

The Short Baseline Near Detector at Fermilab: First data and the exciting neutrino physics program ahead

Michelle Stancari



U.S. DEPARTMENT
of **ENERGY**

Fermi National Accelerator Laboratory is managed by
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No.	No.	No.	No.	No.	No.	No.	No.	No.
H 1	F 8	Cl 15	Co & Ni 22	Br 29	Pd 36	I 42	Pt & Ir 50	
Li 2	Na 9	K 16	Cu 23	Rb 30	Ag 37	Cs 44	Os 51	
G 3	Mg 10	Ca 17	Zn 24	Sr 31	Cd 38	Ba & V 45	Hg 52	
Bo 4	Al 11	Cr 19	Y 25	Ce & La 33	U 40	Ta 46	Tl 53	
C 5	Si 12	Ti 18	In 26	Zr 32	Sn 39	W 47	Pb 54	
N 6	P 13	Mn 20	As 27	Di & Mo 34	Sb 41	Nb 48	Bi 55	
O 7	S 14	Fe 21	Se 28	Ro & Ru 35	Te 43	Au 49	Th 56	

SCIENTIFIC CHINESE ALPHABETIC & NUMERICAL

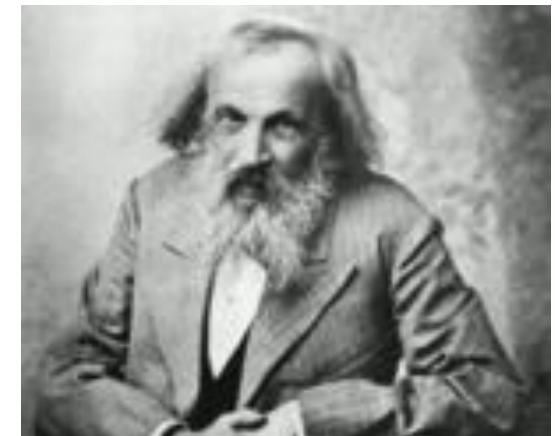
Group	Group I RO	Group II RO ² HO	Group III RO	Group IV RO ² HO	Group V RO	Group VI RO ² HO	Group VII RO	Group VIII RO ² HO	Group IX RO	Group X RO ² HO	Group XI RO	Group XII RO ² HO	Group XIII RO	Group XIV RO ² HO	Group XV RO	Group XVI RO ² HO	Group XVII RO	Group XVIII RO ² HO	Group XIX RO	Group XX RO ² HO	Group XXI RO	Group XXII RO ² HO	Group XXIII RO	Group XXIV RO ² HO	Group XXV RO	Group XXVI RO ² HO	Group XXVII RO	Group XXVIII RO ² HO	Group XXIX RO	Group XXX RO ² HO		
Group I	H-1																															
Group II	Li-7	Be-9	B-11	C-12	N-14	O-16	F-19																									
Group III	Na-23	Mg-24	Al-27	Si-28	P-31	S-32	Cl-35																									
Group IV	K-39	Ca-40		Ti-48	V-51	Cr-52	Mn-55	Fe-56	Co-59	Ni-59	Cu-63																					
Group V		Zn-65			As-75	Se-78	Br-80																									
Group VI	Rb-85	Sr-87		Zr-90	Nb-94	Mo-96		Hu-104	Rh-104	Pd-106	Ag-108																					
Group VII																																
Group VIII																																
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The first periodic tables (1866, 1871)

SCIENTIFIC CHINESE JOURNAL, A. M. 1900

This is a detailed scientific periodic table from a Chinese journal. It features a grid of elements with their symbols and atomic weights. The table is organized into groups labeled 'Group I' through 'Group III' at the top. The elements are arranged in a way that shows their periodic properties. The table includes elements from Hydrogen (H) to Uranium (U). The layout is dense and includes various annotations and symbols in Chinese characters.



		1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
		Periodic Table																		
		Atomic Symbol Name Weight																		
		Metals																		
		Nonmetals																		
		Alkali metals																		
		Alkaline earth metals																		
		Lanthanoids																		
		Actinoids																		
		Transition metals																		
		Post-transition metals																		
		Metalloids																		
		Reactive nonmetals																		
		Noble gases																		
		Pnictogens																		
		Chalcogens																		
		Halogens																		
1	H Hydrogen 1.008																	2	He Helium 4.0026	
2	Li Lithium 6.94	Be Beryllium 9.0122																	10	Ne Neon 20.180
3	Na Sodium 22.990	Mg Magnesium 24.305																	18	Ar Argon 39.948
4	K Potassium 39.098	Ca Calcium 40.078	Sc Scandium 44.956	Ti Titanium 47.867	V Vanadium 50.942	Cr Chromium 51.996	Mn Manganese 54.938	Fe Iron 55.845	Co Cobalt 58.933	Ni Nickel 58.693	Cu Copper 63.546	Zn Zinc 65.38	Ga Gallium 69.723	Ge Germanium 72.630	As Arsenic 74.922	Se Selenium 78.971	Br Bromine 79.904	Kr Krypton 83.798		
5	Rb Rubidium 85.468	Sr Strontium 87.62	Y Yttrium 88.906	Zr Zirconium 91.224	Nb Niobium 92.906	Mo Molybdenum 95.95	Tc Technetium (98)	Ru Ruthenium 101.07	Rh Rhodium 102.91	Pd Palladium 106.42	Ag Silver 107.87	Cd Cadmium 112.41	In Indium 114.82	Sn Tin 118.71	Sb Antimony 121.76	Te Tellurium 127.60	I Iodine 126.90	Xe Xenon 131.29		
6	Cs Caesium 132.91	Ba Barium 137.33	57-71	Hf Hafnium 178.49	Ta Tantalum 180.95	W Tungsten 183.84	Re Rhenium 186.21	Os Osmium 190.23	Ir Iridium 192.22	Pt Platinum 195.08	Au Gold 196.97	Hg Mercury 200.59	Tl Thallium 204.38	Pb Lead 207.2	Bi Bismuth 208.98	Po Polonium (209)	At Astatine (210)	Rn Radon (222)		
7	Fr Francium (223)	Ra Radium (226)	89-103	Rf Rutherfordium (267)	Db Dubnium (268)	Sg Seaborgium (269)	Bh Bohrium (270)	Hs Hassium (277)	Mt Meitnerium (278)	Ds Darmstadtium (281)	Rg Roentgenium (282)	Cn Copernicium (285)	Nh Nihonium (286)	Fl Flerovium (289)	Mc Moscovium (290)	Lv Livermorium (293)	Ts Tennessine (294)	Og Oganesson (294)		
For elements with no stable isotopes, the mass number of the isotope with the longest half-life is in parentheses.																				
6	La Lanthanum 138.91	Ce Cerium 140.12	Pr Praseodymium 140.91	Nd Neodymium 144.24	Pm Promethium (145)	Sm Samarium 150.36	Eu Europium 151.96	Gd Gadolinium 157.25	Tb Terbium 158.93	Dy Dysprosium 162.50	Ho Holmium 164.93	Er Erbium 167.26	Tm Thulium 168.93	Yb Ytterbium 173.05	Lu Lutetium 174.97					
7	Ac Actinium (227)	Th Thorium 232.04	Pa Protactinium 231.04	U Uranium 238.03	Np Neptunium (237)	Pu Plutonium (244)	Am Americium (243)	Cm Curium (247)	Bk Berkelium (247)	Cf Californium (251)	Es Einsteinium (252)	Fm Fermium (257)	Md Mendelevium (258)	No Nobelium (259)	Lr Lawrencium (266)					

ptable.com

mass →	$\approx 2.3 \text{ MeV}/c^2$	$\approx 1.275 \text{ GeV}/c^2$	$\approx 173.07 \text{ GeV}/c^2$	0	$\approx 126 \text{ GeV}/c^2$
charge →	$2/3$	$2/3$	$2/3$	0	0
spin →	$1/2$	$1/2$	$1/2$	1	0
QUARKS	u up	c charm	t top	g gluon	H Higgs boson
	$\approx 4.8 \text{ MeV}/c^2$	$\approx 95 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$	0	
	$-1/3$	$-1/3$	$-1/3$	0	
	$1/2$	$1/2$	$1/2$	1	
	d down	s strange	b bottom	γ photon	
	$0.511 \text{ MeV}/c^2$	$105.7 \text{ MeV}/c^2$	$1.777 \text{ GeV}/c^2$	$91.2 \text{ GeV}/c^2$	
	-1	-1	-1	0	
	$1/2$	$1/2$	$1/2$	1	
	e electron	μ muon	τ tau	Z Z boson	
LEPTONS	$< 2.2 \text{ eV}/c^2$	$< 0.17 \text{ MeV}/c^2$	$< 15.5 \text{ MeV}/c^2$	$80.4 \text{ GeV}/c^2$	
	0	0	0	± 1	
	$1/2$	$1/2$	$1/2$	1	
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	
				GAUGE BOSONS	

“Recent” Discoveries

Charm quark – 1974

Tau lepton – 1978

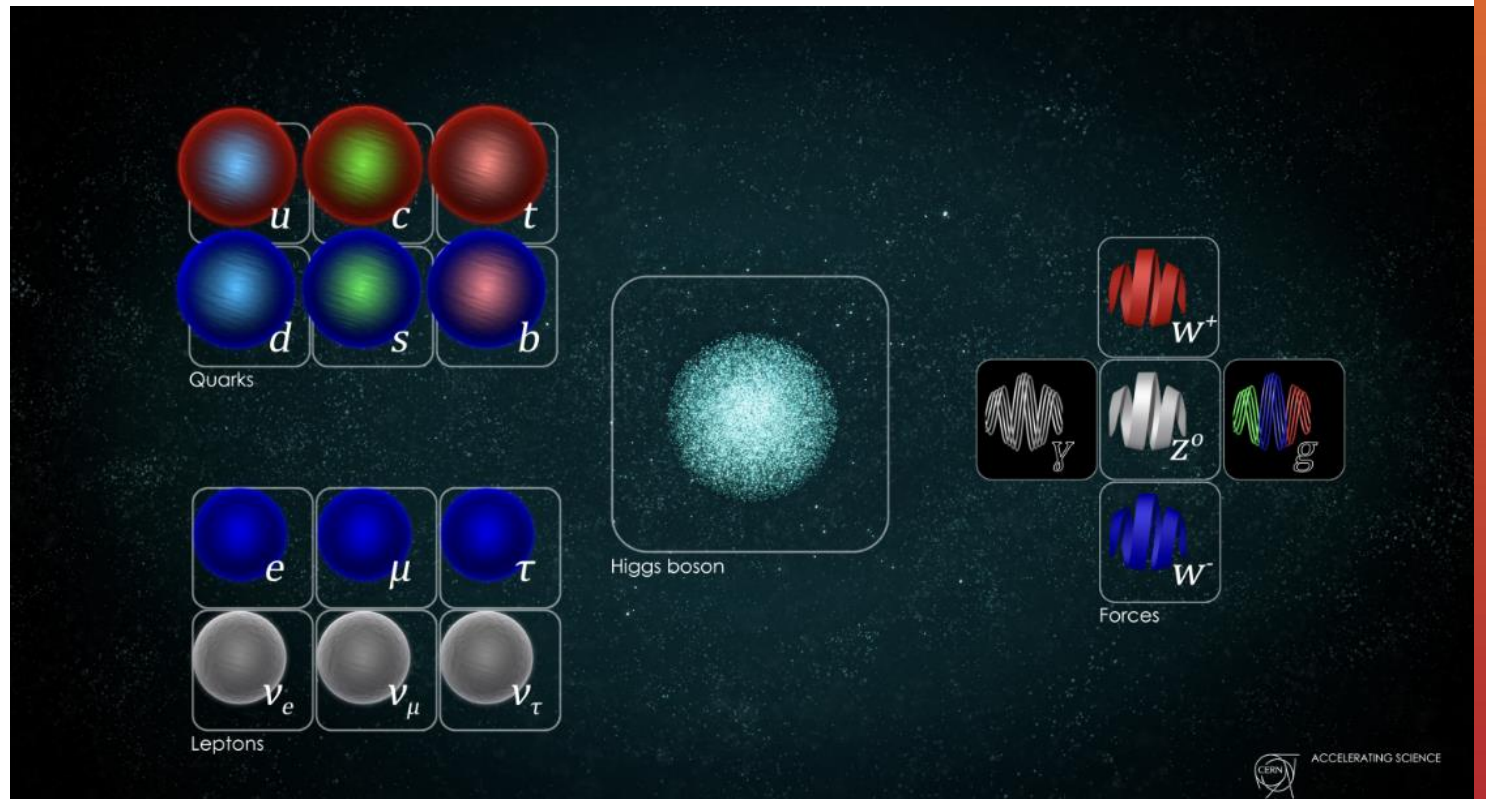
Bottom quark – 1977

W,Z bosons - 1983

Top quark – 1995

Tau neutrino – 2000

Higgs boson – 2012



Intriguing properties of neutrinos

- Tiny cross section – interact only through the weak force
- Mass so tiny that it has not yet been measured
- Oscillate from one flavor to another

Studying oscillations with experiments

CONCEPT: Find or make a neutrino source. Get two detectors. Just count leptons in each!

However . . .

- Oscillation probability depends on neutrino flavor and energy
- Nuclear target complicates extracting the neutrino energy from the observed final state particles
- Preferred final states . . . a hadron leaving the interaction improves vertex resolution and background rejection.

REALITY: precision oscillation measurements require exquisite understanding of BOTH the neutrino source and the neutrino-nucleus interactions

The Short Baseline Neutrino Program at Fermilab

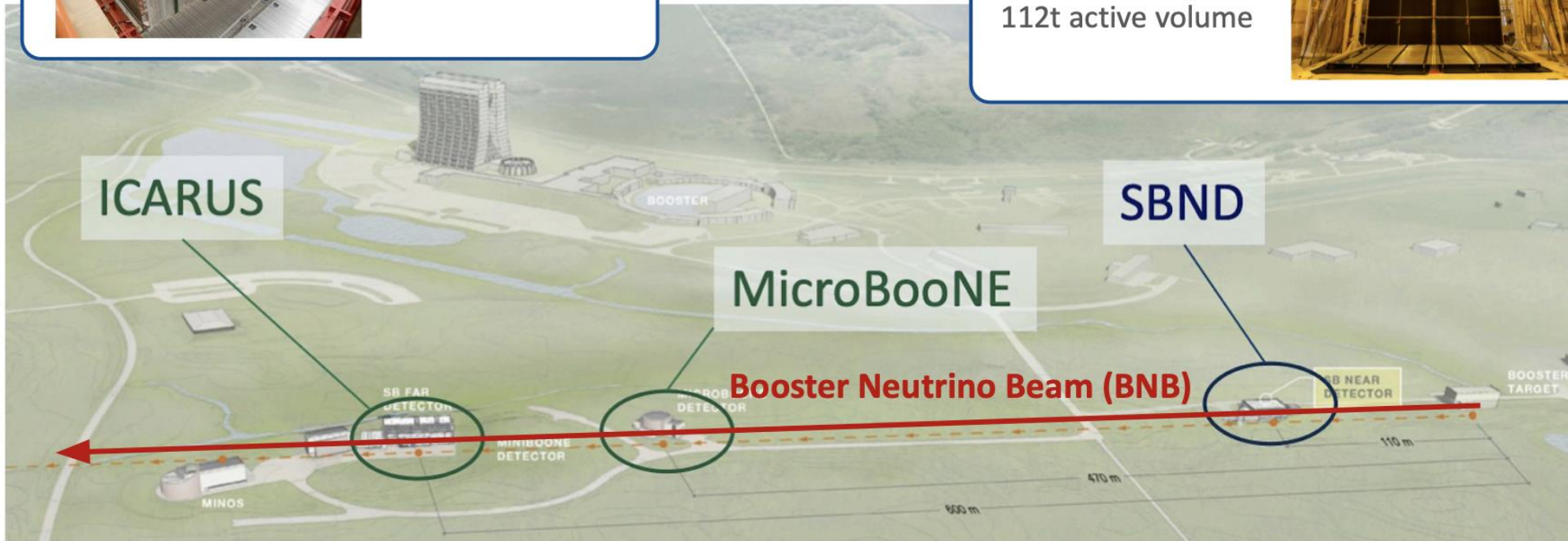


ICARUS

600m baseline
470t active volume

SBND

110m baseline
112t active volume



SBND will perform a robust measurement in the search of sterile neutrinos, while also investigating a broad spectrum of other new physics beyond the standard model.

Strategy to reduce uncertainties:

- Three detectors sampling the **same neutrino beam at different distances** (BNB)
- **Same nuclear target** (Ar) and **detector technology** (LArTPC: liquid argon time projection chambers)

The Short Baseline Neutrino Program at Fermilab



ICARUS

600m baseline
470t active volume

SBND

110m baseline
112t active volume



Active mass: 112 t
Distance: 110 m
Operation: 2024-

Active mass: 476 t
Distance: 600 m
Operation: 2021-

SBN will perform a robust measurement in the search of sterile neutrinos, while also investigating a broad spectrum of other new physics beyond the standard model.

Strategy

- Three detectors sampling the **same neutrino beam at different distances** (BNB)
- **Same nuclear target** (Ar) and **detector technology** (LArTPC: liquid argon time projection chambers)

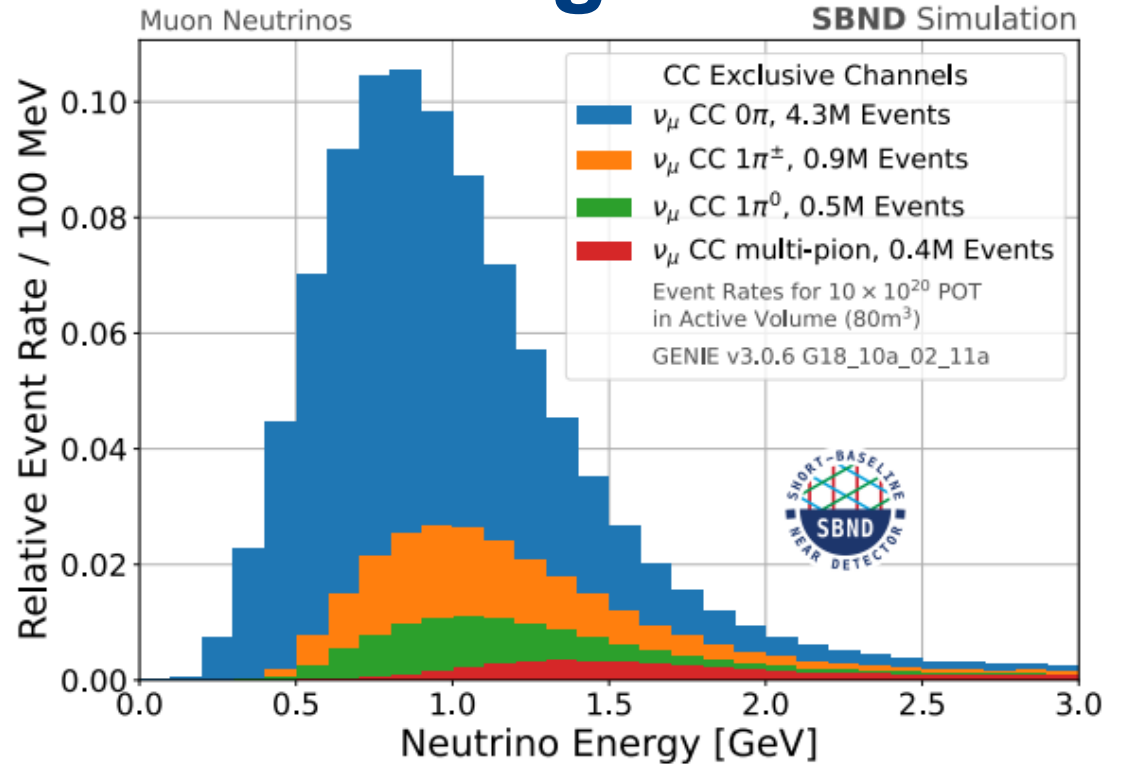
Near Detector – only 110 m from the target!

