

RTE–cNSF: A Hybrid Physics–ML Architecture for Accurate Photon Timing in JUNO

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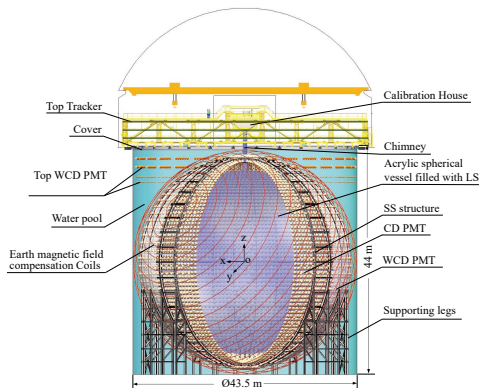
16, June 2026

The Jiangmen Underground Neutrino Observatory (JUNO)

The Detector

- 20 kton liquid scintillator (LS) target
- Acrylic sphere of 17.7 m radius, with a thickness of 12 cm.
- $\sim 18,000$ 20-inch PMTs.

Main physics goal: measure neutrino mass ordering via precise reactor $\bar{\nu}_e$ energy spectrum measurement.



Challenging Photon Timing

Photon timing is critical

- Precise energy and vertex reconstruction relies on accurate photon timing.
- Timing affected by complex optical processes: scattering, absorption, re-emission, and reflection.
- GEANT4 provides high-fidelity simulations but is computationally expensive to “fit” optical parameters.

The gain from differentiable simulation

With a differentiable photon timing model, we can do backpropagation through the likelihood to:

- extract the optical parameters directly from Calibration data (laser, ^{68}Ge , ...)
- perform end-to-end fits of physics parameters (vertex, energy) within the same framework.

The gain will be transformative, but how do we build this.

The Radiative Transfer Equation (RTE)

Instead of tracking individual photons, we model the **photon radiance** and its evolution in space, time:

Time-dependent RTE / Boltzmann Transport Equation

$$\underbrace{\frac{1}{c} \frac{\partial L}{\partial t} + \hat{s} \cdot \nabla L}_{\text{transport}} + \underbrace{\mu_t L}_{\text{loss}} = \underbrace{S(\mathbf{r}, t, \hat{s})}_{\text{source}} + \underbrace{\mu_s \int f(\hat{s} \cdot \hat{s}') L(\mathbf{r}, t, \hat{s}') d\hat{s}'}_{\text{in-scatter}}$$

$$\mu_t = \mu_a + \mu_s, \quad \mu_a = 1/L_{\text{abs}}, \quad \mu_s = 1/L_{\text{scat}}.$$

Source term S

Scintillation emitted *isotropically* from vertex V :

$$S = \frac{Y_{\text{eff}}}{4\pi} f_{\text{scint}}(t) \delta^3(\mathbf{r} - V)$$

f_{scint} : time PDF (bi-exponential $\otimes$ Gaussian).

In-scatter term

Photons in direction $\hat{s}'$ redirected into $\hat{s}$ with probability $f(\hat{s} \cdot \hat{s}')$:

- Rayleigh: $f \propto 1 + \cos^2 \vartheta$
- Re-emission (isotropic): $f = 1/(4\pi)$

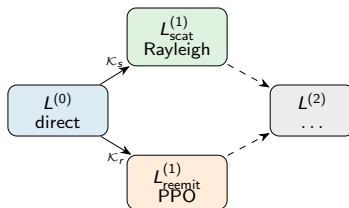
Key property

The in-scatter term is a **linear** operator on $L \Rightarrow$ the solution is a convergent **Neumann series**.

