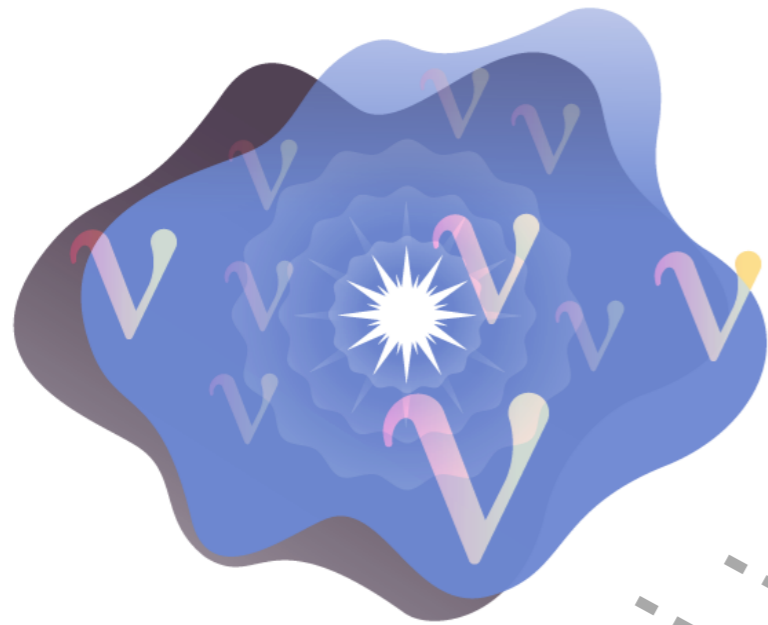


SNeND Measurements: “Supernova Neutrino constrained by Neutrino Data” for DUNE

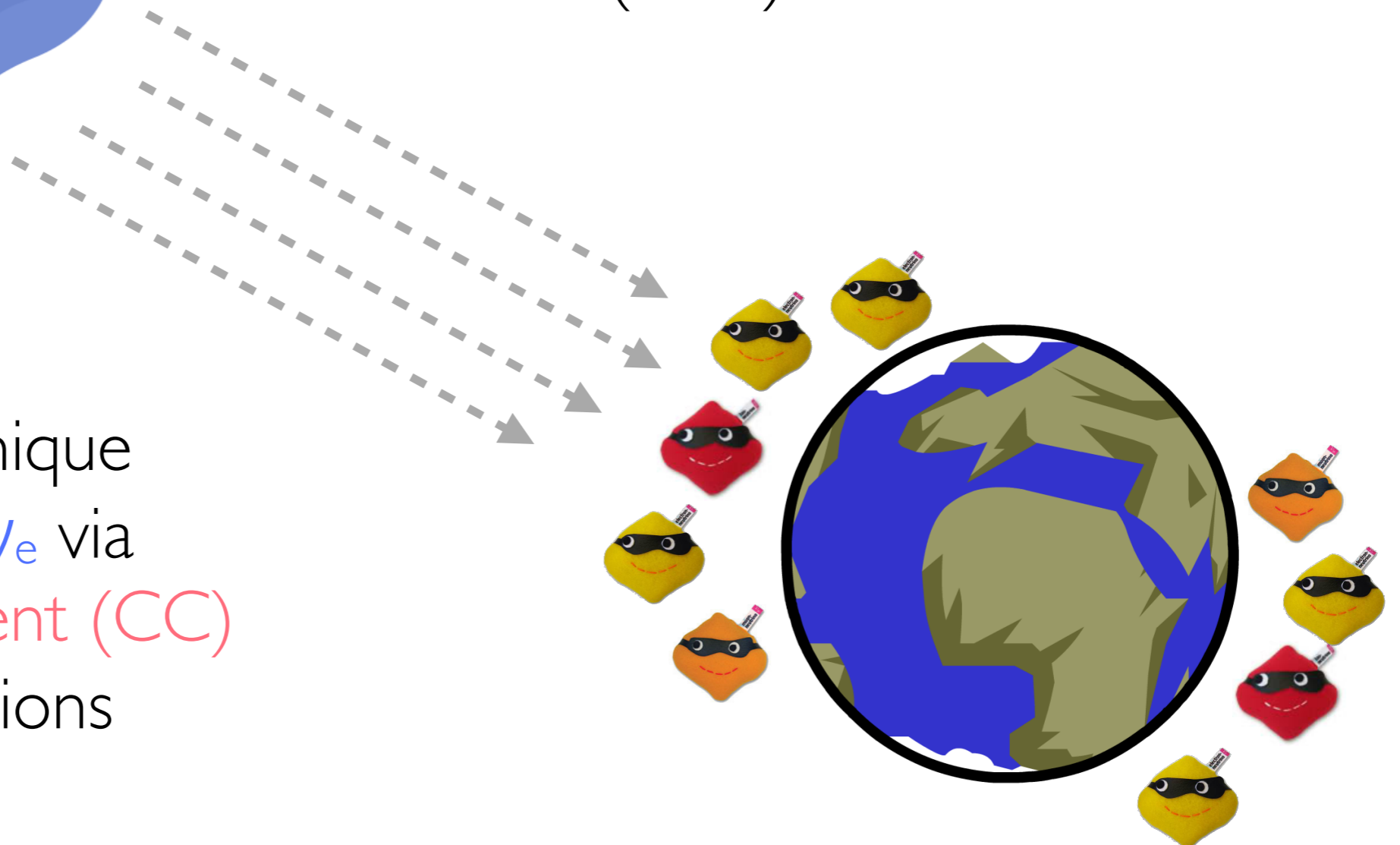
Yun-Tse Tsai
PoU IF Brainstorming@SLAC
September 26th 2025

Supernova Neutrino



Core-collapse supernova
neutrino energy:
 $O(1-10)$ MeV

DUNE has unique
sensitivity to ν_e via
charged-current (CC)
 ν_e -Ar interactions