Unique opportunities:

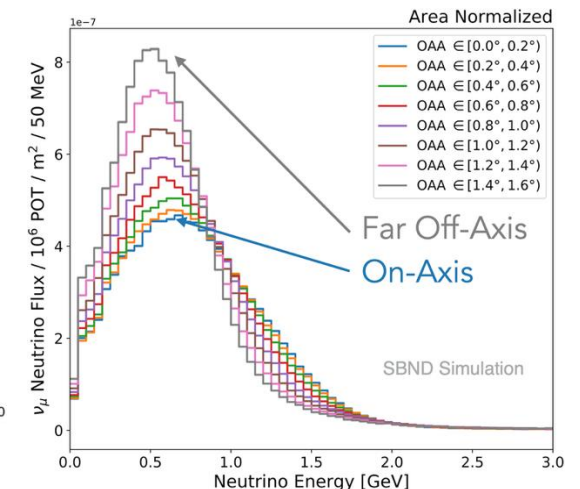
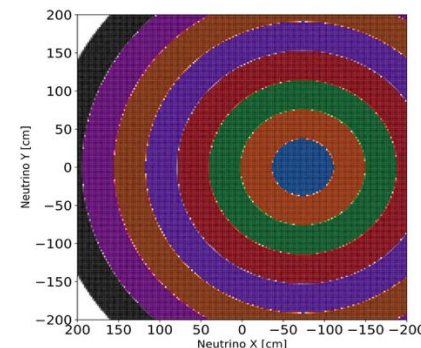
- Up to 7000 neutrino interactions per day – unprecedented statistics
 - Cross sections, cross sections, cross sections
 - rare process/BSM searches
- Prism effect – sample off axis fluxes

Unique challenges:

- Neutrino pile-up
- Backgrounds from the beam
- Sheer data volume ~100 MB/event

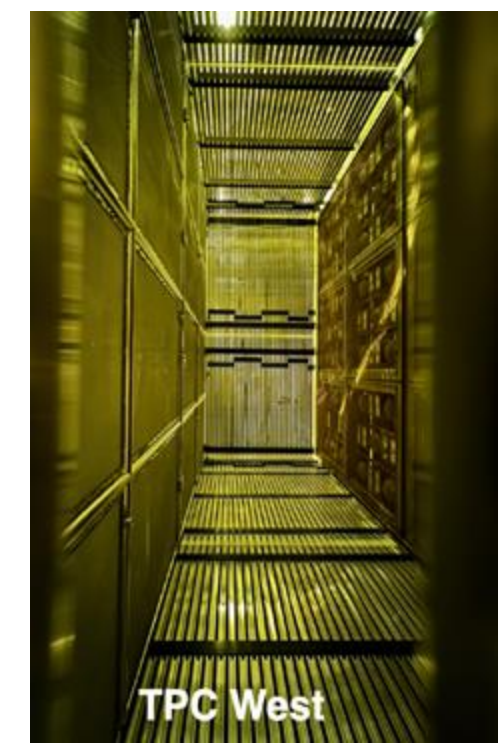
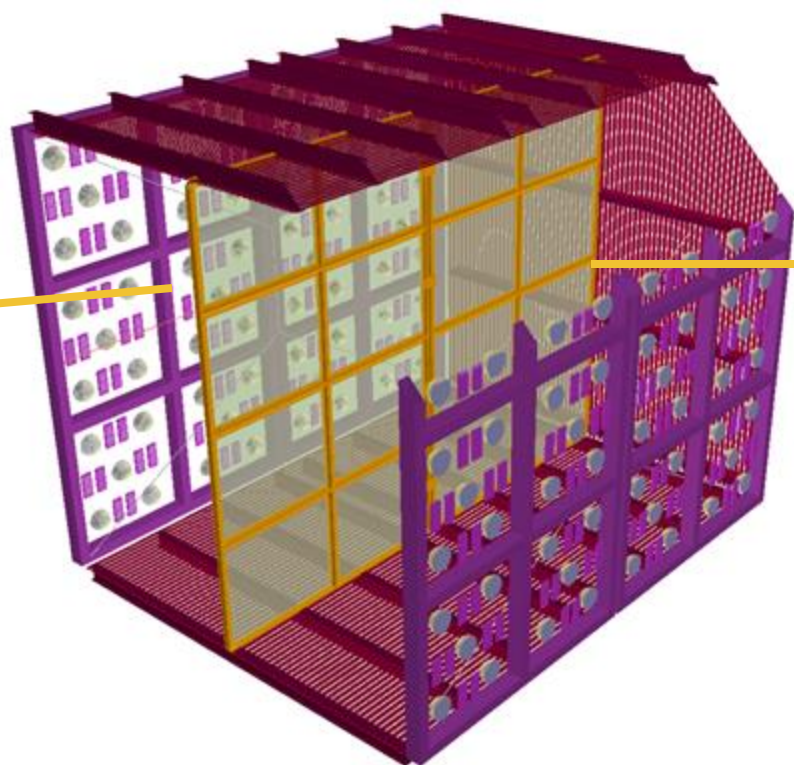


SBND Simulation



SBND TPC

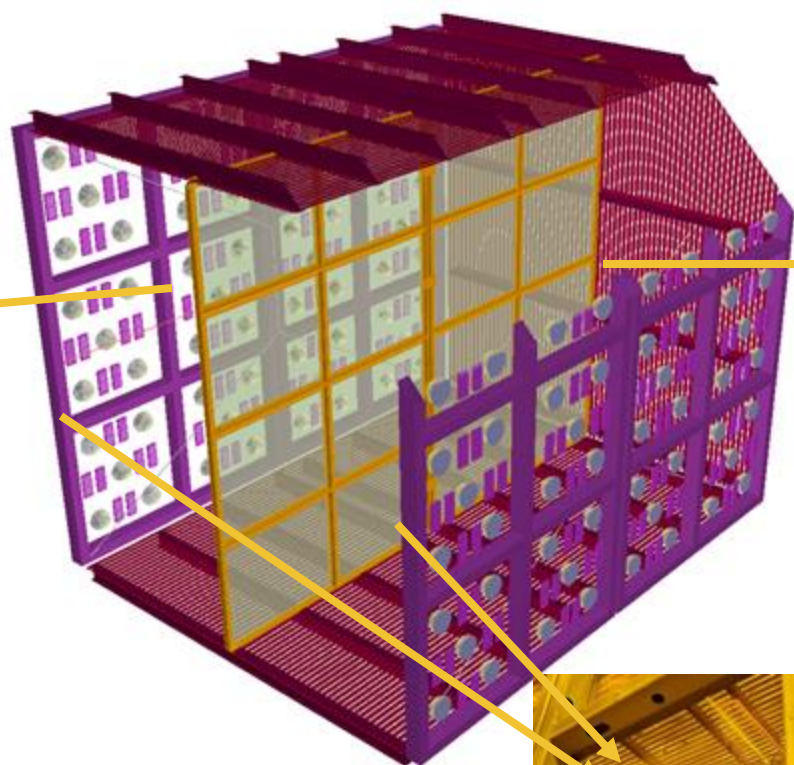
2 drift volumes
[total 4 X 4 X 5 m]



SBND TPC

2 drift volumes

[total 4 X 4 X 5 m]



Anode Plane
on either side. Each consists
of 3 planes of wires with 3
mm spacing and different
angle per plane. Total of
11,260 wires

SBND TPC

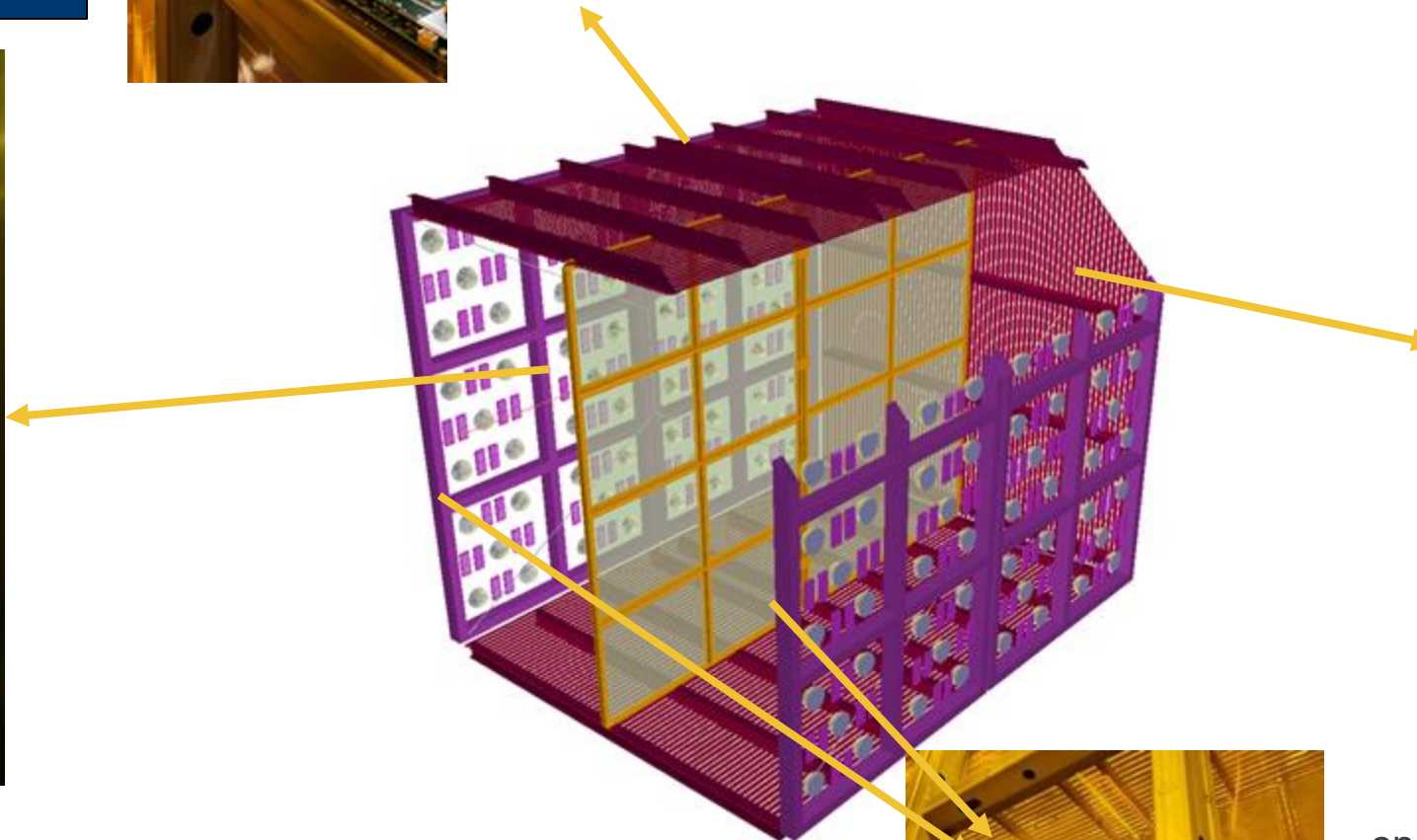
2 drift volumes
[total 4 X 4 X 5 m]



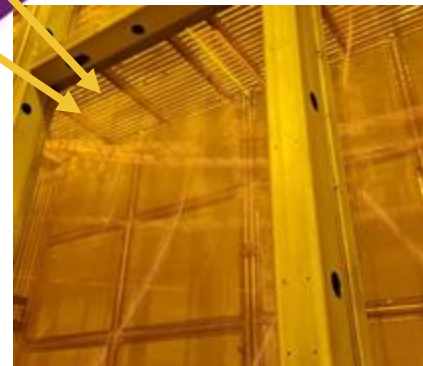
Cold (89K)
Electronics to
pre-amplify and
digitize signals



TPC East



TPC West



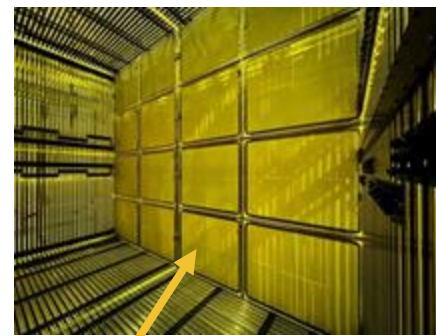
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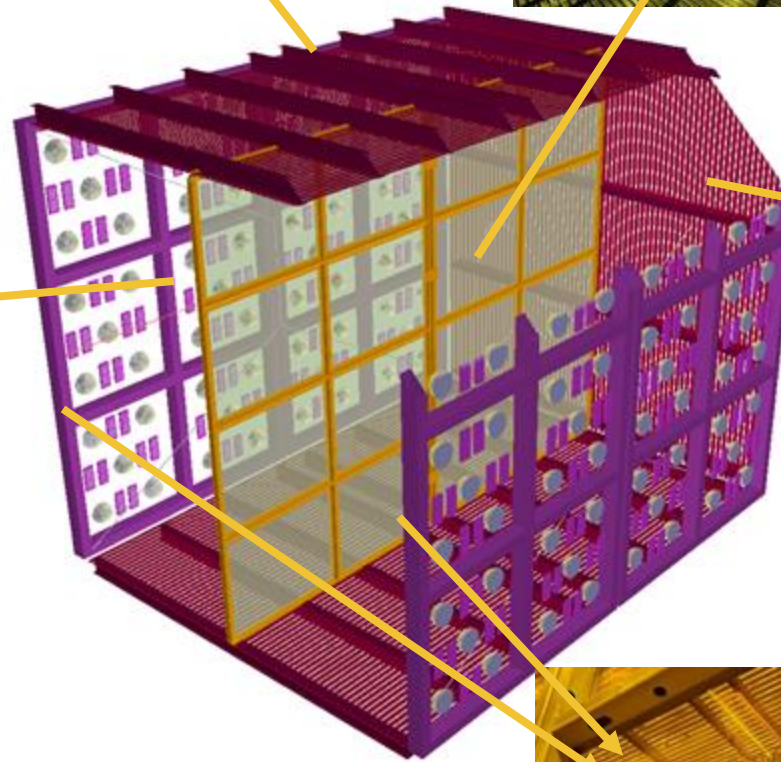
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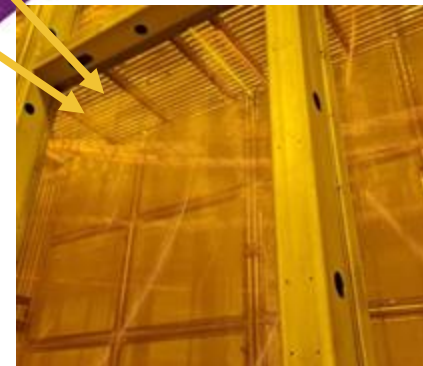
Cathode Plane
in the middle; divides the detector
into 2 TPCs. Will be supplied with
-100 kV.



TPC East



TPC West



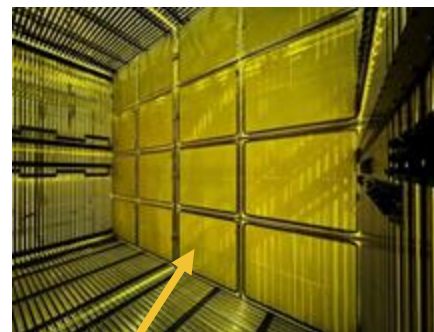
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SBND TPC

2 drift volumes
[total 4 X 4 X 5 m]



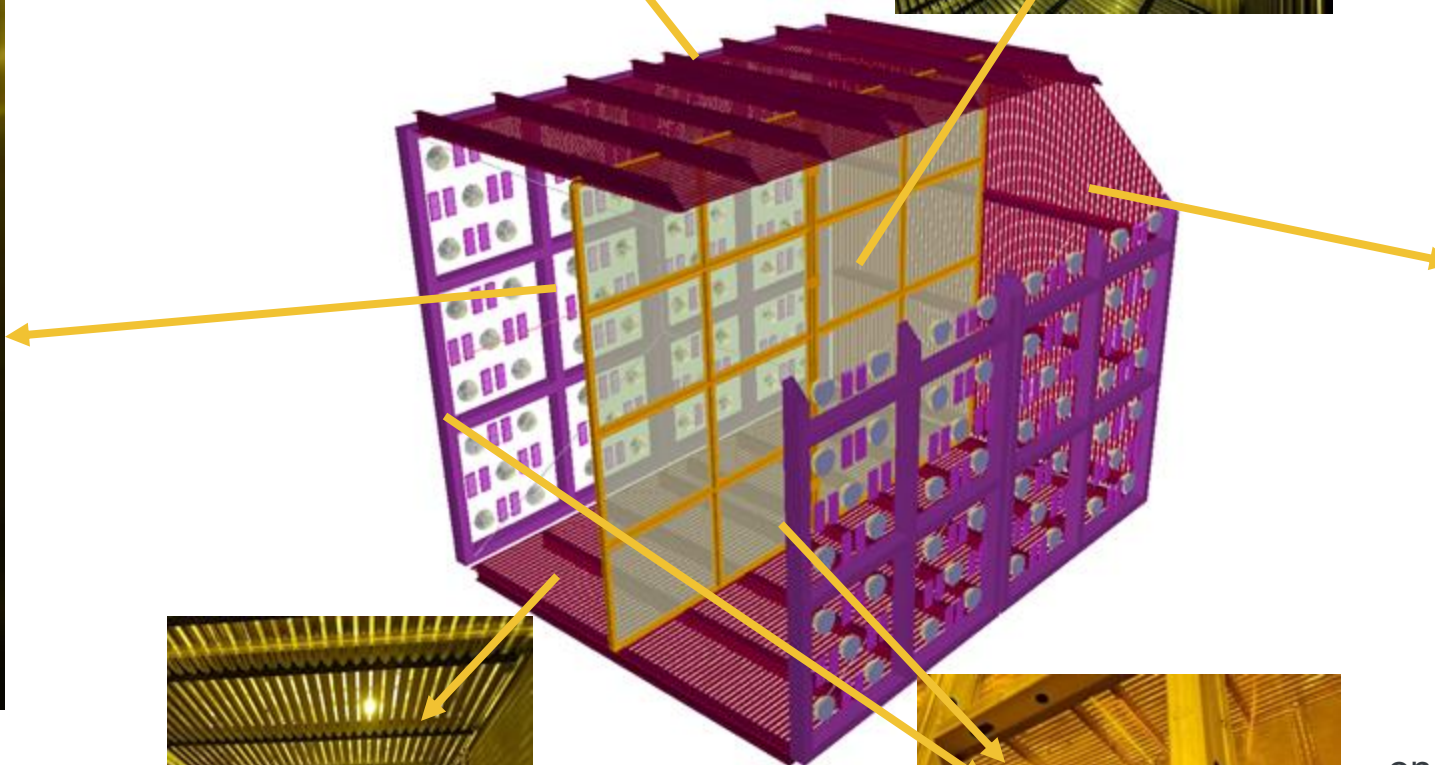
Cold (89K) Electronics to pre-amplify and digitize signals



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TPC East



TPC West

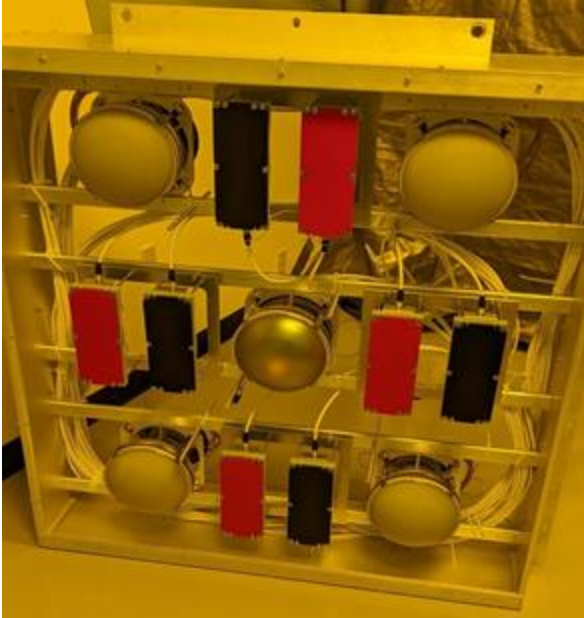


Field Cage that wraps around the 2 LArTPCs to step down the voltage & ensure uniform electric field of 500 V/cm.



