 10 to the minus 10 or so.

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00:10:55.890 --> 00:11:06.690

Tim Tait: All right, so still neutrinos actually decay, like I just said. So for example, the five and diagram, you see here on the lower left shows you how you can have a sterile neutrino come in.

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00:11:07.530 --> 00:11:14.430

Tim Tait: Interact with through a loop of who zones and leptons and end up producing an ordinary active neutrino.

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00:11:15.210 --> 00:11:30.690

Tim Tait: Right, which would have to be lighter. We know their masters or less than an EV, much less than anybody and a photon. So this decay process will happen, but because it's a loop diagram because it involves weak interactions, because the the startling Trina was so light, it probably

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00:11:31.950 --> 00:11:37.530

Tim Tait: Would be slow enough, at least for the mixing angles. We're talking about that you wouldn't expect it to happen yet.

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00:11:37.890 --> 00:11:47.430

Tim Tait: However, if it's happening. Slowly, then you'll be able to observe the gamma ray that's being produced, just like the case of the gamma ray lines from women isolation that we saw yesterday.

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00:11:47.910 --> 00:11:57.210

Tim Tait: This will produce a gamma ray always with the same energy, right, because you have a massive particle arrested case into two particles that are both basically mass less by comparison.

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00:11:57.690 --> 00:12:11.280

Tim Tait: So each particle will therefore take away half of the energy of the parent and the gamma ray or sorry, the x ray in this case would actually be m sterile divided by two. So here's another plot which isn't the same

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00:12:12.720 --> 00:12:21.630

Tim Tait: parameter space, but with the axes flipped so that mixing angle is plotted on the x axis. The mass is plotted on the y axis, you can see constraints from x ray searches.

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00:12:22.740 --> 00:12:28.470

Tim Tait: You can see some regions that might be favored by theory, trying to explain why pulsars seem to get kicked out of the plane of the galaxy.

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00:12:29.070 --> 00:12:38.400

Tim Tait: And you can also see an interesting little red blob here and this was extracted from a possible x ray signal that was observed in 2014 as processed by

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00:12:39.090 --> 00:12:45.960

Tim Tait: Dissolved in so he's the one who actually provided this content that's pretty exciting. That means that maybe there is actually even a signal of dark matter.

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00:12:46.680 --> 00:12:54.990

Tim Tait: However, of course, this is something that needs to be confirmed and the situation is kind of very confusing at the moment because there have been lots of follow up studies.

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00:12:55.350 --> 00:13:01.290

Tim Tait: Some of them actually see something like this signal, though not none of them and see it as strongly as the original observations did

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00:13:02.040 --> 00:13:10.110

Tim Tait: But some of them don't see anything at all. And so it's possible that maybe this is just something that's an artifact of the particular place that the original authors were looking

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00:13:11.340 --> 00:13:19.860

Tim Tait: Or just maybe telling us something more interesting about astrophysics. So this is something that we're going to keep our eye on.

But we can't really declare victory and say that we've seen dark matter.
Yes.

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00:13:23.100 --> 00:13:28.530

Tim Tait: So now moving on. I'm going to go back to something that's a little bit more Wimp like and tell you a little bit about simplified models.

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00:13:28.860 --> 00:13:42.120

Tim Tait: simplified models have emerged as the primary way in which the LSD experiments are communicating the results of their searches for dark matter, and in some cases are being also adopted and some of the plots, you see from

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00:13:43.170 --> 00:13:50.640

Tim Tait: Searches for dark matter scattering with nuclei or indirect detection searches, too. So let's understand how they work.

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00:13:51.210 --> 00:14:04.950

Tim Tait: A simplified model is just taking what might be a complete theory, but then cutting it down to just the essential parts that contain the dark matter, as well as the most important particles that mediate interaction with the standard model. So maybe instead of

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00:14:07.680 --> 00:14:18.570

Tim Tait: Maybe instead of looking at the entire minimal supersymmetric standard model. You might just look at the MSP and then maybe one of the next lightest particles, that's important when you calculate it's cross sections.

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00:14:19.320 --> 00:14:26.040

Tim Tait: So I'll show you a couple examples of these as the first one. If we're interested in a dark matter.

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00:14:26.670 --> 00:14:36.690

Tim Tait: Of a theory were dark matters interacting with corks, you could imagine that you include just the dark matter and some SU three charge scale or particle which mediate the interaction.

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00:14:37.290 --> 00:14:42.570

Tim Tait: So in other words, we have the dark matter. And if I look at the little diagram down at the bottom that's indicated by chi.

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00:14:42.990 --> 00:14:49.620

Tim Tait: Interacts with the Standard Model cork, but there has to be something else to make this whole vertex gauge invariant. And that thing is.

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00:14:50.160 --> 00:14:59.430

Tim Tait: A scale are charged and both SU three charged and electrically charged particle. So, this thing would look kind of like a squirt from supersymmetry.

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00:14:59.850 --> 00:15:09.270

Tim Tait: But supersymmetry would have actually dictated what the structure of this interaction would be, and in this more general simplified model. You're not really constrained to live in the supersymmetric region.

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00:15:12.060 --> 00:15:19.410

Tim Tait: Trying to conserve flavor in the standard model. In other words, not have this kind of vertex mediates a flavor violation of loop level.

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00:15:20.280 --> 00:15:27.180

Tim Tait: Usually suggest that we include the mediator particle to also have a flavor index. So in other words, if there are three quirks that can couple to

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00:15:27.600 --> 00:15:33.750

Tim Tait: You know, let's say the right handed up charm and top cork, then there would be three of the scale or mediators, the

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00:15:34.350 --> 00:15:38.070

Tim Tait: The partner right of the right handed up the right hand of Charmin the right hand to top

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00:15:38.760 --> 00:15:47.070

Tim Tait: And that they would all have the same message interactions. And if that's true, we get kind of like a super Jim mechanism which prevents us from getting large flagrant violation.

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00:15:47.550 --> 00:15:55.590

Tim Tait: So this theory. It looks like a little bit of Susie, but with more freedom in terms of choosing the coupling in mass, etc. And there are only three parameters which makes it much more

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00:15:56.490 --> 00:16:05.850

Tim Tait: Easy to deal with in the supersymmetric standard model, right. And those are the mass of the dark matter, the mass of the mediator particle. And then the overall coupling strength of each one of the mediators of the quirks.

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00:16:09.150 --> 00:16:15.060

Tim Tait: Alright, so the ELYSEE experiments have placed very interesting limits on theories like these.

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00:16:15.630 --> 00:16:20.220

Tim Tait: A couple of them are shown here and I even labeled it scores, because that's what I'm translating here.

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00:16:20.880 --> 00:16:37.140

Tim Tait: Um, so, for example, you could imagine that the la chica Bruce A pair of the scale or particles and then that vertex we saw before would let them decay into a dark matter particle and a jet have had Ron's so this process repair production would produce a net of

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00:16:38.370 --> 00:16:48.180

Tim Tait: Two jet's have had runs and missing transverse momentum and then plus maybe there could be some also initial state radiation, you have to allow for that it happens very often feel he

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00:16:49.800 --> 00:16:53.700

Tim Tait: Said, so they devise a series of searches both that list and CMS due

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00:16:54.750 --> 00:17:06.120

Tim Tait: To the search for for events like those may appear with the background expectations so far all the measurements have been consistent with the background. And so you end up with

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00:17:06.570 --> 00:17:14.760

Tim Tait: Regions a parameter space that you can exclude so in the upper left, you see a plot that's plotted in the plane of the mass of the mediator and the mass of the dark matter.

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00:17:16.320 --> 00:17:26.130

Tim Tait: So the LLC, for some reason, likes to put the mass of the mediator on the x axis and the mass of the dark matter on the Y axis. That's kind of the opposite of what direct detection experiments. Do I don't know why.

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00:17:27.270 --> 00:17:44.940

Tim Tait: But it's fine. You just turn your head and you understand what it's telling you. So, and you see that it takes a bite out of the parameter space here for the lower masses of mediator and dark matter, you can sort of robustly exclude dark matter masses up to you know about 400 TV or so.

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00:17:46.170 --> 00:17:54.210

Tim Tait: And you can you're sensitive to dark matter masses, at the moment, up to about 600 GB or so, whereas for the mediator. You can robustly exclude them.

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00:17:54.900 --> 00:18:02.280

Tim Tait: Only actually down to about 400 but you can actually, in some cases where the dark matters very light excluded up to one and a half TV.

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00:18:02.940 --> 00:18:06.570

Tim Tait: Since you're pretty interesting limits that put a pretty tight constraints on the

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00:18:07.320 --> 00:18:22.230

Tim Tait: On the theory, the blue region is actually just showing you an older result. And so you can see how much progress the hell he made in going from having 3% of our earnings to 36% of words. And of course there's more data already collected and a lot more to come in.

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00:18:24.840 --> 00:18:29.430

Tim Tait: So you could also imagine looking for a modern jet signature, like the one that we discussed last time.

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00:18:29.820 --> 00:18:37.560

Tim Tait: We're here we have a cork and a glue on which fuses actually into a standard model of cork. But then that splits into the mediator particle and the dark matter.