Supernova Neutrino Flux

Pinched-thermal form: to fit simulated flux

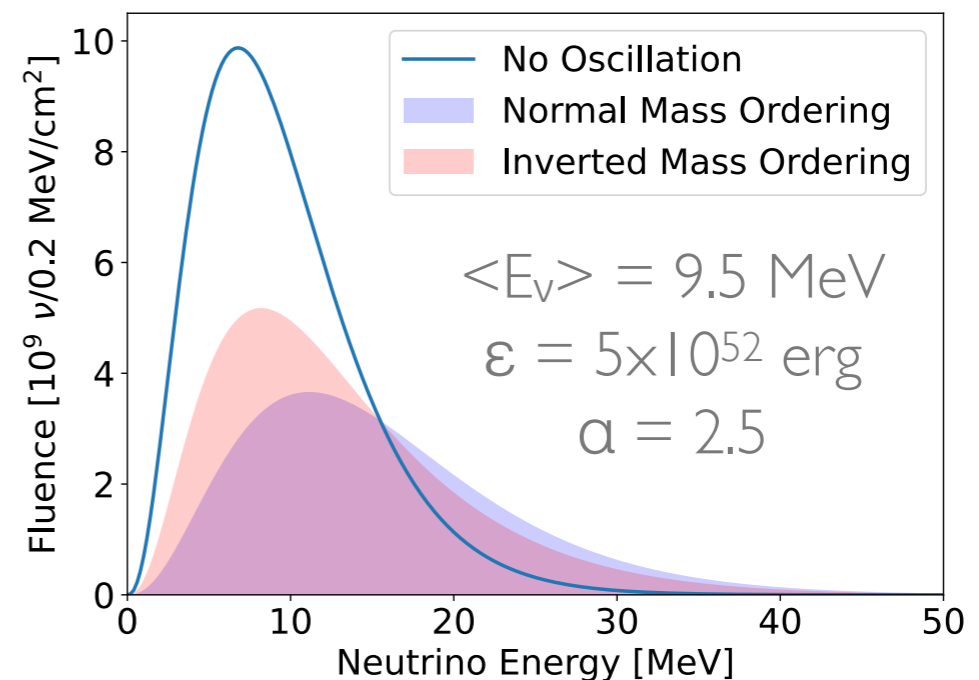
$$\phi(E_\nu) = \mathcal{N} \left(\frac{E_\nu}{\langle E_\nu \rangle} \right)^\alpha \exp \left[-(\alpha + 1) \frac{E_\nu}{\langle E_\nu \rangle} \right]$$

E_ν : neutrino energy

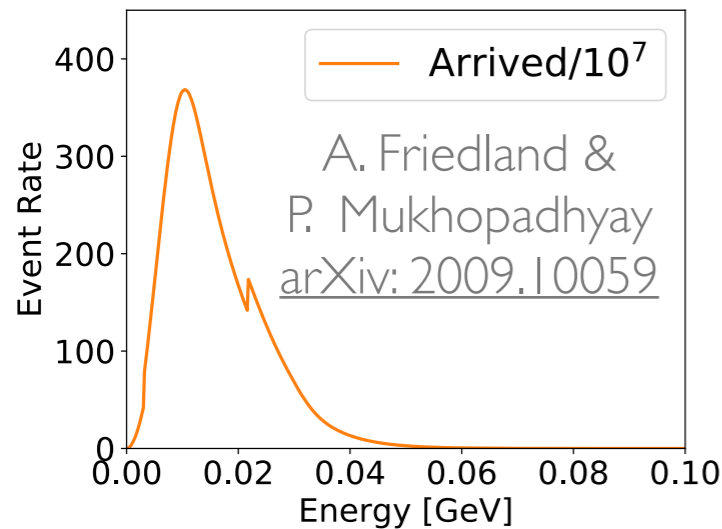
$\langle E_\nu \rangle$: average E_ν

$N \propto \nu$ luminosity, ε

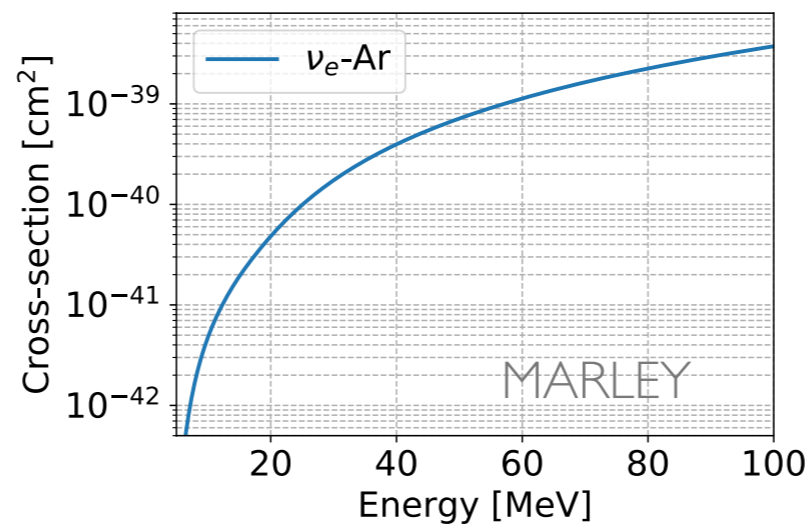
α : pinching parameter



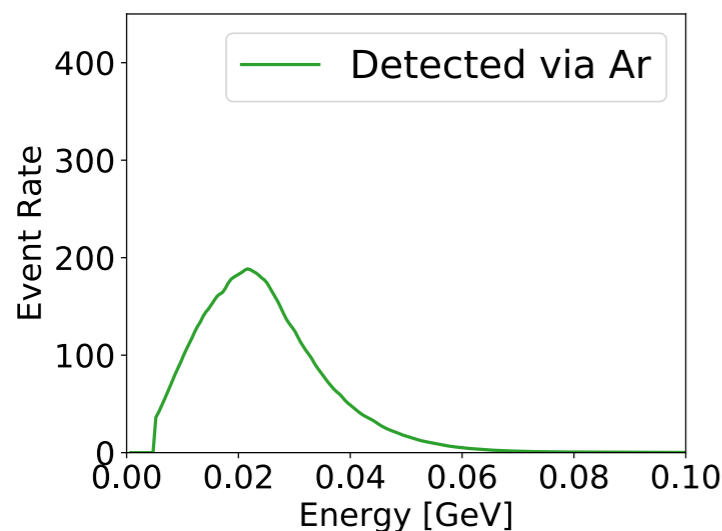
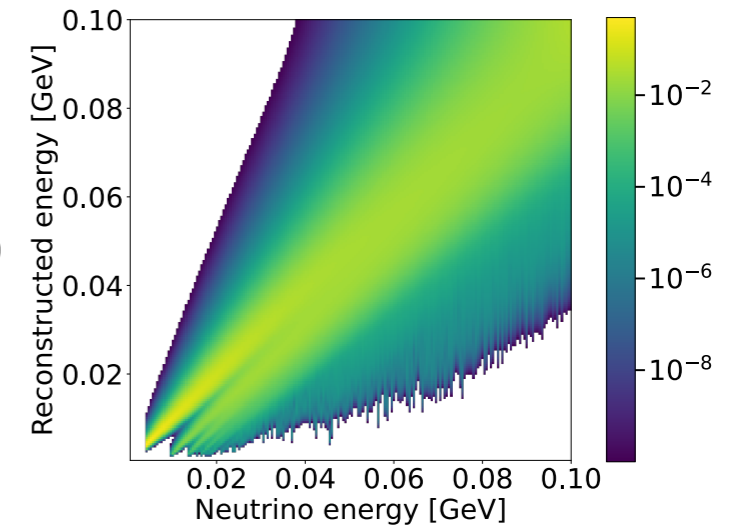
SN Neutrino Detection