Neumann Series Solution

Rewrite as $(\mathcal{T} - \mathcal{K})L = S$; formal solution:

$$L = \sum_{n=0}^{\infty} L^{(n)} = \mathcal{T}^{-1}S + \mathcal{T}^{-1}\mathcal{K}\mathcal{T}^{-1}S + \mathcal{T}^{-1}\mathcal{K}^2\mathcal{T}^{-1}S + \dots$$



$n = 0$

direct (ballistic)

$n = 1$

Rayleigh scatter

$n = 1$

PPO re-emission

Implementation strategy

Each term is a separate **precomputed kernel** convolved with the scintillation PDF. Truncate at $n = 1$; Only model single scatter now.

Boundary — When things gets complicated

Boundary conditions

The acrylic sphere boundary causes:

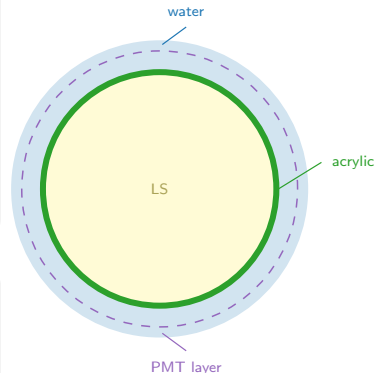
- Total internal reflection (TIR) for $\sin \theta_i > n_w/n_{LS}$.
- Partial Fresnel reflection and refraction for $\sin \theta_i < n_w/n_{LS}$.
- Reflection at the PMT photocathode, steel frame, $\dots$.

Introduce late photons, re-distribute photon directions — Need to be modelled.

TIR is a game changer

Sphere chord identity: angle of incidence preserved on every bounce:

TIR conditions hold for every bounce $\Rightarrow$ photon trapped indefinitely until scatter.



Schematic cross-section of the JUNO detector (not to scale).

The “dirty” geometry

To model a full sphere detector response like JUNO, we need to capture more than in LS:

When geometry gets “dirty”

- Photon path intersects the acrylic sphere and water outside $\Rightarrow$ Fresnel reflection/refraction, TIR, multiple bounces.
- Exact ray-tracing beyond the first boundary is messy: many surfaces, many bounces.

Approach 1: Pure RTE surrogate

- Consider boundary as a special scattering, introducing $L_{\text{boundary}}^{(1)}$ term in the Neumann series.
- Track TIR with a cut-off time, then consider the photon as uniformly leaving the detector.

Approach 2: Two-stage architecture

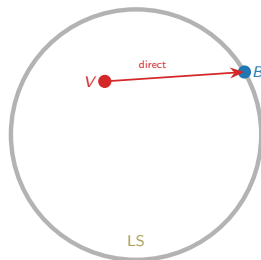
- **Our strategy:** Integrate the RTE surrogate without boundary physics, up to the first boundary hit.
- Post-boundary: use machine learning to model the “dirty” photon distribution.

LS-side: RTE Surrogate

Neumann series truncated at $n = 1$

$$L = L^{(0)} + L_{\text{Rayleigh}}^{(1)} + L_{\text{re-emit}}^{(1)}$$

- $L^{(0)}$: ballistic — photon travels straight to boundary without scattering.
- $L_{\text{Rayleigh}}^{(1)}$: single Rayleigh scatter en route.
- $L_{\text{re-emit}}^{(1)}$: absorbed then re-emitted by PPO (isotropic).



Ballistic

Evaluated *directly* at the unique first-hit state:

$$L^{(0)}(B, \hat{s}) \propto \frac{e^{-\mu_t d_{VB}}}{d_{VB}^2}$$

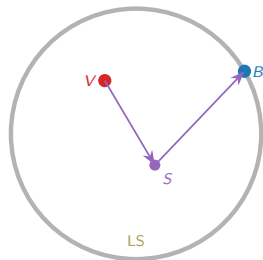
Direction $\hat{s}$ is fixed by geometry $\widehat{VB}$.