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00:18:37.950 --> 00:18:43.170

Tim Tait: So the mediator will decay into a jet have had runs and another dark particle. So, I'll have one hard jet.

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00:18:43.650 --> 00:18:51.090

Tim Tait: And missing transverse momentum, you know, plus whatever the additional state radiation looks like that has a somewhat different topology and so

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00:18:51.480 --> 00:19:01.680

Tim Tait: There's no dedicated search for this theory that I could easily look at, but there was some work I did with some other authors, where we actually analyze the available searches in the context of the signal.

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00:19:02.280 --> 00:19:13.680

Tim Tait: And what you can see here is shown in this plot in the lower right part because this process for mono jet actually involves the vertex and involving the cork and the dark matter and the mediator.

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00:19:14.370 --> 00:19:21.420

Tim Tait: You can actually put a bound in the plane of these masses on the size of that interaction. So, see that's somewhat different than what happens

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00:19:22.140 --> 00:19:33.000

Tim Tait: When you look at pair production because pair production. The production rate only depends on the QC decoupling, and you can really just actually put a limit on the masses themselves without worrying about the coupling.

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00:19:33.600 --> 00:19:40.650

Tim Tait: Coupling has to be a big enough for it to decay inside the detector. But that's a very mild constraint that the coupling is bigger than like 10 to the minus 10 or something.

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00:19:41.520 --> 00:19:48.600

Tim Tait: So basically any coupling is ruled out any reasonable coupling is ruled out for masses in this region, whereas for the mono jet signature.

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00:19:49.650 --> 00:19:54.540

Tim Tait: The coupling that you can rule out varies very dramatically as you change the masses of the particles.

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00:19:57.270 --> 00:20:00.870

Tim Tait: Okay. You can also talk about the direct detection of theories like this.

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00:20:01.800 --> 00:20:10.290

Tim Tait: If the case where it's a micron a particle, it's actually dominated by one loop effects where the dark matter comes in and scatters off of the glue ones that are part of the nuclei.

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00:20:11.040 --> 00:20:21.600

Tim Tait: And you can see some. So there are three local spin dependent interactions, but the loop level spin independent interaction, even though it's one loop is actually a much more constraining

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00:20:22.290 --> 00:20:30.090

Tim Tait: Limit than the tree level is just because of the fact that it's been independent. So you can see the spin dependent is putting bounds on couplings that are, you know, around

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00:20:30.840 --> 00:20:43.170

Tim Tait: Two or three or four right they're big enough, you start to worry about perturbation theory, whereas the spin independent scattering is putting downs of the coupling that are much less than one and that looks pretty good point of view perturbation theory.

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00:20:46.740 --> 00:20:51.540

Tim Tait: Okay, another important classes simplified models are models where a vector is exchanged.

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00:20:52.230 --> 00:21:03.360

Tim Tait: So we could have a corks they're shown here is the blue lines in the diagram. They a couple of some kind of new vector particles and had a dark force carrier. So this could be considered a case of a

144

00:21:05.100 --> 00:21:10.290

Tim Tait: Dark sector theory and then that mediator is the thing that itself couples to the dark matter.

145

00:21:11.340 --> 00:21:18.360

Tim Tait: So there's a lot of parameters and principle that would be needed to describe a theory like this, you need the mass of the dark matter, the mass of the mediator.

146

00:21:18.930 --> 00:21:25.770

Tim Tait: Which I call this the prime here. You need the overall coupling strength of the Z prime and then actually there could be different charges for all of the

147

00:21:26.070 --> 00:21:34.470

Tim Tait: Left handed corks right handed up quirks right handed down corks left handed leptons right hand to charge leptons, etc, etc. Even the Higgs.

148

00:21:35.580 --> 00:21:38.490

Tim Tait: That could also be kinetic mixing what you're going to talk about a little bit later.

149

00:21:39.720 --> 00:21:47.310

Tim Tait: So there are theoretical considerations such as the dark sector more particles to cancel gauge anomalies, etc.

150

00:21:47.790 --> 00:21:51.390

Tim Tait: They are important, but they may not actually matter very much for the LSC searches.

151

00:21:51.750 --> 00:22:01.230

Tim Tait: What's typically done when they present results is they just assume that all the couplings are the same. So all of the Standard Model couplings are the same. And then they treat the dark matter coupling is differently.

152

00:22:01.890 --> 00:22:13.050

Tim Tait: Or maybe they assume that all the couplings are the same, but there are no couplings. The leptons and then they look at the dark matter of coupling as a separate parameter that gets you down to about three or four parameters and it becomes much more tractable that point.

153

00:22:14.670 --> 00:22:21.840

Tim Tait: And so again, they do similar searches. These are now mostly dominated by the model jet search which is cartoon to get over here where you see corks coming in.

154

00:22:22.290 --> 00:22:31.050

Tim Tait: They emit so it's labeled x here, but the dominant boundaries. You come from the case where x is a jet they produce the mediator particle which then the case in the dark matter.

155

00:22:32.280 --> 00:22:39.390

Tim Tait: And the next slide, we'll see the case where it can dictate even back into the standard model and they place limits on the parameter space that are shown.

156

00:22:39.870 --> 00:22:47.850

Tim Tait: Here on these two images. So you can see again the mass of the mediator, which is labeled as Z_b for the Atlas plot and

157

00:22:48.810 --> 00:23:06.510

Tim Tait: M. Ed or for mediator on the CMS blocks are both on the X axis, the mass of the dark matter is plotted on the y axis, the regional

parameter space to gets rolled out is this kind of wedge shaped curve sort of a triangle. You can see that that structure is there for both of the

158

00:23:07.590 --> 00:23:08.550

Tim Tait: Both of the plots.

159

00:23:09.780 --> 00:23:20.610

Tim Tait: They've also plotted regions where the relic density would be expected to work out just to guide your eye. So that's the line that's read here on atlas or blue and CMS.

160

00:23:21.270 --> 00:23:26.370

Tim Tait: And you can understand pretty easily. The structure of the limits. If we think about the sort of kinematic regimes.

161

00:23:27.240 --> 00:23:38.610

Tim Tait: For this process. So first of all, at some point, as I make the mediator. Let's say the dark matter mass is very low. If I make the mediator heavier and heavier at some point the mediator is so heavy. I don't typically have enough energy to produce it.

162

00:23:39.150 --> 00:23:52.020

Tim Tait: The LLC stops producing anything and I get a limit that peters out at about that mass. And you can see that for Atlas. This was about one and a half TV and CMS is very similar. It's like maybe 1.6 or seven or so.

163

00:23:54.030 --> 00:24:03.930

Tim Tait: Now the structure with the dark matter mass is a little bit more complicated. But what's been done here on the Atlas plot to guide your eye is they plotted the line. This is this dotted line here.

164

00:24:04.590 --> 00:24:13.650

Tim Tait: Where the mediator mass is equal to twice the dark matter mass. So this is the line where if I live on this line or below it. The mediator can actually decay into the dark matter directly

165

00:24:14.250 --> 00:24:23.340

Tim Tait: And that gives me a huge enhancement of this process because now I don't have to pay the price of an off shell mediator. I just produced the mediator on shell and then wait for educating the dark matter.

166

00:24:24.300 --> 00:24:34.500

Tim Tait: And so that's why the limits, more or less. Hug this line. And then, of course, eventually they cut off because the mediator just gets too heavy. So that explains why these shapes are sort of these branches.

167

00:24:35.850 --> 00:24:43.200

Tim Tait: Another thing to point out that sort of a hidden assumption. You can see often in these plots is written on the plot, you'll see

168

00:24:43.770 --> 00:24:51.570

Tim Tait: The value of the coupling the quirks and the value of the coupling the dark matter, right, because this process clearly depends on both of those things at least

169

00:24:52.230 --> 00:24:56.370

Tim Tait: It depends on the coupling the quirks in the region where the mediator is on shell.

170

00:24:57.030 --> 00:25:03.810

Tim Tait: And even then, it's still kind of depends on the couple into dark matter because there's a branching ratio for the mediator to go into dark matter versus to go into corks

171

00:25:04.380 --> 00:25:17.940

Tim Tait: That's important. So you have to look at these couplings and make sure that they match the kind of situation that you're thinking about or these plots can actually look very different. So these are important caveats, and you can see that there were slight differences.

172

00:25:19.140 --> 00:25:19.740

Tim Tait: Here.

173

00:25:20.850 --> 00:25:26.970

Tim Tait: In the sense that Atlas decided to do a vector mediator this plot and CMS that an axial vector mediator.

174

00:25:27.660 --> 00:25:37.890

Tim Tait: But they chose the same couplings for the quirks and the dark matter. And that's why the limits you get are very similar. They're not exactly the same nor slight difference between vector an accelerator.

175

00:25:40.710 --> 00:25:47.190

Tim Tait: You can also map these into the plane of direct detection. So this sort of shows you how to compare this type of search

176

00:25:47.670 --> 00:26:00.510

Tim Tait: For this kind of model at a collider versus this kind of search for the same model in the direct detection experiments so translating those wedges gives you a limit that looks like the red line, you see here in the CMS plot.