Anode Plane on either side. Each consists of 3 planes of wires with 3 mm spacing and different angle per plane. Total of 11,260 wires

SBND Photon Detection System



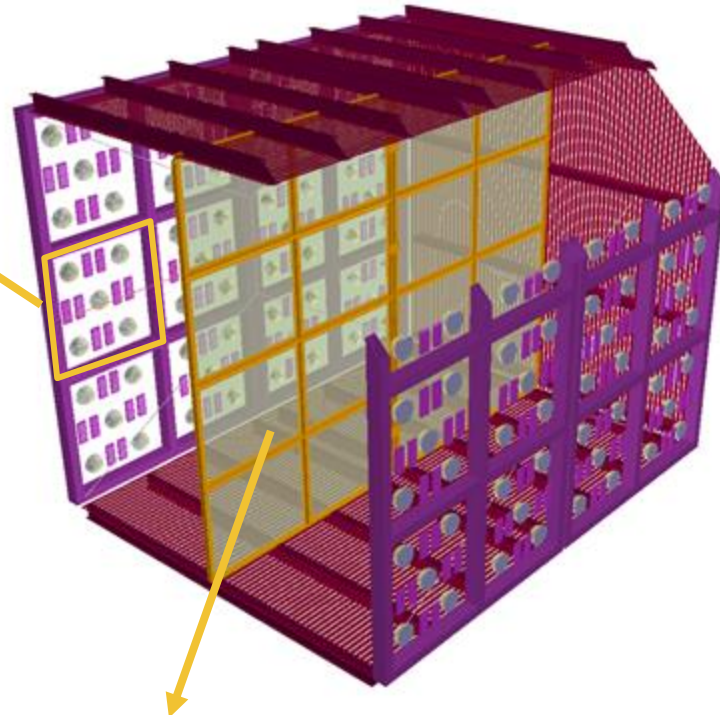
24 Anode Plane boxes

4x24 = 96 **PMTs** (TPB coated)

1x24 = 24 **PMTs** (uncoated)

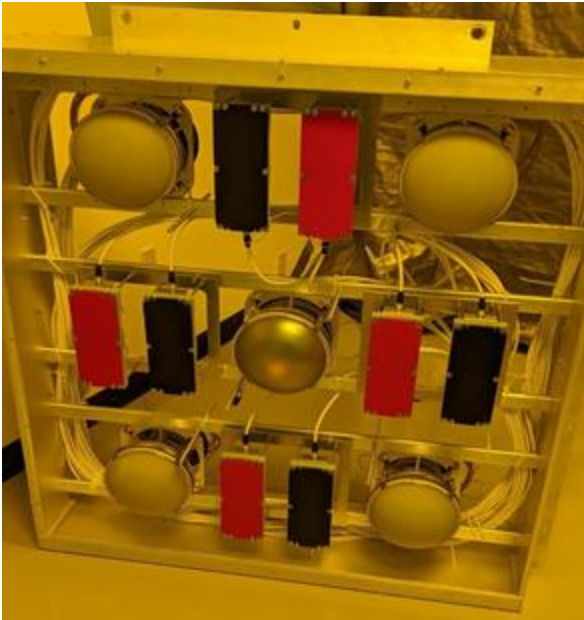
8x24 = 192 **X-ARAPUCAs***

*sensitive to UV
+ visible light



Cathode Plane
with TPB coated
reflective foils
mounted between
mesh panels.

SBND Photon Detection System



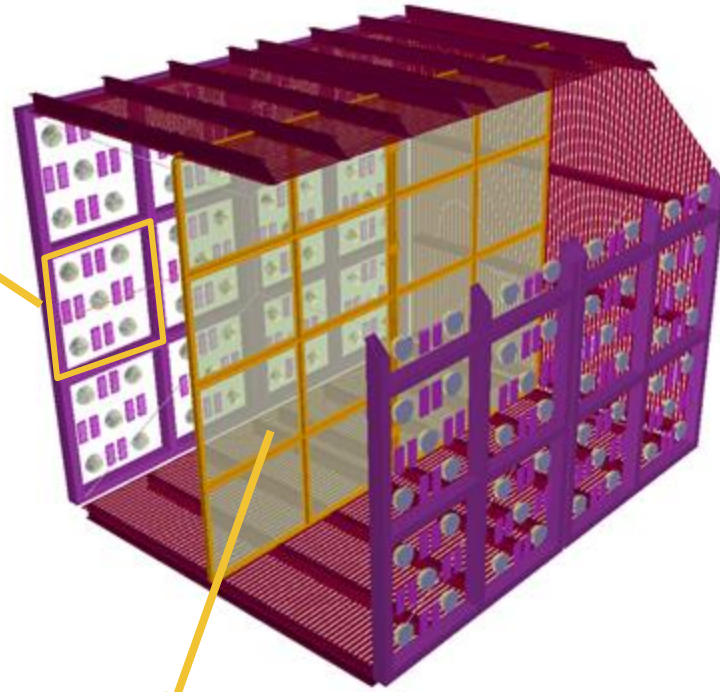
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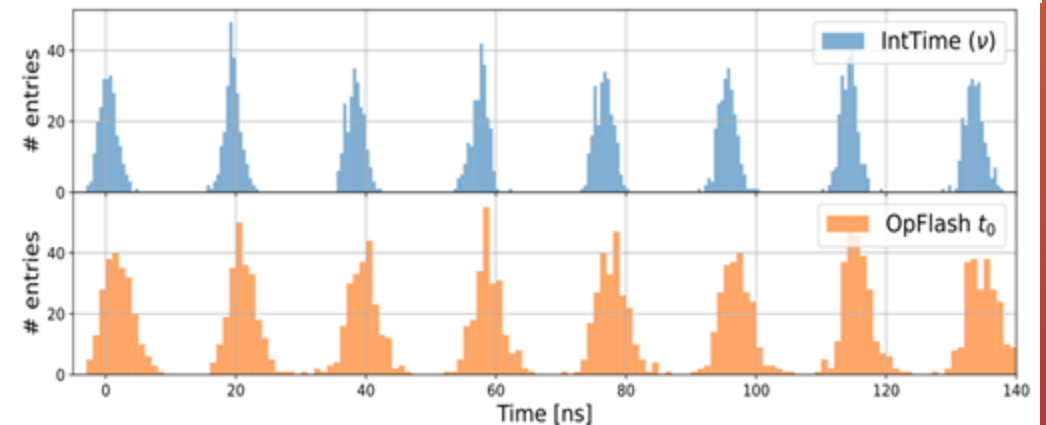
Cathode Plane
with TPB coated
reflective foils
mounted between
mesh panels.

While ionization electrons drift slowly at ~ 1.6 mm/us, scintillation light is fast.

- triggering
- cosmic background rejection

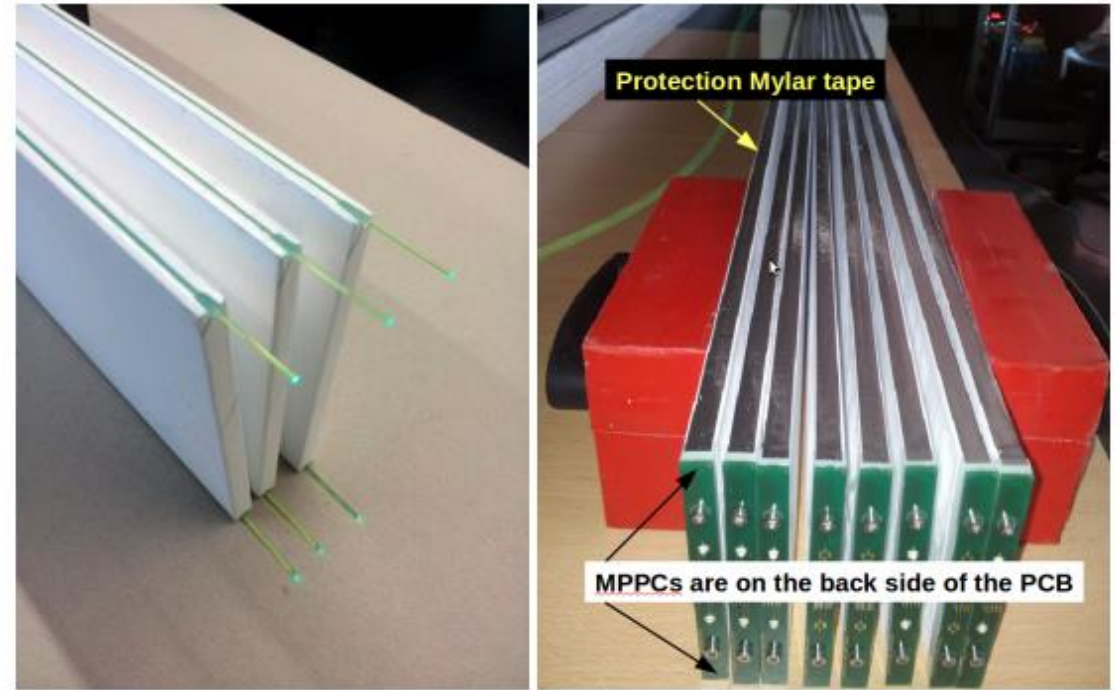
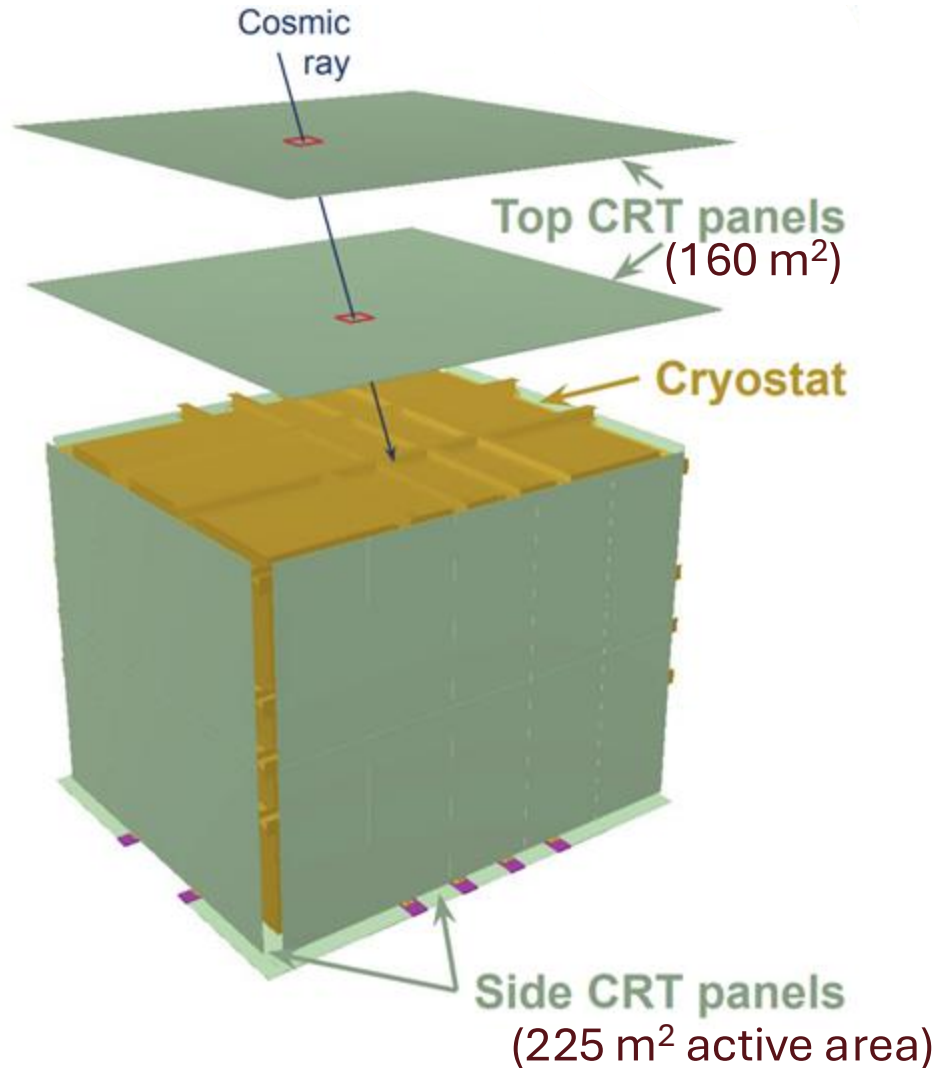
Unprecedented coverage enables SBND to explore using light for calorimetry

Experiment	Average light yield (PE/MeV)	Uniform light collection?
MicroBooNE	~ 5	no
LArIAT	~ 18	yes
pDUNE-SP	1.9 at 3.3m	no
SBND	~ 80 (> 50 min)	yes
DUNE: Vertical Drift	~ 38 (> 16.5 min)	yes



Simulated (top) and reconstructed (bottom) light flashes showing the neutrino beam structure.

SBND Cosmic Ray Tagger



Designed to tag muons

- Each plane has two layers: one "x" strips and one "y" strips
- Each strip has 2 SiPMs, one per fiber
- Single strip (1 cm thick) absorbs radiological decays
- Readout triggered by 4-fold time coincidence of one "x" strip and one "y" strip
- Strips are 1 cm x 11 cm x 2.5-4.5 m
- position resolution: <3 cm
- time resolution: < 2 ns



SBND Trigger System



Two custom electronics boards.

- “MTC-A” - Analog summer
- “PTB” - FPGA-based Gating and Logic

Trigger inputs

- PMT multiplicity
- Beam “early warning” signals
- CRT activity per wall/plane

Why do we need a trigger?

- Beam spill rate is 5 Hz
- There is 1 neutrino interaction every ~20 beam spills at SBND
- Events are large (~100 MB). Data management and processing isn't trivial.

Only save events when scintillation light is detected during the beam spill*

*plus other trigger streams for calibration, monitoring, background evaluation and modeling

- “Flash” triggered PDS waveform acquisition, roughly 14 flashes in 3 ms window per event

SBND Installation Photos



March 2019

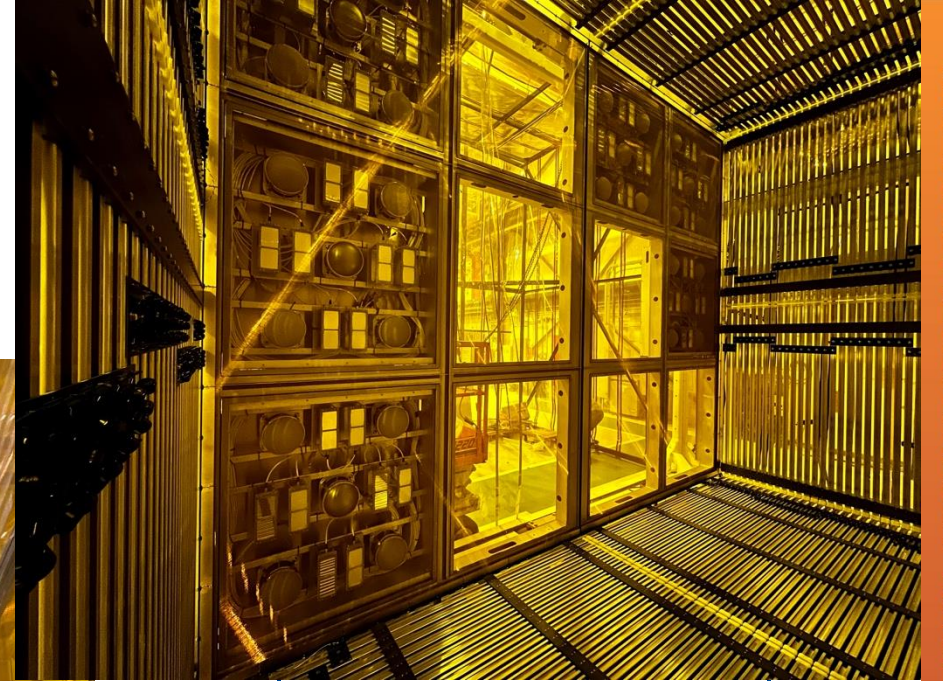


October 2019

January 2022



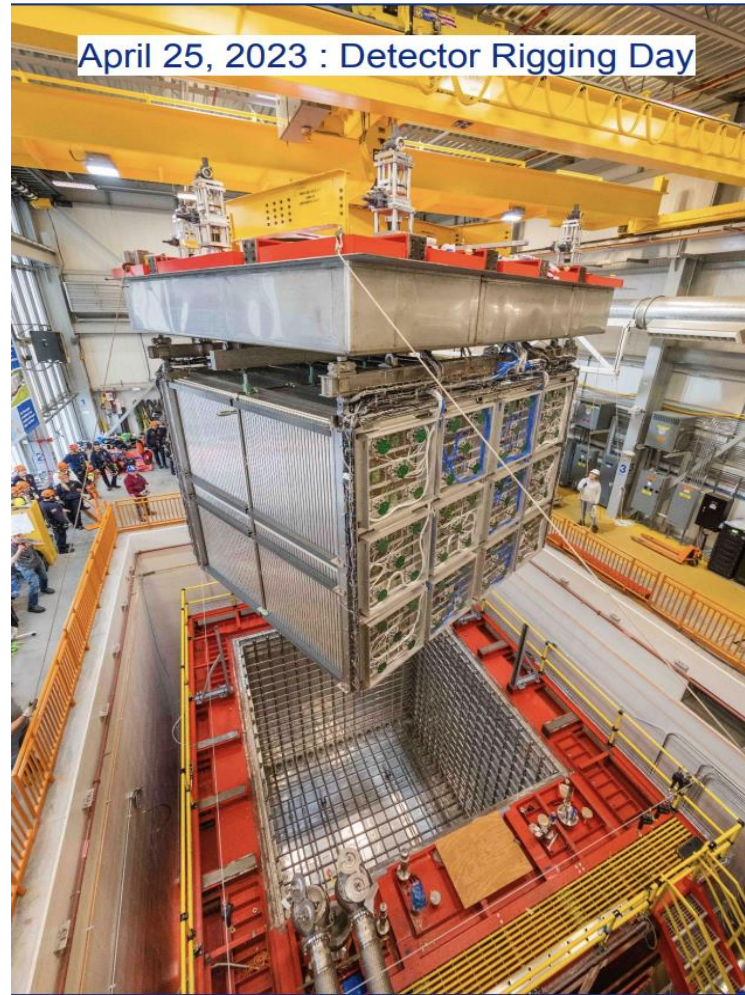
Summer 2022



SBND Installation Photos



December 2022



April 2023



February 2024

SBND Installation Photos

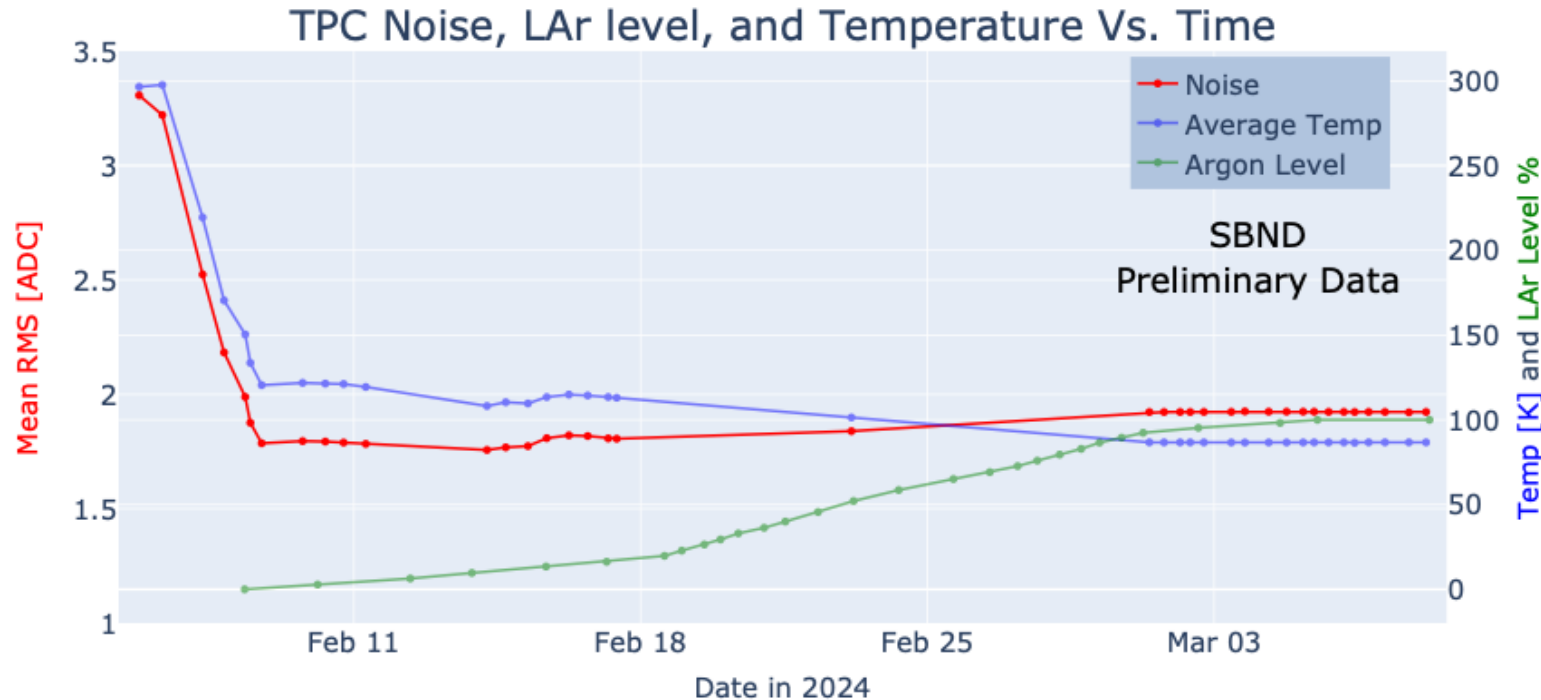


~May 2024



August 2024

TPC performance - noise

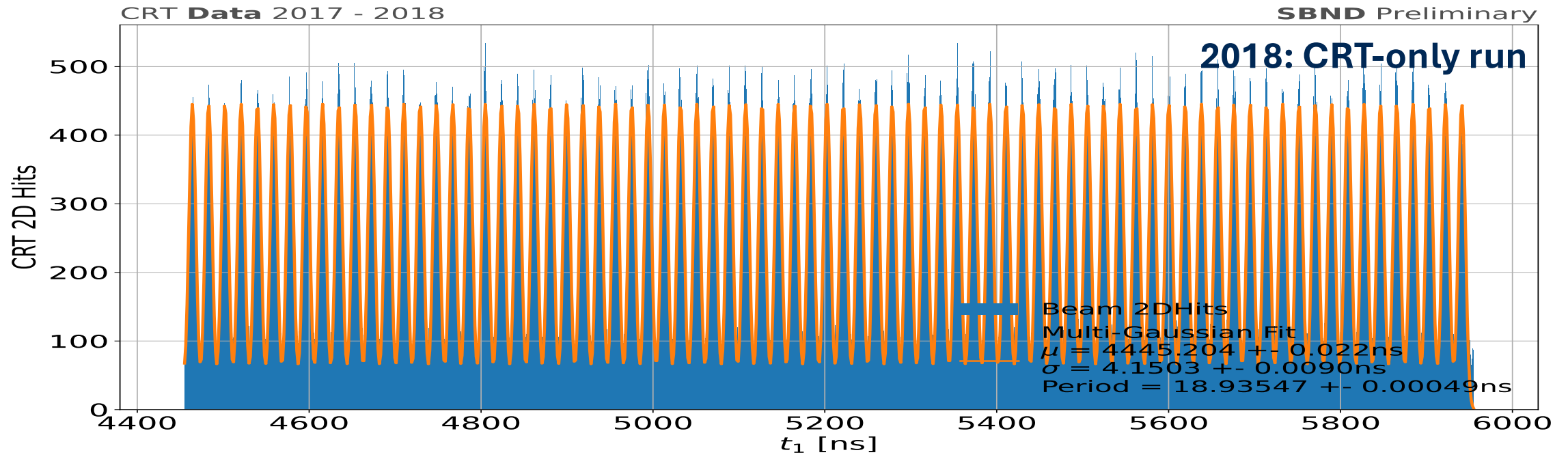


- **Raw waveform RMS from 4 m long collection plane wires**
- **From bench testing: lowest achievable is 380 ENC / 1.96 ADC with no external or coherent noise**

