



# UCIRVINE

FERMILAB-SLIDES-26-0101-V

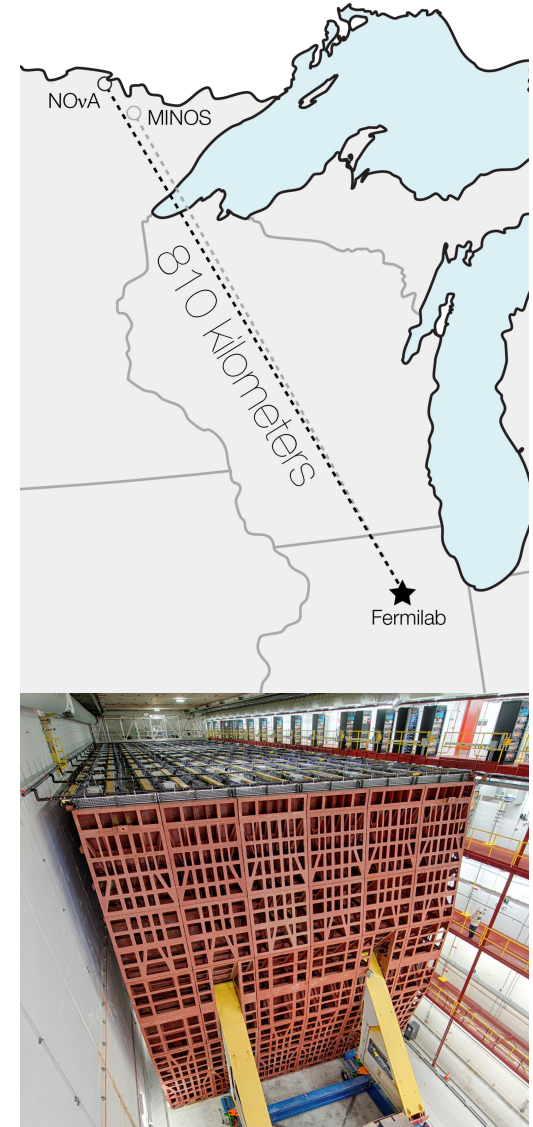
## Regression Convolutional Neural Network for Energy Estimation in NOvA

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University of California, Irvine  
19 June 2026



# NOvA Experiment Overview

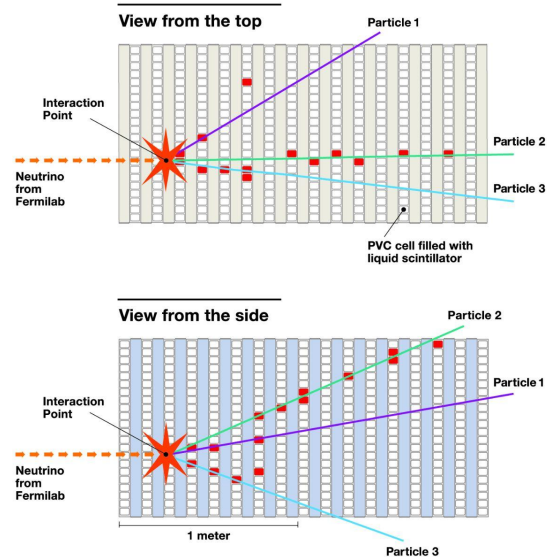
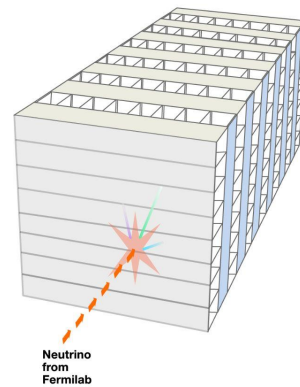
- Long-baseline neutrino oscillation experiment
- NuMI beam at Fermilab produces muon (anti-)neutrinos
- Functionally equivalent near and far detectors
  - 810 km apart
- Observe  $\nu_\mu$  disappearance and  $\nu_e$  appearance
- Measure  $\Delta m_{32}^2, \theta_{23}, \delta_{CP}$



# NOvA Detector

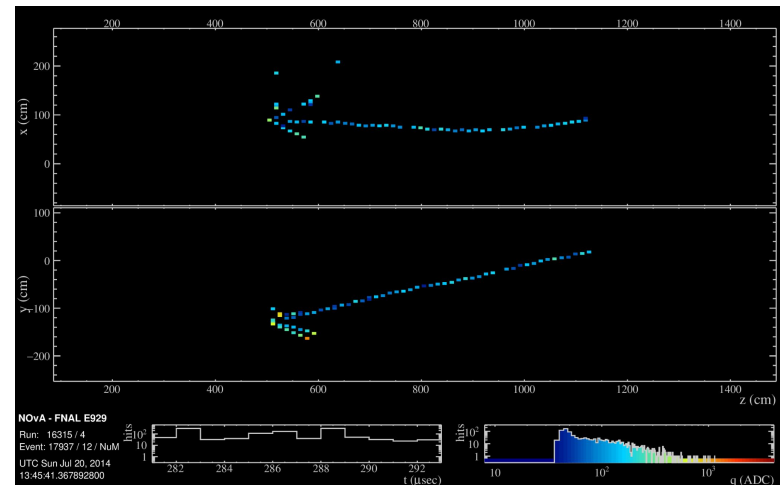
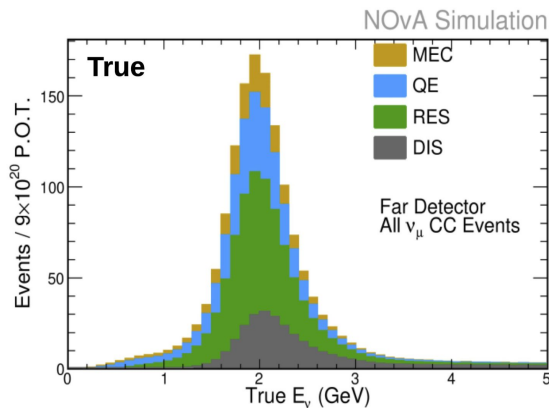
- Pixels: long 4 cm x 6 cm PVC cells
  - Filled with liquid scintillator
  - Wavelength shifting optical fiber
  - Avalanche photodiode for detection
- Alternate horizontal and vertical planes → 2 detector views