⊗



⊗

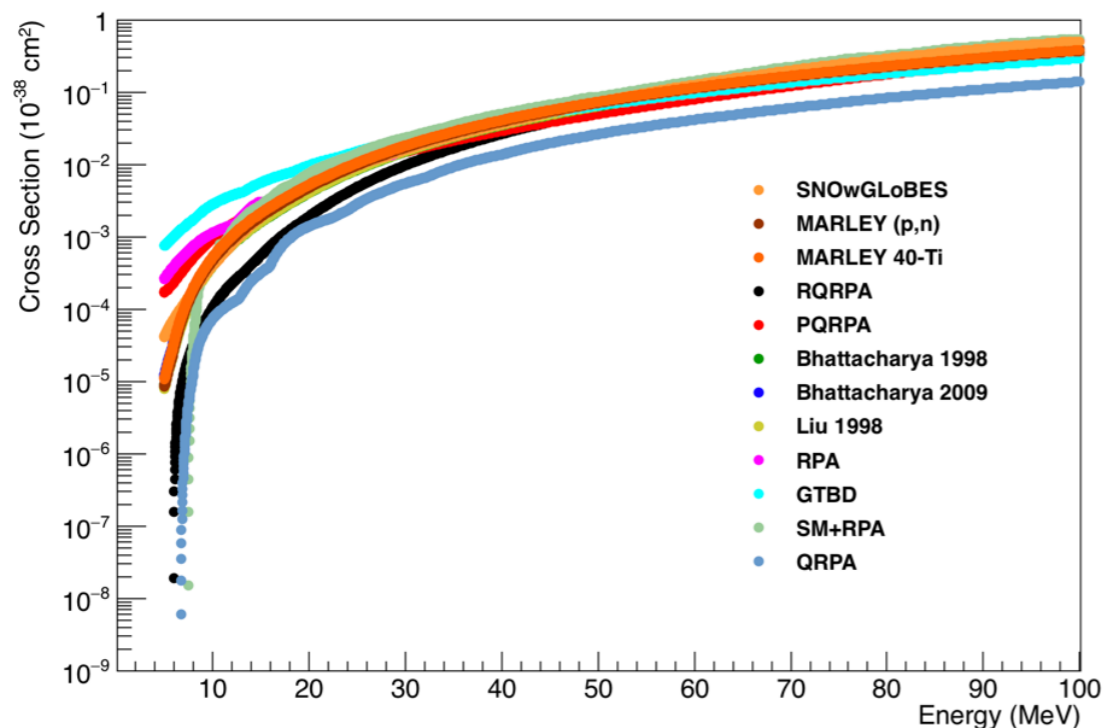


- Detect convolved ν flux and interaction cross sections
- Disentangled ν fluxes are desired
- These ν_e -Ar CC cross sections have never been measured
- Uncertainties from cross section models are relevant

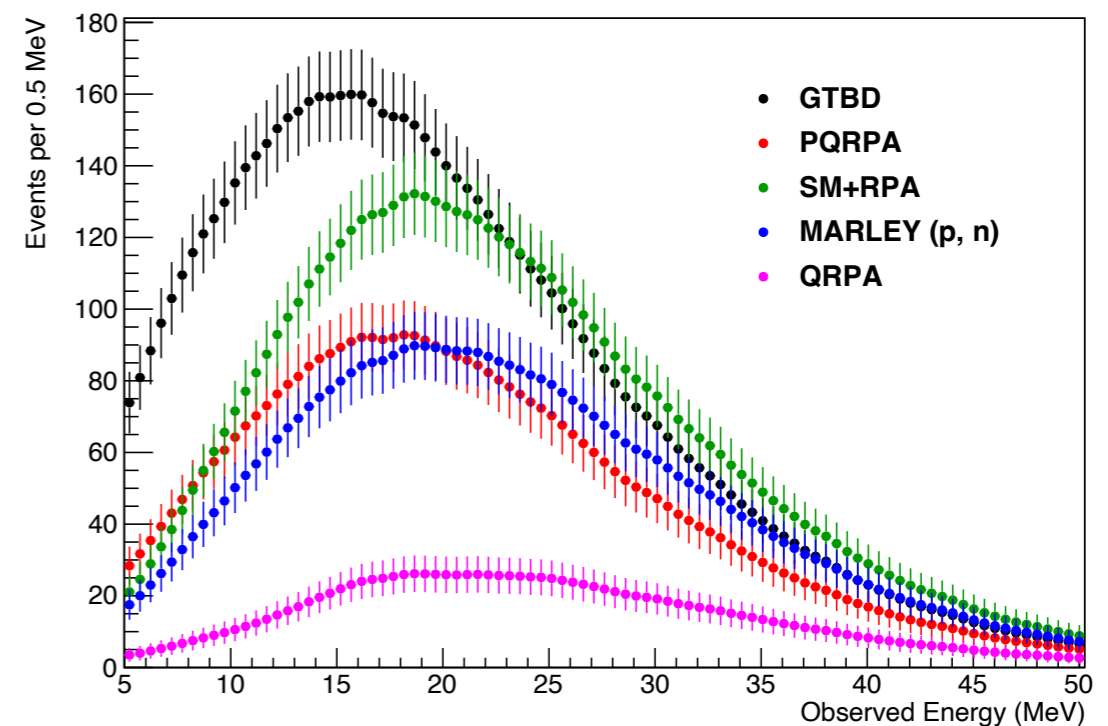
Impact from ν_e -Ar σ

- ν_e -Ar charged-current (CC) interaction cross section with $E_\nu < 100$ MeV has **never been measured**
- Theoretical models vary $> O(10\%)$
- Highly significant impacts on DUNE SN ν measurements, particularly on ϵ , biased from -94% to $+1400\%$ in extreme scenarios (PRD 107, 112012 (2023))