LArTPCs have no electron multiplication for ionization electrons, thus low noise is essential. For 3 mm path length, MIP muon signal is $\sim 15,000$ collectable electrons per wire, 2-4 nA current pulse

- Low noise grounding strategy started with the SBND building design
- Grounding monitored throughout installation and cryo commissioning with active mitigation
- Additional filters added to wire bias voltage connections after protoDUNE-VD lesson learned

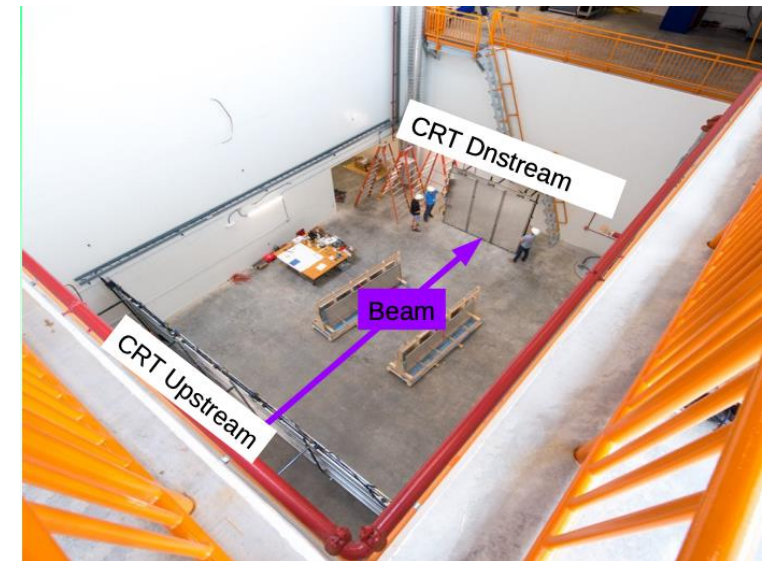
Beam substructure measured by the CRT



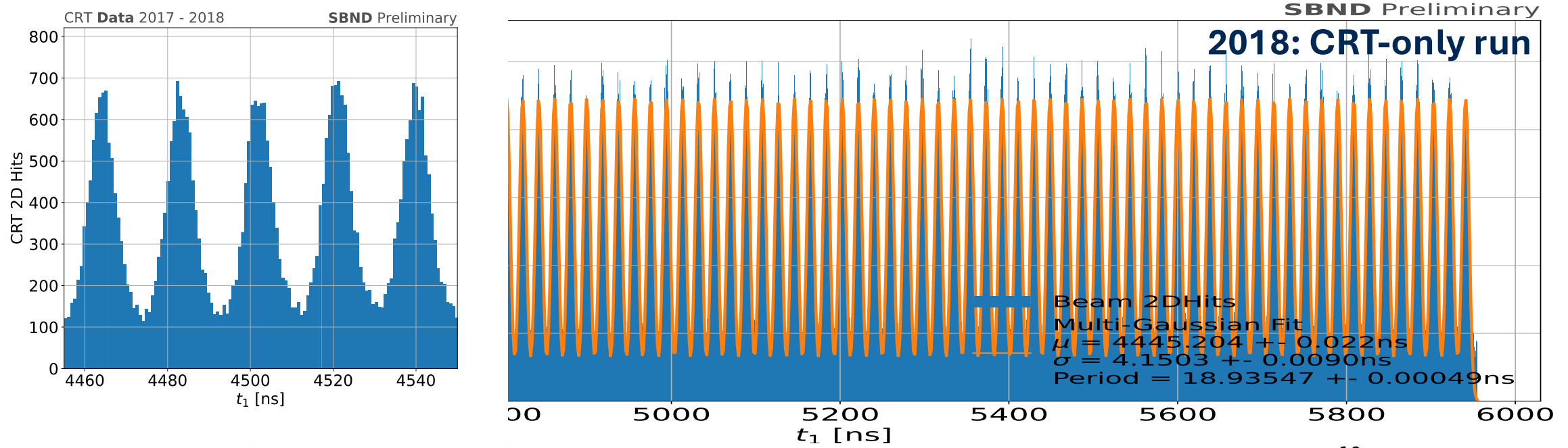
CRT detects muons from neutrino interactions in the dirt upstream of the detector

The beam spill substructure can be seen

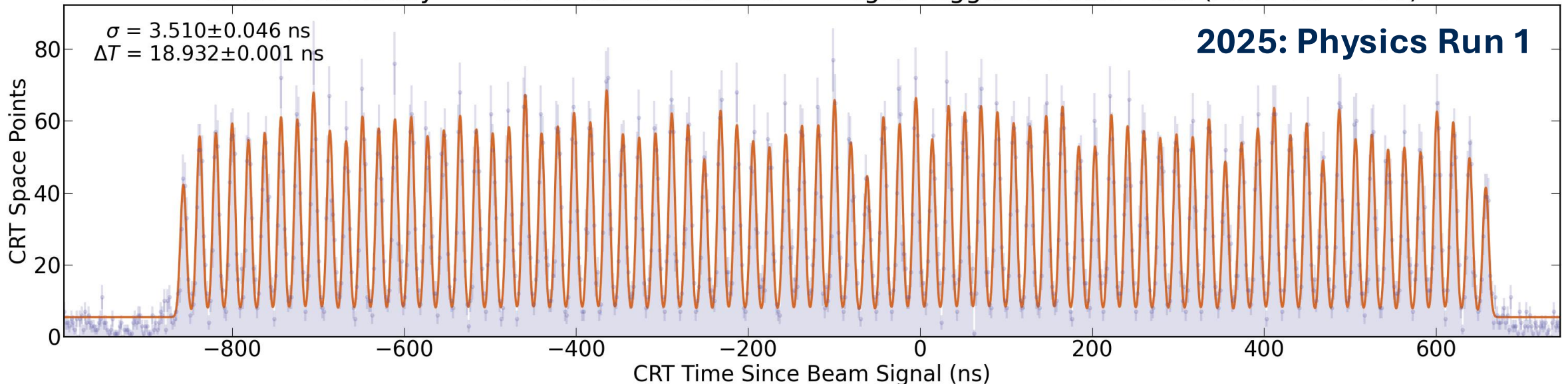
- 81 bunches with 19 ns spacing
- spill duration of $1.6 \mu\text{s}$



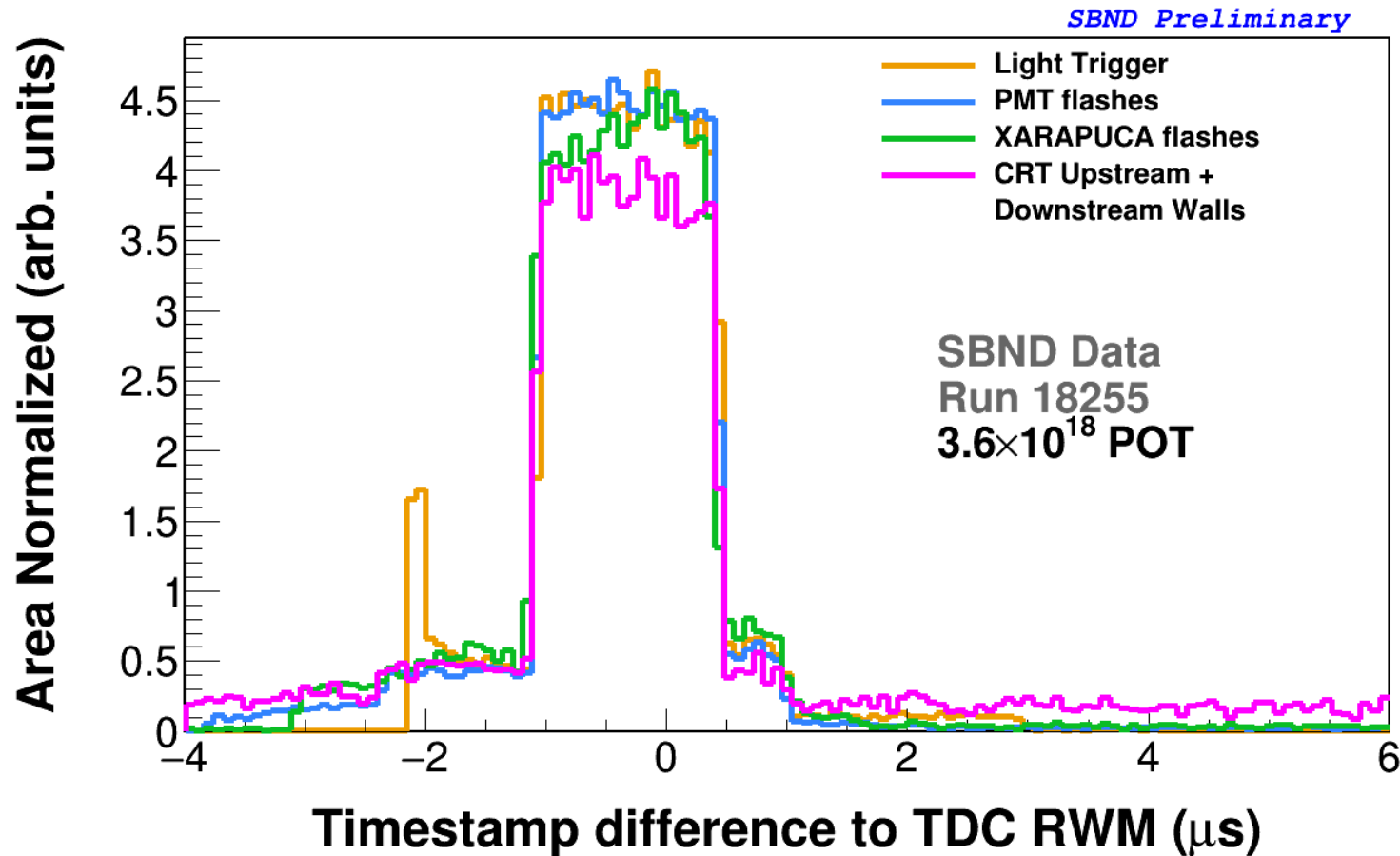
Beam substructure measured by the CRT



SBND Preliminary - CRT Downstream Wall - Light Triggered Beam Data (1.58×10^{19} POT)



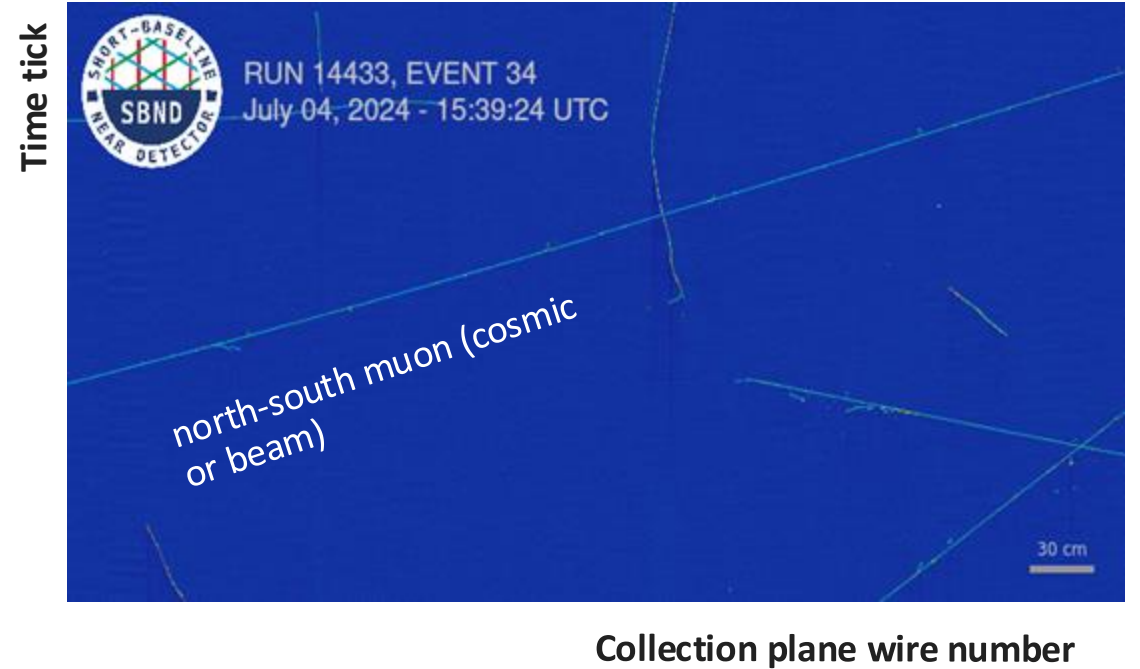
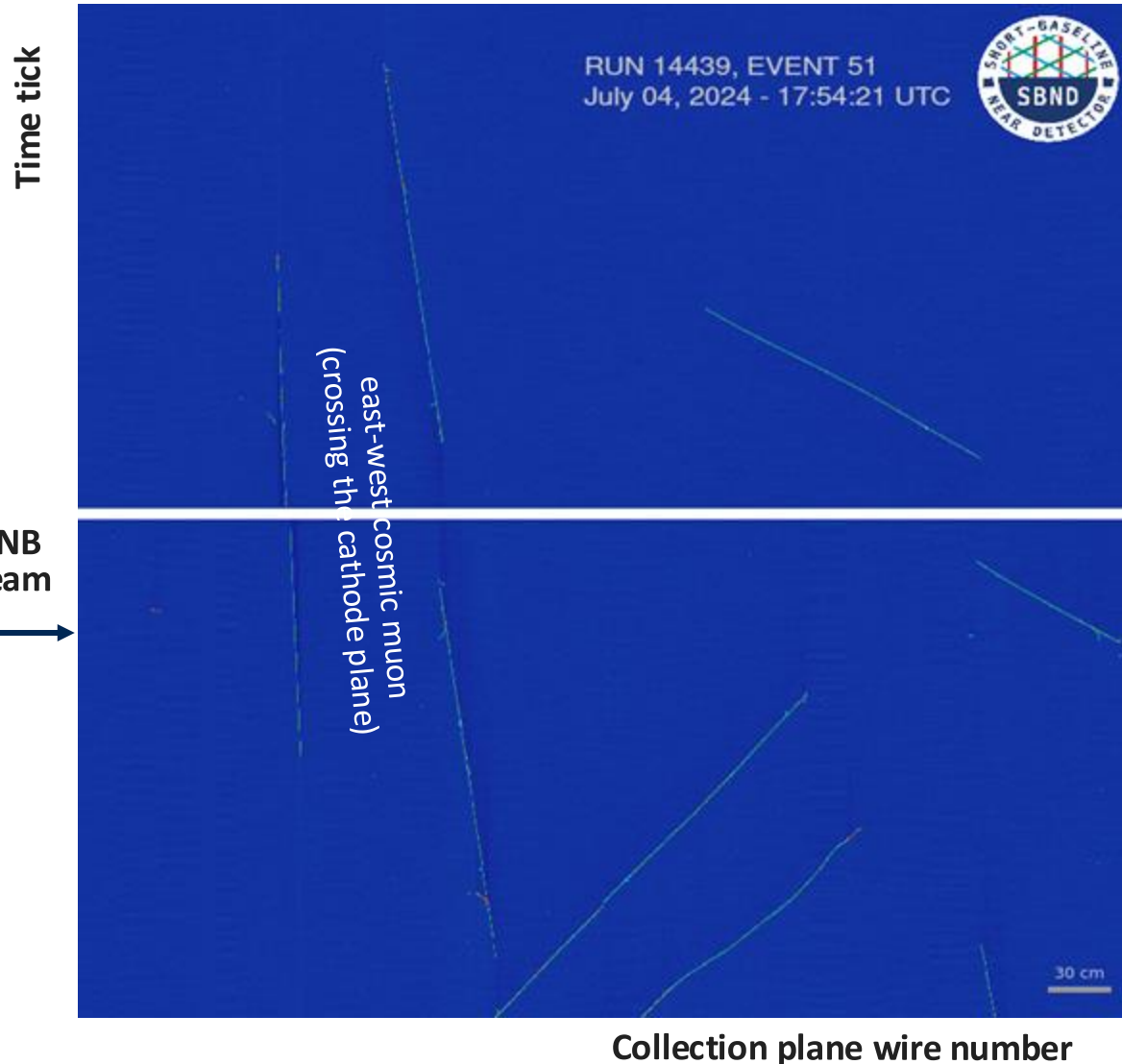
Data synchronization and timing validation



Remember . . . light is fast, charge is slow

- Independent TPC reconstruction of beam spill timing is in progress.
- 1 TPC clock tick = $0.5 \mu\text{s}$