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Single-scatter

Integrated numerically over S in LS using **QMC** (Sobol):

$$L^{(1)}(R_{\text{src}}, \theta, \alpha_s, \phi_s)$$

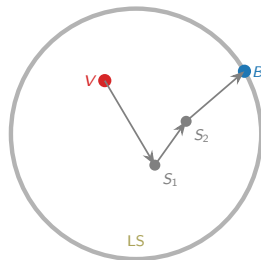
Sampled in (θ, ϕ) on sphere and (α_s, ϕ_s) exit angles.

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note: double scattering not modelled for now.

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Results from Approach 1

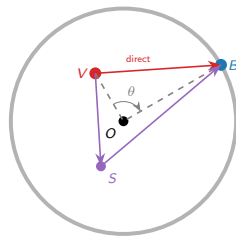
Data

Using GhostHunter 2025 competition data¹: Geant4 simulation with JUNO optical parameters and geometries.

The RTE parameters are fit to the Geant4 MC data.

Symmetry reduction

Since all geometries are spherically symmetric, the light field only depends on the radial source position R_{src} and the polar angle θ of the boundary hit.

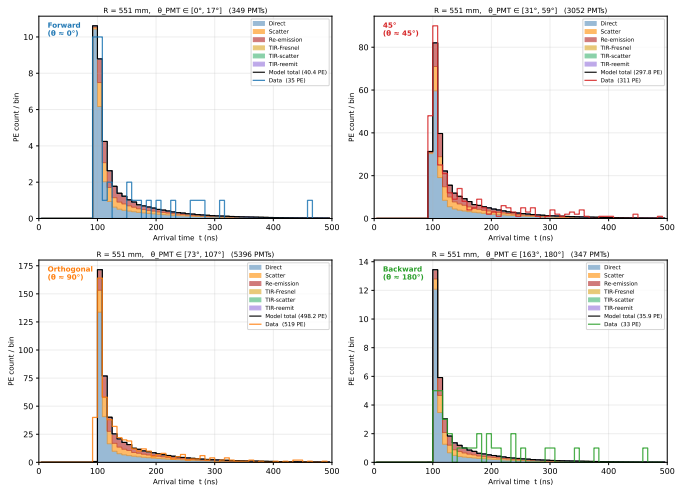


Comparison of the RTE surrogate with Geant4 data at different source positions R_{src} and boundary angles θ will be shown in the next slides.

¹<https://ghosthunter.thu.de/p/2025/>

Results from Approach 1

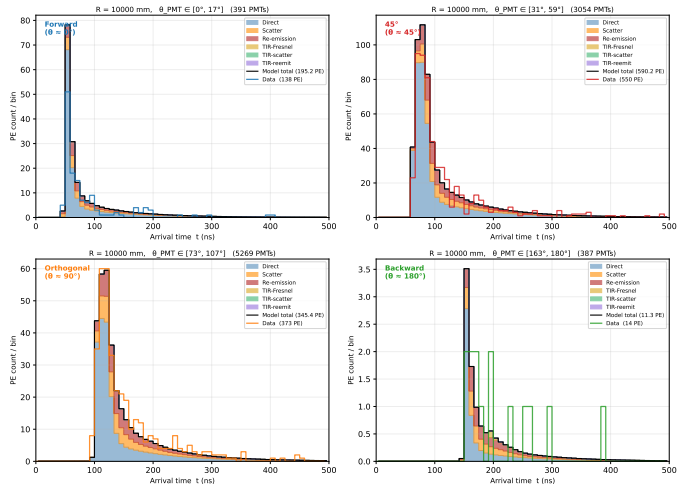
JUNO — $R = 551$ mm — PE Arrival Time by Angular Region



For $R = 551$ mm, the RTE surrogate (stacked) shows good agreement with Geant4 (histogram), single scatter term comprehensively captures the late tail.