Cross Section Models

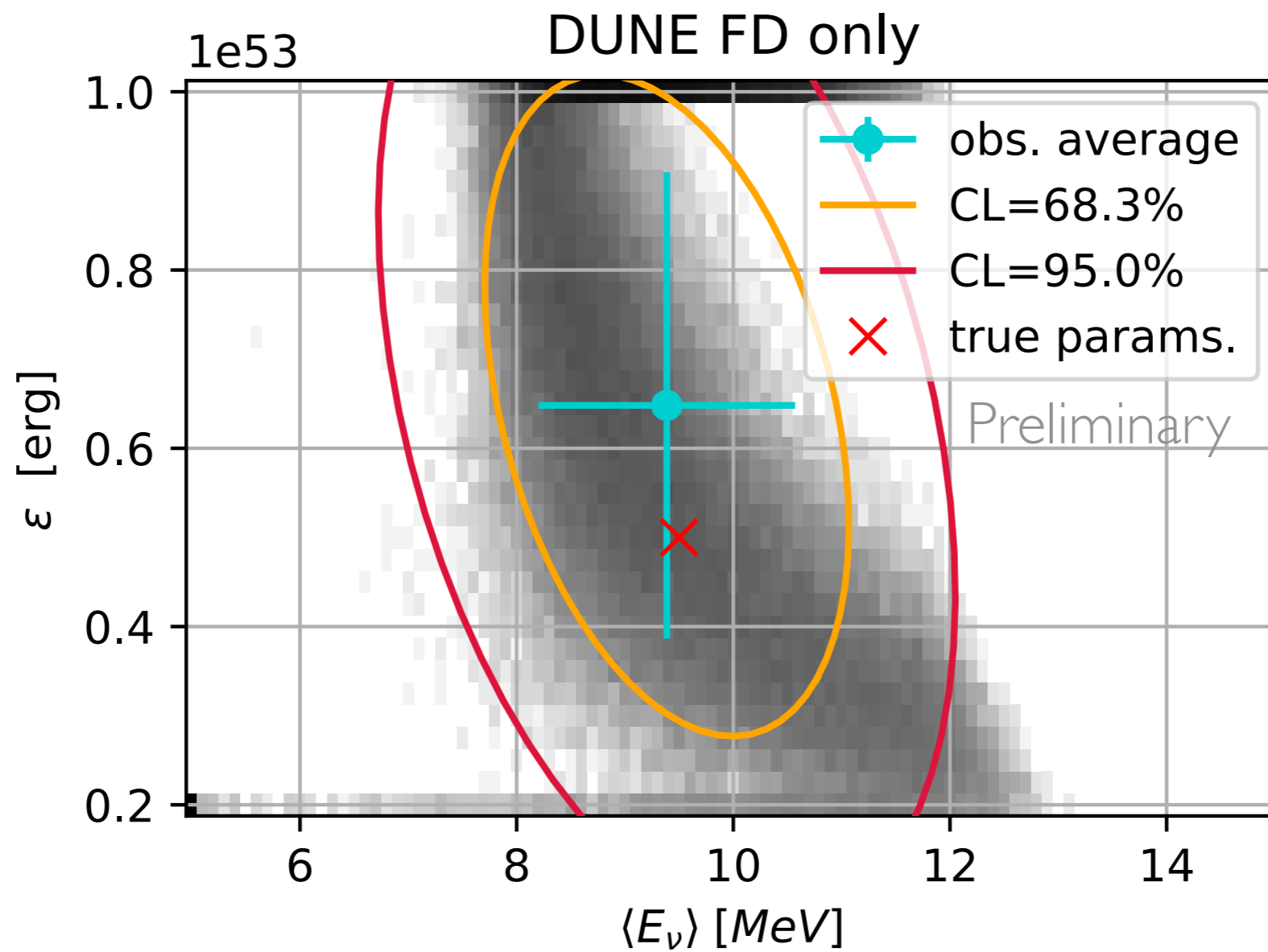


Test spectra for different ν_e - ^{40}Ar CC cross section models



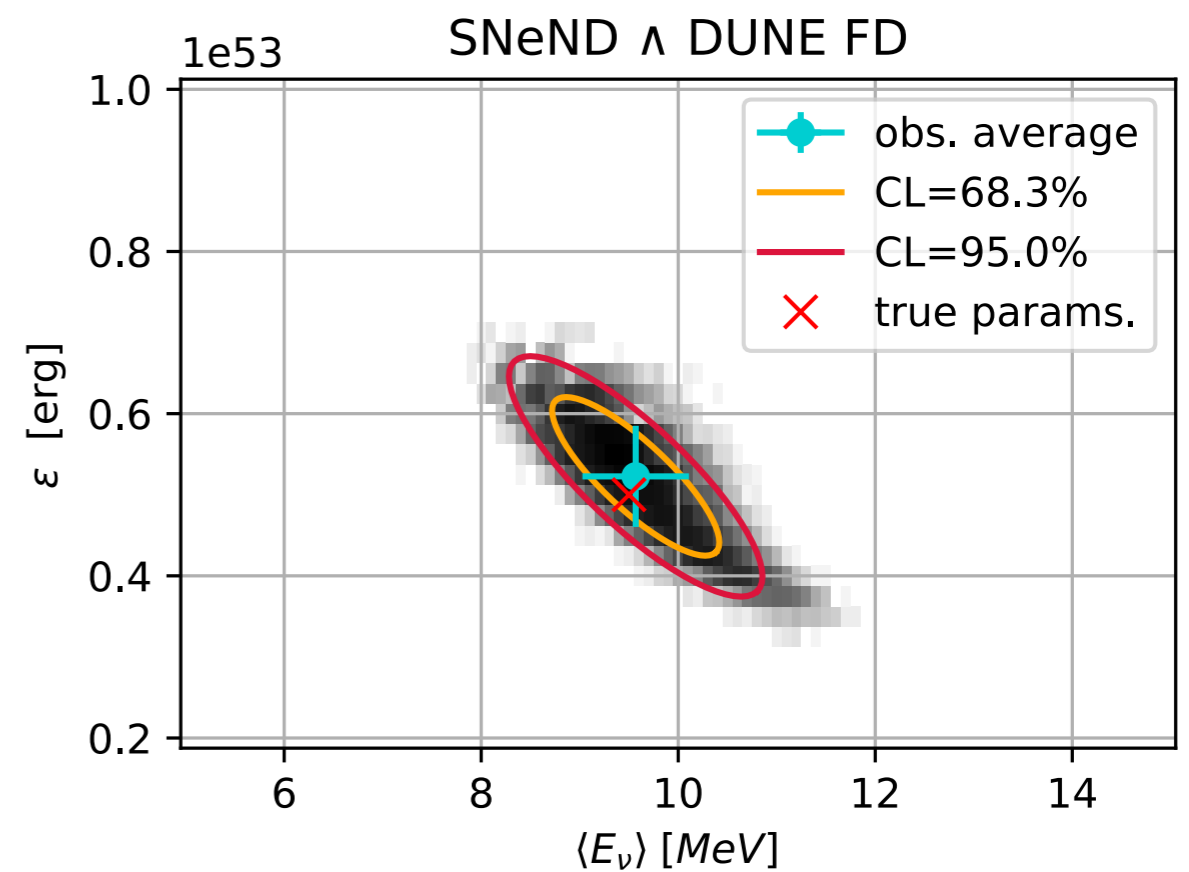
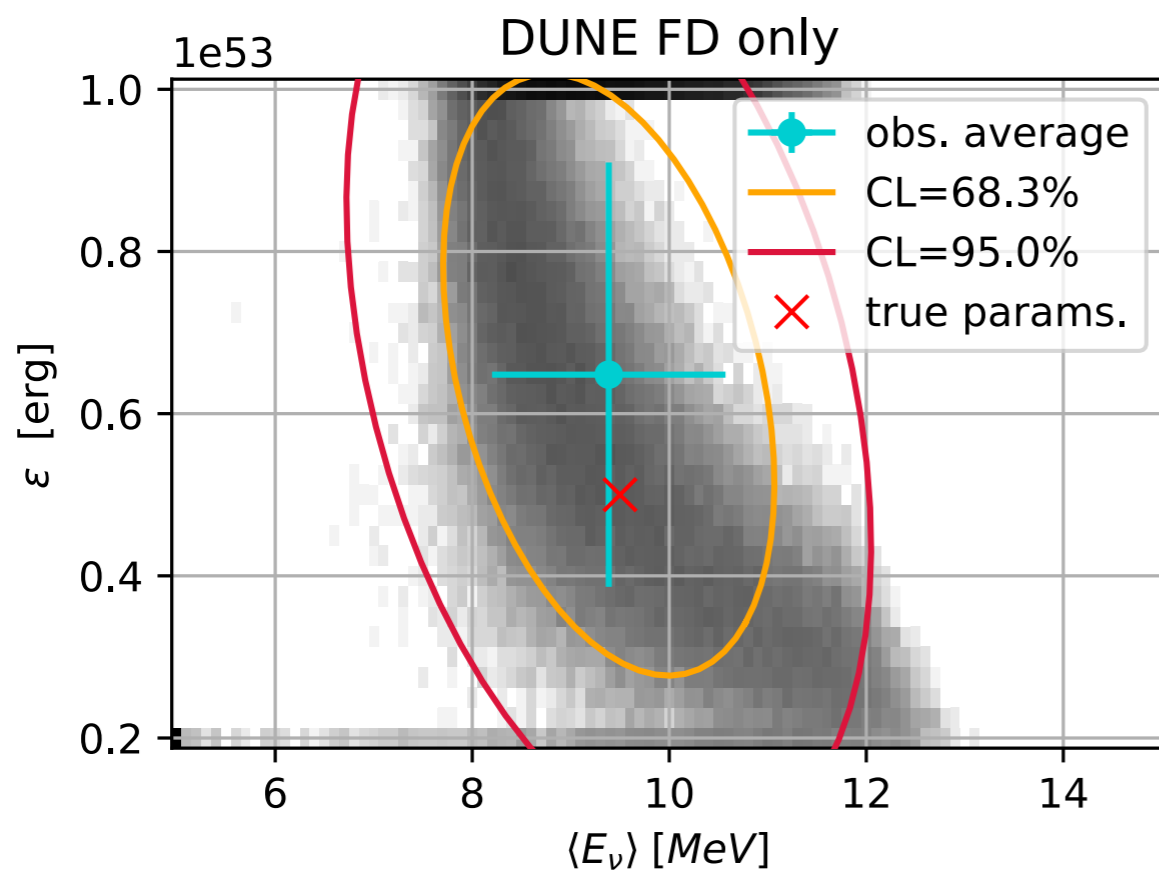
Measure ν_e -Ar CC σ