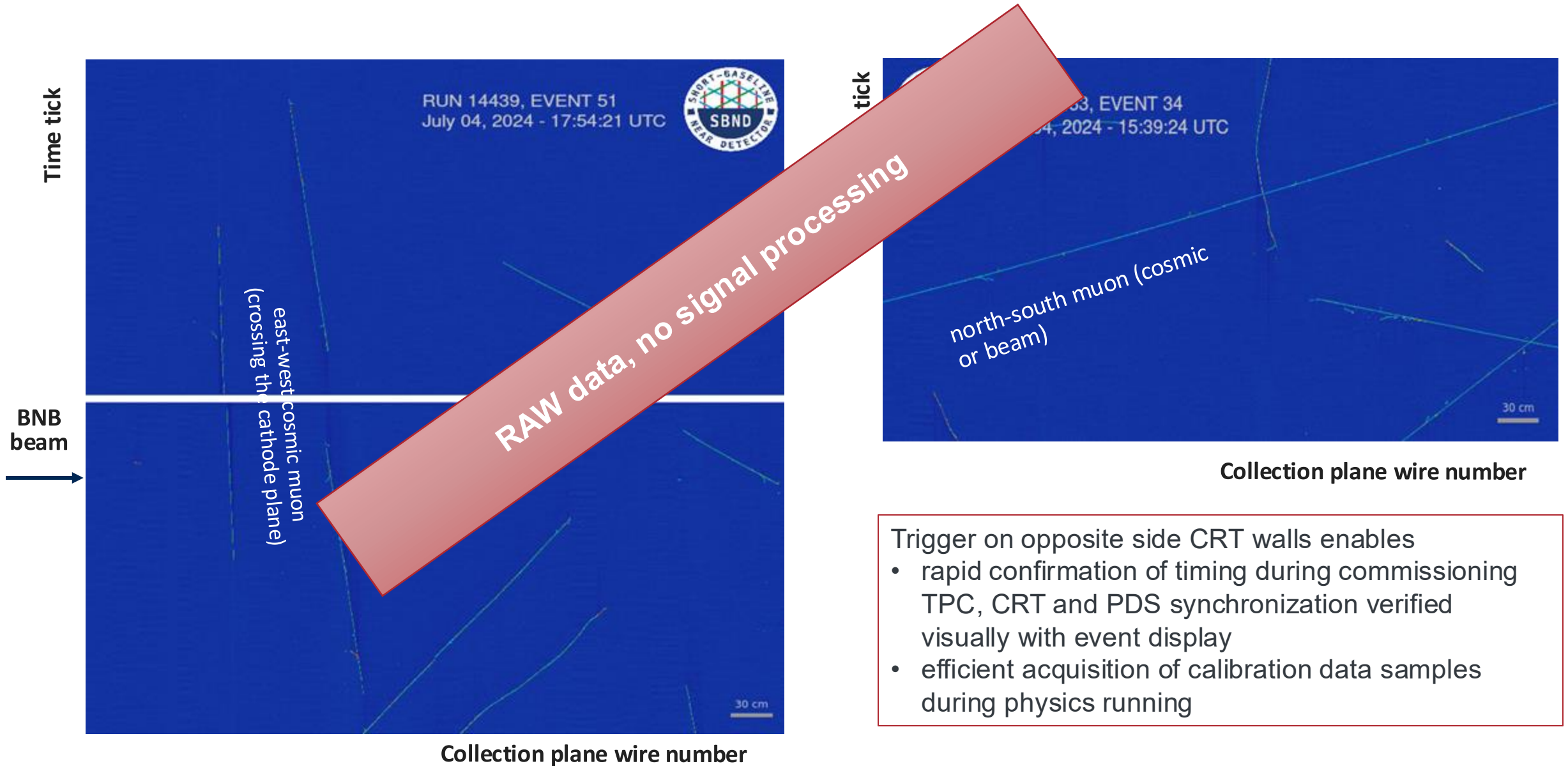
Data synchronization and timing validation



Trigger on opposite side CRT walls enables

- rapid confirmation of timing during commissioning
- TPC, CRT and PDS synchronization verified visually with event display
- efficient acquisition of calibration data samples during physics running

Data synchronization and timing validation

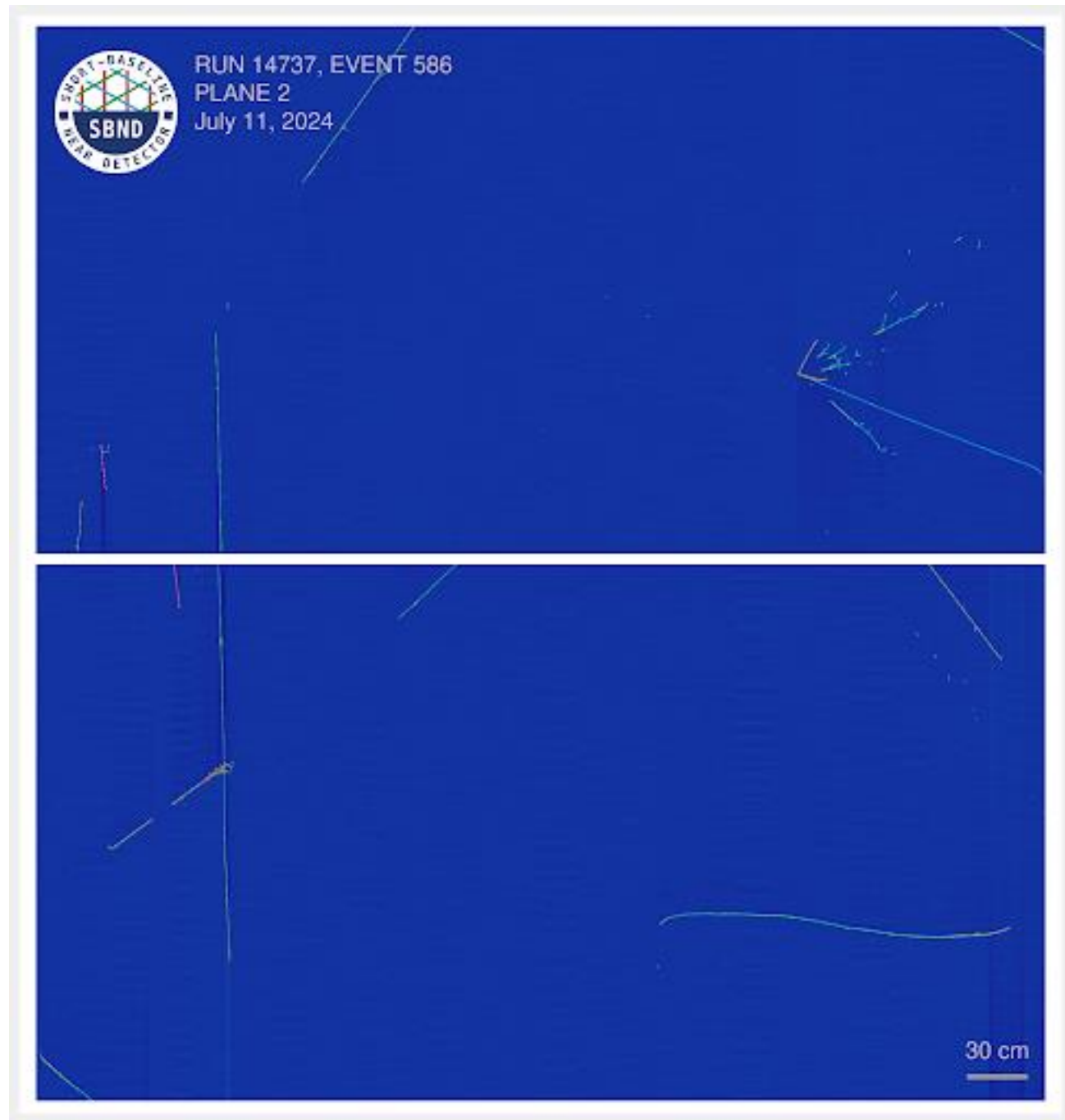


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- rapid confirmation of timing during commissioning
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July 4-11

Initial physics run

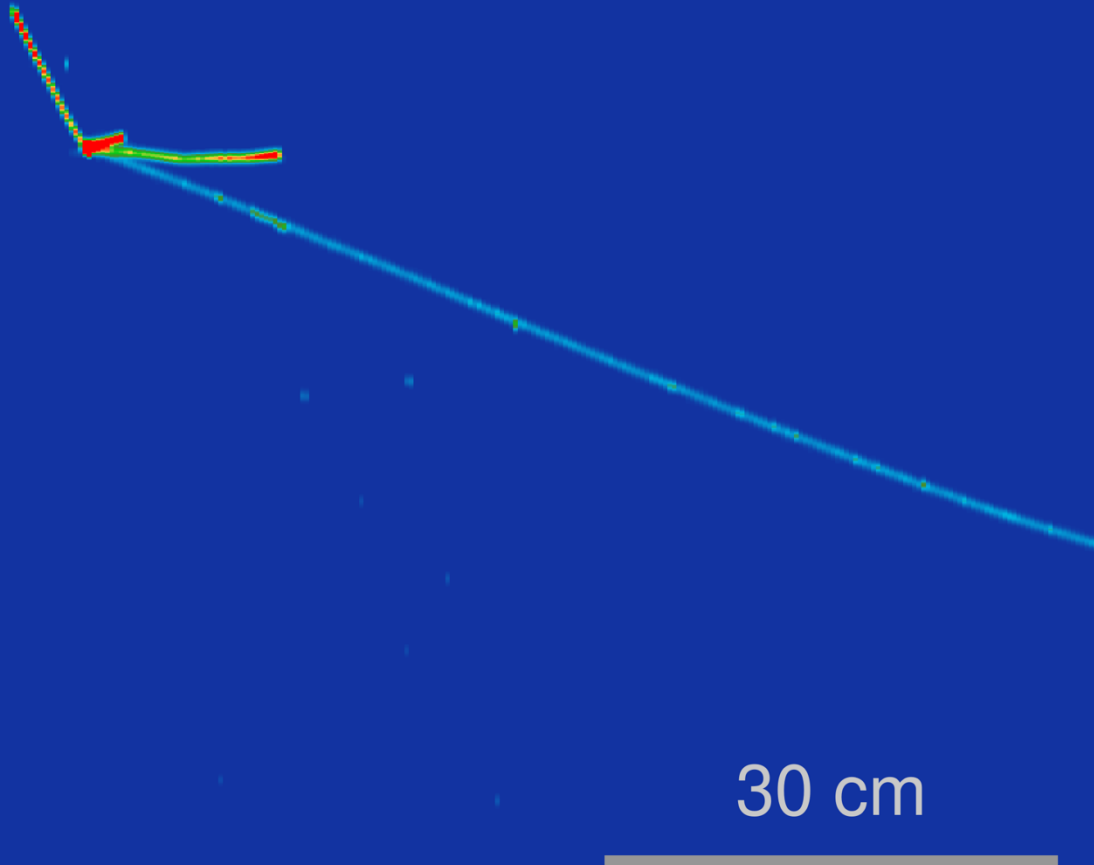
Time tick



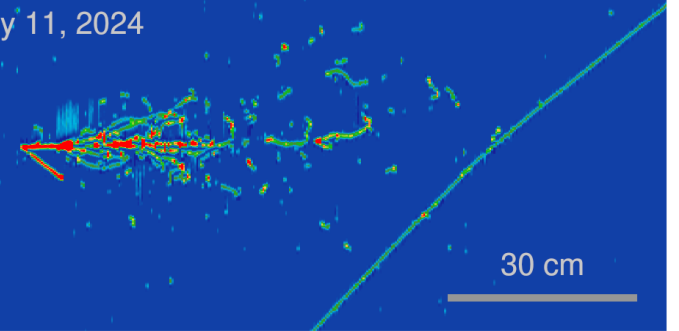
Wire
number



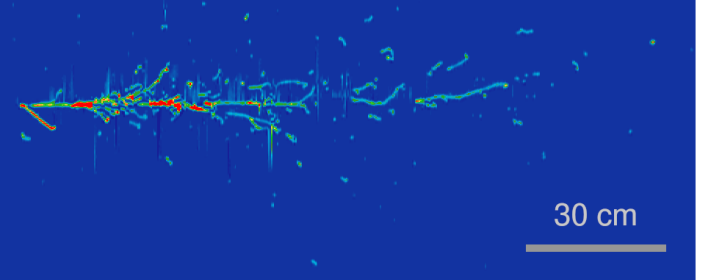
RUN 14737, EVENT 1881
July 11, 2024



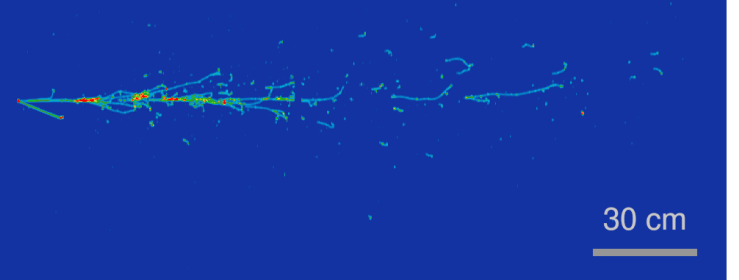
RUN 14729, EVENT 41
PLANE 0
July 11, 2024



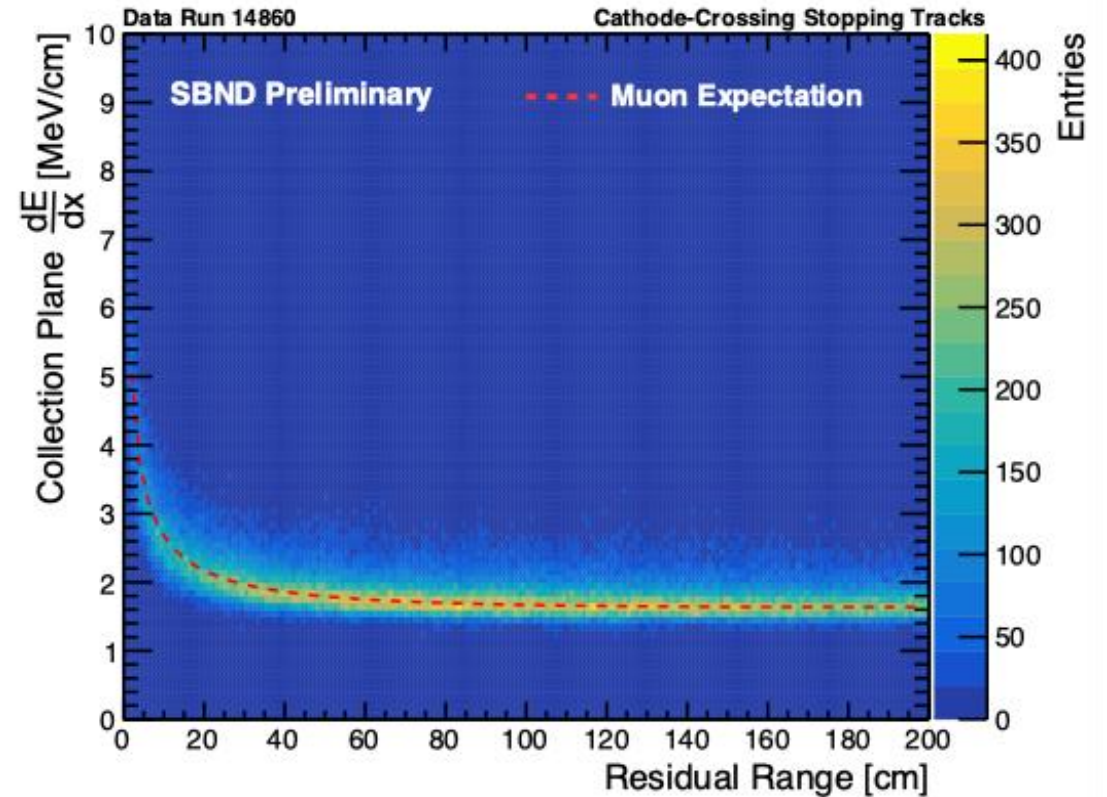
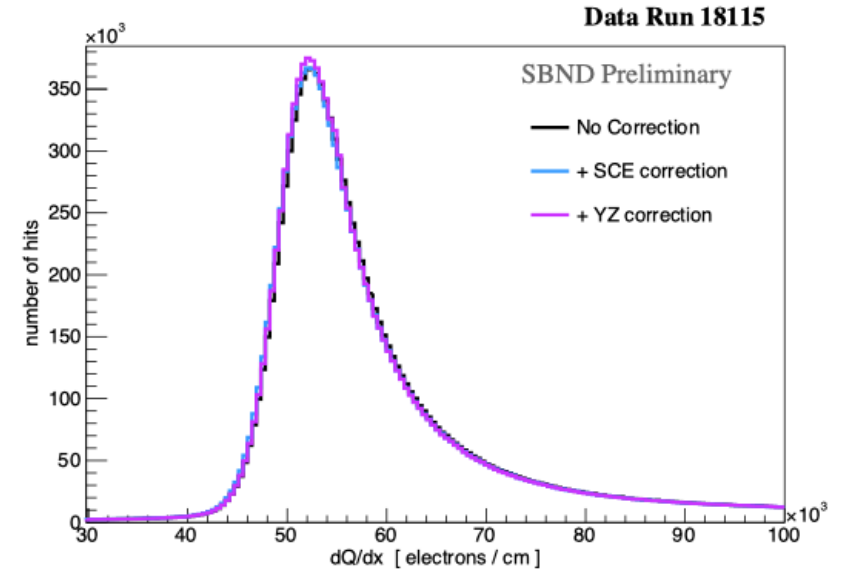
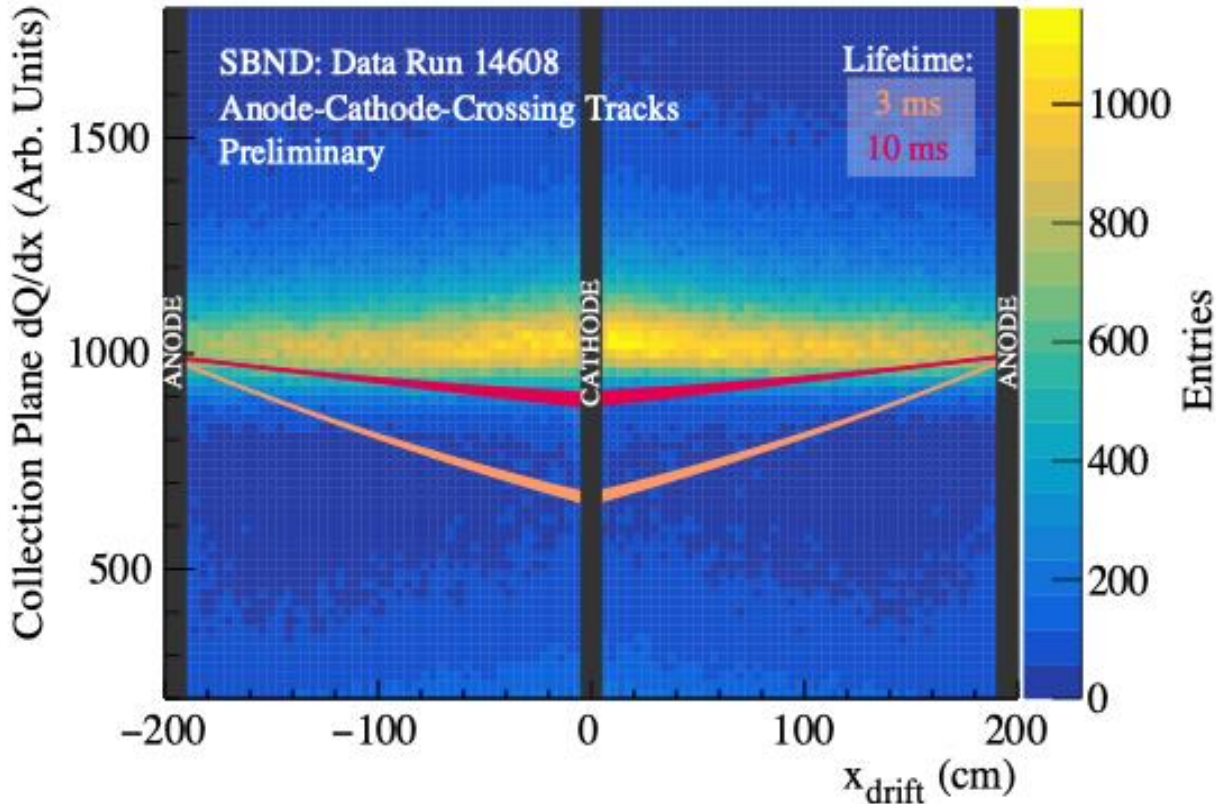
RUN 14729, EVENT 41
PLANE 1
July 11, 2024