177

00:26:01.740 --> 00:26:10.740

Tim Tait: This is for a spin independent coupling. So you can see that it takes a bite out of the parameter space that is fairly independent of the mass, up until a mass of

178

00:26:11.310 --> 00:26:16.260

Tim Tait: A few hundred GV and then at some point, it kind of rolls over and closes off.

179

00:26:17.190 --> 00:26:29.730

Tim Tait: The reason for this roll over this roll back is actually because you don't really trust your calculations up here because the couplings get too big and your, your simplified model doesn't work. So then for comparison. They've also shown the plots from

180

00:26:31.110 --> 00:26:40.080

Tim Tait: direct detection experiments like Lux xenon and Panda X. These are actually hard to resolve on this particular scale that they're showing, but you can see them all down here.

181

00:26:40.470 --> 00:26:47.160

Tim Tait: As these blue lines. So you can see that direct detection experiments are much more sensitive. If the mass of the dark matter is around

182

00:26:47.580 --> 00:26:57.330

Tim Tait: Say 10 or 100 GB. But actually, much less sensitive down below that. So the, the colliders are really exploring this particular theory very well in this low mass region.

183

00:26:58.230 --> 00:27:09.960

Tim Tait: So for Atlas I've chosen a spin dependent comparison. And so they're actually the spin dependent limits the best ones on on this channel or from Pico

184

00:27:10.500 --> 00:27:22.980

Tim Tait: That they're shown as the magenta line. You can see that the colliders are actually doing better for a wider range of masses. So the colors are really filling an information at low mass and for interactions that produced independent scattering but no spin independent scatter

185

00:27:25.560 --> 00:27:34.260

Tim Tait: And we can also look directly for the mediators. So you can imagine that the corks fuse and to one of the mediators, but then the mediator instead of decaying the dark matter to case back into corks

186

00:27:34.950 --> 00:27:44.100

Tim Tait: There's actually a wide variety of different strategies to search for these particles. You can see the Atlas summary plot kind of has all these different channels put together.

187

00:27:44.700 --> 00:27:55.830

Tim Tait: And you can see that the values of coupling that you're probing are sort of typically around point one or point two, but for specialized searches, they cannot get all the way down to close to point out one or so.

188

00:27:56.550 --> 00:28:06.720

Tim Tait: And it changes a lot with the mass, because the different search strategies require you to trigger on the particles are producing and as you make the mass of immediate or very light. It gets very hard to

189

00:28:07.800 --> 00:28:12.990

Tim Tait: To trigger. So you have to change the way you search and as a result you get a different limit and

190

00:28:13.440 --> 00:28:25.440

Tim Tait: I have to really have to say these limits, especially low masses are extremely creative use of LDC data their limits that I think maybe you would have said they shouldn't be able to get at all. And it's really quite remarkable that they do so well.

191

00:28:26.730 --> 00:28:40.620

Tim Tait: Of course, that high energy is the energy is also the king because it's the only man made source of high energy collisions that we have, but there you'd expect it to be doing well. And so while that's still amazing it maybe took a little bit less effort on the analysis side.

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00:28:41.670 --> 00:28:49.950

Tim Tait: And over here, there's a CMS plot that shows you a similar way of looking at it in terms of the masses of the mediator in the dark matter and shading the couplings.

193

00:28:51.150 --> 00:28:52.500

Tim Tait: So I'm just gonna leave it at that.

194

00:28:54.630 --> 00:28:55.170

Tim Tait: Okay.

195

00:28:56.400 --> 00:29:06.930

Tim Tait: So I'm going to take a couple minutes and talk about dark matter, a couple of the glue ons. So another possible variation. This is something that I worked on with rotini got ballet and

196

00:29:07.890 --> 00:29:18.270

Tim Tait: Government data is the idea of the dark matter might just interact directly with the glue ons and not really bother with the quirks accepted higher level.

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00:29:18.780 --> 00:29:24.030

Tim Tait: So for example, you could imagine that there's a dark matter. That's a scam or particle it couples to some kind of

198

00:29:24.420 --> 00:29:37.050

Tim Tait: stealer mediator. So that would be a normalized double interaction that involves to dark matter particles and to color to S3 color and scale or mediator particles, it would look like sort of the left half of the Simon diagram here.

199

00:29:39.030 --> 00:29:44.820

Tim Tait: But this scale or mediator particle doesn't actually have to be stable. It doesn't have to have the two symmetry, it might decay.

200

00:29:45.510 --> 00:29:48.090

Tim Tait: Because it also appears squared and the interaction.

201

00:29:48.690 --> 00:29:55.140

Tim Tait: So then in this case I can actually make a loop of the scale or particles and I can couple the dark matter, primarily the blue ones.

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00:29:55.380 --> 00:30:01.440

Tim Tait: So of course the dark matter is really coupling primarily to to mediators, but because it's two mediators and not one Mediator.

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00:30:01.920 --> 00:30:10.080

Tim Tait: The way to get to the standard model and only the standard model is to close the meat eaters into a loop and attach the blue ones. And that's basically what the idea here is

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00:30:10.950 --> 00:30:20.760

Tim Tait: So actually the color and the flavor representation of the mediator particles are completely free to choose. It's not like the previous case of the skin color mediator.

205

00:30:21.030 --> 00:30:28.230

Tim Tait: Where I had to choose something that would give me a gauge invariant interaction with corks any representation would actually be a gauge and area here.

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00:30:30.090 --> 00:30:35.130

Tim Tait: So then you can do all of the things that we've looked at so far, you can ask about the relic density

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00:30:35.940 --> 00:30:42.690

Tim Tait: It turns out that for preservative values of this coupling λ a thermal relic actually favors the case where the

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00:30:43.230 --> 00:30:48.120

Tim Tait: Mass of the mediator is less than the mass of the dark matter. So the dark matter can annihilate into the mediator.

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00:30:48.510 --> 00:30:53.040

Tim Tait: So that was kind of a surprise because normally we would have thought, okay, the scale or should we have yours on the dark matter.

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00:30:53.460 --> 00:31:04.920

Tim Tait: But again, since it doesn't become to the dark matter. There's really no restriction say that that says that that should be the case. So here it shows in the plot it shows you the value of λ times

211

00:31:06.180 --> 00:31:11.310

Tim Tait: The square root of the representation of the color scale or

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00:31:12.420 --> 00:31:15.780

Tim Tait: That just happens to be the thing that scales out of the cross section conveniently

213

00:31:16.830 --> 00:31:27.000

Tim Tait: It shows you how big that would have to be in order to get the right relic density in the plane of the mass of the mediator in the mass of the dark matter. And so you can get small masses.

214

00:31:27.450 --> 00:31:32.910

Tim Tait: And then you require smallish couplings as you go to much larger masses, like a few TV you need larger couplings.

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00:31:33.300 --> 00:31:43.020

Tim Tait: One thing is that because of the square root of NF times are this is not really a coupling. And so actually you can stay perturb of roughly all the way up to here. If you choose the representation to be large enough

216

00:31:45.900 --> 00:31:56.520

Tim Tait: Then you can search for the mediators of the LSC you can produce them in pairs. Now you can actually just let them to K directly into corks by adding interactions between them. The quirks that depends on what the representation as you choose.

217

00:31:57.360 --> 00:32:09.030

Tim Tait: The bounds are actually relatively weak if they just became the ordinary untangle jets. The bound is just a few hundred GV so that's much less than this, the specialized searches for sports that we saw.

218

00:32:10.890 --> 00:32:18.930

Tim Tait: I we just did a specific example where we considered the mediators to have charged electric charge four thirds and couple to too uptight corks that are right handed.

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00:32:19.950 --> 00:32:23.520

Tim Tait: There are many different variations, though, that you can play with that assumption.

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00:32:24.600 --> 00:32:29.520

Tim Tait: In the case where they a couple to the top cork, then you do get some more specialized searches and

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00:32:30.180 --> 00:32:38.760

Tim Tait: So what's shown in the plot is actually as a function of the coupling of the mediator to the top corks, which is why one and the mass of the mediator.

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00:32:39.240 --> 00:32:46.050

Tim Tait: There are two different channels, you can use to search for this theory below masses are excluded by looking for two tops plus two jets.

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00:32:46.470 --> 00:32:57.270

Tim Tait: And then there's this blue region which is excluded by two tops plus one jet all of these constraints are really relatively weak. If you think about it, they're sort of tell you the mass has to be bigger than a few hundred GB at most.

224

00:33:00.360 --> 00:33:11.400

Tim Tait: And then you can also look at the direct searches direct detection generally provides a strong bound unless the dark matter masses, particularly small you can see projections here.

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00:33:12.450 --> 00:33:13.560

Tim Tait: For the

226

00:33:14.610 --> 00:33:24.270

Tim Tait: Week. Sorry. You can see the Lux experiments results here. That was actually the best experiment at the time that we did this work and you can see that it takes interesting bite out of the parameter space.