If we know the total σ and $\sigma(45 \text{ MeV})$ at the precision of 20%, as suggested by [PRD 107, 112012 \(2023\)](#)

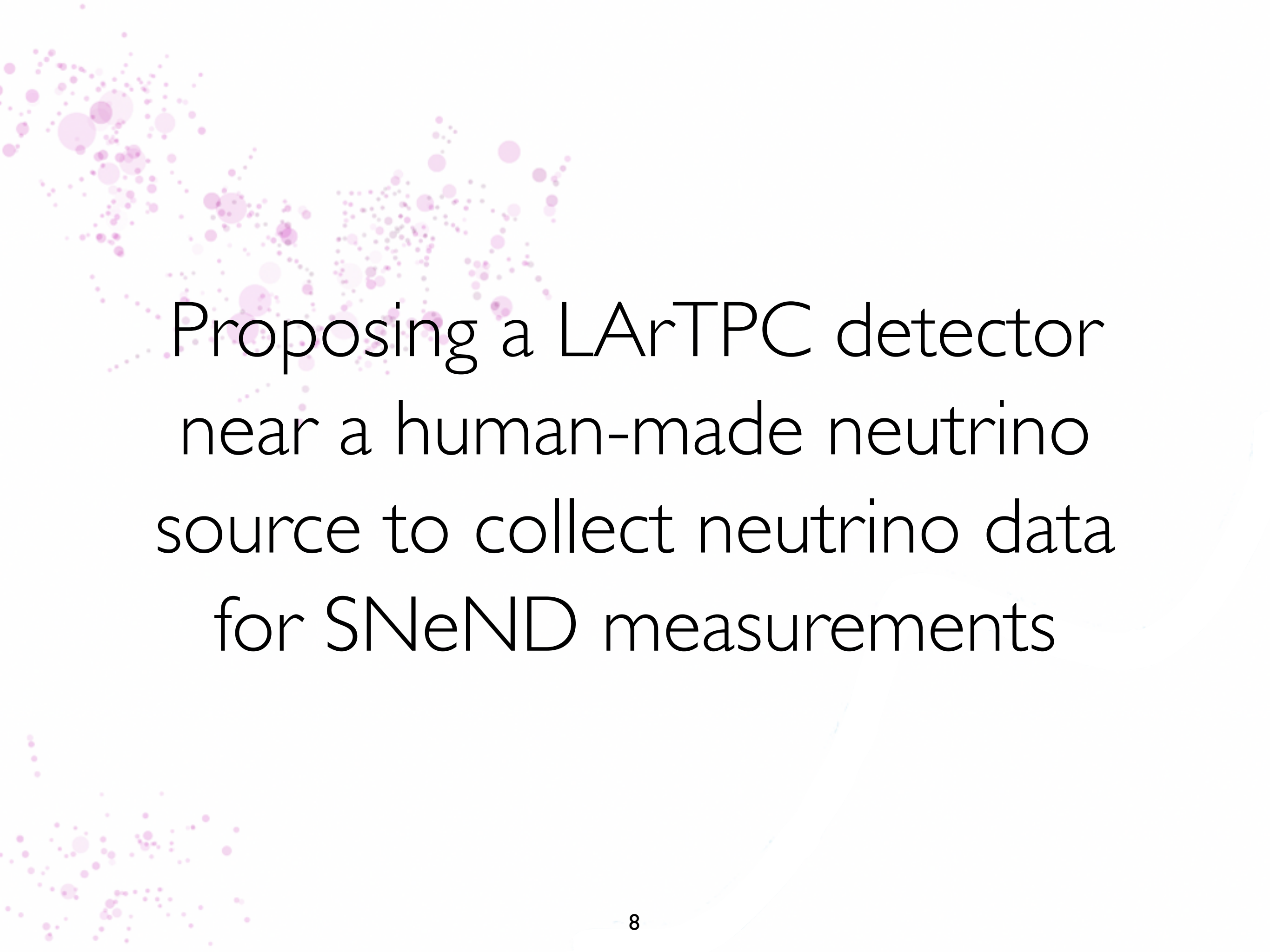


SNeND Constraints

...and with the “neutrino data” constraints, SNeND,



Preliminary result. In collaboration with Gianluca Petrillo (SLAC), Yen-Hsun Lin (NCTS)