RUN 14729, EVENT 41
PLANE 2
July 11, 2024



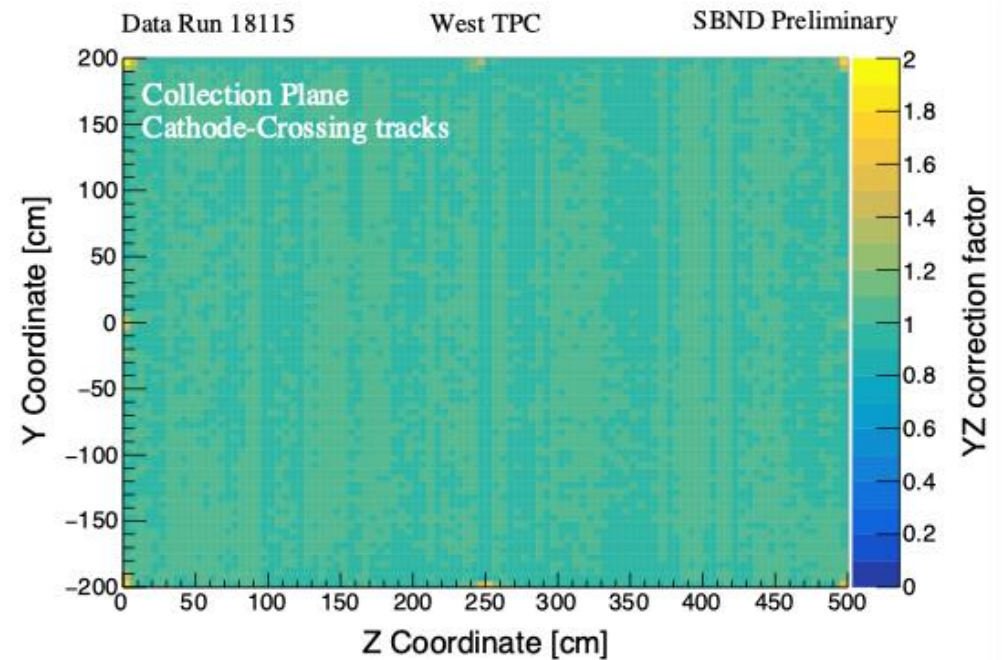
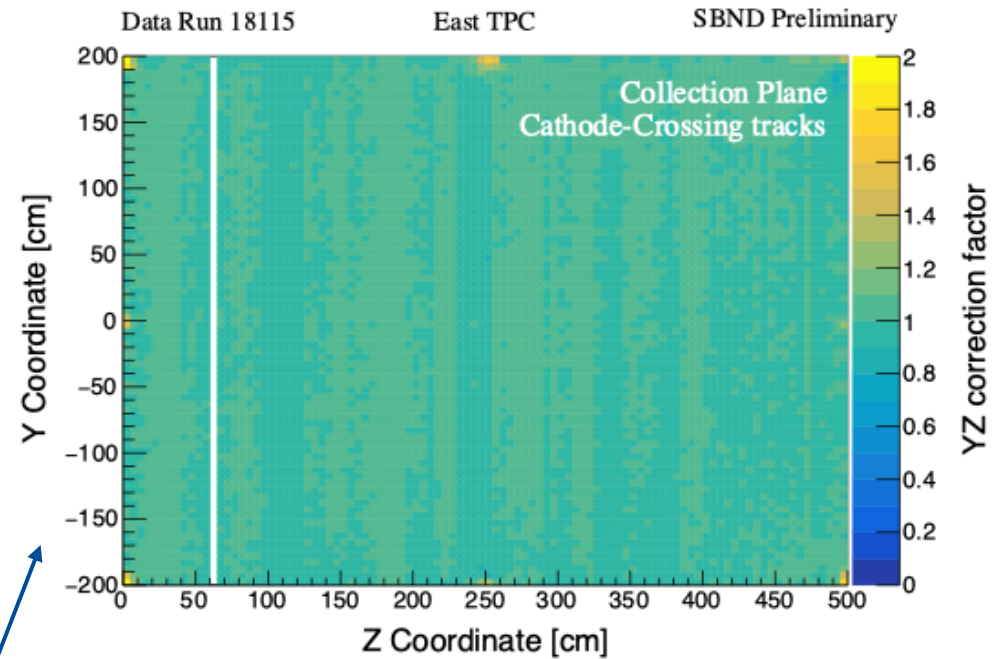
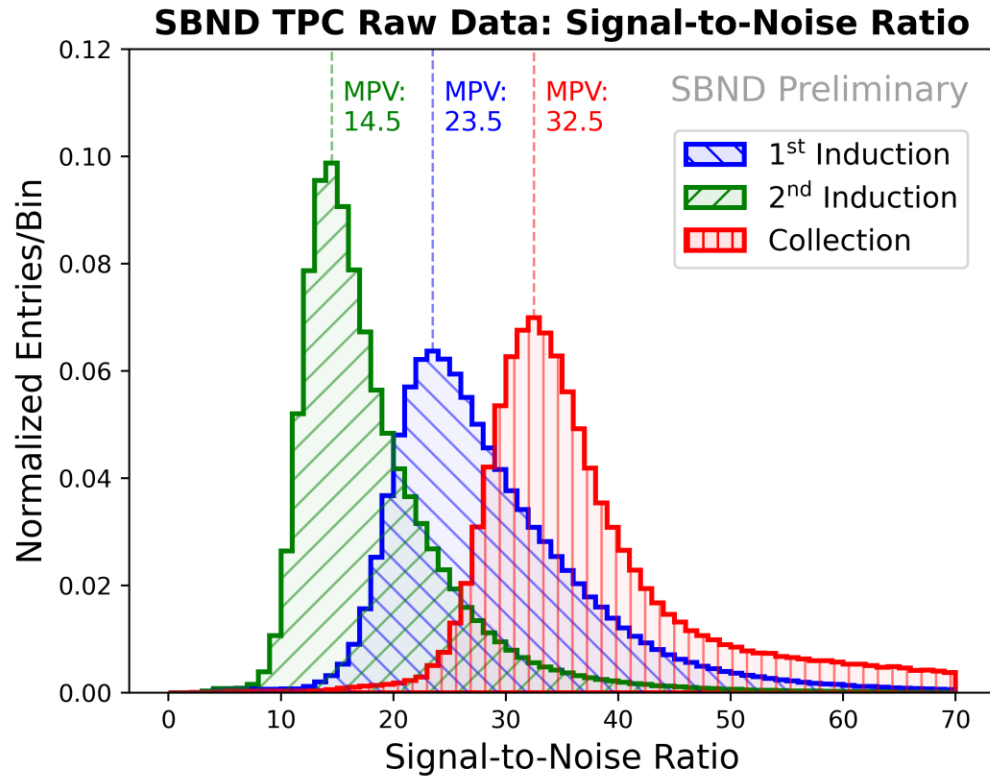
Argon Purity and TPC calorimetry



Electron lifetime (ms)	3	10	25	125
Surviving fraction after 1.25 ms drift	66%	88%	95%	99%

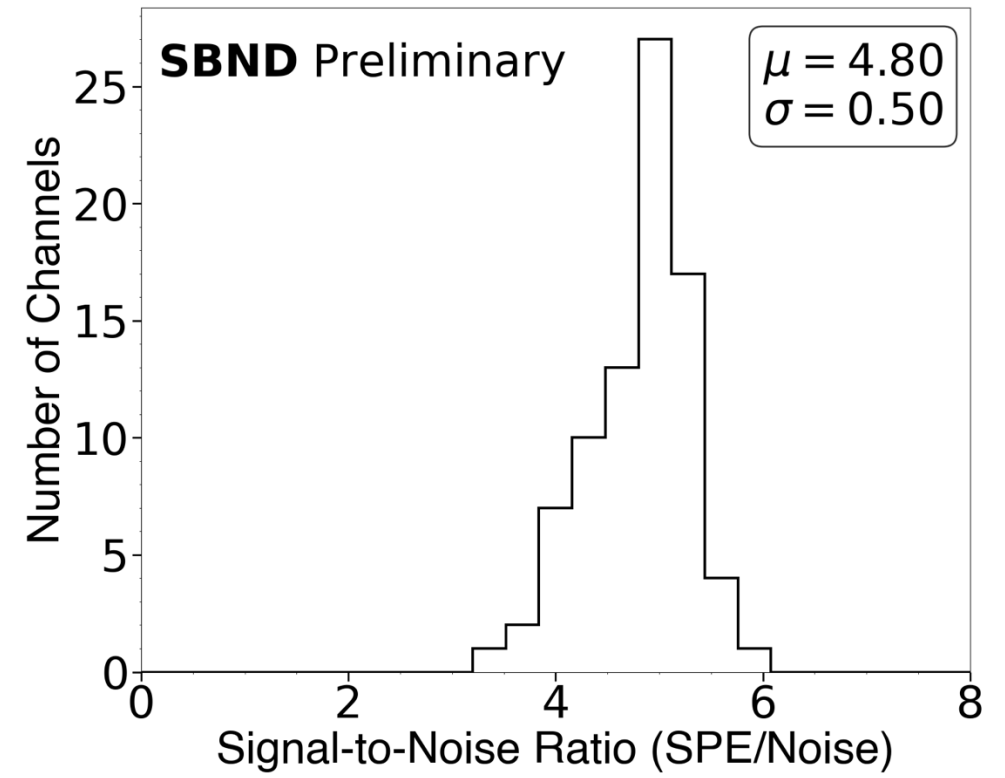
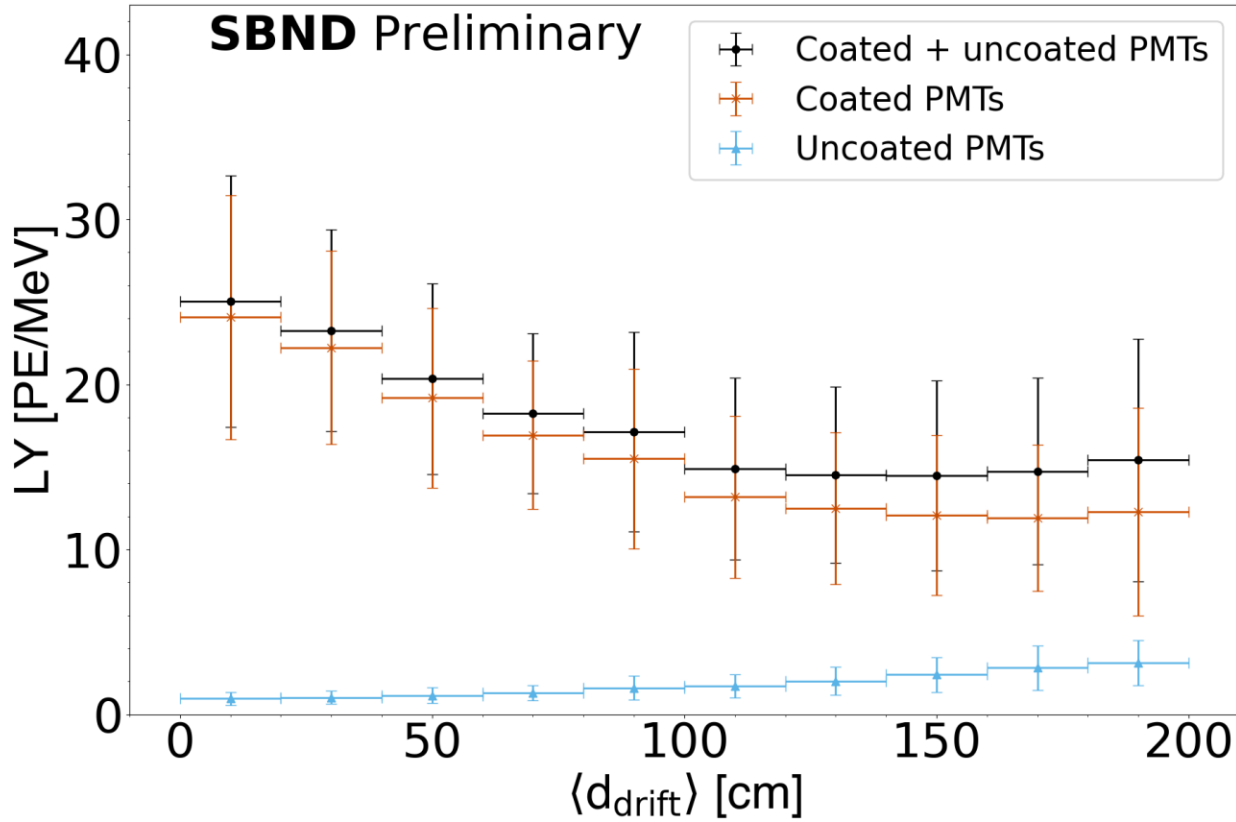
DESIGN REQUIREMENT

TPC performance



Current uniformity of detector response to MIP muons across both wire planes

PMT performance

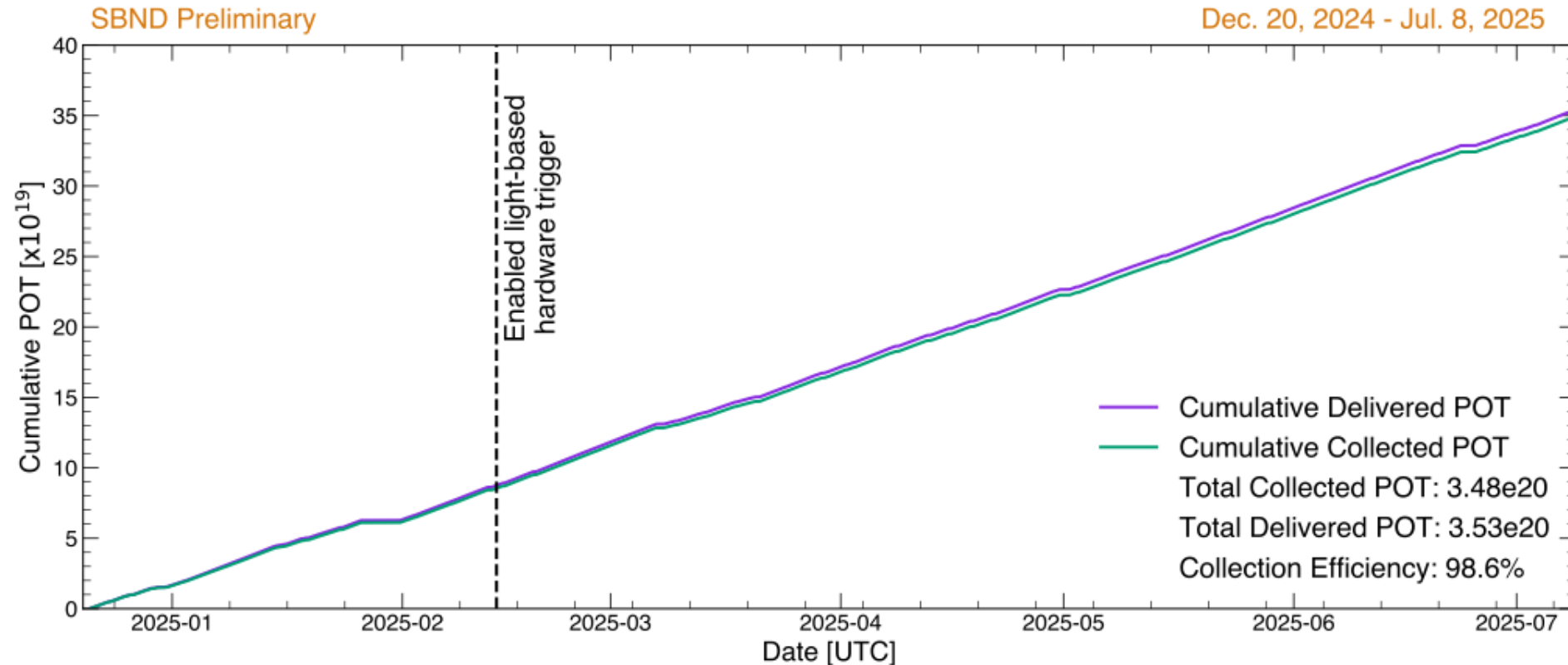


XArapuca readout installation ongoing this summer shutdown



6+ months of BNB data so far . . .

SBND Run 1 Cumulative POT



. . . two more years of running ahead!

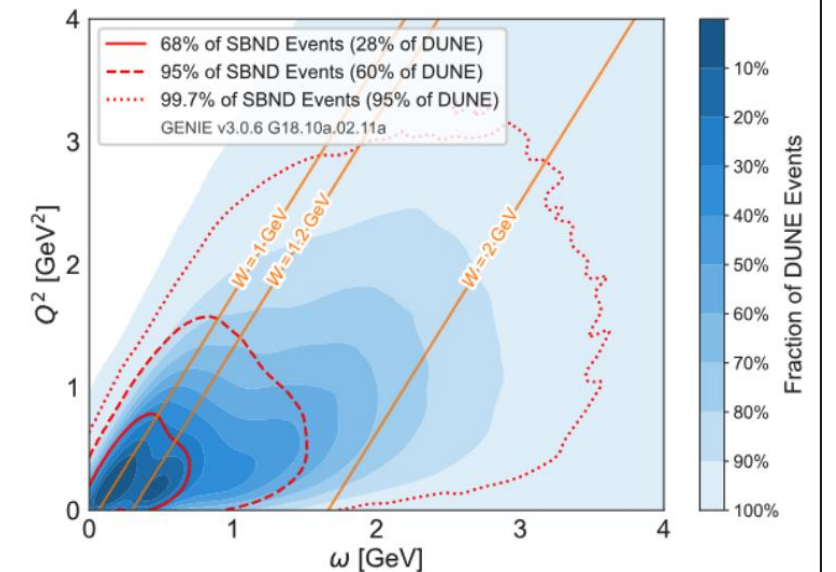
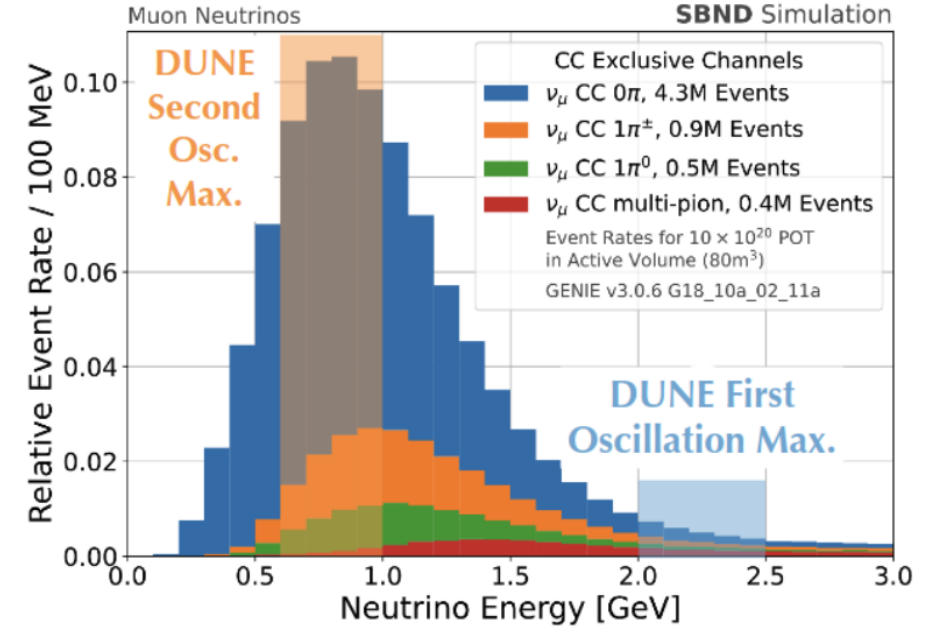
Where is SBND pushing the envelope?

On the road to DUNE

- 3rd iteration of DUNE cryostat design – instrumented with strain gauges and temperature sensors to evaluate performance
- 5th generation of TPC cold electronics
 - Low noise without software mitigations/removal
 - Important for DUNE far detector “triggering”
- Side by side PMTs and XArapucs for detailed studies of performance
- Neutrino-Ar cross section measurements
- PRISM

Near detector challenges

- Tools for addressing modest neutrino pileup in light system and exposure accounting
- Tools to understand beam-related backgrounds



Conclusions

- The high statistics and exceptional performance of the SBND detector will deepen our understanding of neutrino-argon interactions
- A robust search for sterile neutrinos is on the horizon, combining data from all three detectors in the SBN program

Watch the fall conferences for exciting new results from SBND!





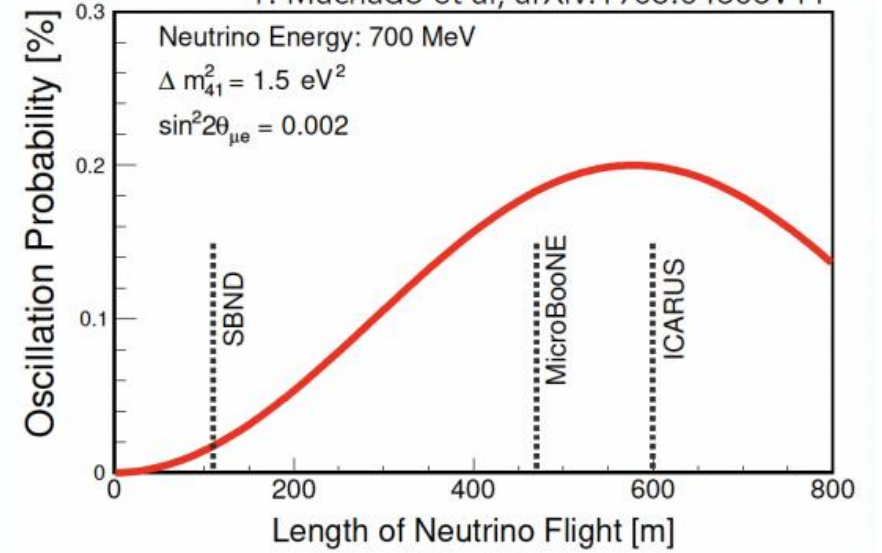
02

Backup Slides

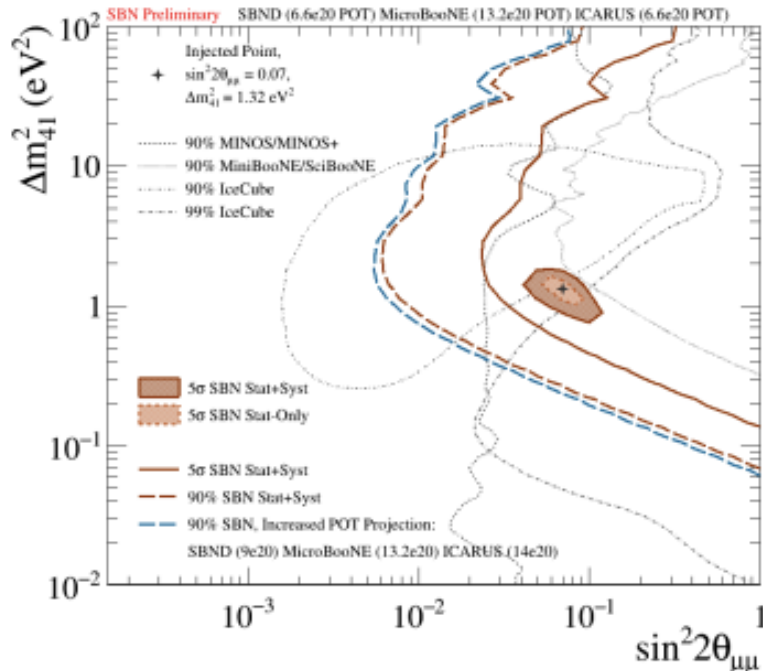
SBN Oscillation Program

The SBN program tests the sterile neutrino hypothesis by covering the parameter regions allowed by past anomalies at 5 significance.