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00:33:26.130 --> 00:33:34.830

Tim Tait: The mono jet signature occurs at one loop we actually did all the one loop calculation, because we were curious to see how it would work with all the different momentum dependent for the loop integrals.

228

00:33:35.460 --> 00:33:49.530

Tim Tait: And that was interesting. The results of the ELYSEE, were not particularly helpful though for large enough, our λ , you can see something with large data set, but something like 100 TV programs aren't on collider would do a lot better if we ever get one of those

229

00:33:52.440 --> 00:34:05.280

Tim Tait: So all of these simplified models are sketches of dark matter not complete theories of dark matter. And the idea is, of course, and if we start to see them, we would put the sketch together and have it come to life. So I believe this is showing us how that would work.

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00:34:07.470 --> 00:34:13.260

Tim Tait: And so this brings me to the end of the second part of my lectures which

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00:34:14.520 --> 00:34:19.710

Tim Tait: We've already talked about the first set of conclusions. There are many specific detailed models of dark matter and about Susie

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00:34:20.460 --> 00:34:29.610

Tim Tait: Now today, this morning, so far we've seen that axons and sterile neutrinos are well motivated extension to the standard model they explained the strong CP problem and neutrino masses.

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00:34:30.150 --> 00:34:35.790

Tim Tait: And we have more specialized types of searches to find them. But we do have searches. We know how to do and we're looking for them.

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00:34:36.660 --> 00:34:42.750

Tim Tait: And then we talked a little bit about simplified models that have been adopted as a bridge between searches for dark matter and theoretical constructions.

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00:34:43.230 --> 00:34:49.590

Tim Tait: They don't contain the full distracting details of a complete theory, but they hopefully summarize the broad impacts of searches on many of them.

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00:34:51.330 --> 00:34:59.640

Tim Tait: So this brings me to the last section of the lectures which are more like examples of dark sectors and I have a few different examples I've chosen to go through

237

00:35:00.480 --> 00:35:10.740

Tim Tait: I think we should roughly get through all of them, but I may have to speed up a little bit. At one point the person I'm going to talk about is the idea that dark matter could be composite so dark matter could be made out of other things.

238

00:35:12.120 --> 00:35:22.230

Tim Tait: So dark matter, of course, must be weakly interacting with the standard model, but could still be part of a dark strongly interacting system, which plays an important role in terms of what it actually looks like in the galaxy.

239

00:35:23.010 --> 00:35:29.670

Tim Tait: So one example is dark matter could have a light force carrier, leading to it to form dark atoms, or even dark molecules.

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00:35:30.090 --> 00:35:36.690

Tim Tait: And I've got a reference here to one of the first papers that took this idea seriously, there's actually been a lot of work that's followed up since then.

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00:35:37.500 --> 00:35:43.680

Tim Tait: The cartoon here, you see that there are too dark matter particles. One has positive dark charge one has negative dark charge

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00:35:44.070 --> 00:35:54.240

Tim Tait: They exchanged the dark force carriers and they've been bound together to form something like a dark, Adam. In this case, you can imagine more complicated theories. You might even get dark molecules.

243

00:35:55.800 --> 00:36:07.050

Tim Tait: You could have a dark matter which fields are very strong interaction that can find like SU three confines and it could form dark hadron's. So here I've got a cartoon that shows

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00:36:07.680 --> 00:36:13.590

Tim Tait: A case where the dark matter is actually just a dark set of gluons now by themselves, they would be massless.

245

00:36:14.010 --> 00:36:24.570

Tim Tait: But because they're confined to they're stuck together in a little ball of energy. Right. The same way that the proton is much heavier than the quarks of make it up. And so the dark matter could be something like that.

246

00:36:25.650 --> 00:36:31.290

Tim Tait: And of course, if I take this idea a little bit further, there could be a whole range of possible stable states like dark nuclei.

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00:36:31.710 --> 00:36:41.550

Tim Tait: So if I take something like dark hadrons. They may have nuclear interaction, similar to the nuclei that we know about. And they may form actually larger composite objects that are made out of the dark

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00:36:42.600 --> 00:36:49.320

Tim Tait: The dark nuclei and you can read more about that. There. So these are all actually areas of active research.

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00:36:49.740 --> 00:36:54.510

Tim Tait: The parameter space is not really entirely mapped out and all the phenomena haven't really been discovered.

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00:36:54.810 --> 00:37:04.800

Tim Tait: The references, give you a place to start to learn about the ideas, but they're not complete, and I'm sorry, I'd be happy to talk about this offline in more detail and point you to other directions. You can go

251

00:37:06.600 --> 00:37:18.750

Tim Tait: So I'm going to tell you a little bit about a very simple idea. This is something that I worked on a few years ago where actually the theory is just a pure gauge theory, consisting of a hidden sector $su(2)$ so basically this is just like you take

252

00:37:20.340 --> 00:37:31.830

Tim Tait: You add a set of dark blue ones of some kind to your theory and nothing else just nothing else at all and maybe even imagine that they don't interact with the Standard Model except through gravity.

253

00:37:33.360 --> 00:37:38.640

Tim Tait: So you can think about this is kind of like a hidden gauge group which just is totally separate from the standard model.

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00:37:39.480 --> 00:37:45.600

Tim Tait: And of course there are variations, where I would include things like dark quarks, but just to keep things really simple at the beginning. Let's not bother to do that right

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00:37:46.800 --> 00:37:56.340

Tim Tait: So we know that this theory is defined by the number of dark colors. There are, in other words, what is n and the confinement scale, right, which tells me it which energy scale.

256

00:37:56.910 --> 00:38:06.540

Tim Tait: That Dark Age theory confines and you wouldn't be a few just that value what you're really adjusting at high energies is the value of that strong coupling constant

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00:38:06.930 --> 00:38:13.710

Tim Tait: Because then it's going to run down in a way that you can predict based on and and it will be confined at some scale and you could calculate at that point.

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00:38:14.190 --> 00:38:24.060

Tim Tait: So rather than going through all of that calculation. I'm just going to talk about the confinement scale itself. And I know I can always adjust what the value of that coupling is to get whatever confinement scale.

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00:38:25.650 --> 00:38:33.060

Tim Tait: So this can find it scale characterizes the masses of the of the dark composite the particles that form.

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00:38:34.260 --> 00:38:39.990

Tim Tait: Now, since all I have our glow lines. The only had Ron's that exists in this theory are glue balls.

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00:38:40.500 --> 00:38:46.800

Tim Tait: So the massive the lowest glue ball is, you know, something like two pi times lambda. If you look at lattice calculations.

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00:38:47.370 --> 00:38:57.930

Tim Tait: And also there are more glue balls. They have all different kinds of spins and parity assignments and they are split by some number that's also similar to lambda, between the two of them.

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00:38:59.580 --> 00:39:08.250

Tim Tait: So from here on. I'm going to stop saying darker hidden. Most of the time, and I'm just going to say glue and glue ball. But remember, I'm talking about the dark stuff, not, not the ordinary live ones.

264

00:39:10.140 --> 00:39:15.840

Tim Tait: So one interesting feature of this type of theory is the fact that blue balls interact strongly with another with one another.

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00:39:16.260 --> 00:39:23.670

Tim Tait: Right, and that's obvious because they're made of the strongly interacting glue ons when they get close enough to each other. They actually realize that they like to to interact

266

00:39:24.180 --> 00:39:33.210

Tim Tait: And so if you get to glue balls close to each other. They actually scatter very strongly because of the strong dynamics. You can't really compute this very reliably and perdition theory.

267

00:39:34.080 --> 00:39:39.210

Tim Tait: That you could actually do some calculations on the lattice and they're even have been a few calculations have been done.

268

00:39:39.690 --> 00:39:45.690

Tim Tait: In this case, making the theory. So simple actually makes the calculation. A little easier because pure glue is much easier to deal with.

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00:39:46.440 --> 00:39:52.020

Tim Tait: So you can cartoon, the self interaction of the glue balls as just like a geometric geometric cross section, you basically just say

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00:39:52.710 --> 00:40:02.910

Tim Tait: Let's imagine you know that the cross section between these two blue balls that are scattering with each other is equal to the actual area that they see. So that would be, you know, πr^2 .

271

00:40:03.480 --> 00:40:11.970

Tim Tait: Where r is actually set by this confinement scale and then you just ask how likely it is for them to overlap and you get a geometric cross section.

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00:40:12.690 --> 00:40:20.400

Tim Tait: That cross section obviously struggling interacting. That's where the four five comes from. It goes like one over the confinement scale squared.

273

00:40:22.200 --> 00:40:29.040

Tim Tait: And right and that's, again, just the dimensional analysis of saying that my objects have roughly that size because that's the scale of the interaction that holds them together.

274

00:40:29.550 --> 00:40:43.110

Tim Tait: And then the one over in square. It's a little harder to understand and it basically is something that comes out from large and based on a different kinds of theoretical considerations I can answer questions about that by email, but I'm not going to go into it here.

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00:40:44.280 --> 00:40:51.390

Tim Tait: But the upshot is really for small in any way a single parameter λ controls both the mass and the cross section.