Proposing a LArTPC detector
near a human-made neutrino
source to collect neutrino data
for SNeND measurements

Proposing a LArTPC detector near a human-made neutrino source to collect neutrino data for SNeND measurements

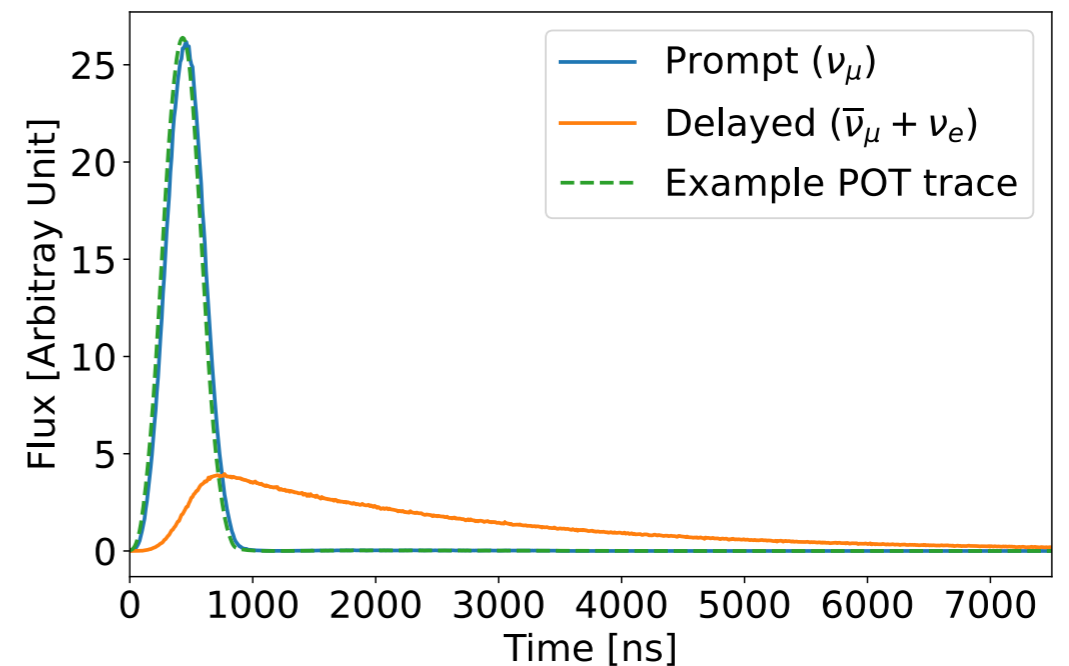
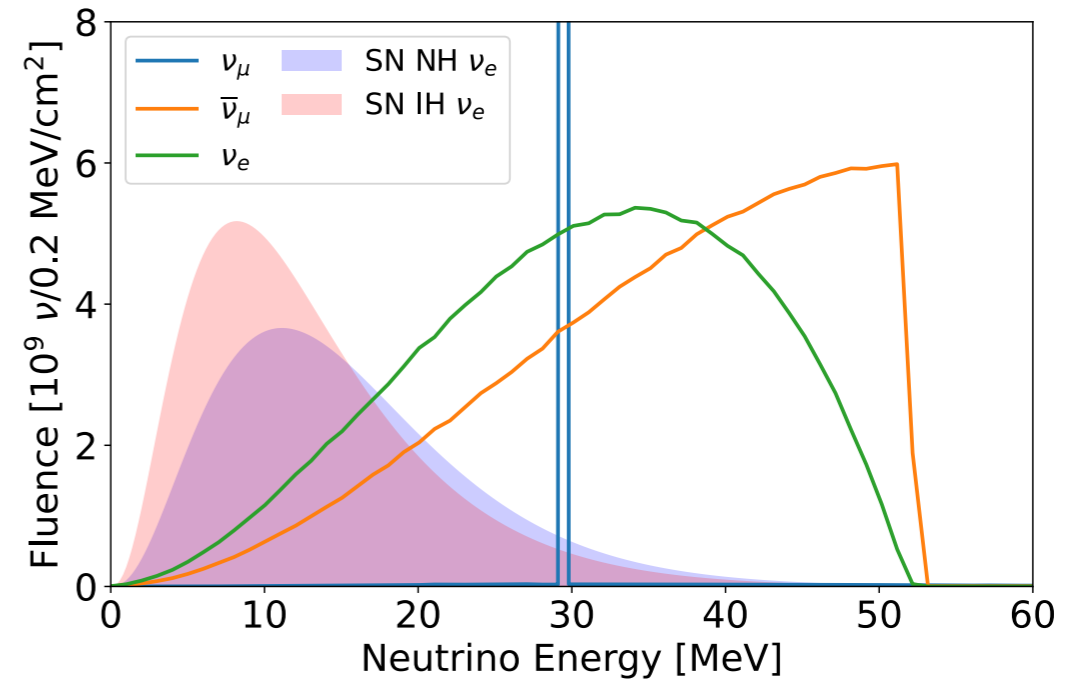
$$\rho(E_{\nu,j}) = \frac{\sum_k \phi_{\text{SN}}(E_{\nu,k}) \times \sigma(E_{\nu,k}) \times R_{j,\text{DUNE}}(E_{\nu,k})}{\sum_k \phi_{\text{SNeND}}(E_{\nu,k}) \times \sigma(E_{\nu,k}) \times R_{j,\text{SNeND}}(E_{\nu,k})}$$