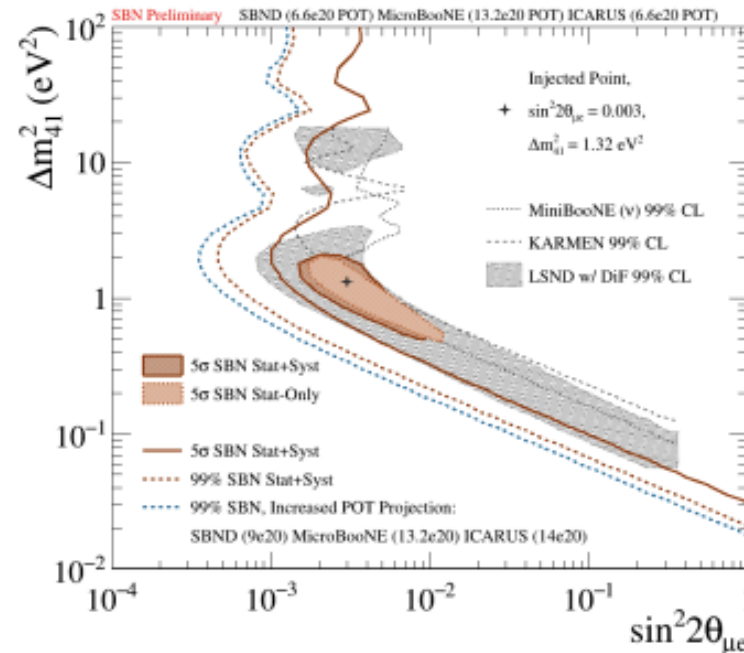
Complementary measurements in different modes are important for interpretation in terms of sterile neutrino oscillation



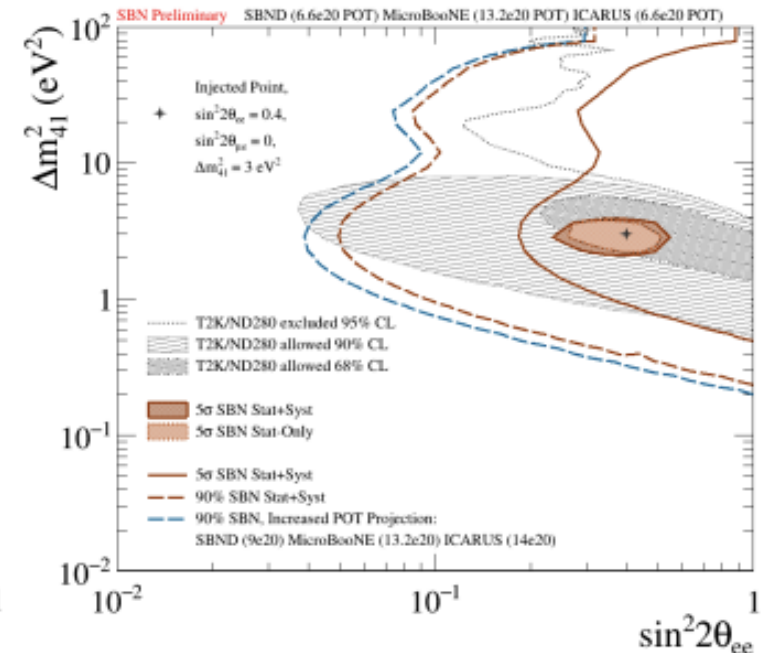
ν_μ disappearance



ν_e appearance



ν_e disappearance



Alternative Explanations

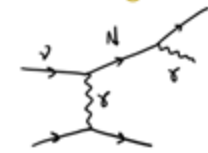
- There are many alternative models to explain the MiniBooNE excess that are not in tension with experimental results (including MicroBooNE)
- SBN can exploit these and other BSM scenarios using the strengths of each detector
 - SBND:** proximity to beam target
 - ICARUS:** observing the off-axis NuMI beam in addition to the BNB

Dark Neutrinos



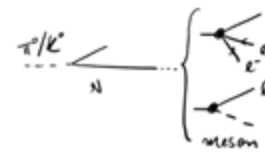
Light Z_D - Bertuzzo, Jana, Machado, Zukanovich PRL 2018
 Bertuzzo, Jana, Machado, Zukanovich PLB 2019
 Argüelles, Hostert, Tsai PRL 2019
 Heavy Z_D - Ballett, Pascoli, Ross-Lonergan PRD 2019
 Ballett, Hostert, Pascoli PRD 2020

Transition Magnetic Moment



Griener PRL 2009
 Coloma, Machado, Soler, Shoemaker PRL 2017
 Atkinson et al 2021
 Vergani et al PRD 2021

Heavy Neutral Leptons



Long list, ex.
 Ballett, Pascoli, Ross-Lonergan JHEP 2017
 Kelly, Machado PRD 2021

Higgs Portal Scalars



Pat Wilczek 2006
 Batell, Berger, Ismail PRD 2019

Courtesy of P. Machado

not an exhaustive list

Cathode plane

Drift

Three anode wire planes

BNB $CC \nu_e$ simulation

18 cm

Investigating such models uses the unique capabilities of the LArTPC technology, with *high track and shower kinematic resolution, very good particle ID, calorimetric information on electro- magnetic & hadronic activity*

Why do we need a trigger? (the beam rate is only 5 Hz after all)

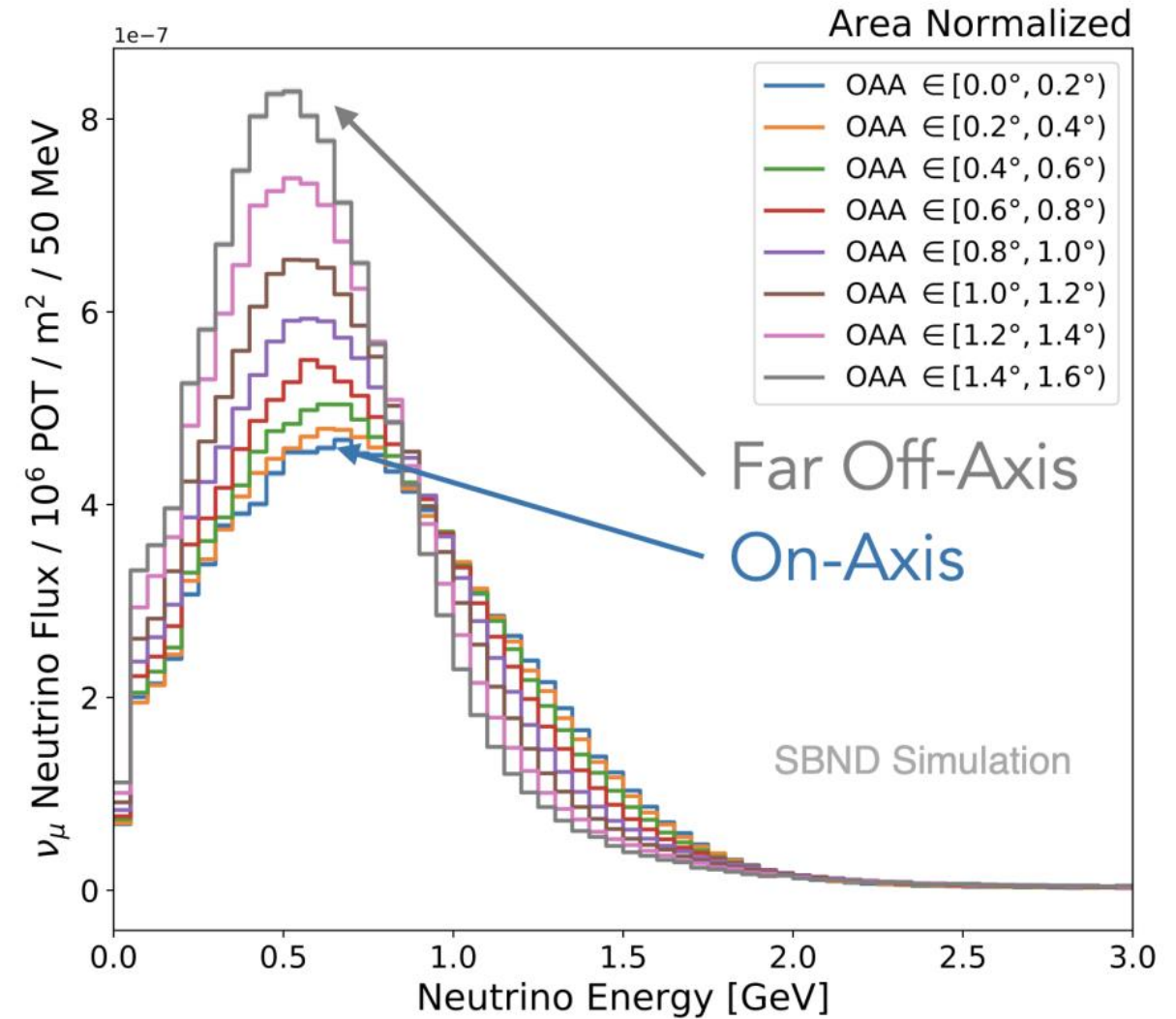
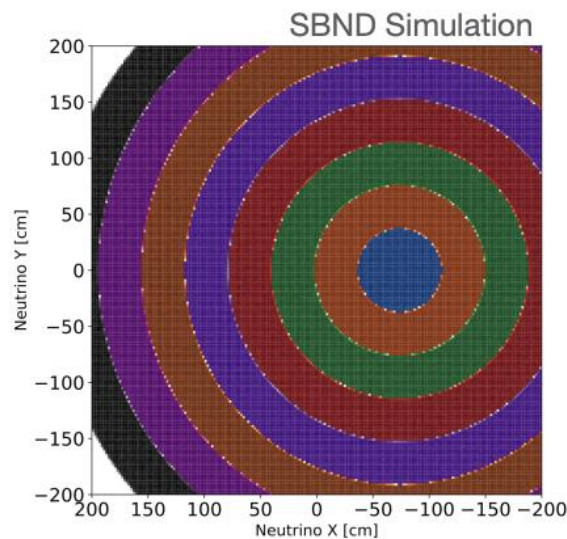
- There is 1 neutrino interaction every ~ 20 beam spills at SBND
- Events are large, processing the data offline isn't trivial.

Only save events when scintillation light is detected during the beam spill

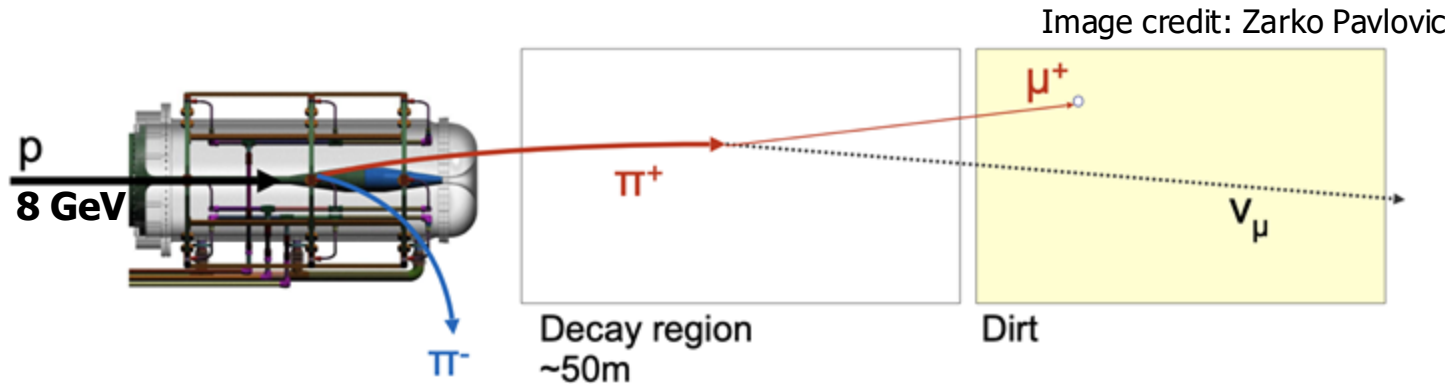
Aside on why events are so large and what we can do about it

- LArTPCs are slow, the detailed 3D image is a lot of numbers. SBND drift time is 1.25 ms and the TPC waveforms are sampled at 2 MHz
- Precision calorimetry without electron multiplication requires signal processing on raw waveforms.
 - with 3mm wire spacing, MIP muons deposit only 25,000 collectable electrons per wire in 1-2 us \rightarrow $O(1)$ nA signals 😬
- New tools such as real time signal processing/ROI finding are in development across the community . . . exciting!
- Today's pixel (instead of wires) anode planes integrate hit charge on the chip . . . game changing!

- Off-axis angle directly corresponds to the neutrino interaction vertex position
- The flux spectrum evolves as a function of the off-axis angle
 - Further off-axis fluxes peak lower and tighter
- Allows SBND to leverage PRISM concept



The Booster Neutrino Beam @ SBND



Beam Composition :

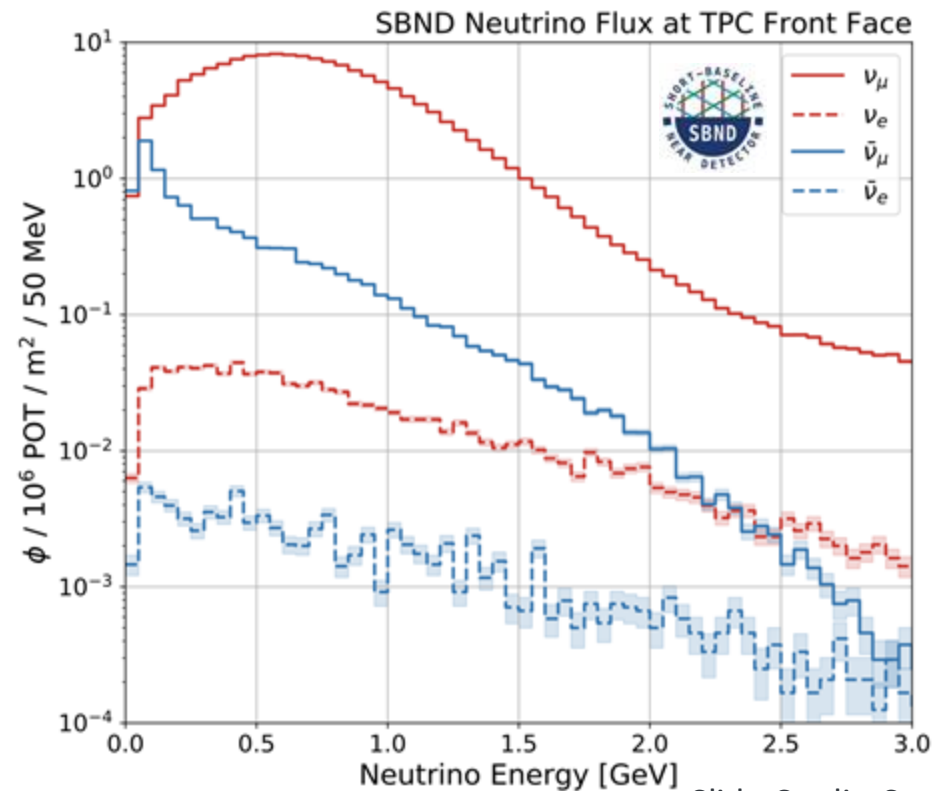
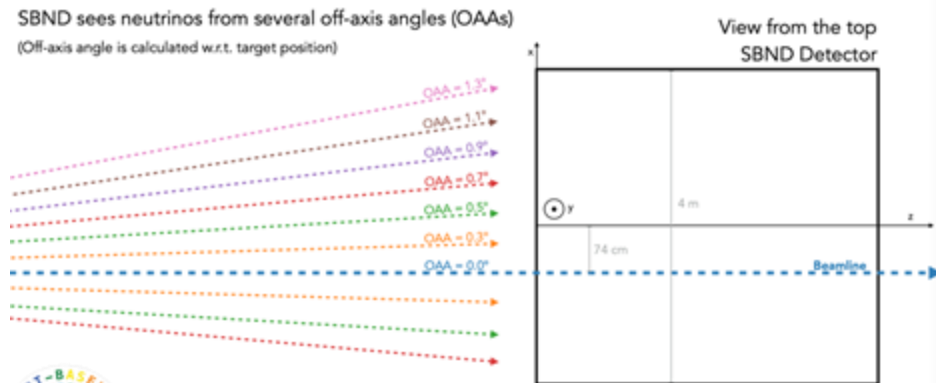
$$\bar{\nu}_\mu = 93.6\%$$

$$\nu_\mu = 5.9\%$$

$$\nu_e + \bar{\nu}_e = 0.5\%$$

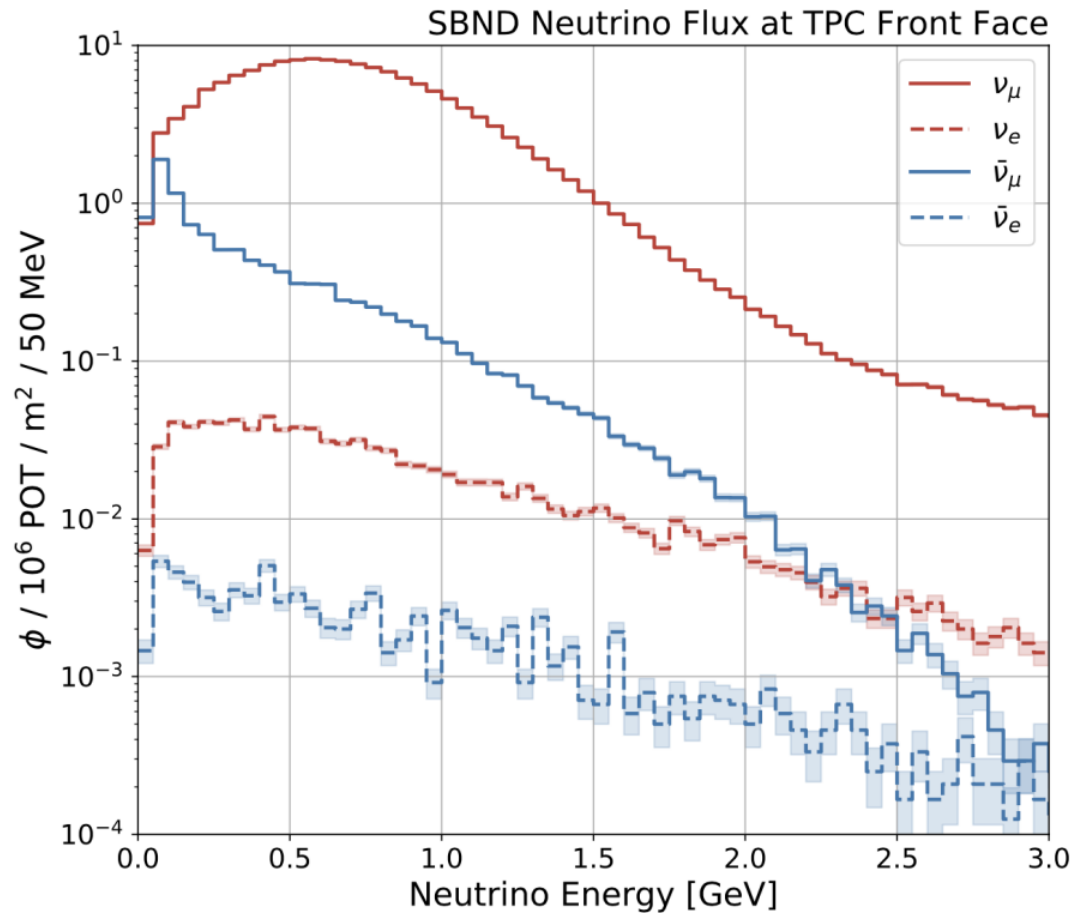
- $\langle E_\nu \rangle \sim 800$ MeV.
- Projected to take **10-18 X 10²⁰ POT** of data in total => large statistics on Argon.
- Close to the target + slightly off-axis => SBND can sample **off-axis fluxes** (hear more about this in the next talk by Lauren Yates).