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00:40:52.170 --> 00:41:01.800

Tim Tait: For large and becomes important to and so you can actually just ask yourself, how much would I what λ . What I need in order to get a cross section that would be observable in terms of

277

00:41:02.310 --> 00:41:08.160

Tim Tait: The dark matter scattering doing something to the galaxies and structure that I observed in the universe.

278

00:41:08.580 --> 00:41:19.710

Tim Tait: And the answer turns out to be around 580 or so, which is pretty amusing when close to the actual value of the strong coupling

constant you know the real glue on have lambda up about two or 300 and maybe

279

00:41:20.880 --> 00:41:24.120

Tim Tait: That's of course a coincidence. But it's, it's an interesting one.

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00:41:25.350 --> 00:41:31.290

Tim Tait: You can estimate the relevancy of the glue balls by tracking the relevancy of nuance to temperatures where the theory combines

281

00:41:31.650 --> 00:41:40.800

Tim Tait: So that's actually what I've done over here is a quick back of the envelope calculation basically all I do is I just say the blue balls are the blue ones have some temperature

282

00:41:41.700 --> 00:41:46.290

Tim Tait: At any point in the universe that tells me what their energy density is based on the equilibrium density

283

00:41:46.650 --> 00:41:56.730

Tim Tait: Then when I get to the confinement scale a temperature around lambda, the energy of the dark blue ones gets converted into the blue balls. So I just conserve the energy and turn that much glue on into that much blue ball.

284

00:41:58.200 --> 00:42:00.480

Tim Tait: And so then you estimate the relevancy of them that way.

285

00:42:02.370 --> 00:42:11.850

Tim Tait: And this is where it's important that if there are no relevant connectors between the visible the hidden sectors than the temperature in the hidden sector t_h and the temperature, the visible sector t_v

286

00:42:12.270 --> 00:42:21.240

Tim Tait: Could actually generically be different from each other and you parameters. This possibility with the ratio of the temperatures, which is fine to be t_h divided by T .

287

00:42:21.960 --> 00:42:28.860

Tim Tait: So they're interesting corrections and the usual thermal distributions. And that's because your processes were saved three blue balls.

288

00:42:29.580 --> 00:42:36.630

Tim Tait: All annihilate to produce to blue balls. So this is a process that was called cannibalization by the first reference listed here.

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00:42:36.990 --> 00:42:45.480

Tim Tait: Because what it does is it basically takes the energy of one of the glue balls and the mass energy, one of the blue balls and turns it into kinetic energy of the two that come out.

290

00:42:46.020 --> 00:42:52.680

Tim Tait: Right. Or in other words, it's cannibalization because these three blue balls have gotten together. Two of them have eaten their friend and now they have more energy.

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00:42:55.110 --> 00:43:06.420

Tim Tait: And so here, if you want to look at this parameter space for blue ball only dark matter I plotted it in the parameter space of this relative size of the temperatures. So let's see parameter on the x axis.

292

00:43:07.530 --> 00:43:19.680

Tim Tait: Right, that's T hidden divided by T standard model and then what the confinement scale is we already calculated that you got interesting self interactions where you might see something funny going on in galaxy formation.

293

00:43:20.370 --> 00:43:25.560

Tim Tait: Is sort of down around a few hundred N_{ev} for the confinement scale.

294

00:43:26.100 --> 00:43:35.430

Tim Tait: And then for different values of the number of colors and you know to is reasonable 10s seems a little larger hundred seems kind of crazy, but just to get an idea how it scales.

295

00:43:36.240 --> 00:43:42.960

Tim Tait: You can see the lines in this parameter space of the relative temperature and λ , where you get the right relic density

296

00:43:43.500 --> 00:43:48.060

Tim Tait: So, you know, if you live down here, say, on this point, you can have two colors of dark blue ones.

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00:43:48.630 --> 00:44:00.360

Tim Tait: The relative temperature between the hidden in Standard Model sector would be something like one over 100 and you'd have interesting

cross sections that might lead to observable changes and say small galaxies.

298

00:44:01.980 --> 00:44:13.380

Tim Tait: Okay I think just to make sure I get through some other important things. I'm going to skip the case where I add something like a dark cork. In this case, it was in the joint representation. So it was more like a dark gloomy know

299

00:44:14.400 --> 00:44:18.030

Tim Tait: You can go through those. And you can always send me an email. If you want to talk about them some more.

300

00:44:19.860 --> 00:44:33.240

Tim Tait: But I want to actually finish up by saying something about dark about light, dark force carriers. In other words, cases where the dark force carrier is a light being low mass particle and you'll have to forgive me for indulging in the absurd workplace.

301

00:44:35.430 --> 00:44:41.400

Tim Tait: So recently, there's been a lot of attention given to the parameter space where the mediating particles have low masses that will be couplings.

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00:44:41.910 --> 00:44:46.680

Tim Tait: This is a natural parameter space that complements high mass searches that colliders

303

00:44:47.670 --> 00:44:56.550

Tim Tait: For Dark Matter of masses below around 10 GB there are strong constraints which favorite cases in which the dark matter annihilate directly into some kind of light, dark mediator.

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00:44:57.270 --> 00:45:04.860

Tim Tait: So if we would like to realize the abundance of dark matter through the freeze out mechanism we need light mediators for dark matter masses below the GV scale.

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00:45:05.370 --> 00:45:12.450

Tim Tait: And so here are some plots that were made in the reference. You can see here that came out just this year what they've done is they've looked at

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00:45:13.530 --> 00:45:26.580

Tim Tait: They were very clever and analyzing the parameters and figuring out which combinations work, which means that they have plots that are really applicable to a lot of parameter space, but it also makes your head hurt a little bit. When you try to analyze it. So,

307

00:45:27.630 --> 00:45:38.070

Tim Tait: Just rest assured that these were a very good choice, but they're just not an easy choice. So what's been done here is the x axis contains the ratio of the mediator mass to the dark matter mass.

308

00:45:38.610 --> 00:45:44.670

Tim Tait: Right, so pick the dark matter mass and the mediator mass is the number on the x axis times that aren't matter mass.

309

00:45:45.420 --> 00:45:55.950

Tim Tait: Then the y axis has got a combination of parameters that includes a kinetic mixing parameter. So in other words, how much this dark mediator mixes with the photon. We're going to see what that means. On the next slide.

310

00:45:56.670 --> 00:46:06.090

Tim Tait: Alpha D is the coupling of that mediator to the dark particle. And then the ratio of the dark matter mass to the mediator mass, right, which is in the opposite.

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00:46:07.290 --> 00:46:13.050

Tim Tait: Order here to the fourth power. Right. And that's because the annihilation cross section skills like this quantity

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00:46:13.590 --> 00:46:23.460

Tim Tait: And so you can actually plot what the thermal target is for a freeze out relic here. And so what they've done is they plotted it for a couple different choices of the dark coupling constant

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00:46:24.090 --> 00:46:27.720

Tim Tait: You can see the thermal target has got this interesting structure.

314

00:46:28.200 --> 00:46:42.870

Tim Tait: It this this dip here right is the same kind of resonant annihilation that we saw could happen for supersymmetric particles. Right. It's what happens basically when r is equal to two and therefore, the, the mediator mass is equal to twice the dark matter.

315

00:46:44.730 --> 00:46:58.800

Tim Tait: So this was a case for something like a dark photon and the upper plot in the lower plot you see a case where you're only a couple of baryons right so you don't couple the leptons that changes the systematics of things a little bit, especially the way that you would search for it.

316

00:47:00.330 --> 00:47:00.750

Tim Tait: Okay.

317

00:47:01.950 --> 00:47:09.750

Tim Tait: So like mediators can also help with some of the puzzling measurements that we wonder might be indications and new physics that we know about.

318

00:47:10.230 --> 00:47:16.230

Tim Tait: The most famous one of these is G minus two of the Milan. So there's a long standing discrepancy between theory and data.

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00:47:17.070 --> 00:47:27.360

Tim Tait: This plot here shows you the experimental measurements from you, plus minus and their average. You can see they're consistent with each other so doesn't look like a premier plus the vinyl have a different t minus two.

320

00:47:27.990 --> 00:47:30.330

Tim Tait: That's going to do this quantum field theory doesn't really let that happen.

321

00:47:31.830 --> 00:47:38.430

Tim Tait: And the theory theoretical determinations based on some inputs from either a plus or minus data or tower decay data.

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00:47:38.790 --> 00:47:45.690

Tim Tait: And these are really just trying to actually understand some of the hydraulic contributions to the process because at this point we're sensitive to loop corrections.

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00:47:46.410 --> 00:47:55.500

Tim Tait: And so there's actually some preference for the plus and minus result, both of them though are pretty discrepant with the experiment by a few sigma

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00:47:56.130 --> 00:48:08.940

Tim Tait: And so the idea is if you have a light, dark force carrier, like say a dark photon. You could imagine that it also gives you a one

loop contribution to the TG minus two of them you on and it could help fix up this

325

00:48:09.990 --> 00:48:19.320

Tim Tait: This discrepancy so somewhat less established and less. Definitely. Sorry. One last thing of a d minus two of them, the one

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00:48:19.980 --> 00:48:29.370

Tim Tait: We should actually get updates about measurements of G minus two of them you on now experiment is running a Fermilab and I believe that that should really happen anytime

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00:48:31.170 --> 00:48:40.080

Tim Tait: So then moving on to a couple other interesting things to keep on your radar. There's a nuclear physics transition which I've got a cartoon here that was made by flip tomato.