Results from Approach 1

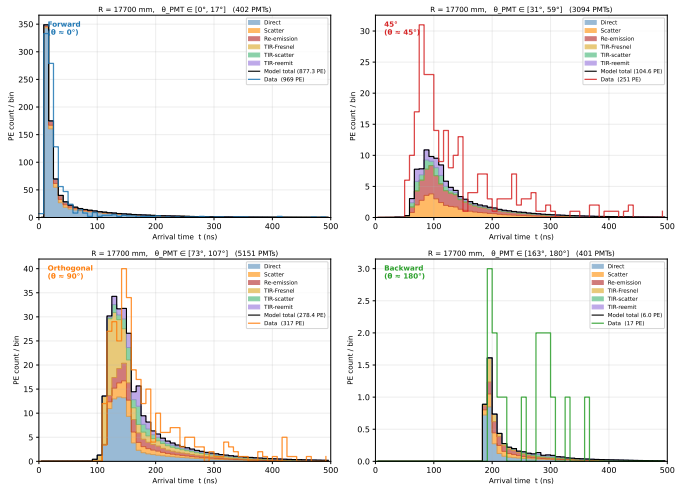
JUNO — $R = 10000$ mm — PE Arrival Time by Angular Region



For $R = 10000$ mm, no significant discrepancy is observed, the RTE surrogate still performs well.

Results from Approach 1

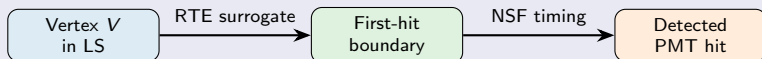
JUNO — $R = 17700$ mm — PE Arrival Time by Angular Region



For $R = 17700$ mm, the RTE surrogate fails to capture the photon contribution in the TIR dominated region ($\theta \approx 45^\circ$), leading to a large discrepancy with Geant4.

Two-Stage Architecture (for Approach 2)

Core idea: split at the first boundary



Stage 1: LS-side (RTE)

- Track photon transport *up to* the first boundary (Acrylic-LS) hit.
- Neumann-series surrogate: ballistic + single scatter.
- Output: first-hit boundary state $(B, \hat{s})$.

Stage 2: Outer-layer (NSF)

- Models all post-boundary physics: Fresnel, TIR, PMT optics.
- Frozen neural spline flow (NSF) learned from Geant4 data.
- Output: timing T at a specific PMT.

Key interface convention

The LS-side model is responsible only for the *first* hit on the effective boundary. All later reflection / TIR / re-hit histories are absorbed by the outer-layer kernel.

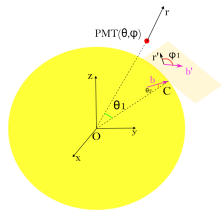
Outer-layer: cNSF Timing Model

Frozen conditional kernel

A pre-trained conditioned neural spline flow (cNSF) learns the conditional timing density:

$$p_{\text{cNSF}}(T \mid \theta_1, \theta_2, \phi_1, \text{detected})$$

Trained on Geant4 data; **frozen** during downstream fitting.



Condensed geometric variables

3 degrees of freedom due to spherical symmetry:

- θ_1 : PMT direction vs. incidence point from detector center.
- θ_2 : exit-angle quantity (Snell-mapped from α_s).
- ϕ_1 : tangent-plane azimuth of exit direction.

Key empirical result

When the *true (i.e., detected)* $(\theta_1, \theta_2, \phi_1)$ values are given, the frozen cNSF model can produce **excellent** timing distributions.

⇒ The main challenge is *getting the right boundary measure* into cond-space.

Full Detected-Hit Model

Factorized form

The detected-hit distribution factorizes as:

$$p_{\text{detected}}(\theta_1, \theta_2, \phi_1, T) = q_{\text{RTE}}(\theta_1, \theta_2, \phi_1) \cdot \eta(\theta_1, \theta_2, \phi_1) \cdot p_{\text{cNSF}}(T \mid \theta_1, \theta_2, \phi_1)$$

- q_{RTE} : first-hit boundary pushforward measure (from LS-side QMC).
- η : learned cond-space reweighting / selection factor (i.e., the probability of detection).
- p_{cNSF} : frozen outer-layer timing kernel conditioned on detection.