SBND sees neutrinos from several off-axis angles (OAAs)
(Off-axis angle is calculated w.r.t. target position)



Slide Credit: Supraja B

Booster Neutrino Beam Flux at SBND

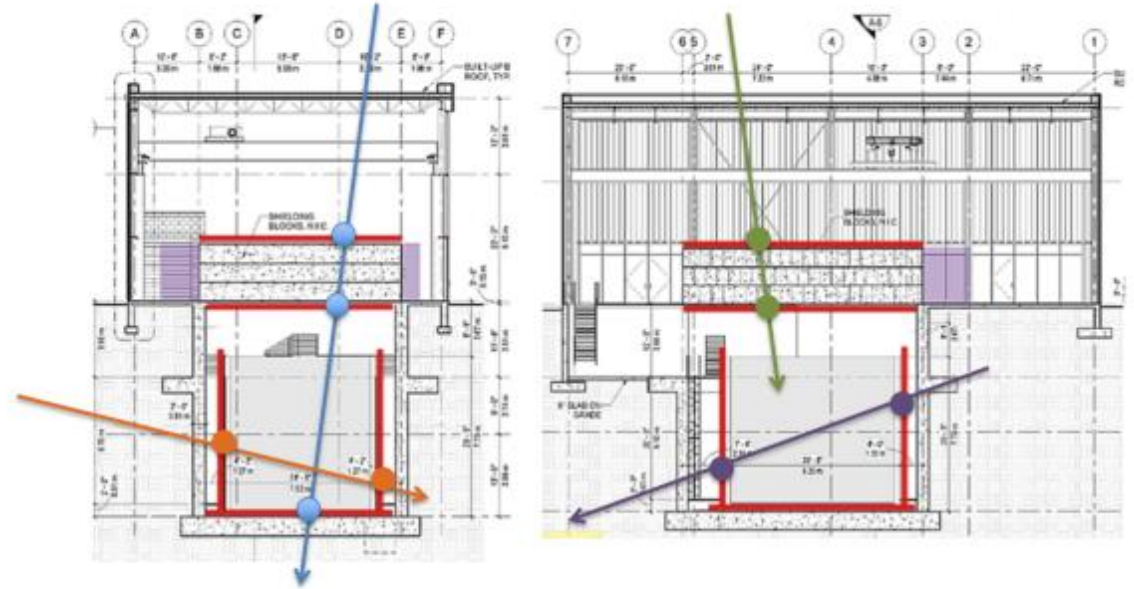
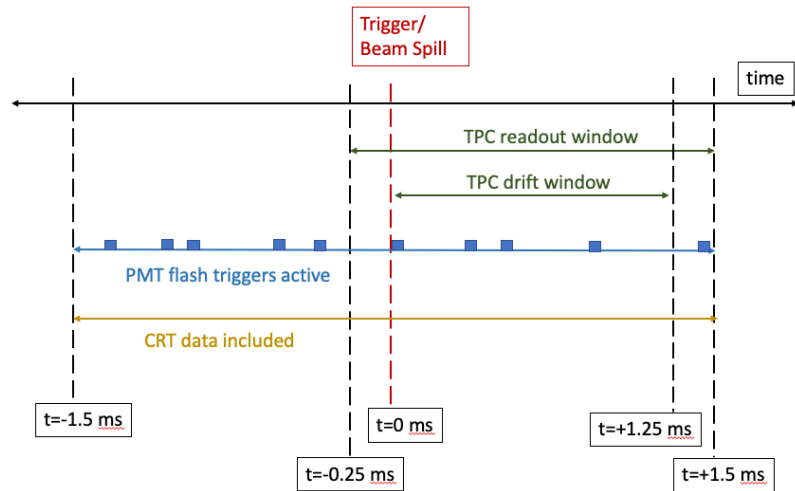


- The primary beam of interest at SBND is the Booster Neutrino Beam (BNB)
- The mean energy for muon neutrinos is about 0.8 GeV
- Beam composition by neutrino flavor:
 - 93.6% ν_μ
 - 5.9% $\bar{\nu}_\mu$
 - 0.5% $\nu_e + \bar{\nu}_e$
- Plan to collect data corresponding to 10^{20} – 18×10^{20} protons on target (POT) over the course of a 3–4 year run

CRT based triggers

Long muon tracks are a critical commissioning tool and the basis of many calibration measurements.

The ability to trigger on the presence of muon tracks means that pure calibration samples can be acquired quickly and/or written to separate data streams. The calibration process is considerably faster when processing calibration data is decoupled from processing physics data

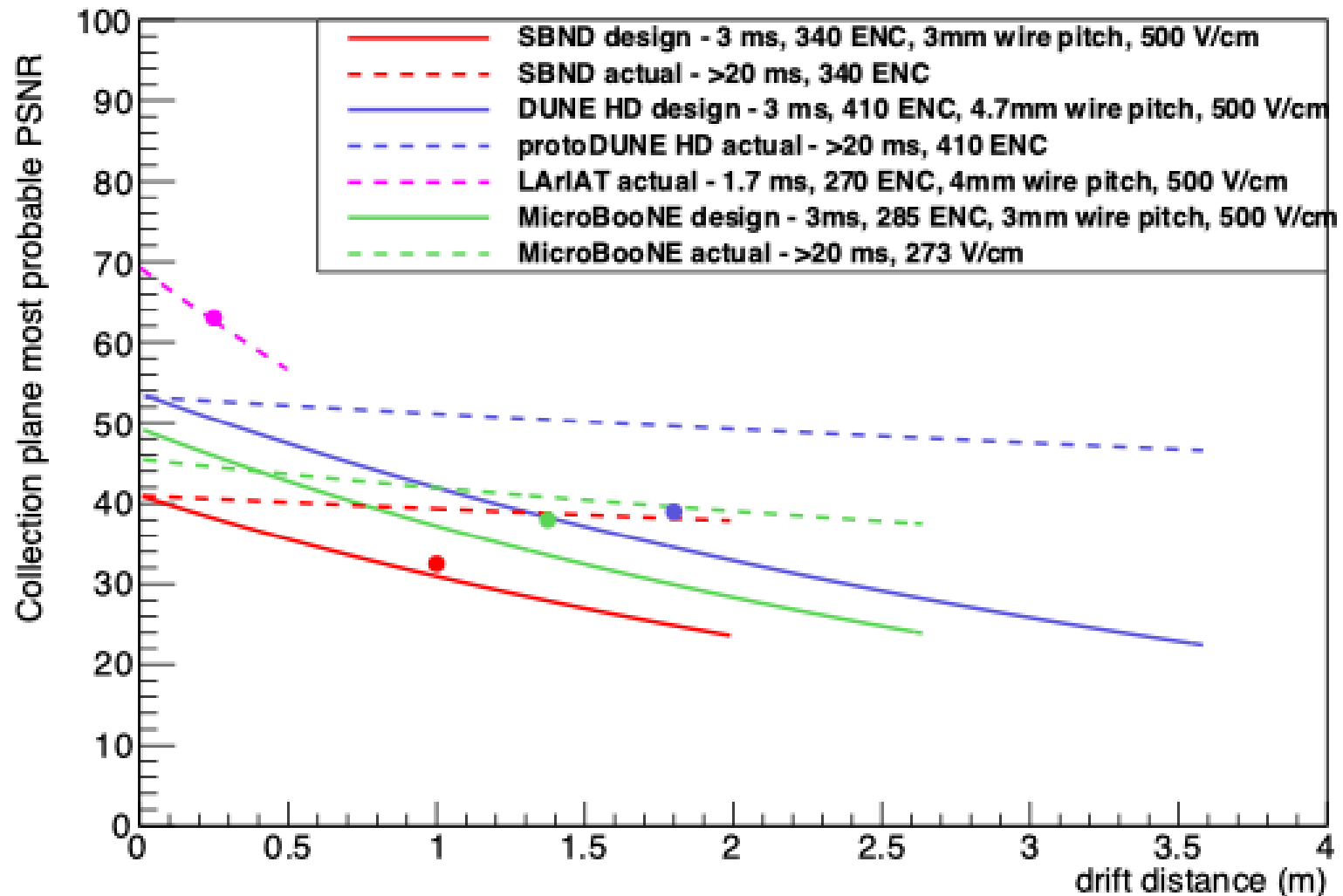


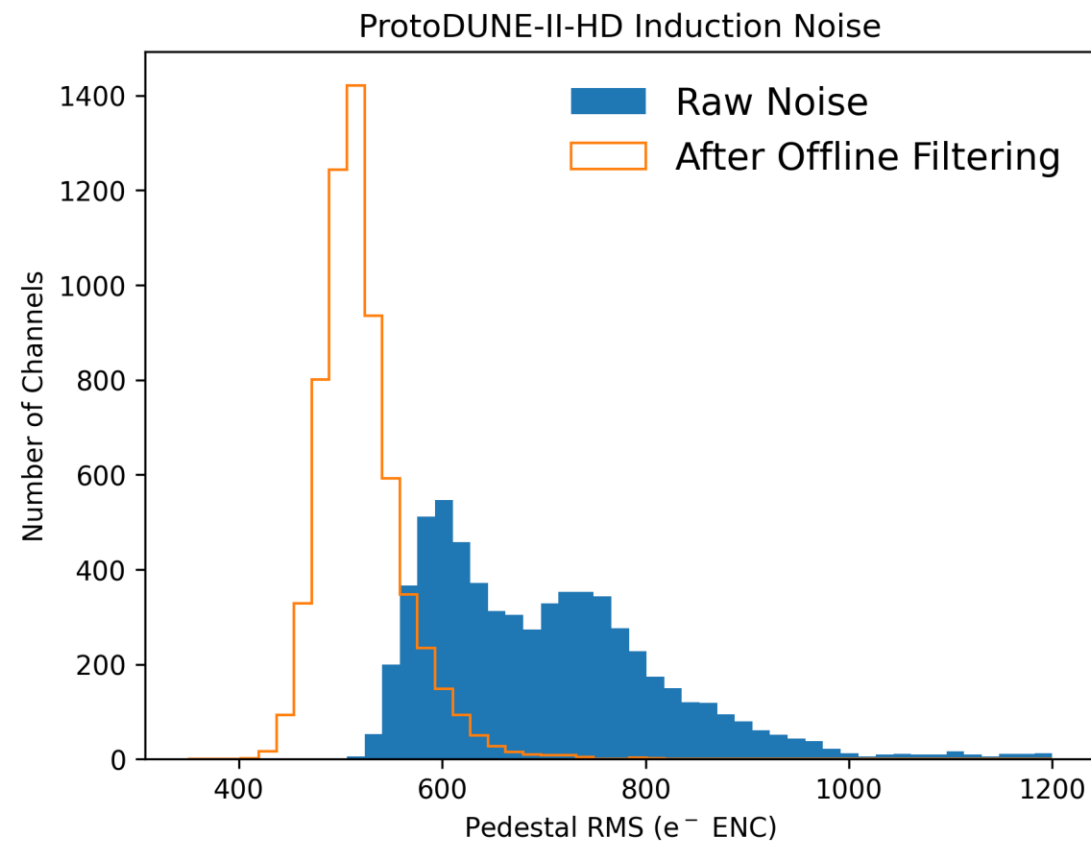
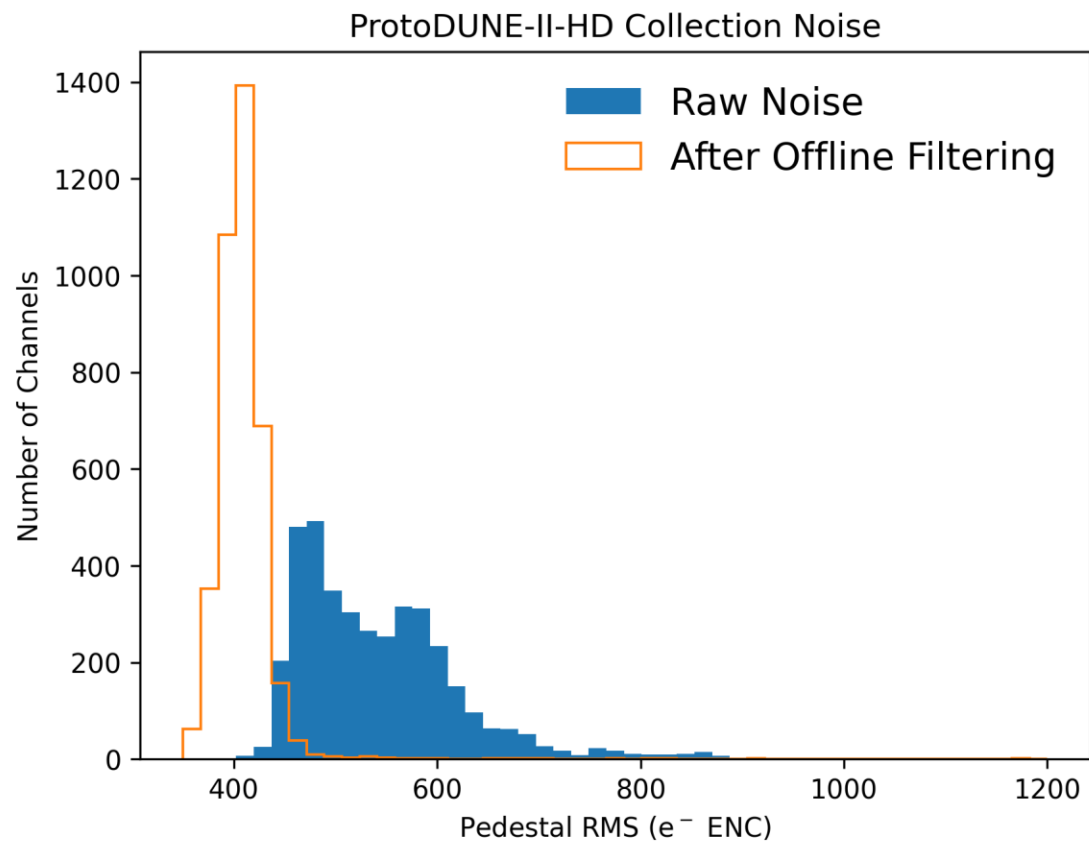
- CRT-A: Vertical through-going
- CRT-B: Stopping muons (Michel sample)
- CRT-C: Horiz. through-going Anode-Cathode crossing muons
- CRT-D: Horiz. through-going “parallel” muons

All triggers shown here also require high PMT multiplicity to ensure the muon track enters the active TPC volume

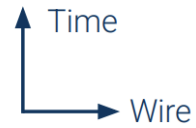
For dedicated runs, a subset of the CRT strips can be activated, restricting the trigger to certain track angles/topologies

Evolution of LArTPC cold electronics





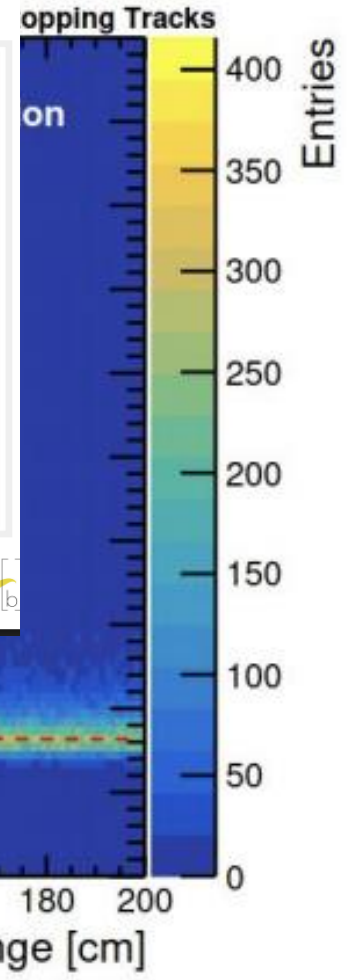
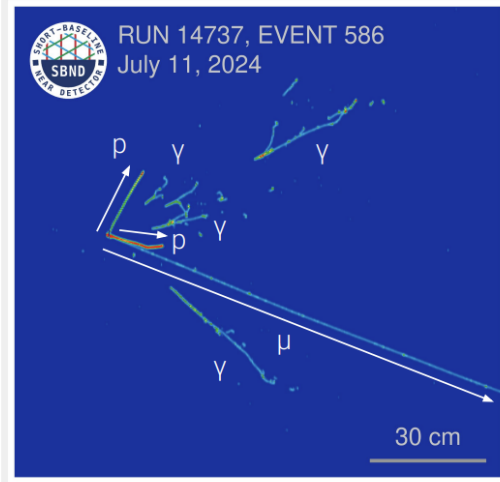
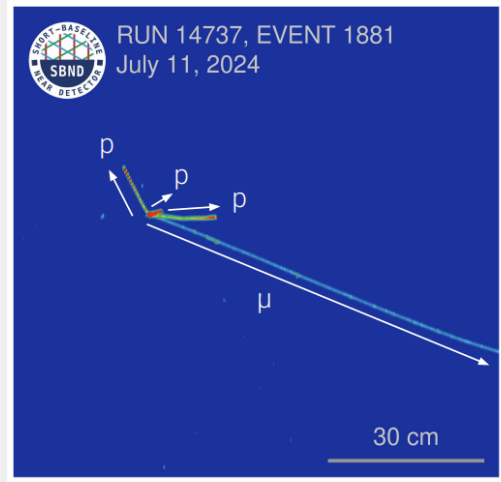
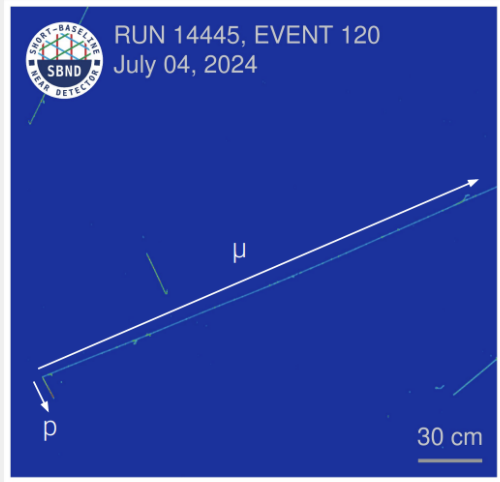
Captured First Neutrinos in July 2024



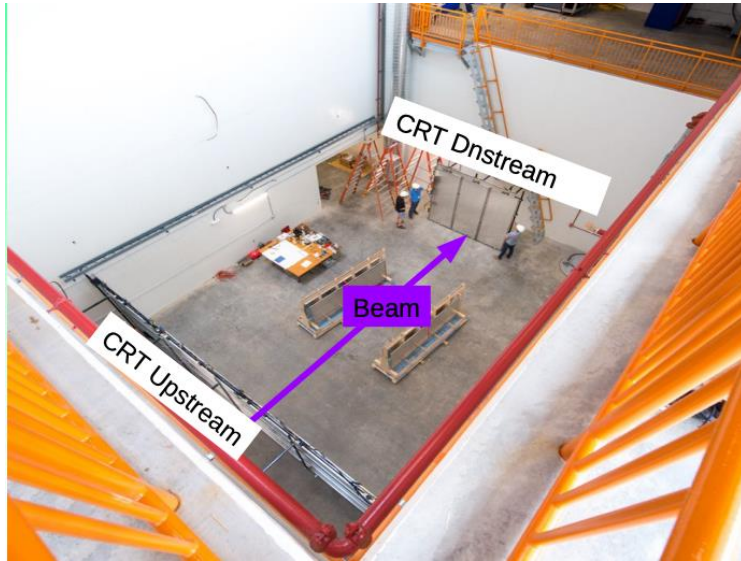
Candidate: $\mu + p$

Candidate: $\mu + 3p$

Candidate: $\mu + 2p + 2\pi^0$



CRT Only Run in 2018



Will we be able to see the beam spot in the upstream CRT wall?
YES!