328

00:48:40.620 --> 00:48:46.950

Tim Tait: Where an experiment produces and excited state of brilliant. He watches it decay into the ground state plus

329

00:48:47.820 --> 00:48:54.270

Tim Tait: A plus A minus and they see what looks like a resonance structure in the policy minus which might be the sign of a new particle around 17

330

00:48:55.260 --> 00:49:01.770

Tim Tait: So the main thing here with this experiment is that it's really only done by one experimental group and you'd really like to see it reproduced by another one.

331

00:49:02.130 --> 00:49:10.320

Tim Tait: Though they have done more measurements with other nuclei and they seem to see similar effects. So really, though, that just makes me say once again.

332

00:49:11.040 --> 00:49:18.180

Tim Tait: You really would like to see it done by another experimental group because you're really worried a little bit about whether it's something funny about their experiment.

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00:49:19.350 --> 00:49:26.340

Tim Tait: Though no one has found any problem with that so far. And then the last one is an older measurement of pi zero d plus A minus that was

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00:49:26.970 --> 00:49:30.480

Tim Tait: Mentioned by Katana it's off by about two and a half sigma or so.

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00:49:30.870 --> 00:49:40.440

Tim Tait: And it might point to axial couplings to quirks and electrons. So here's a diagram where you have the two quirks and make up the pi zero they fuse into a massive mediator.

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00:49:40.860 --> 00:49:46.890

Tim Tait: And then that gives you extra he plus a minus in your decay and you need to have axial couplings, at least with the cork side.

337

00:49:51.540 --> 00:50:02.130

Tim Tait: Alright light mediators also open up the possibility to invoke a summer felt like enhancement of the dark matter small dark matter of velocities. So this could enhance annihilation.

338

00:50:03.210 --> 00:50:08.280

Tim Tait: I'm guessing Tracy probably told you a little bit about this yesterday, but unfortunately didn't get a chance to take a look at our lecture.

339

00:50:10.020 --> 00:50:19.200

Tim Tait: Summing up the effect of the mediator on the scattering can lead to a large enhancement compared to the leading order annihilation rate and right the diagram is something like this one here on the lower right side.

340

00:50:19.710 --> 00:50:31.650

Tim Tait: Where you have your dark matter coming in. And the idea is because of the attractive force between the mediator. It's actually more likely to get together and annihilate than it would if there was no such interaction pulling the incoming particles together.

341

00:50:32.040 --> 00:50:34.260

charlie young: Excuse me five minutes.

342

00:50:34.710 --> 00:50:36.000

Tim Tait: Yeah. Absolutely. Great. Thanks.

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00:50:37.110 --> 00:50:43.320

Tim Tait: The impact is enhanced for small velocities and also for me to your immediate or masses that are much smaller than the mass of the dark matter.

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00:50:43.650 --> 00:50:48.300

Tim Tait: You can actually see the calculation here, which was done by Tracy and some collaborators, a while ago.

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00:50:48.960 --> 00:50:54.480

Tim Tait: And this is plotted in the plane of ϵV , right, which is velocity divided by the coupling α

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00:50:55.170 --> 00:51:03.330

Tim Tait: On the x axis and the y axis is ϵf , which is defined to be a mass of the mediator divided by α times the mass of the dark matter.

347

00:51:03.960 --> 00:51:15.030

Tim Tait: And you see, you get big enhancements like as big as even factors of 1000 or 10,000 if you go to small values of both of those things. In other words, small velocities and small mediator mass compared to the dark matter less

348

00:51:16.770 --> 00:51:25.380

Tim Tait: So 400 GB dark matter if you want to take advantage of this kind of enhancement. This naturally argues the mediator mass should be down to be somewhere at or below the GPA scale.

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00:51:27.240 --> 00:51:34.320

Tim Tait: And another interesting thing that gets opened up. If I talk about like mediators is something I alluded to previously when I talked about the relative density.

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00:51:34.770 --> 00:51:41.010

Tim Tait: You can imagine the dark matter, instead of annihilating in the standard model particles also just annihilate directly into mediators themselves.

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00:51:41.490 --> 00:51:57.120

Tim Tait: So the rate for this to happen is fixed by the coupling of the dark matter to the mediator and so for example if I have dark matter annihilating into mediators, like in the top diagram that annihilation cross section is going to go like that coupling of mediator to dark matter.

352

00:51:58.860 --> 00:52:07.380

Tim Tait: To the fourth power. So that's what this annihilation cross section go like, and the media, you're going to have much much smaller coupling to the Standard Model particles like called that epsilon

353

00:52:07.770 --> 00:52:12.570

Tim Tait: And that's because you know if I'm worried about seeing the the annihilation products of the dark matter.

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00:52:13.410 --> 00:52:23.520

Tim Tait: The mediator can travel for killer parsecs before it needs to decay. In other words, this coupling this lifetime could be incredibly long and I would still get the same signal, but I would expect to see in my telescopes.

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00:52:24.510 --> 00:52:34.320

Tim Tait: So it's a generic way of getting an indirect signal while suppressing direct and collider constraints, because if I look for example at the direct scattering dark matter with a nuclear

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00:52:36.180 --> 00:52:45.240

Tim Tait: Exchanging a mediator. Right. Well, now I have a vertex of mediator with dark matter. So that's G and the vertex of mediator with sorry Meteor dark matters up here that's G.

357

00:52:45.660 --> 00:52:52.650

Tim Tait: Mediator with the nucleus that's down here is epsilon. Now this cross section goes like epsilon squared d squared. And so

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00:52:53.490 --> 00:52:58.470

Tim Tait: It's a parametric Lee smaller than the annihilation cross section and obviously

359

00:52:59.130 --> 00:53:05.130

Tim Tait: If I make epsilon small I can have the annihilation. Be very prominent while having the direct section be switched off.

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00:53:05.700 --> 00:53:14.160

Tim Tait: And we've seen, actually this idea to get some application in terms of the GV access of gamma rays from the direction of the galactic center to construct theories in which

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00:53:14.700 --> 00:53:23.430

Tim Tait: You get the strong annihilation signal you would need to explain that observation. But while still satisfying the lack of evidence for dark matter scattering with having a clear

362

00:53:25.980 --> 00:53:37.260

Tim Tait: Okay then. Last thing I want to mention or the last topic I want to mention is how you would build a theory of light mediators, they're very important theoretical constraints on theories with light particles.

363

00:53:37.740 --> 00:53:43.050

Tim Tait: So of course, there are the experimental constraints on searches for them or indirect contributions other observable.

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00:53:43.650 --> 00:53:50.700

Tim Tait: And such constraints are most easily avoided if the mediator carries no standard model charge, but I also have to worry about the Standard Model gauge cemetery.

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00:53:51.210 --> 00:53:59.970

Tim Tait: The mediator gauge cemetery, if I write down a mediator. That's a vector particle, I need to actually have a gauge symmetry to protect its interactions or generically my theory won't be self consistent

366

00:54:00.720 --> 00:54:07.110

Tim Tait: We need to be able to write down the standard model you call interaction. So when when I introduce new kinds of charge, I run the risk of

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00:54:07.560 --> 00:54:15.630

Tim Tait: Not being able to write down those interactions and I have to figure out how to engineer them. And of course there could be gauged anomalies, which mean the theory is not self consistent

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00:54:16.530 --> 00:54:23.490

Tim Tait: All of these summarize actually in this plot here where you see a bunch of different constraints on an axial vector particle

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00:54:24.600 --> 00:54:29.340

Tim Tait: You see a bunch of constraints experimentally. And then if you

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00:54:29.790 --> 00:54:37.920

Tim Tait: In particular, if you're looking at what's allowed by g minus two of them Yuan, right. So, g minus to the new one favors a certain range of region region a parameter space.

371

00:54:38.250 --> 00:54:42.210

Tim Tait: It likes this orange band here in terms of the mass. The mediator and coupling.

372

00:54:43.110 --> 00:54:56.190

Tim Tait: To leptons and g minus two of the electron forces you to live in this little tiny sliver that goes like this lightens up a little bit larger masses. So actually, it's very hard to get things to be totally self consistent

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00:54:59.550 --> 00:55:10.020

Tim Tait: And even if you have canceled. Your anomalies, there are still some contributions that come from processes involving longitudinal modes of your vector particle

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00:55:11.610 --> 00:55:25.080

Tim Tait: At high energies. So I'm really out of time so I'm actually just going to leave that at that. But these are important constraints that you have to look at that are not obvious because they are really can be thought of at some level was one loop processes, but they're still very important

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00:55:27.030 --> 00:55:37.050

Tim Tait: So, and then an idea which is a received a lot of attention is the postulate. The new light force carriers, a dark photon and the idea is you add a new vector boson with a small mass

376

00:55:37.680 --> 00:55:43.080

Tim Tait: And that's probably a whole dark Higgs sector under which the dark matter is charged with the standard model is not

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00:55:43.530 --> 00:55:50.400

Tim Tait: But then you include this last term in the long run, which is the kinetic mixing term between the dark photon and the hypercharged goes on.