3D schematic of  
NOvA particle detector



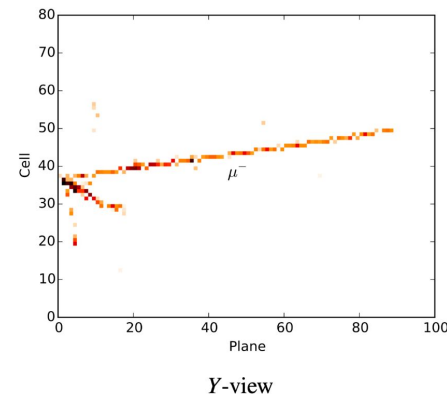
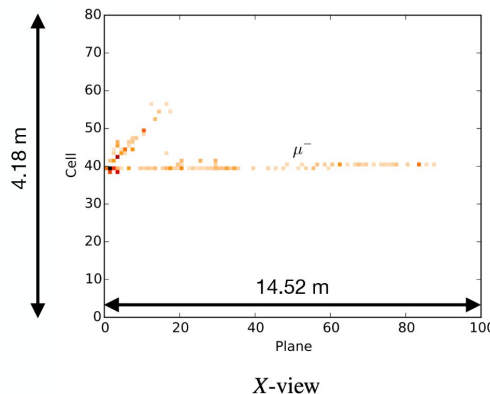
# Energy Estimation

- High energy estimation fidelity required to measure oscillation parameters
  - Oscillation probability is a function of neutrino energy
$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E}\right) \quad (\text{Simplified 2-Flavour Case})$$
  - Improve energy resolution  $\rightarrow$  improve oscillation parameter sensitivity
- Model must perform for many interaction modes and event topologies in NOvA



# Regression CNN (RegCNN)

- Model Inputs:
  - Muon track length
  - “Pixelmap” - detector event readout in 2 views
    - 2 Standard 100 x 80 images
    - Cropped around neutrino interaction
    - Train with random calibration scaling of pixelmap values
- Simultaneous prediction of total muon neutrino energy and muon energy



# Model Training

- Reweight true neutrino energy spectrum in training data
  - Flatten peak region, increase representation of tail regions
- Train separate models for each neutrino flavor, detector, beam configuration
- For  $\nu_\mu$  and  $\nu_e$ :
  - Far Detector (FD)
  - Near Detector (ND)
  - $\nu_\mu$  Beam Configuration
  - $\bar{\nu}_\mu$  Beam Configuration
- Present validation studies on  $\nu_\mu$  ND sample in this talk

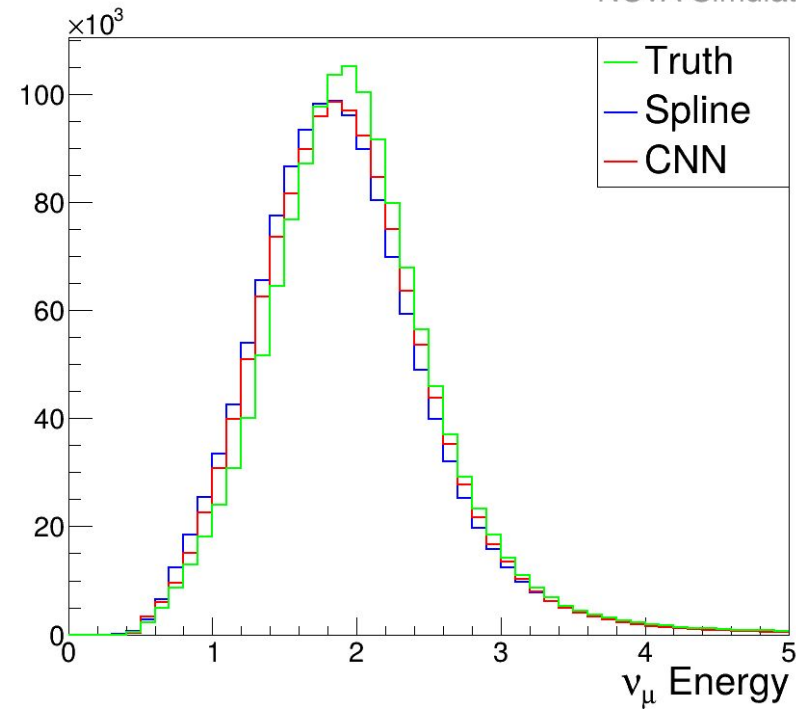
# Model Validation

- Compare RegCNN performance to Spline energy estimator
  - Spline fit for NOvA's 10-year dataset analysis
- Primary benchmarks:
  - Energy spectrum
  - Energy spectrum “quartiles”
    - Calculate ratio of hadronic energy to total energy
    - Split dataset into 4 subsets based on this ratio
  - Resolution: Standard Deviation $[(\text{Reconstruct}-\text{True})/\text{True}]$
  - Bias: Mean $[(\text{Reconstruct}-\text{True})/\text{True}]$

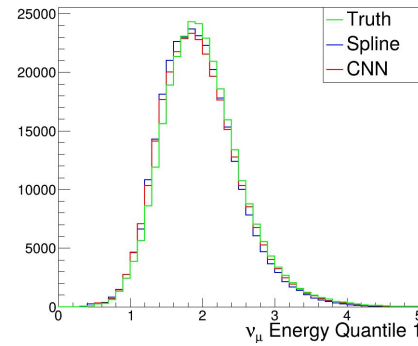
3rd Energy Estimation Method: Transformer  $\nu_{\mu}$  Energy Estimator developed by University of Minnesota.  
See Neutrino 2026 Poster by Leon Tong for Transformer performance