ϕ_{SNs} at 2-3% precision

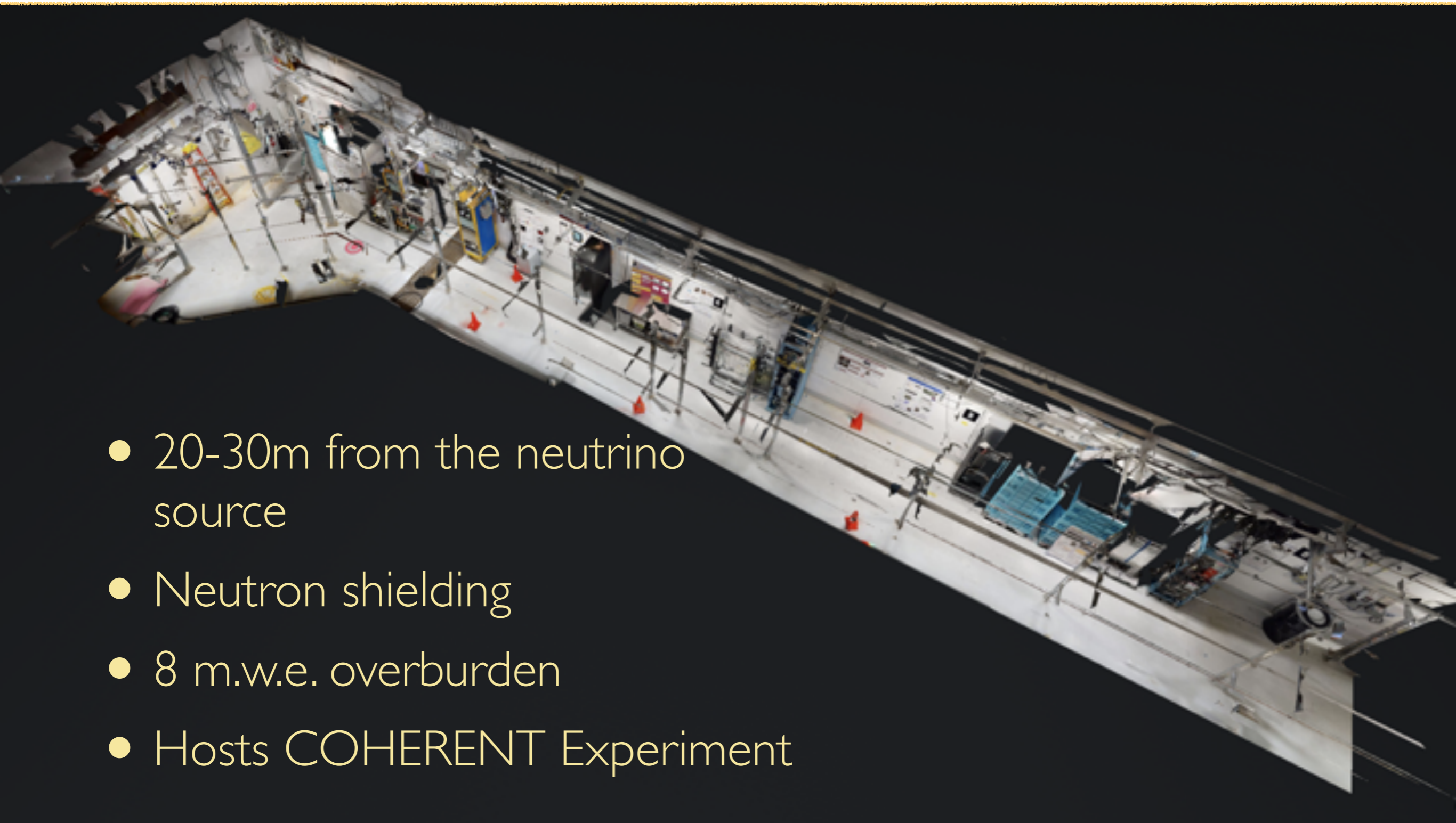
COHERENT-LArTPC

Measure ν_e -Ar CC σ @SNS

- Neutrinos produced from π^+ decay at rest, $E_\nu \sim \mathcal{O}(10\text{MeV})$
 $\pi^+ \rightarrow \mu^+ \nu_\mu \rightarrow (e^+ \nu_e \bar{\nu}_\mu) \nu_\mu$
- H- LINAC: 1 GeV @ 1.4 MW, 60 Hz; mostly pions
 - Will be upgraded 2024!
- Liquid mercury target
 - Minimize pions decay-in-flight
- Operate ~ 5000 hours/year
- $2.81 \times 10^{14} \nu/\text{cm}^2/\text{flavor}/\text{year}$ @ 20m



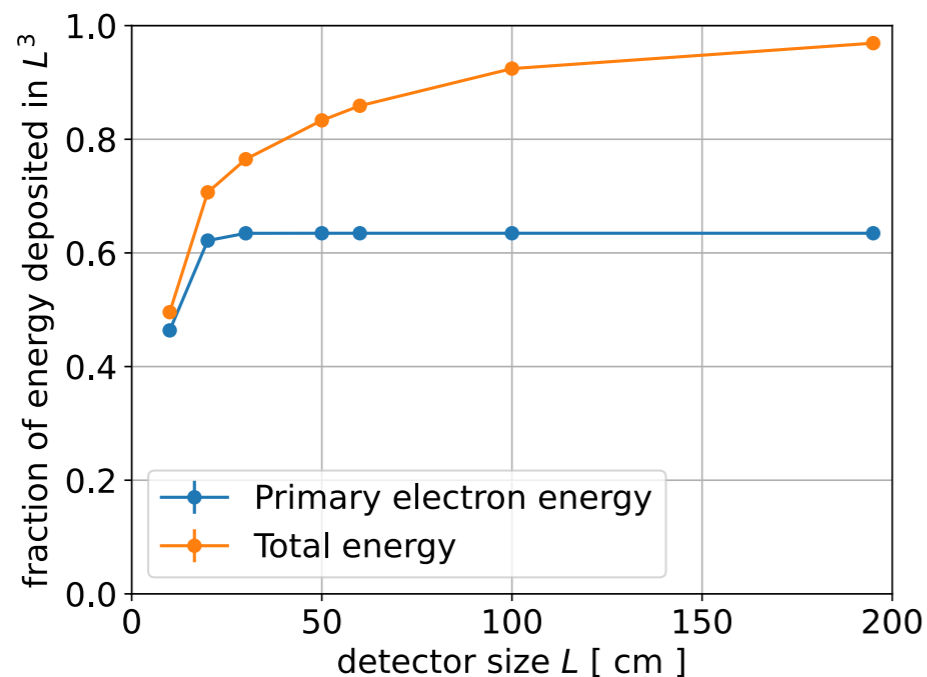
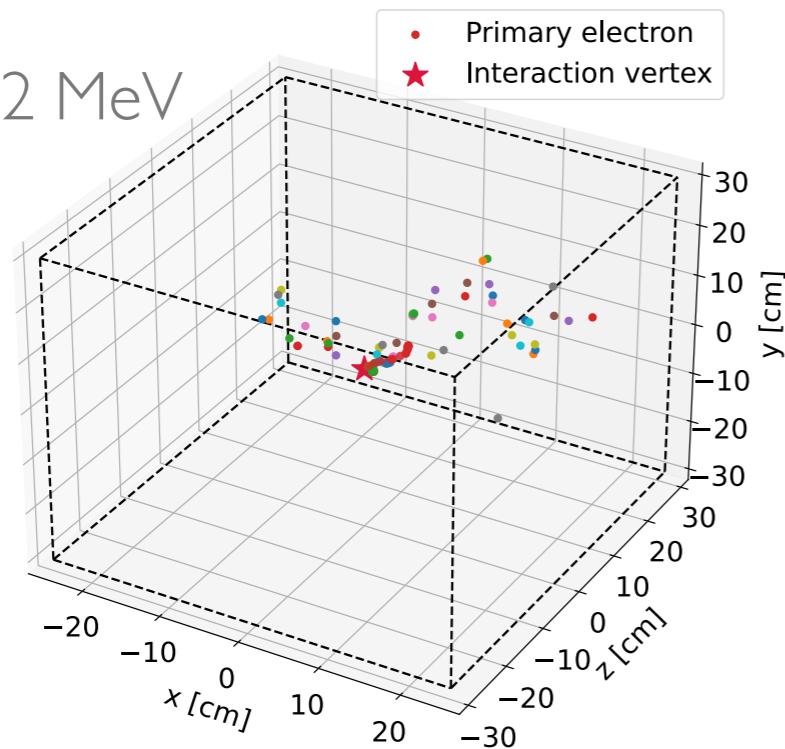
First Target Station