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00:55:50.790 --> 00:55:56.730

Tim Tait: And then when you die, analyze the kinetic terms, you end up with an effective coupling of g to the dark matter.

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00:55:57.150 --> 00:56:03.360

Tim Tait: But a coupling to Standard Model particles that looks like the electromagnetic coupling scaled down by the parameter ϵ

380

00:56:04.230 --> 00:56:12.540

Tim Tait: So that's one is small you naturally get small coupling to the standard model. I know. Jonathan things spent some time talking about this case, I'm not going to go through in great detail.

381

00:56:13.830 --> 00:56:20.280

Tim Tait: And of course, he told you all about all the great experimental searches that are currently being plotted to find particles like once

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00:56:21.450 --> 00:56:27.000

Tim Tait: So to recap, part three, a complex dark sector expands the complexity of the physics of dark matter.

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00:56:27.570 --> 00:56:33.870

Tim Tait: Given the complexity of that visible sector, it seems natural to imagine that the dark sector could impact could encompass a morally rich physics.

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00:56:34.290 --> 00:56:41.310

Tim Tait: Dark Matter could be a composite built out of more fundamental ingredients and held together. You either by weak or strong self interactions.

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00:56:41.910 --> 00:56:47.820

Tim Tait: A light mediator allows for several different phenomena and has become a standard element of the dark matter model, model building toolkit.

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00:56:48.450 --> 00:56:53.670

Tim Tait: It could make contact with other low energy anomalies that are currently experimentally observed

387

00:56:54.300 --> 00:57:01.530

Tim Tait: They open up new processes for dark matter to participate in, and they can be searched at lower energy but higher luminosity experimental facilities.

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00:57:02.100 --> 00:57:12.720

Tim Tait: So while we're still exploring the space of possible theories of dark matter. There's a rich experimental program that will explore these interesting regions of parameter space. And of course, there's a lot more to come, both from the theory and from the experiment.

389

00:57:13.650 --> 00:57:17.850

Tim Tait: So I'll leave you with my favorite diagram of dark matter. And thank you very much for listening.

390

00:57:19.500 --> 00:57:32.580

charlie young: Thank you very much. Tim, I really enjoy your series of three lectures on that matter tech sector. Great. So let's go to Q AMP. A now and Greg. Will you take over, please.

391

00:57:33.030 --> 00:57:34.590

grzegorz madejski: Sure. Um, can you hear me.

392

00:57:35.280 --> 00:57:46.770

grzegorz madejski: Yes, yeah. Great. Okay. So let me get cracking on questions, quite a few of them came in, even after your last lecture, we have a grand total of it but 658 for all three like

393

00:57:48.000 --> 00:58:05.010

grzegorz madejski: Yeah, so I certainly will not be able to get through all of them, but let me start with the very first one slide for the question is, is there any special relationship between the accident and the Python. Could you also relate the mass of the axiom to the massive another Meza

394

00:58:06.510 --> 00:58:08.970

Tim Tait: Um, there is actually a special relationship.

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00:58:10.050 --> 00:58:20.430

Tim Tait: In some sense. I mean, it's not that special. But what happens is the accion mixes have very little bit with all of the pseudo scale or means on right the accion is a pseudo Skyler and because

396

00:58:20.940 --> 00:58:27.030

Tim Tait: By virtue of its existence CP is a good symmetry. So therefore it can only mix with the pseudo scale or maisons

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00:58:27.540 --> 00:58:37.710

Tim Tait: Now it can mix with all of them. And it doesn't mix with all of them. However, the mixing goes something like the accion mass divided by the pie on, Matt. Sorry, by the mass of the maze on it's mixing with

398

00:58:38.340 --> 00:58:46.470

Tim Tait: And so it mixes the strongest with the lightest pseudo scale or means ons and therefore it mixes most strongly with the pion which is the lightest muscular

399

00:58:48.210 --> 00:58:50.940

Tim Tait: But there's nothing really special about the pie on other than that.

400

00:58:52.410 --> 00:59:04.290

grzegorz madejski: OK, and now the one that has to do with Slide six would the existence of a sterile neutrino in play the existence of a sterile leptons counterpart.

401

00:59:06.360 --> 00:59:15.420

Tim Tait: I mean, in some sense, a sterile neutrino is a sterile leptons. I guess if you're asking. I think probably the question is asking, is there a star. I'll charge left on

402

00:59:16.530 --> 00:59:18.390

Tim Tait: Which doesn't quite make you know

403

00:59:19.530 --> 00:59:26.730

Tim Tait: It doesn't make Samantha semantic sense because if it's sterile. It can't have electric charge. And therefore, it wouldn't look like a charge left on

404

00:59:27.300 --> 00:59:38.940

Tim Tait: But I think, to go back to the real question no. So the standard model left on ones that are left handed coming doublets. And so the neutrino that's left handed comes with an electron that's left handed or some other charge left on

405

00:59:40.380 --> 00:59:46.230

Tim Tait: The sterile neutrino isn't electroweak singlet and so it comes all by itself without the need for anything else.

406

00:59:48.180 --> 00:59:56.580

grzegorz madejski: Okay, the next one has to do Western neutrinos can sterile neutrino dark matter also be produced the ref result.

407

00:59:58.170 --> 01:00:02.400

grzegorz madejski: Can they also be much heavier than the key scale. That's a good question.

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01:00:03.000 --> 01:00:12.300

Tim Tait: It is a good question. So, in principle, the answer is yes, a sterile neutrino could be produced by freeze out if it had you know sufficiently weak interactions.

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01:00:13.260 --> 01:00:22.650

Tim Tait: The thing is that the sterile neutrino at least the one that's responsible for producing neutrino masses has this decay diagram that I showed here, which means that

410

01:00:23.280 --> 01:00:32.520

Tim Tait: In general, as I make it heavier and the rate of that decay goes up. Right. Remember, because of phase space your particles are more likely to decay to the lighter particles.

411

01:00:34.200 --> 01:00:41.520

Tim Tait: It becomes more tricky to have the neutrino live long enough to be the dark matter and other words to have a lifetime. That's the order of the age of the universe.

412

01:00:42.270 --> 01:00:52.950

Tim Tait: So I think once you get to the point where you realize this, you'd have to make that mixing angle so tiny that it wouldn't probably be able to explain the neutrino masses that we see.

413

01:00:53.580 --> 01:01:03.150

Tim Tait: So it would still be a sterile neutrino. It could still be dark matter, it could still be produced freeze out but it probably would not be the explanation for why we see the neutrino masses that we do in the standard

414

01:01:05.940 --> 01:01:22.830

grzegorz madejski: Alright, the next one is probably there's sort of two questions that are relevant here that are very similar. Could you speak a little bit more on why the x ray signal is only possible one, I think that they're referring to the three and a half keV line.

415

01:01:24.060 --> 01:01:25.230

grzegorz madejski: The original one at

416

01:01:26.490 --> 01:01:30.450

Tim Tait: Right, so the signal is a three and a half keV.

417

01:01:32.340 --> 01:01:42.780

Tim Tait: It was observed actually by two different groups. So what's been plotted here is the one from the Bible at all. But there's also another group in Europe, which is seemed very similar results. And actually, in a very similar energy

418

01:01:45.060 --> 01:01:53.670

Tim Tait: The signal is seen by looking at stacks of galaxies. So you take a whole bunch of galaxies and, you know, they all have different they've all got different stuff in them.

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01:01:55.140 --> 01:01:57.510

Tim Tait: Greg, actually you know a lot about this yourself personally

420

01:01:58.740 --> 01:02:10.260

Tim Tait: I think right that's right. It was the clusters, rather than galaxies, but the dark matter should always occur at the same energy. And so you put them all together all the data together and you hope the dark matter signal when peak true compared to the other stuff.

421

01:02:12.090 --> 01:02:23.520

Tim Tait: And even if you look at certain sets of data, you find a very nice signal. It doesn't seem to be explained by known atomic physics and it seems like it's distinct from the background noise.

422

01:02:24.690 --> 01:02:38.010

Tim Tait: Other groups, however, have looked in other locations. So for example, if you look at the Milky Way, our own galaxy. Right. It's got a lot less dark matter in it, but it's a lot closer to us. So any signal of dark matter, you know, is relatively a lot brighter.

423

01:02:39.150 --> 01:02:43.740

Tim Tait: You see what maybe a hint of a signal, but it's really not very convincing coming from our galaxy.

424

01:02:44.340 --> 01:02:49.650

Tim Tait: If you look at different sets of galaxy clusters, you actually find you don't really see the signal at all.

425

01:02:50.280 --> 01:03:01.920

Tim Tait: So that's the real reason we have sort of two things going on. Experimentally, it doesn't seem like we see the signal absolutely everywhere that we might hope to and so that that cast some doubt as to whether it's really dark matter.