# Neutrino Energy

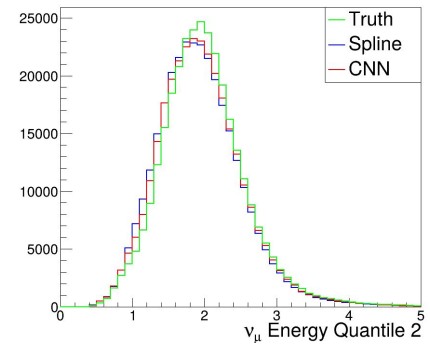
NOvA Simulation



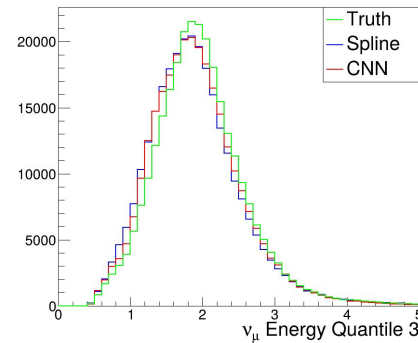
NOvA Simulation



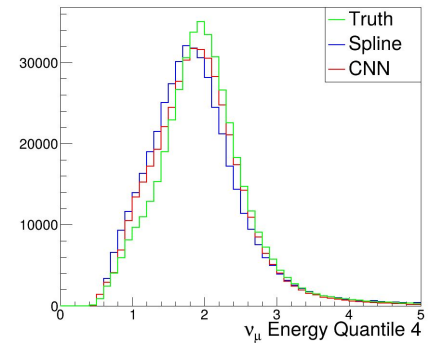
NOvA Simulation



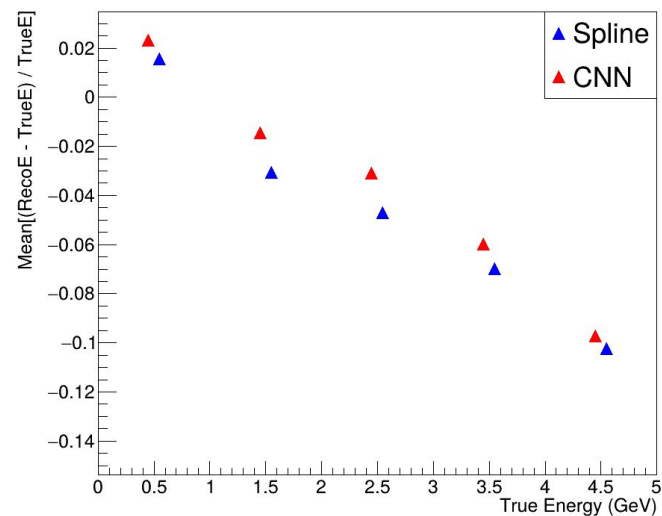
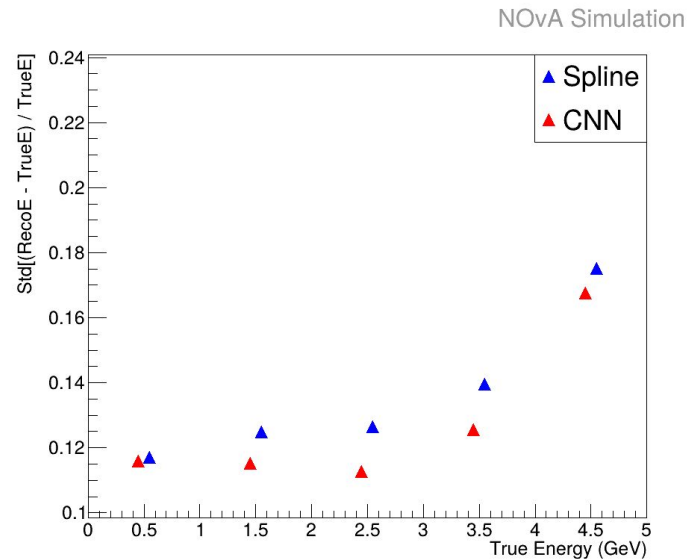
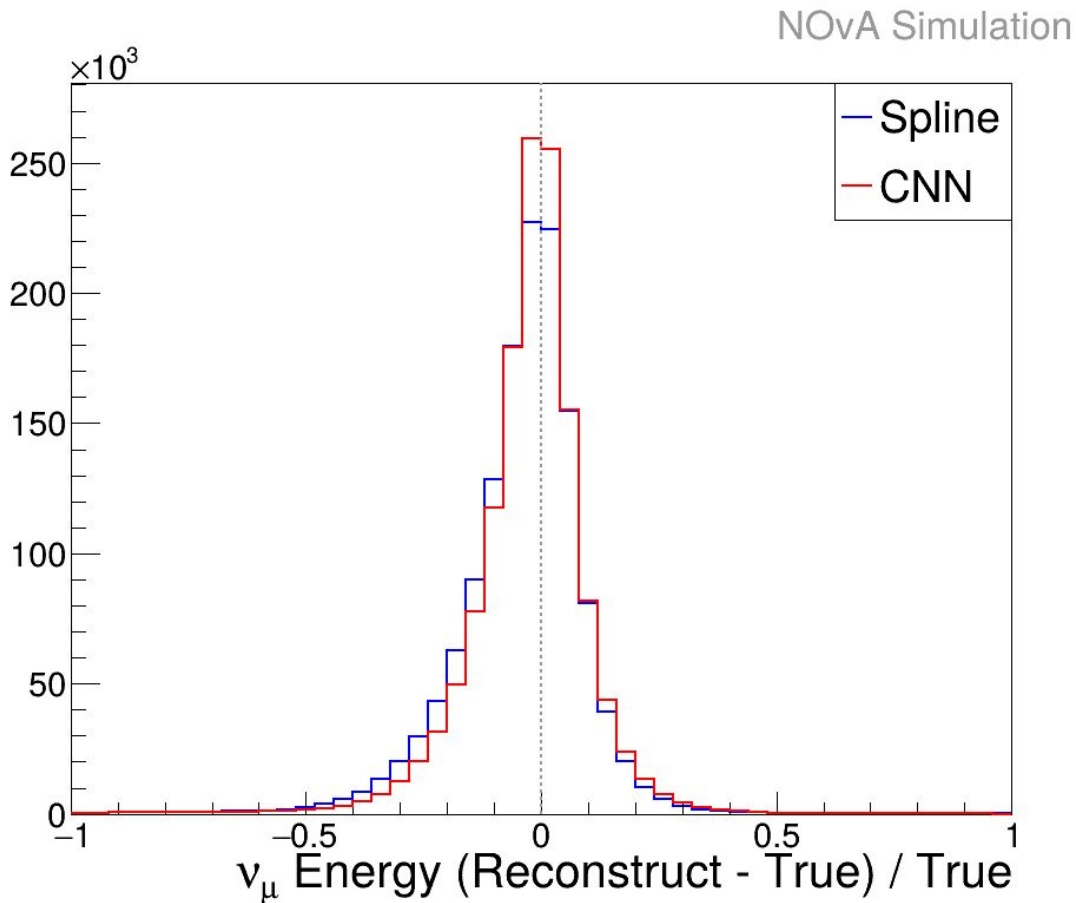
NOvA Simulation