Role of η

η is *not* an extra timing model — it is a learned reweighting in cond-space. Current evidence: cNSF kernel is strong but only reveals distribution shape; the gap is the RTE-to-cond pushforward.

η Network: Learning the Detection Probability

Formulation

η is a binary classifier trained on the detected vs. undetected first-hit photon samples, and thus (implicitly) depends on the distribution of first-hit photon roots given by inner layer (RTE or G4)

$$\eta(\theta_1, \theta_2, \phi_1) = P(\text{detected} \mid \theta_1, \theta_2, \phi_1)$$

Data imbalance

The samples exhibit a **significantly large** negative-to-positive ratio. According to our G4 dataset:

- Out of 202M photons on LS boundary, only 33M are detected by PMTs, 169M are undetected.
- #Positive samples: 33M, #negative samples: $N_{\text{PMT}} \times 2.02 \times 10^8 - 3.3 \times 10^7 \approx 3.55 \times 10^{12}$.
- Negative-to-positive ratio $\approx 1.1 \times 10^5$.

Training and adjustment

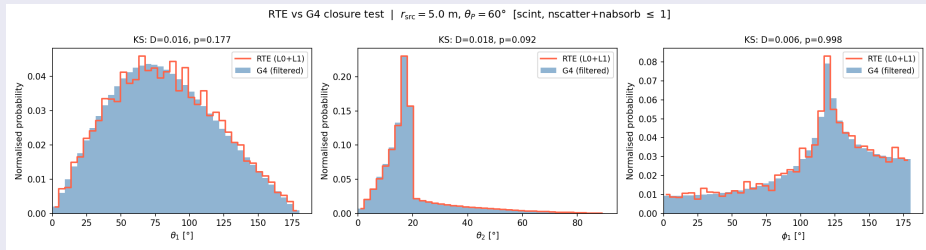
We adjust the actual data ratio to 1 : 1 by under-sampling the negative samples and train a Gradient Boosted Neural Network (GBNN). The final logits in η is adjusted accordingly.

Current Status: The RTE solution

The Baseline

GEANT4 simulation with selection of single-scatter photons that hit the boundary, record its $(\theta_1, \theta_2, \phi_1)$ distribution as the “pushforward” target. Direct comparison without fitting. Liquid scintillator optical parameters fixed to JUNO official MC values.

Preliminary results (RTE)



Good consistency between RTE surrogate and Geant4 for the first-hit boundary distribution.

Current Status: Outer-layer cNSF / η Results

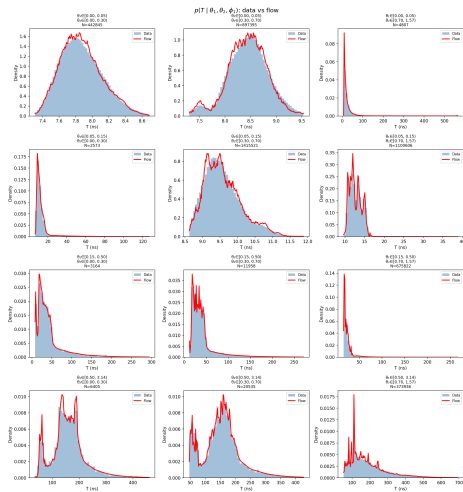


Figure: Conditional histograms of Geant4 detected data (blue shade) v.s. cNSF (red; outlines).