426

01:03:02.550 --> 01:03:10.230

Tim Tait: And from a more theoretical side. There's also been some work that suggests that maybe there are even ordinary Standard Model lines that go to current that energy

427

01:03:10.680 --> 01:03:20.250

Tim Tait: That just you know correspond to kind of exotic atomic physics you know say involving potassium or something that we just don't really have very good measurements of or very good theoretical understanding of

428

01:03:21.480 --> 01:03:34.770

grzegorz madejski: Yeah, I think I can add here that there is a pretty strong evidence that settling does not exist in the observations of the prestigious cluster which we're done with the very high resolution energy resolution instrument on board if he told me before. Anytime you failed.

429

01:03:36.300 --> 01:03:43.380

grzegorz madejski: So hopefully we will have a lot more cluster data and in your future from the Athena and even before that from the prison missions.

430

01:03:45.030 --> 01:04:00.870

grzegorz madejski: Okay, so another question. Where does the Zenon one t experiments axiom hypothesis recently published fit into this plot on slide five, and is it intention with the current astrophysical limits.

431

01:04:02.130 --> 01:04:12.480

Tim Tait: Okay, so it doesn't really fit onto this plot at all the xenon does a report in excess of something scattering with electrons actually

432

01:04:12.930 --> 01:04:22.860

Tim Tait: Relativistic Lee, so something that carries a lot of kinetic energy. And as a result, it doesn't look very much like you'd expect. Dark matter to write dark matter should be non relativistic in order to be part of our Halo.

433

01:04:24.090 --> 01:04:24.840

Tim Tait: So,

434

01:04:27.150 --> 01:04:38.400

Tim Tait: So yeah, it doesn't really fit onto this plot. It is, however, in some tension with astrophysical limits for how astrophysical objects should be cooling. Right. Once you have a light.

435

01:04:38.850 --> 01:04:48.360

Tim Tait: Particle that couples to something say that your, your star is made out of your star can admit those particles and cool down and if it cools down too fast right faster than

436

01:04:48.810 --> 01:04:54.810

Tim Tait: Then is allowed, given the population of those objects you see around you, then you can put a limit on to the existence of that particle

437

01:04:55.770 --> 01:05:11.160

Tim Tait: There has been some discussion that these limits have some fairly large uncertainties that are not really very well quantified and so I think the situation is not totally clear definitely if you do things, you know, very simply, and just make a simple estimate its intention.

438

01:05:12.630 --> 01:05:18.240

Tim Tait: So that would argue, you have to do some work to try to explain your way around that tension by making your theory more complicated.

439

01:05:19.080 --> 01:05:27.000

Tim Tait: However, there is probably enough inherent uncertainty in the estimates that I think there's also a very healthy process going on right now where people are revisiting them and trying to understand

440

01:05:27.600 --> 01:05:40.350

Tim Tait: You know how we can make a bit more informed estimate of the limit, and that also might end up giving a may end up revealing that there's some corner of parameter space that's allowed but you know that hasn't been finished yet.

441

01:05:42.270 --> 01:05:59.820

grzegorz madejski: Okay, there is this actually relevant to another question which you probably answered already in this this last statement on slide 12 the questioner at the person who asked the question says for these plots are the colored areas allowed are prohibited by ATLAS and CMS.

442

01:06:00.840 --> 01:06:14.670

Tim Tait: Yeah, no, that's a good question. The two plots actually use color somewhat differently. So let's start with the Atlas plot basically the Atlas plot is that at 95% competence limit everything enclosed inside this solid black line is excluded.

443

01:06:16.290 --> 01:06:23.640

Tim Tait: And everything outside of it is allowed. Right. So everything that everything, starting with the 00 mass point and you know inside the wedge is excluded.

444

01:06:24.420 --> 01:06:34.080

Tim Tait: The colors here are really just showing what the expected limit was compared to the observed limit. And that's something you always want

to do because you want to make sure that something has not gone wrong with your analysis.

445

01:06:34.560 --> 01:06:46.620

Tim Tait: So, you know, if you get a limit that looks wildly different than one of the one you expected to get, then chances are you forgot to do something or you didn't take some effect into account when you made your estimation and that calls into question kind of the whole thing.

446

01:06:47.820 --> 01:06:54.210

Tim Tait: The CMS actually limit us as something usually uses color in a very different way. So,

447

01:06:55.560 --> 01:07:02.520

Tim Tait: In general what it's doing is it's allowing that the production cross section for this process can be scaled up and down.

448

01:07:04.080 --> 01:07:11.370

Tim Tait: Compared to what you would expect to get for the couplings that they chose so right they chose these couplings to quarks and couplings. The dark matter.

449

01:07:11.820 --> 01:07:18.540

Tim Tait: And if you make that choice, you would be at one on their color scheme. And that's why, if you look at where that color occurs. It's sort of this

450

01:07:19.290 --> 01:07:30.990

Tim Tait: Greenish in between yellow and you know turquoise color. It's the red line and that's their observed 95% confidence limit. So again, what's excluded is inside here. If these couplings are true.

451

01:07:32.130 --> 01:07:42.150

Tim Tait: If you were to scale the couplings down though, you would produce skewer particles and you get a weaker limit. And so then the shading of the colors shows you how to do that. So if you were to

452

01:07:42.660 --> 01:07:56.760

Tim Tait: Scale them down scale, the whole cross section down by a factor of 10 you'd go from this greenish color to this kind of like blue color on the boundary between turquoise and deeper blue and you get a limit that would look I guess something more like

453

01:07:58.200 --> 01:08:00.960

Tim Tait: Like a few hundred GB to maybe a TV.

454

01:08:02.520 --> 01:08:04.290

Tim Tait: I don't know if you can see me trying to

455

01:08:05.910 --> 01:08:07.740

Tim Tait: sketch it with the cursor.

456

01:08:09.060 --> 01:08:20.040

Tim Tait: So yeah, the color on the CMS plot is trying to give you some idea of what would happen if you change your assumptions about the color couplings, the color on the Atlas plot is really just trying to compare what they saw versus what they expected to see.

457

01:08:21.990 --> 01:08:31.170

grzegorz madejski: Okay, now there's a fairly complex question which sort of is listed in two parts. So let me read the first part and then I'll say the second thing I'm not sure.

458

01:08:31.260 --> 01:08:42.300

grzegorz madejski: If I understand completely. Is this quote simplified model basically searching through a more general parameter space of dark matter mass media mass or cap link.

459

01:08:42.930 --> 01:08:52.950

grzegorz madejski: And seeing which overall combinations of values are allowed so far by data or not. Instead of trying to test one particular theory at a time.

460

01:08:53.730 --> 01:09:08.070

grzegorz madejski: And then they elaboration of this is for full models, what is typical range of VR of Z underscore V going to Chi Chi and how do they influence the results.

461

01:09:09.780 --> 01:09:11.610

Tim Tait: So the original

462

01:09:12.660 --> 01:09:13.470

Tim Tait: Question.

463

01:09:16.500 --> 01:09:21.210

Tim Tait: I think the original statement is a fair summary, I would just put the caveat on it that

464

01:09:22.200 --> 01:09:31.110

Tim Tait: You still need to specify the structure of the simplified model. So it's not totally generic in mass, mass and coupling, you have to add the structure, like for example.

465

01:09:31.500 --> 01:09:46.800

Tim Tait: In this case, the dark matter interacts with a vector mediator, which also interacts with corks in the other case, we had a case where the dark matter interacted with a scale or particle that carried us to three color but up to that quibble the statement is correct.

466

01:09:48.930 --> 01:09:51.930

Tim Tait: This question about the branching ratio.

467

01:09:53.280 --> 01:09:55.560

Tim Tait: I'm not sure.

468

01:09:57.960 --> 01:10:08.880

Tim Tait: I mean, it's a model specific question. And so in general a full theory will make a prediction for what that branching ratio is even if a simplified model doesn't. Maybe that's the best way to answer it.

469

01:10:10.770 --> 01:10:27.300

grzegorz madejski: Alright, the next one is in the vector and scale our dark mediator theories are this mediators, then the bulk of what we see as dark matter or are they in addition to the dark matter that makes up 25% of the universe pie.

470

01:10:28.740 --> 01:10:39.570

Tim Tait: Right. So actually, in both of these cases, the mediator particle is not actually the dark matter, it would have dictated into the dark matter by today and so

471

01:10:40.590 --> 01:10:48.960

Tim Tait: It makes up, none of the dark matter and they are some particles that are in addition that are not currently present in the universe. But we're at early times

472

01:10:50.700 --> 01:11:00.450

grzegorz madejski: Okay, I think I'll have time for questions. So we should thank you very much them for great lecture today and series of three excellent lecture so I

473

01:11:01.110 --> 01:11:12.330

grzegorz madejski: Will send you all our questions are in some kind of form for you to put together answers for and then we're going to post them on indigo Charlie, do you have anything else to add before we take a break.

474

01:11:12.870 --> 01:11:20.250

charlie young: No, thank you very much. Tim thank you Greg for doing the Q AMP. A so I think I'll stop the recording now.