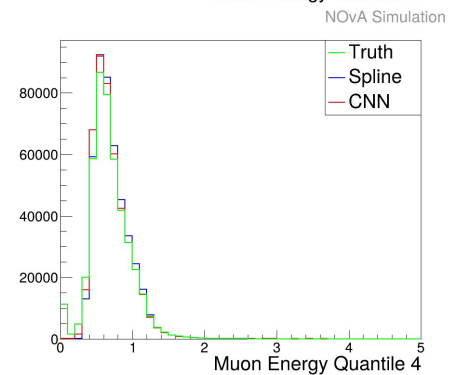
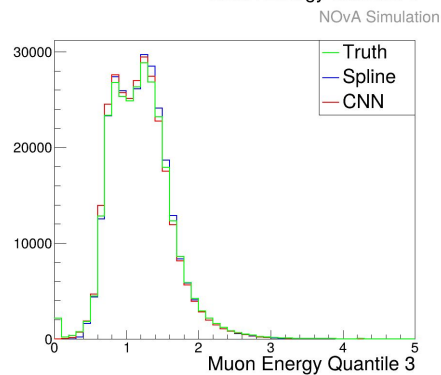
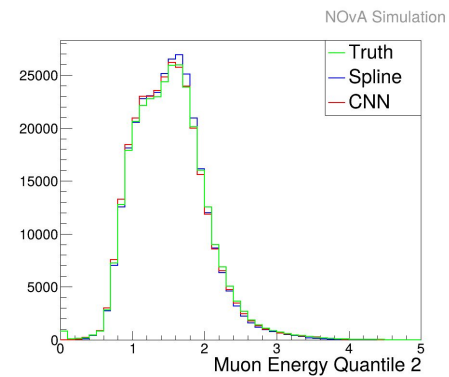
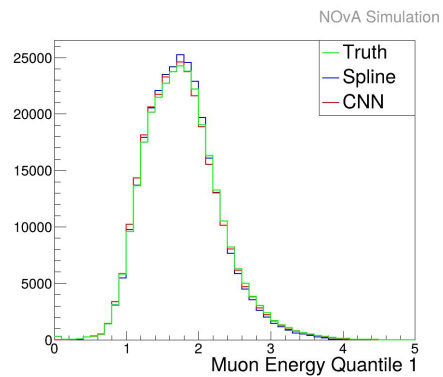