- 20-30m from the neutrino source
- Neutron shielding
- 8 m.w.e. overburden
- Hosts COHERENT Experiment

Proposed LArTPC

$E_\nu = 42 \text{ MeV}$

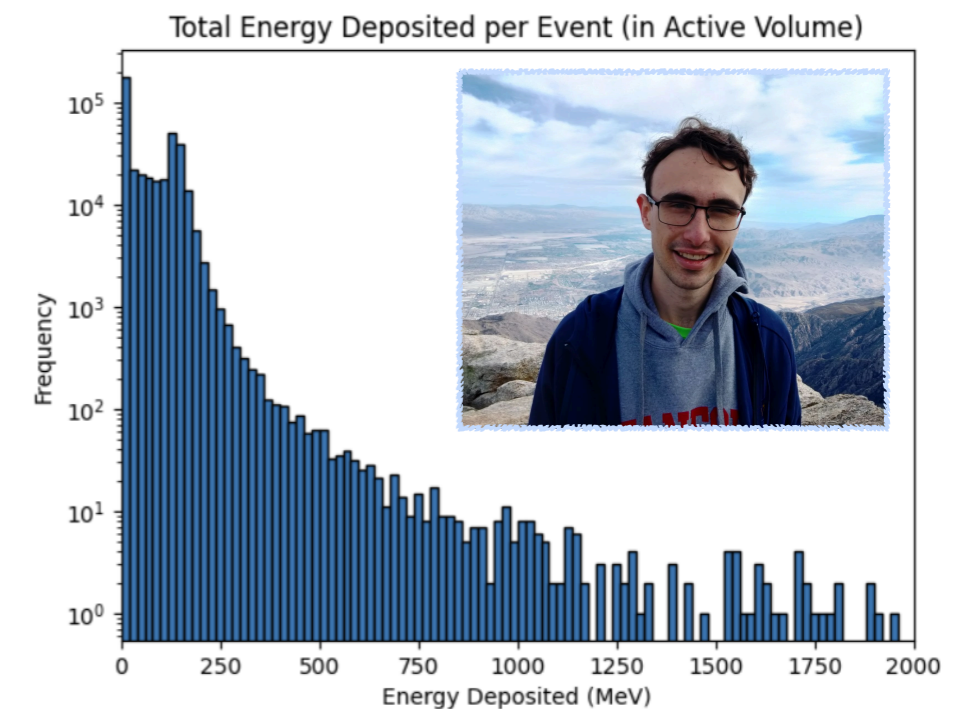
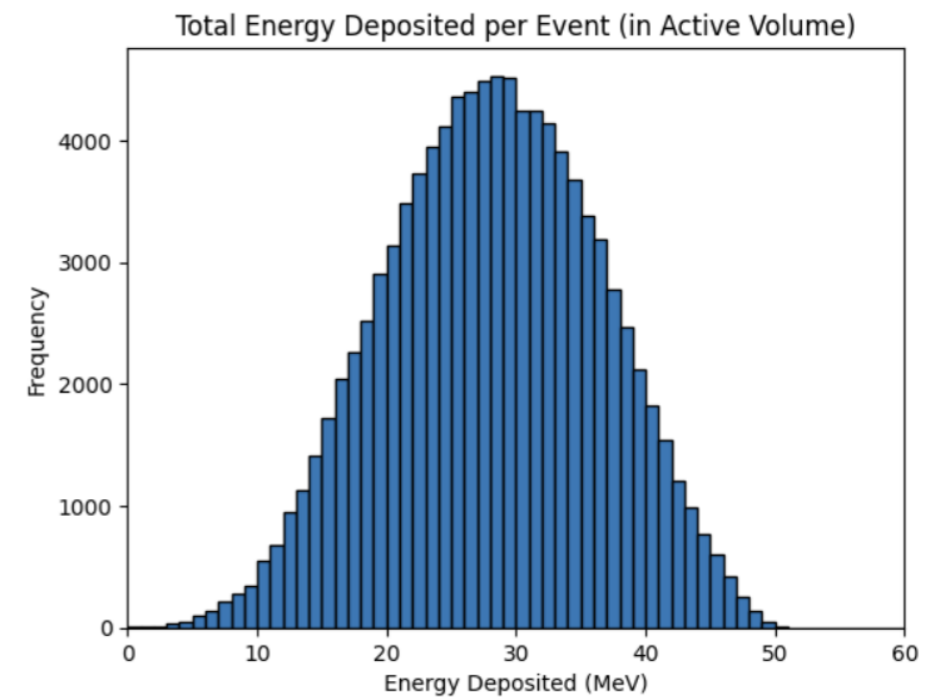


- $50 \times 60 \times 60 \text{ cm}^3$, 250 kg LAr in the active volume
- Pixelated charge readout
- 30cm drift distance: 2 TPCs
- Expected signal: $\sim 150/\text{year}$
- Expected background
 - Cosmic rays (main)
 - Beam related neutrons
 - ν interactions outside
 - Hot-off gas (radiological waste)

ν_e -Ar CC Sensitivity

GEANT4-level study by John Vaccaro (Stanford)

	Location 1	Location 2
Signal	176.4	86.9
Cosmics	27	27
Beam related neutrons	0.1	3.82
ν outside	44.5	21.9
Likelihood to observe 3 S.D. in 1 year	100%	93.3%
Likelihood to observe 3 S.D. in 3 year	100%	97.5%



Opportunities & Synergy

- Synergy in the MeV-scale physics LArTPC R&D, e.g. GAMPix, and DUNE ND-LAr, and theorist efforts at SLAC
- Collaborators from Hawaii, Napoli (Italy), Campinas (Brazil), Alabama, UT Arlington, Wellesley, NYCU (Taiwan) expressed interests
- Future opportunities
 - 10-ton-scale for differential cross section measurements (ASTAE-scale)
 - The Second Target Station at Oak Ridge and European Spallation Source



Backup

20% Uncertainty on σ