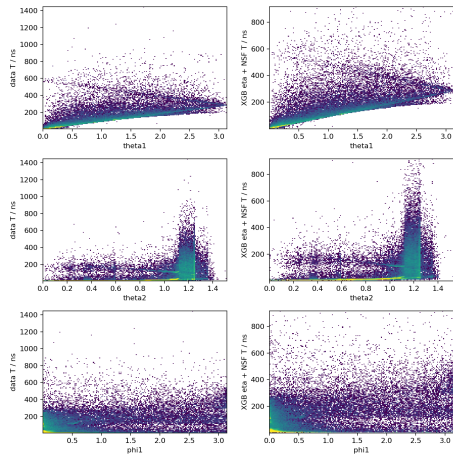


Figure: Marginal distribution of G4 detected data (left) v.s. first-hit samples (from G4) + GBNN η + cNSF (right). Shape is similar but scale still mismatches.

Summary

- Pure RTE surrogate (without NSF) already shows good agreement with Geant4 for the first-hit boundary distribution.
- Using effective boundary condition to model the acrylic sphere is a good first step, but it fails to capture the TIR-dominated region.

Next steps

- Further train and benchmark the NSF model / η network on Geant4 data.
- Integrate and test the full two-stage architecture with the frozen NSF and the RTE surrogate on both Geant4 and real calibration data.

Thanks!

Term $L^{(0)}$: Direct (Ballistic) Photon

Geometry

Vertex V , PMT P at R_{PMT} , opening angle θ :

$$d_{\text{VP}} = \sqrt{R_{\text{PMT}}^2 + r_v^2 - 2R_{\text{PMT}}r_v \cos \theta}$$

$$\cos \alpha = (R_{\text{PMT}} - r_v \cos \theta) / d_{\text{VP}}$$

d_{LS} = path inside LS (Snell)

Scintillation PDF

Bi-exponential $\otimes$ Gaussian (analytic erfc form):

$$f_{\text{scint}} = f_f G(\tau, \tau_f) + (1 - f_f) G(\tau, \tau_s)$$

$$G(\tau, \tau_d) = \frac{1}{2\tau_d} e^{\sigma^2/(2\tau_d^2) - \tau/\tau_d} \text{erfc}\left(\frac{\sigma/\tau_d - \tau/\sigma}{\sqrt{2}}\right)$$

Time-integrated yield

$$\lambda^{(0)} = Y_{\text{eff}} \cdot \frac{\cos \alpha}{d_{\text{VP}}^2} \cdot e^{-d_{\text{LS}}/L_{\text{abs}} - d_{\text{LS}}/L_{\text{scat}}} \cdot T_{\text{fr}}$$

9 free parameters

$Y_{\text{eff}}, L_{\text{abs}}, L_{\text{scat}}, n_{\text{LS}}, \tau_f, \tau_s, f_f, f_{\text{bkg}}, t_0$. ($T_{\text{fr}} = 0$ at TIR: $\sin \theta_i > n_w/n_{\text{LS}}$.)

Rate

$$R^{(0)} = \lambda^{(0)} \cdot \left[(1 - f_{\text{bkg}}) f_{\text{scint}}(t - t_{\text{flight}} - t_0) + \frac{f_{\text{bkg}}}{T_{\text{MAX}}} \right]$$

Term $L_{\text{scat}}^{(1)}$: Single Rayleigh Scatter

Physics

Photon travels $V \rightarrow S \rightarrow P$. At scatter point S :

- Rayleigh phase: $f(\vartheta) \propto 1 + \cos^2 \vartheta$
- Scatter prob.: $\mu_s e^{-\mu_t d_{VS}}$
- Attenuation: $e^{-\mu_t d_{SP}} \cdot T_{\text{fr}}$
- Extra delay: $\Delta t = (d_{VS} + d_{SP} - d_{VP}) n_{\text{LS}} / c$

Key insight (RTE paper [2401.15698])

For fixed (V, P, t) : all scatter points with $d_{VS} + d_{SP} = ct/n_{\text{LS}}$ lie on an **ellipsoid**. The integral reduces to a **1-D integral** over the angle φ in the V - P plane.