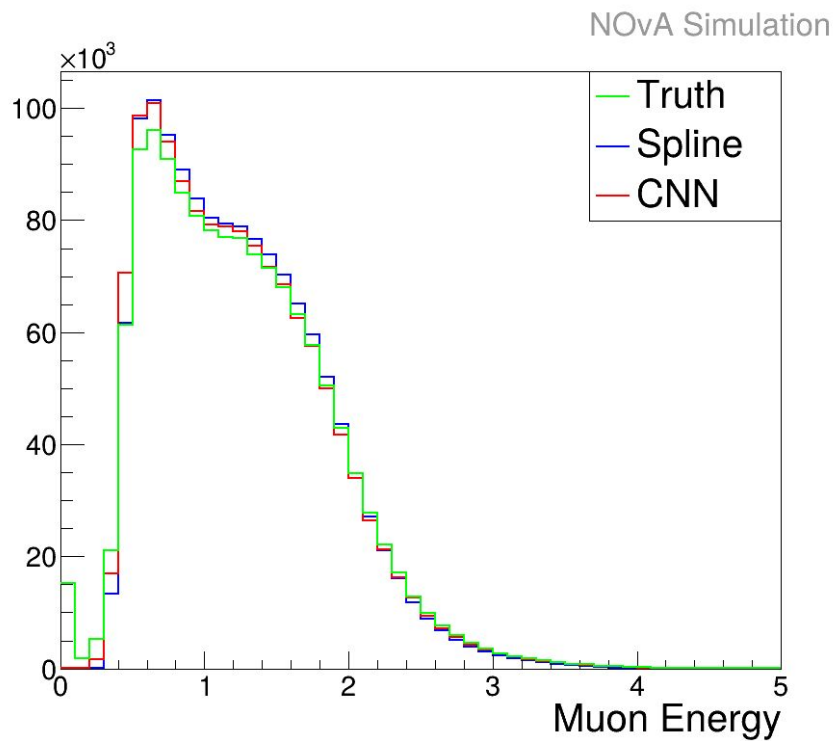
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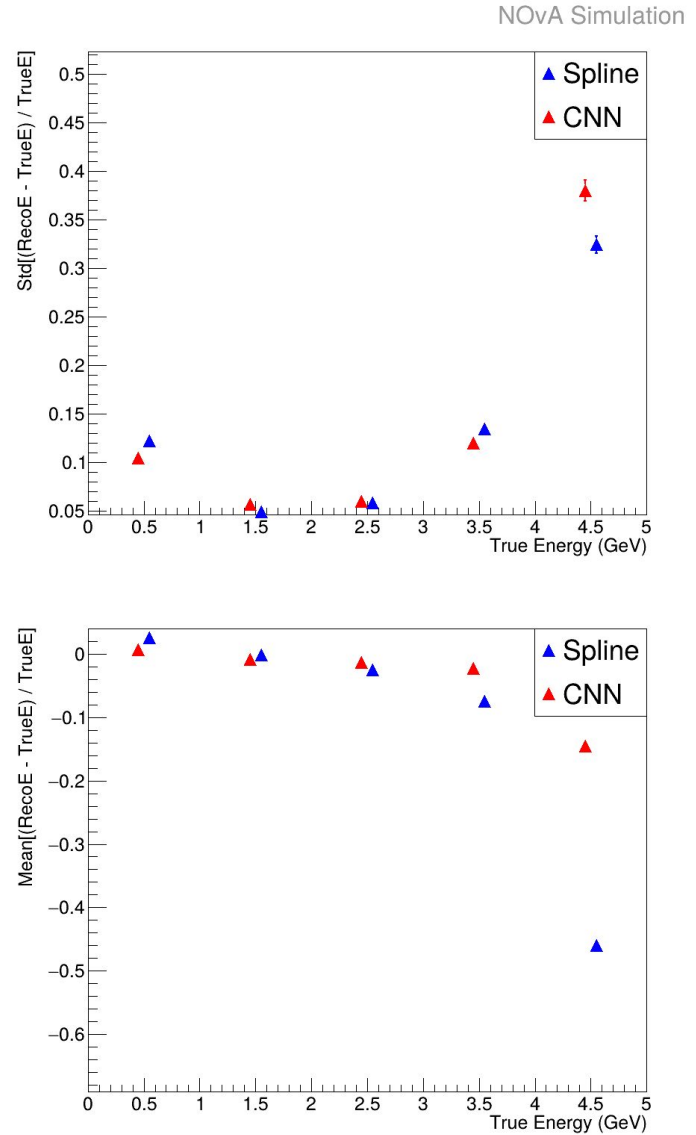
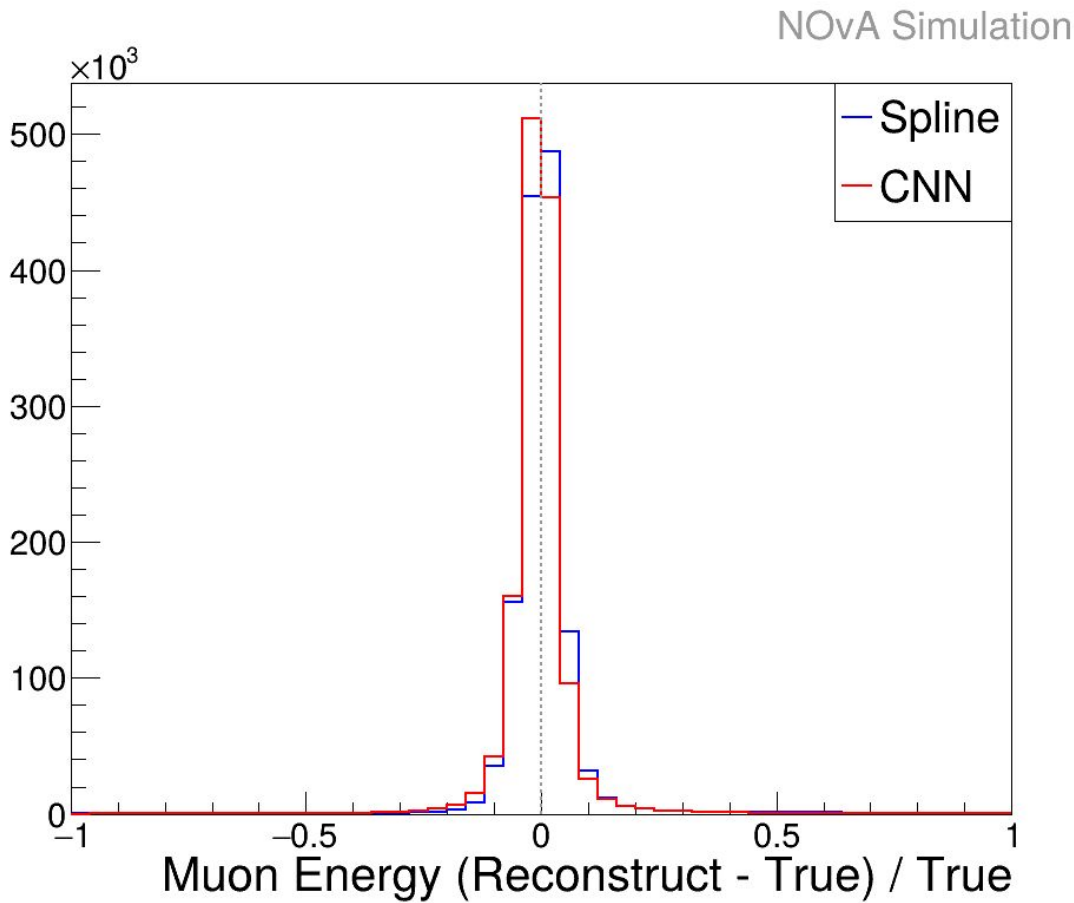
# Neutrino Energy



# Muon Energy

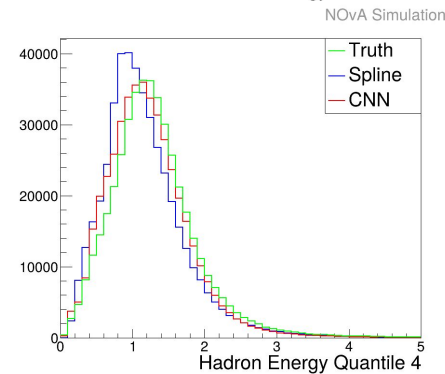
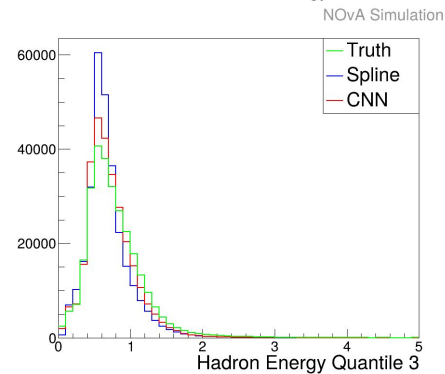
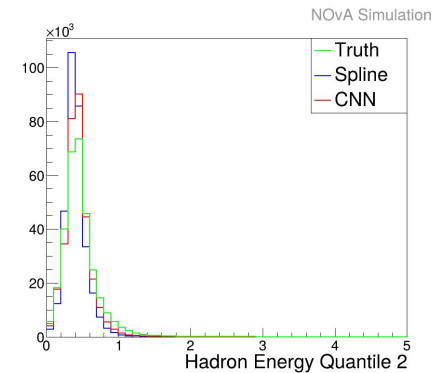
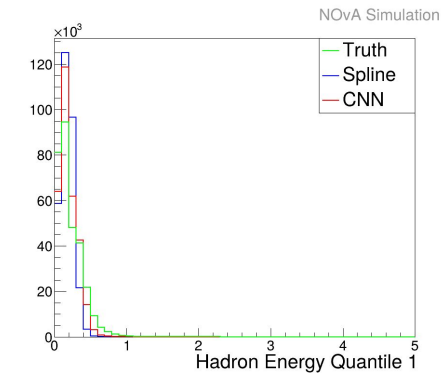
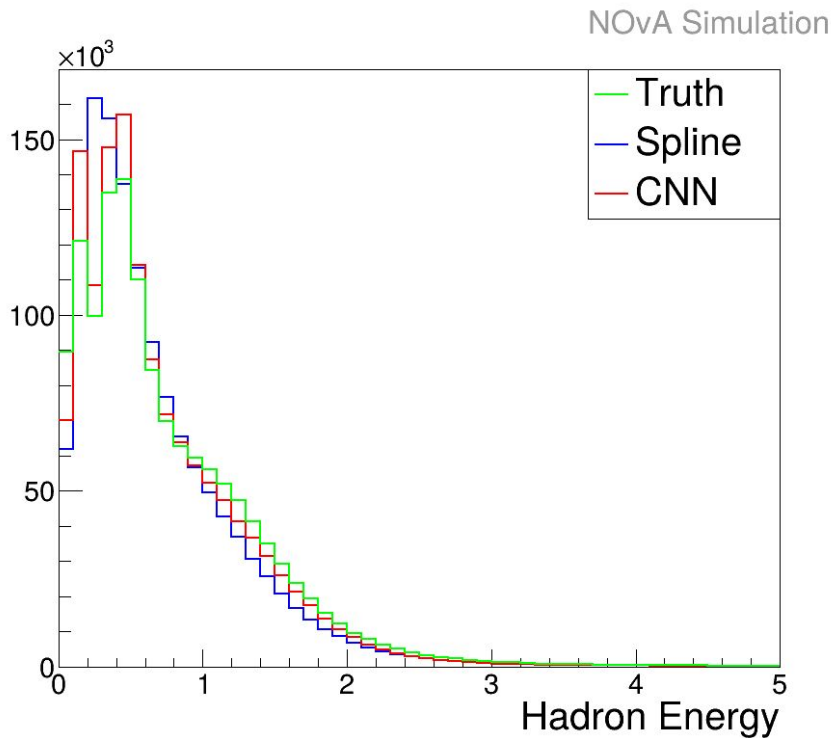