Precomputed kernel & gradient

Stored: geometry only, binned by $L = d_{VS} + d_{SP}$:

$$g(r_v, \theta_{vp}, L) = \int f(\vartheta) \frac{\cos \alpha_S}{d_{SP}^2} \delta(d_{VS} + d_{SP} - L) d\varphi$$

No μ_s , no attenuation baked in (≈ 0.4 MB).

At fit time (μ_s, ρ are live JAX params):

$$R_{\text{scat}}^{(1)} \propto \mu_s \sum_k g[\dots, k] e^{-\rho L_k} f_{\text{scint}}(t - t_k)$$

$\nabla_{\mu_s}, \nabla_{\rho}$ both flow exactly.

Term $L_{\text{reemit}}^{(1)}$: Absorption & Re-emission

Physics (PPO fluorescence)

Photon absorbed at S , re-emitted isotropically after lifetime τ_{fluor} :

- Absorption rate: $\mu_a = 1/L_{\text{abs}}$
- Quantum yield: Q_{yield}
- Direction: isotropic, $f = 1/(4\pi)$
- Delay: $V \rightarrow S$ travel *plus* τ_{fluor}

Time structure

Re-emission broadens the time distribution:

$$f_{\text{reemit}} = \frac{e^{-\tau/\tau_{\text{fluor}}}}{\tau_{\text{fluor}}} * f_{\text{scint}}(\tau)$$

Explains the heavy late photon tail in data.

Kernel (same $V \rightarrow S \rightarrow P$ geometry)

Replace $\mu_s f(\vartheta) \rightarrow \mu_a Q_{\text{yield}}/(4\pi)$:

$$h_{\text{re}} = \int \frac{\mu_a Q_{\text{yield}}}{4\pi} e^{-\mu_t(d_{VS}+d_{SP})} \frac{\cos \alpha_S}{d_{SP}^2} d\varphi$$

TIR Kernels: $L_{\text{TIR}}^{(1,\text{surf})}$, $L_{\text{scat}}^{(1,\text{TIR})}$, $L_{\text{reemit}}^{(1,\text{TIR})}$

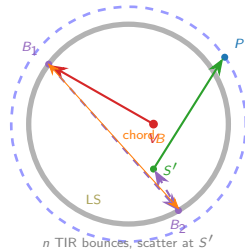
$L_{\text{TIR}}^{(1,\text{surf})}$: sub-TIR single bounce

Path $V \rightarrow B \rightarrow E \rightarrow P$ (bounce at B , exit at E). Bounce points found by Newton iteration on the specular constraint; up to 2 solutions per (r_v, θ_{vP}) . Weight: $R_{\text{fr}}(B) \cdot T_{\text{fr}}(E) \cdot \cos \alpha / d_{EP}^2$.

$L_{\text{scat/reemit}}^{(1,\text{TIR})}$: multi-bounce + scatter

Path $V \rightarrow B_1 \xrightarrow{n} \dots \rightarrow S' \rightarrow P$.

- $n = 0, 1, \dots, n_{\text{max}} = 20$ explicit TIR bounces before scatter/reemit at interior point S' .
- LS path: $d_{VB} + n \cdot \text{chord}_B + s + d_{S'P}$
- **Geometric tail** ($n > 20$): $q = e^{-\mu_t \cdot \text{chord}_B}$, tail weight = $q^{20} / (1 - q)$
- Scatter: prefactor $Y_{\text{eff}} / L_{\text{scat}}$; re-emit: $Y_{\text{eff}} Q_{\text{yield}} / L_{\text{abs}}$



The JUNO Detector

Geometry (radii in mm)

Layer	Boundary
Liquid Scintillator (LS)	$r < 17\,700$
Acrylic sphere	$17\,700 \rightarrow 17\,820$
Water buffer	$17\,820 \rightarrow 19\,500$
PMT photocathode	$R_{\text{PMT}} = 19\,500$

Refractive indices

$n_{\text{LS}} = 1.497$, $n_{\text{acrylic}} = 1.490$, $n_{\text{water}} = 1.340$