# Muon Energy



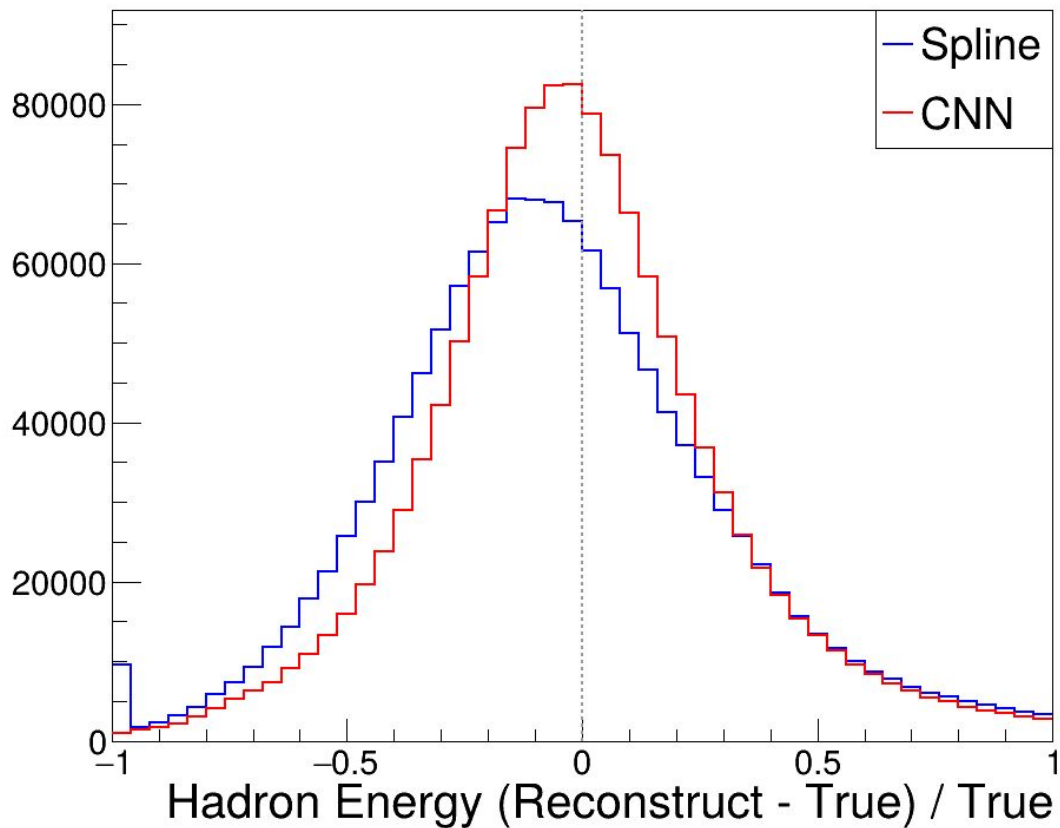
# Hadronic Energy

- Double peak structure in simulated hadron energy spectrum
  - Spline fails to reconstruct double peak without piecewise fit segments
  - RegCNN reconstructs double peak

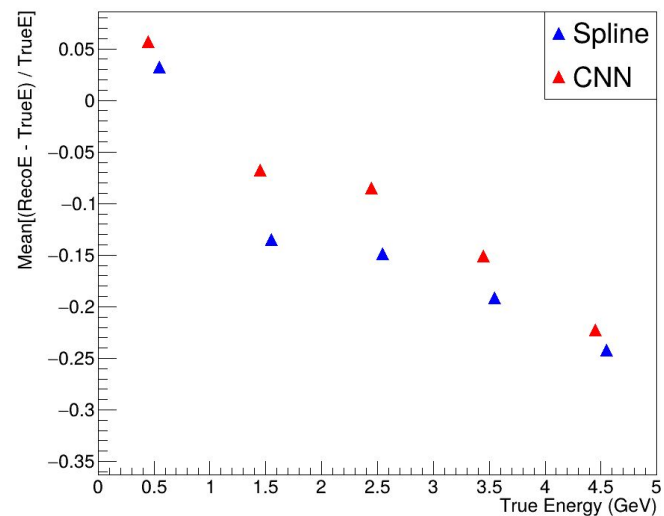
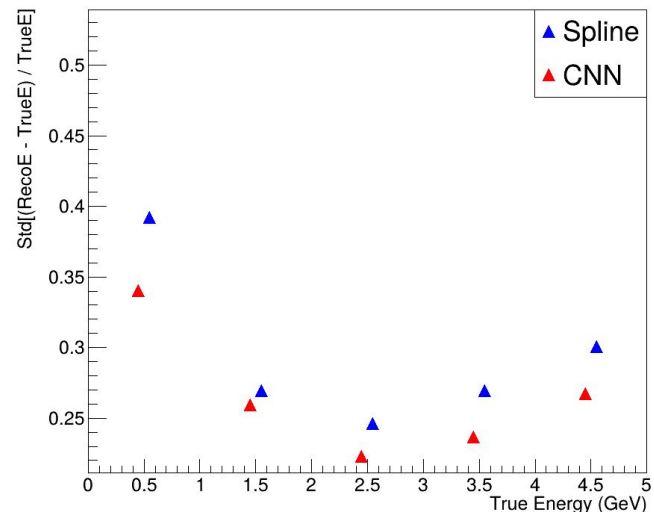


# Hadronic Energy

NOvA Simulation

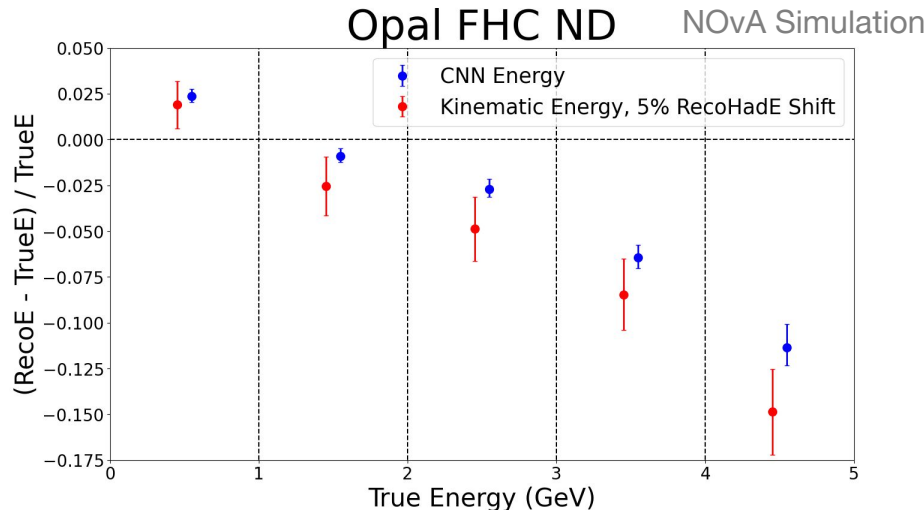


NOvA Simulation



# Calibration Robustness

- Train with calibration shift on input pixelmap
  - Scale by factor from Gaussian distribution (mean=1, std=0.1)
- Evaluate energy estimators on 5% calibration shift samples
  - Spline (Kinematic) method has 5% shift in energy
  - RegCNN demonstrates significant robustness to calibration shifts (~0.5% shift in energy)



(Bars in plot show effect of up/down shift)

# Calibration Systematic in NOvA

- Calibration is largest source of systematic error for NOvA's  $\Delta m_{32}^2$ ,  $\sin^2 \theta_{23}$  measurements
  - NOvA set current world-leading  $\Delta m_{32}^2$  limit

“Precision measurement of neutrino oscillation parameters with 10 years of data from the NOvA experiment”  
(DOI:10.1103/x53y-2b86)

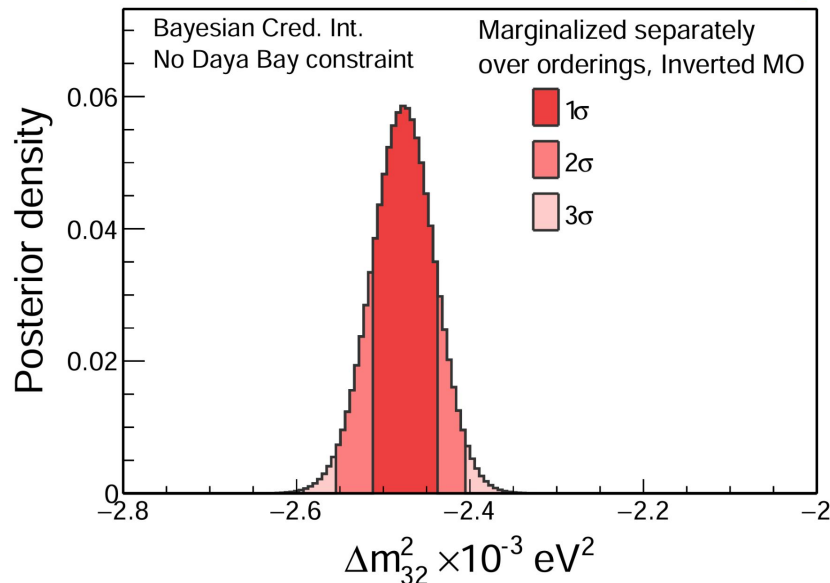
Source	$ \Delta m_{32}^2 $ ( $\times 10^{-5}$ eV <sup>2</sup> )	$\sin^2 \theta_{23}$ ( $\times 10^{-3}$ )	$\delta_{CP}$ ( $\times 10^{-2}$ $\pi$ )
Beam flux	+0.06/-0.12	+0.40/-0.75	+0.16/-1.00
<b>Calibration</b>	<b>+1.40/-1.80</b>	<b>+3.20/-16.0</b>	+1.70/-17.0
Detector model	+0.17/-0.27	+0.36/-3.90	+0.37/-3.30
Lepton Reco.	+0.90/-1.40	+2.70/-4.50	+0.61/-2.90
ND – FD Uncor.	+0.25/-0.34	+2.40/-2.50	+0.79/-4.00
Cross Sections	+0.65/-1.10	+3.10/-5.10	+2.00/-9.70
Neutron model	+0.25/-0.39	+2.70/-0.49	+0.43/-1.30
<b>Systematic Unc.</b>	<b>+1.70/-2.30</b>	<b>+6.60/-18.0</b>	+2.90/-21.0
Statistical Unc.	+3.20/-4.40	+22.0/-85.0	+6.60/-74.0

# Current Oscillation Analyses

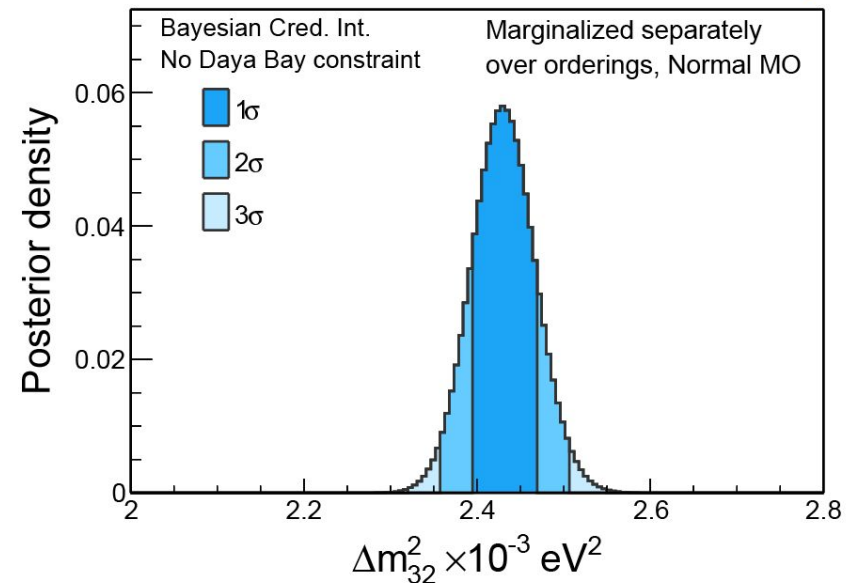
(DOI:10.1103/x53y-2b86): “Precision measurement of neutrino oscillation parameters with 10 years of data from the NOvA experiment”

- Neutrino 2026:
  - Zoya Vallari: “Probing the 3-Flavor paradigm with NOvA”
  - Artur Sztuc (Poster): “Prospects for Mass Ordering Measurements with NOvA”
  - Larry Zhao (Poster): “Bayesian Fit to NOvA Data Subsamples for Three Flavor Oscillation Analysis”

NOvA Preliminary



NOvA Preliminary



# Summary

- Developed simultaneous prediction of neutrino energy and lepton energy in RegCNN
- RegCNN improved energy reconstruction performance compared to Spline used in previous NOvA analyses
- Input pixelmap shifted for robustness to calibration uncertainties
  - RegCVN more robust to calibration shifts than traditional non-ML method
  - Significant reduction of calibration uncertainty in physics analyses

Thank you!